

Rugard Dressler



Formation of Stars



combined visible and near infrared pictures

star-forming pilar with a jet (HH 901)

Sumj

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Mn-

MeaN

Introduction

Carnia nebula: birthplace of stars

Hubble Space Telescope: http://hubblesite.org

<u>Cosmogenic Radio-Nuclides</u> (CoRN)

Produced

- via neutron capture reaction
- during explosive phases of star evolution
- □ in spallation reactions with high energetic cosmic rays

Present in the Solar System

- as constitute of molecular cloud formed our Sun e.g. found in meteorites
- injected continuously to "present days" from super novae
 - e.g. ⁶⁰Fe
- continuously produced at Earth from cosmic rays e.g. ⁷Be, ¹⁴C, ⁵³Mn

Sum Mn-MeaN Introduction







Short Lived Cosmogenic Radio-Nuclides $t_{1/2} < 100 \text{ Ma}$

Sum

parent nuclide	half-life (Ma)	daughter nuclide	estim. init. abundance	Reference	at PSI
⁵³ Mn	3.74	⁵³ Cr	⁵³ Mn / ⁵⁵ Mn = 6.28⋅10 ⁻⁶	Birck & Allègre 1985, Shukolyukov & Lugmair 2006, Trinquier et al. 2008	
¹⁰ Be	1.387	¹⁰ B	¹⁰ Be / ⁹ Be = 7.5·10 ⁻⁴	McKeegan et al. 2000, Chaussidon & Gounelle 2006, 2007	
²⁶ AI	0.717	²⁶ Mg	²⁶ Al / ²⁷ Al = 5.23·10 ⁻⁵	Jacobsen et al. 2008, Lee et al. 1976	
⁶⁰ Fe	2.62	⁶⁰ Ni	⁶⁰ Fe / ⁵⁶ Fe = 5.8·10 ⁻⁹	Quitt´e et al. 2010, Shukolyukov & Lugmair 1993, Tang & Dauphas 2011	Ø
³⁶ Cl	0.301	³⁶ S (~2%) ³⁶ Ar (~98%)	³⁶ Cl / ³⁵ Cl > 17.2·10 ⁻⁶	Jacobsen et al. 2009, Lin et al. 2005	
⁴¹ Ca	0.102	⁴¹ K	⁴¹ Ca / ⁴⁰ Ca = 1.41·10 ^{−8}	Srinivasan et al. 1994, 1996	
⁷ Be	1.46·10 ⁻⁷	⁷ Li	⁷ Be / ⁹ Be = 6.1·10 ^{−3}	Chaussidon et al. 2006	

parent nuclide	half-life (Ma)	daughter nuclide	estim. init. abundance	Reference	at PSI
¹⁴⁶ Sm	103	¹⁴² Nd	¹⁴⁶ Sm / ¹⁴⁴ Sm = 8.4·10 ⁻³	Boyet et al. 2010, Lugmair & Galer 1992, Prinzhofer et al. 1992	
⁹² Nb	34.7	⁹² Zr	⁹² Nb / ⁹³ Nb = 1.6·10 ⁻⁵	Harper 1996, Schönbächler et al. 2002	?
¹²⁹	15.7	¹²⁹ Xe	¹²⁹ / ¹²⁷ = 1.19·10 ⁻⁴	Brazzle et al. 1999, Jeffery & Reynolds 1961	?
¹⁸² Hf	8.9	¹⁸² W	¹⁸² Hf / ¹⁸⁰ Hf = 9.72·10 ⁻⁵	Burkhardt et al. 2008, Kleine et al. 2002, Yin et al. 2002	
¹⁰⁷ Pd	6.5	¹⁰⁷ Ag	¹⁰⁷ Pd / ¹⁰⁸ Pd = 5.9·10 ⁻⁵	Chen &Wasserburg 1996, Schönbächler et al. 2008	?
²⁰⁵ Pb	15.1	²⁰⁵ TI	²⁰⁵ Pb / ²⁰⁴ Pb = 1.0·10 ⁻³	Baker et al. 2010	
¹³⁵ Cs	2.3	¹³⁵ Ba	¹³⁵ Cs / ¹³³ Cs = 4.8·10 ⁻⁴	Hidaka et al. 2001	?
⁹⁷ Tc	4.21	⁹⁷ Mo	⁹⁷ Tc / ⁹² Mo < 3·10 ^{−6}	Dauphas et al. 2002	?
⁹⁸ Tc	4.2	⁹⁸ Ru	⁹⁸ Tc / ⁹⁶ Ru < 2·10− ⁵	Becker &Walker 2003	?
¹²⁶ Sn	0.23	¹²⁶ Te	¹²⁶ Sn / ¹²⁴ Sn < 7.7·10 ^{−5}	Fehr et al. 2006	?

MeaNCoRN

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<u>Mea</u>surement of <u>n</u>eutron capture cross sections and determination of half-live of short-lived <u>co</u>smogenic <u>r</u>adio-<u>n</u>uclides

Rugard Dressler, Dorothea Schumann Laboratory of Radiochemistry and		N
Environmental Chemistry	BIO	ln J
Jost Eikenberg		
Division of Radioprotection and Safety	LOG	Z
Ines Günther-Leopold		eaN
Hot Laboratory Division	NES	CoF
Klaus Kirch		Ž
Laboratory of Particle Physics	NUM	
Gunther Korschinek		C
Physics Department, Technical University Munich, Germany		ion
René Reifarth		
Institute for Applied Physics, Goethe University Frankfurt, Ger	rmany	

Subjects of MeaNCoRN

Neutron Capture Cross Section

- □ thermal neutron: NAA, PNA at PSI
- □ cold neutrons: BOA, ICON at PSI; FRM II in Munich
- □ ultra cold neutrons: UCN at PSI
- □ neutrons at cosmic energies: FRANZ at SGC Frankfurt

Half-Life

- □ ⁶⁰Fe, ⁶³Ni
- □ ¹⁰Be, ⁵³Mn
- □ ²⁶Al, ³²Si, ⁵⁹Ni
- □ ⁴¹Ca, ⁷⁹Se, ¹²⁶Sn, ...

Financial & Manpower Support

- □ SNSF founding (rejected)
- □ PSI Cross Project
- □ ??? other sources ???

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History of ⁵³Mn Half-Life Measurements

Author (Year)	Method	t _{1/2} [Ma]
Wilkinson, et al. (1955)	Compared nuclear reaction yield with ⁵⁴ Mn	0.00014
Sheline, et al. (1957)	Calculated nuclear reaction yield	~2
Kaye, et al. (1965)	Spallation yield of meteorites	1.9±0.5
Hohlfelder (1969)	Mass spectrometry (MS) of meteorites	10.8±4.5
Matsuda, et al. (1971)	MS of 730 MeV proton activation products	2.9±1.2
Honda, et al. (1971)	MS of artificial and meteoritic samples	3.7±0.37
Wölfle, et al. (1972)	Neutron activation of meteoritic samples	3.9±0.6
Heimann, et al. (1974)	decay of meteoritic ⁵³ Mn	3.85±0.4

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CoRN Storehouses



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Fe samples of STIP see talk Maruta Bunka

total material ~ 60 g steel

radio nuclides available ²⁶Al 300 Bq ~ 10^{16} atoms ⁴⁴Ti 100 MBq ≈ 8.4·10¹⁷ atoms ⁵⁴Mn 70 MBq (⁵³Mn ~10¹⁹ atoms) ⁶⁰Co 70 MBq Cu beam dump see talk Marin Ayranov

total material ~ 500 g copper

radio nuclides inventory ²⁶Al 7 kBq $\approx 2.3 \cdot 10^{17}$ atoms ³²Si 10 MBq $\approx 7.8 \cdot 10^{16}$ atoms ⁴⁴Ti 100 MBq $\approx 2.8 \cdot 10^{17}$ atoms ⁵³Mn 500 kBq $\approx 8.4 \cdot 10^{19}$ atoms ⁵⁹Ni 8 MBq $\approx 2.7 \cdot 10^{19}$ atoms ⁶⁰Fe 5 kBq $\approx 5.9 \cdot 10^{17}$ atoms ⁶⁰Co 5 GBq

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ĺ	CoRN Storehouses		
I	STIP samples	Cu beam dump	Z.
	pro low ⁶⁰ Co content separated from matrix purified final product cons	pro storage time ~20 a EoB Dec 1992 low ⁵⁴Mn activity non carrier added 	Mn-53
	 high amount of stable isotopes storage time ~12 a EoB Dec 1999 high ⁵⁴Mn activity 	Cons high activity of ⁶⁰Co not separated from Cu not purified 	oRN ction

Half-Life Measurement $t_{1/2} > 50 a$

Determination of sample activity

• α -/ β -/ γ -Spectroscopy or Liquid Scintillation Counting

Determination of number of atoms

AMS / ICP-MS / HI-ERD

□ ${}^{10}\text{Be}\ t_{1/2} = 1.388 \pm 0.018\ Ma$ G. Korschinek, et al.: Nucl. Instr. and Methods B 268 (2010) 187 □ ${}^{10}\text{Be}\ t_{1/2} = 1.386 \pm 0.016\ Ma$ J. Chmeleff, et al.: Nucl. Instr. and Methods B 268 (2010) 192

⁷⁹Se t_{1/2} = 0.327 ± 0.008 Ma
 G.Jörg, et al.: Applied Radiation and Isotopes 68 (2010) 2339

□ 60 Fe $t_{1/2}$ = 2.62 ± 0.04 Ma G. Rugel, et al.: Phys. Rev. Lett. 103 (2009) 072502

⁶³Ni Half-Life Measurements at PSI

⁶³Ni pure β^{-} emitter $E_{max} = 66.9 \text{ keV}$ ⁶³Ni target for n_ToF production □ 988 mg enriched ⁶²Ni metal pellets neutron activated □ 109 mg ⁶³Ni $\approx 10^{21}$ atoms ≈ 230 GBg □ separation of ⁶³Cu about 0.5% left as waste ≈ 5.10¹⁸ atoms ≈ 1 GBq can be used for other measurements



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Why additional Measurements?

• Collè et al. 1996:

new value based on the revised value from Barnes et. Al own ones using decay measurements with LSC technique
3 data points over 27 years: 101.1 ± 1.4 a

• Collè et al. 2008:

new re-standardization of the used NIST solutions additional data points for the decay measurement 4 points over about 40 years: 101.2 ± 1.5 a

- **1.** Additional measurements using other techniques would be useful to support the result of Collè.
- 2. ⁶³Ni is an excellent play ground to prove the precision of our used instruments and methods.

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Half-Life Measurements at PSI

1. Approach



⁶³Ni / ⁶²Ni

- ICP-MS measurements:
 - **1.** isotopic ratio before purification: $(^{63}Ni + ^{63}Cu) / ^{62}Ni$

2. isotopic ratio after purification:

• Time from EoB to separation:

	sample	^{63/62} Ni (K)	uncertainty	^{63/62} Ni (L)	uncertainty
	after dissolving	0.145366	0.000037	0.124784	0.000266
	after CuS separation	0.125098	0.000010	0.107667	0.000013
	after purification	0.122395	0.000289	0.107846	0.000169
	N _o /N	1.1877	0.0028	1.1590	0.0025
N	Δt	? (25 a)		? (21.5 a)	
	t _{1/2}	(101.2 a)		(101.2 a)	

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No

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