

Is it Possible to Study $^{44}\text{Ti}(\alpha, p)^{47}\text{V}$ Reaction with Radioactive Target?

Tariq Al-Abdullah^(1,2),

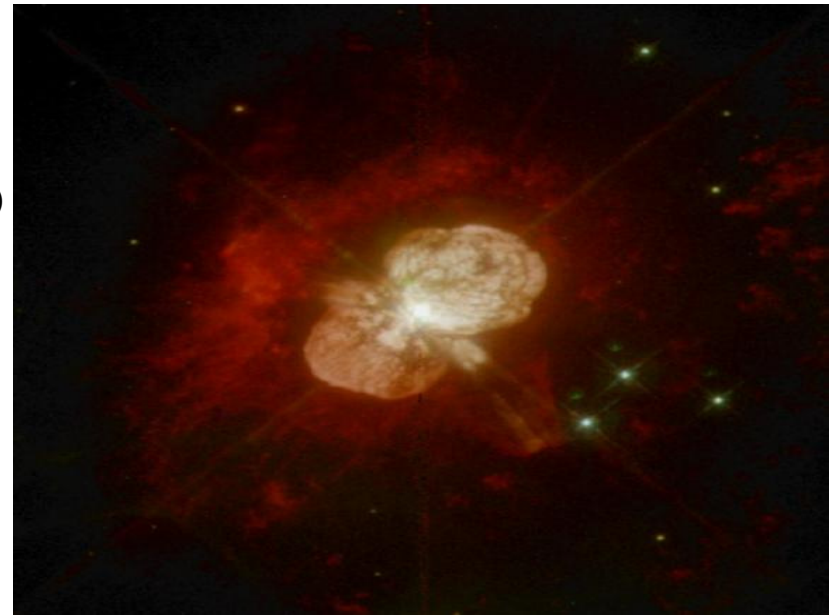
D. Bemmerer⁽²⁾, **K. Schmidt**⁽²⁾,

D. Schumann⁽³⁾, **R. Dressler**⁽³⁾, **M. Ayrarov**⁽³⁾

¹Hashemite University, Zarqa-Jordan,

²HZDR, Dresden-Germany,

³PSI, Villigen-Switzerland



PSI 31.08.2011, Villigen/Switzerland

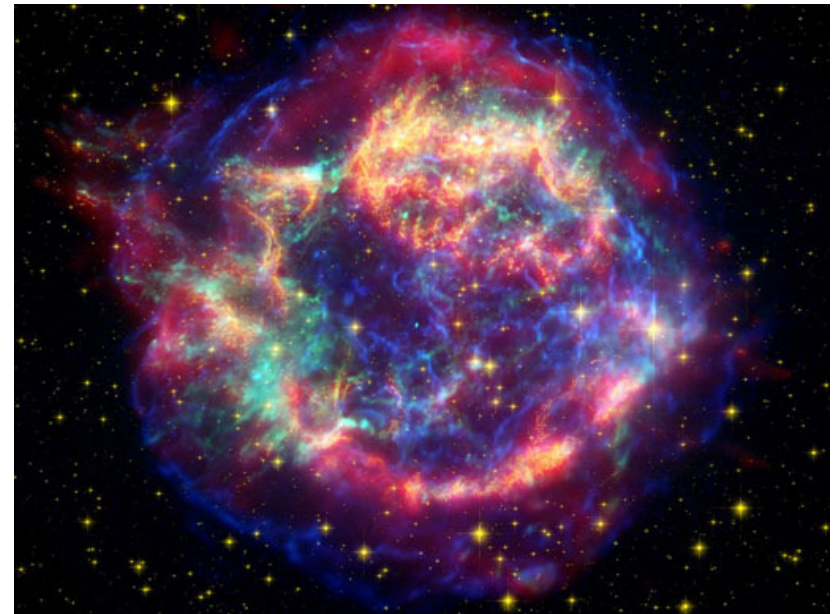


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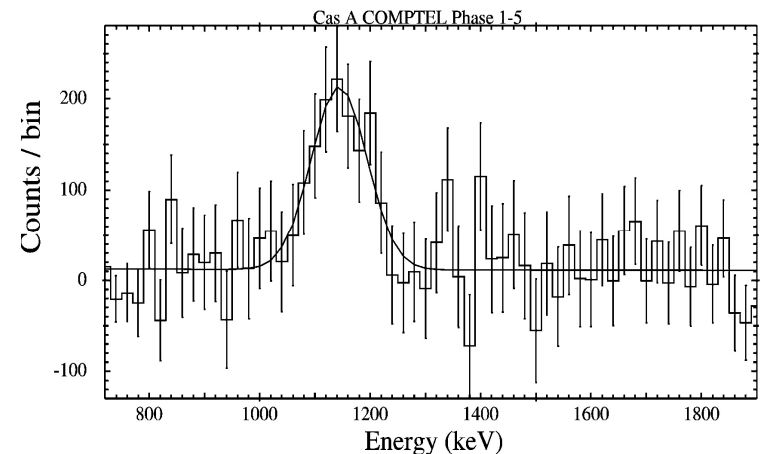
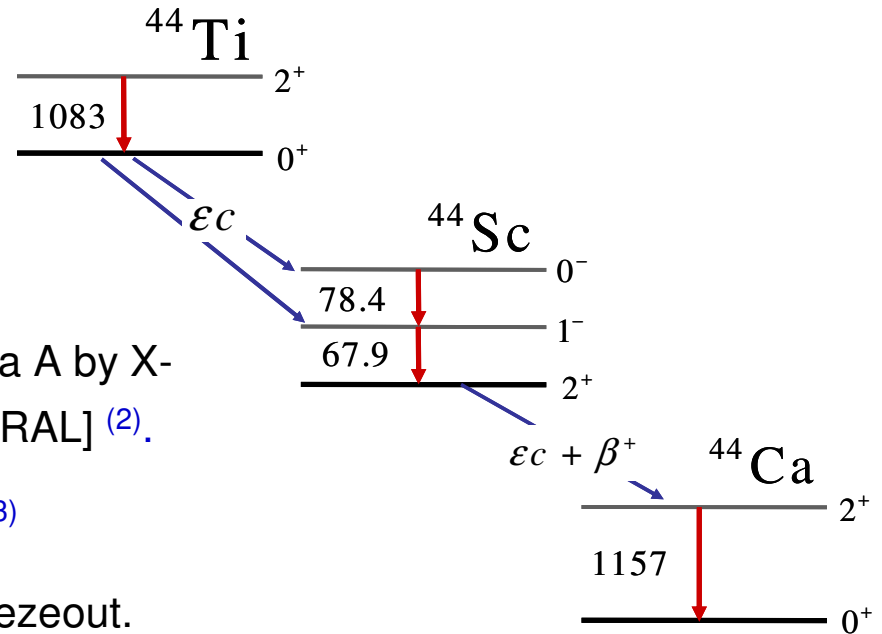
Main Points

- ❑ The Importance of ^{44}Ti in Supernovae.
- ❑ The Reaction Rate for $^{44}\text{Ti}(\alpha, p)^{47}\text{V}$ at Astrophysical Energies.
- ❑ Problems with Radioactive ^{44}Ti Target.
- ❑ The Safety Concerns of the Experiment.
- ❑ Suggestions and Future Plans.



□ The importance of ^{44}Ti in supernovae.

- ^{44}Ti , $t_{1/2} = 58.9 \pm 0.3 \text{ yr}$ ⁽¹⁾.
- Its $^{44}\text{Ti} \xrightarrow{EC} ^{44}\text{Sc}$, 78.4 & 67.9 keV.
- $^{44}\text{Sc} \xrightarrow{\beta^+ + EC} ^{44}\text{Ca}$, γ -lines at 1157 keV.
- 1.157 MeV has been observed from Cassiopeia A by X-rays & γ -rays telescopes [CGRO, RXTE, INTEGRAL]⁽²⁾.
- The ejected ^{44}Ti is estimated: $1.6_{-0.3}^{+0.6} \times 10^{-4} M_{\odot}$ ⁽³⁾
- At $T_9 = 4.3$, its abundance grows after α -rich freezeout.
- The yield is sensitive to: $^{40}\text{Ca}(\alpha, \gamma)^{44}\text{Ti}$, $^{44}\text{Ti}(\alpha, p)^{47}\text{V}$, $^{44}\text{Ti}(\alpha, \gamma)^{48}\text{Cr}$, $^{45}\text{V}(p, \gamma)^{46}\text{Cr}$.
- **L.-S. The**⁽⁴⁾: $^{44}\text{Ti}(\alpha, p)^{47}\text{V}$ is the most important reaction.



- (1) I. Ahmad *et al*, PRC 74, 065803 (2006).
- (2) A. f. Iyudin *et al*, Astron. Astrophys. 284, L1 (1994)
- (3) M. Renaud *et al.*, Astrophys. J. 647, L41 (2006)
- (4) L.-S. The *et al.*, Astropjys. J. 504, 500 (1998)

❑ The reaction rate for $^{44}\text{Ti}(\alpha, p)^{47}\text{V}$ at astrophysical energies.

- This reaction is measured in inverse kinematics by:

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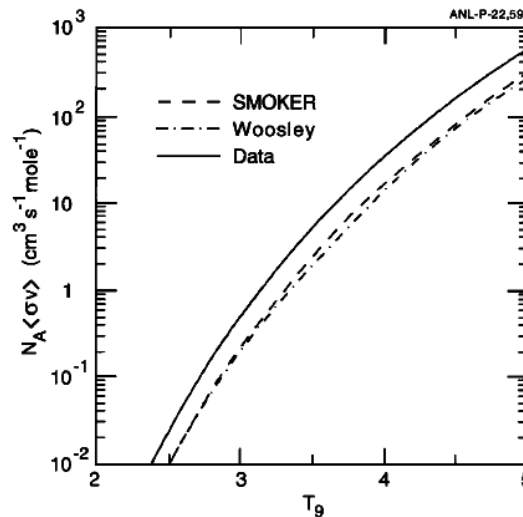
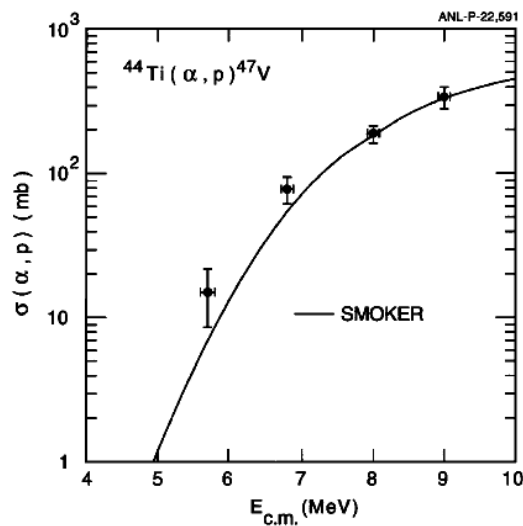
PHYSICAL REVIEW LETTERS

21 FEBRUARY 2000

The $^{44}\text{Ti}(\alpha, p)$ Reaction and its Implication on the ^{44}Ti Yield in Supernovae

A. A. Sonzogni,¹ K. E. Rehm,¹ I. Ahmad,¹ F. Borasi,² D. L. Bowers,¹ F. Brumwell,¹ J. Caggiano,¹ C. N. Davids,¹ J. P. Greene,¹ B. Harss,¹ A. Heinz,¹ D. Henderson,¹ R. V. F. Janssens,¹ C. L. Jiang,¹ G. McMichael,¹ J. Nolen,¹ R. C. Pardo,¹ M. Paul,³ J. P. Schiffer,¹ R. E. Segel,² D. Seweryniak,¹ R. H. Siemssen,¹ J. W. Truran,⁴ J. Uusitalo,¹ I. Wiedenhöver,¹ and B. Zabransky¹

- Energy range: $5.7 \text{ MeV} \leq E_{c.m.} \leq 9 \text{ MeV}$, using a beam of ^{44}Ti



- The estimated reaction rate is a factor of 2 greater than the theoretical calculations, with uncertainty of 37%.

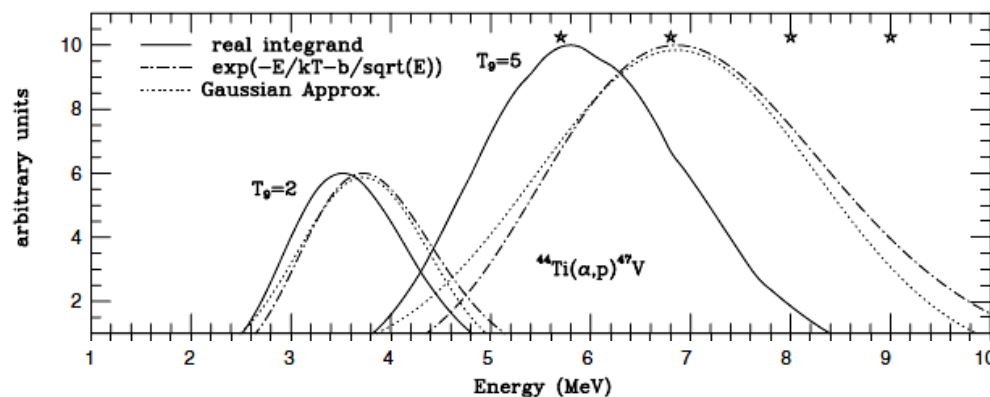
❑ The reaction rate for $^{44}\text{Ti}(\alpha, p)^{47}\text{V}$ at astrophysical energies.

➤ **Old Gamow peak:** An overlap between MB-distribution and Coulomb barrier penetration [Rauscher et al, PRC 81, 045807 (2010).

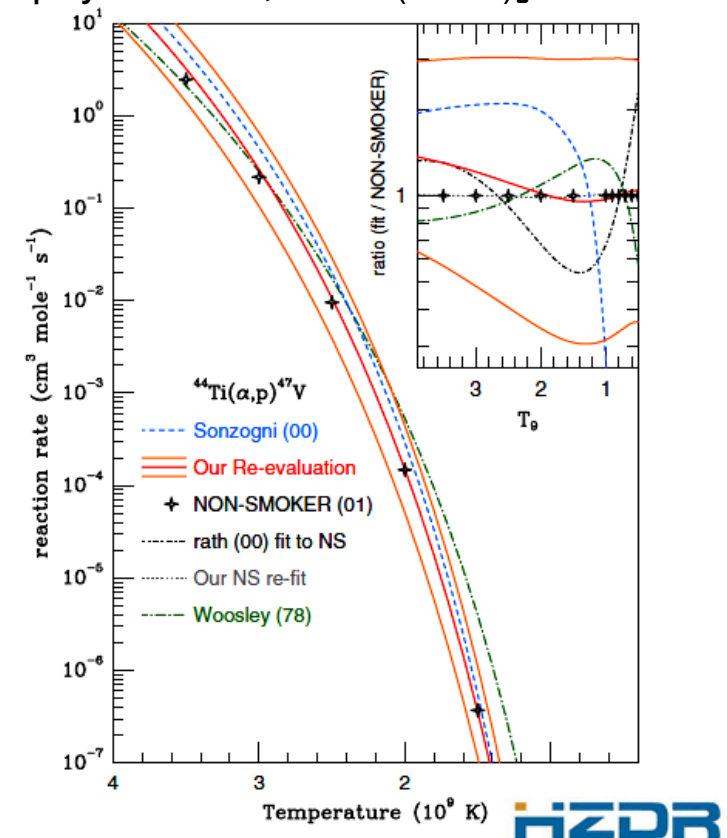
➤ Not good if α -width is changing rapidly with the Coulomb barrier.

➤ **The true Gamow Window** $N_A \langle \bar{\sigma}_{jk}^\mu v \rangle = \frac{3.732 \times 10^{10}}{\hat{A}_j^{1/2} T_9^{3/2}} \times \int_0^\infty \bar{\sigma}_{jk}^\mu E_j^\mu \exp(-11.605 E_j^\mu / T_9) dE_j^\mu, \quad (4)$

➤ The reaction rate is re-evaluated **R. D. Hofmann**; [Astronphys. J. 715, 1383 (2010)].



- The most relevant temp. range is $1 \leq T_9 \leq 5$.
- Cross sections at lower energies dominate both the rates and uncertainties.
- Larger uncertainty in the rate (± 3) $\rightarrow$ large uncertainty in ^{44}Ti synthesis



❑ The planned radioactive ^{44}Ti target.

- A 10-50 MBq radioactive target will be prepared in PSI, [D. Schumann, ERAWAST-PSI].

- For 10 MBq, $D=20$ mm, $T_{1/2}=58.9$ yr, using

$$N = \frac{A}{L}, \quad L = \frac{\ln 2}{T_{1/2}}$$

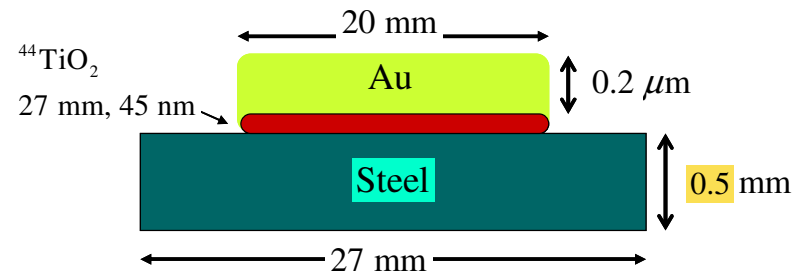
- $\rightarrow N = 2.68 \times 10^{16}$; ^{44}Ti

- The thickness of the ($^{44}\text{TiO}_2$) target is almost 45 nm, or $11 \mu\text{g}/\text{cm}^2$.

- $0.2 \mu\text{m}$ -Gold will coat the radioactive area.

- Stainless steel support: 0.5 mm thick, 27 mm in diameter.

- High purity target is required, no impurities from $^{46-50}\text{Ti}$ isotopes.



Ti (inert)-Target

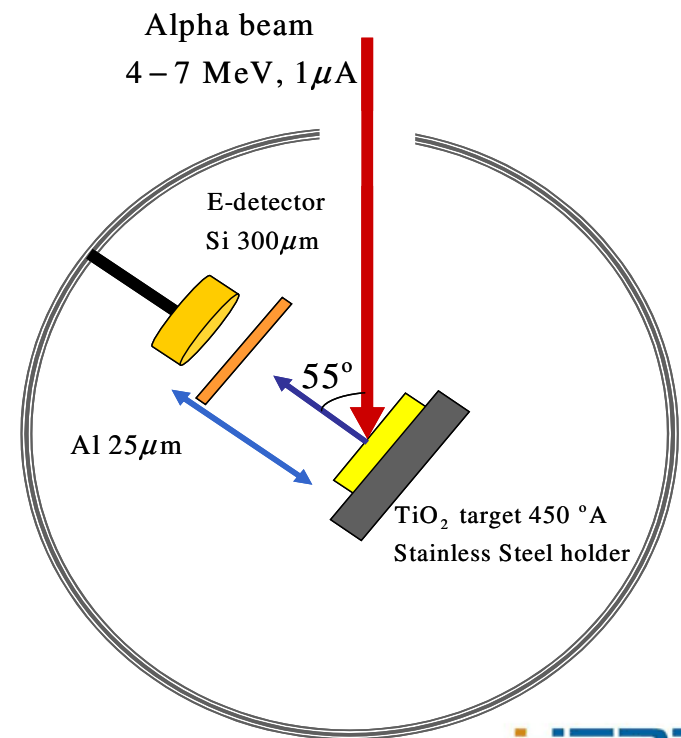
^{44}Ti -Target

Ta-disk



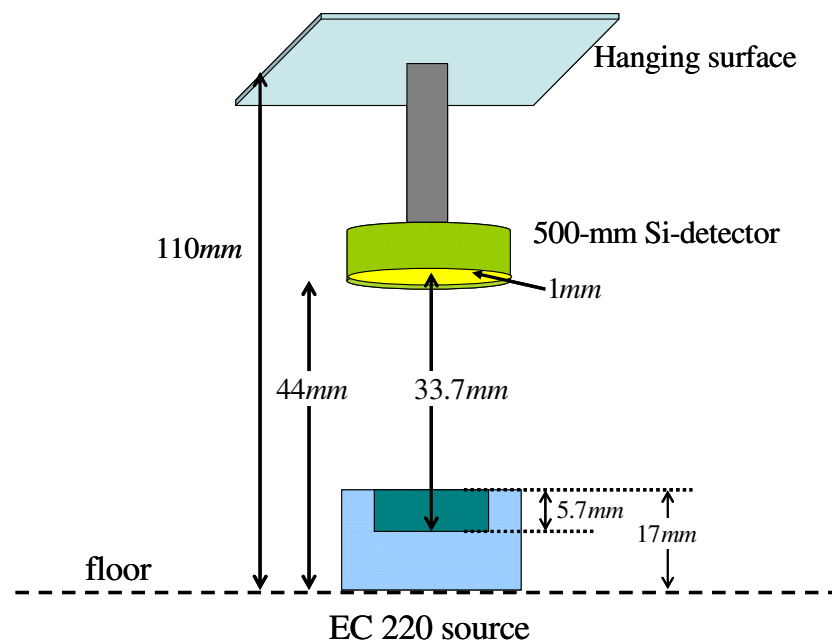
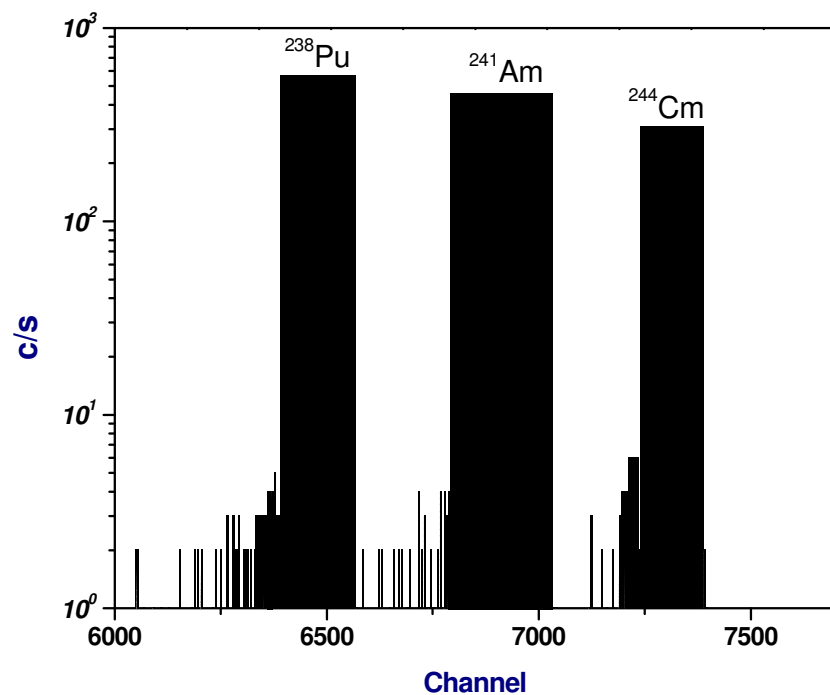
❑ Possible Experimental Setup

- ION-Beam Center at HZDR.
- Q-Value = -- 0.41 MeV.
- α -beam at 4-7 MeV, Angle = 55° . Current = $1\ \mu\text{A}$ for 1-2 weeks.
- $25\ \mu\text{m}$ Al-foil is used to stop α -particles before the detector.
- 100-300 μm PIPS [Partially Implanted Passivated Silicon].



❑ Can $^{44}\text{Ti}(\alpha, p)^{47}\text{V}$ be performed?

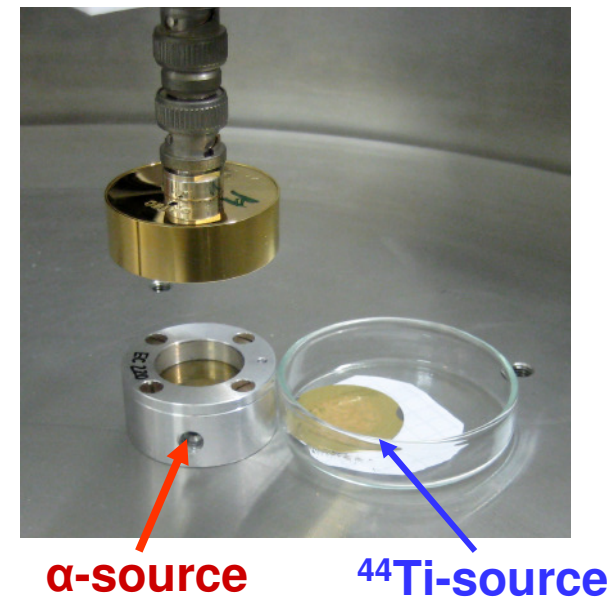
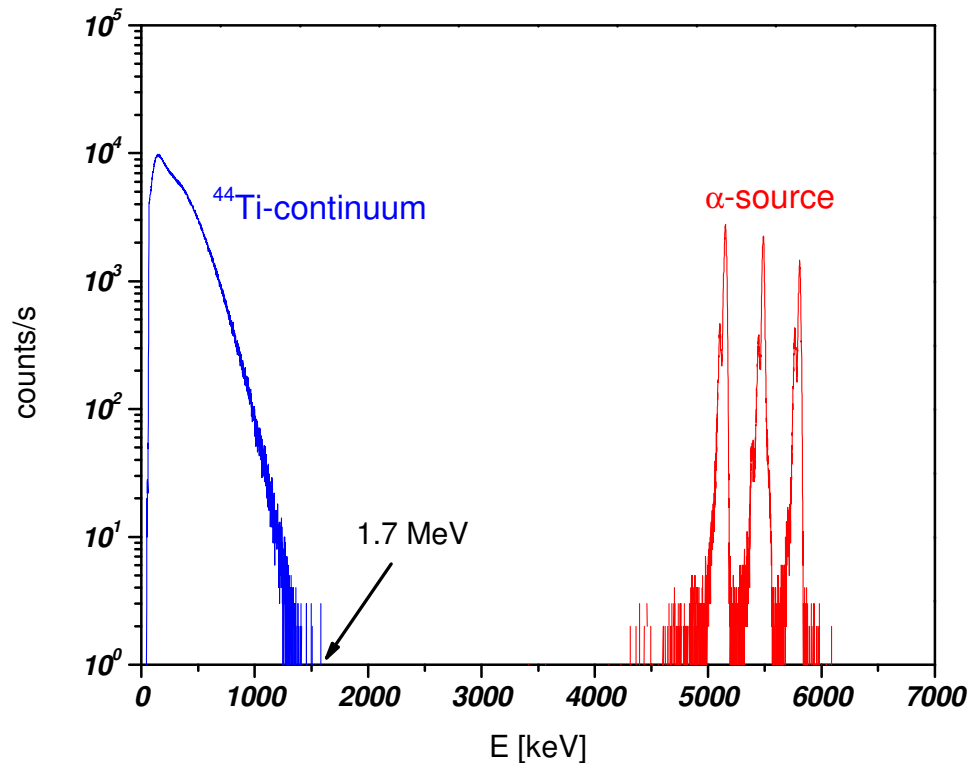
- Starting with simple setup [500 μm -Ortec Si-detector with AmCmPu source].



- For ^{241}Am main peak, $E = 5486 \pm 21$ keV.

❑ Can $^{44}\text{Ti}(\alpha, p)^{47}\text{V}$ be performed?

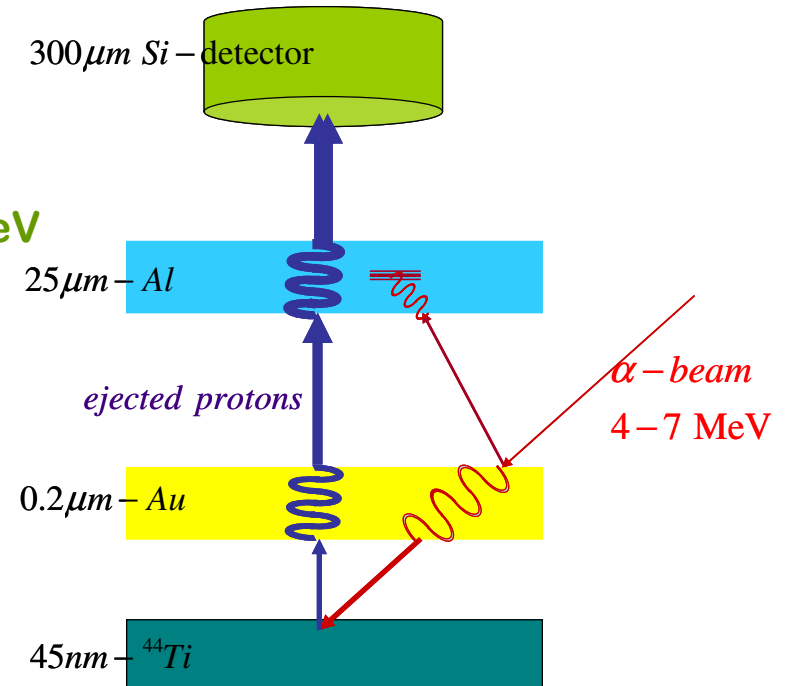
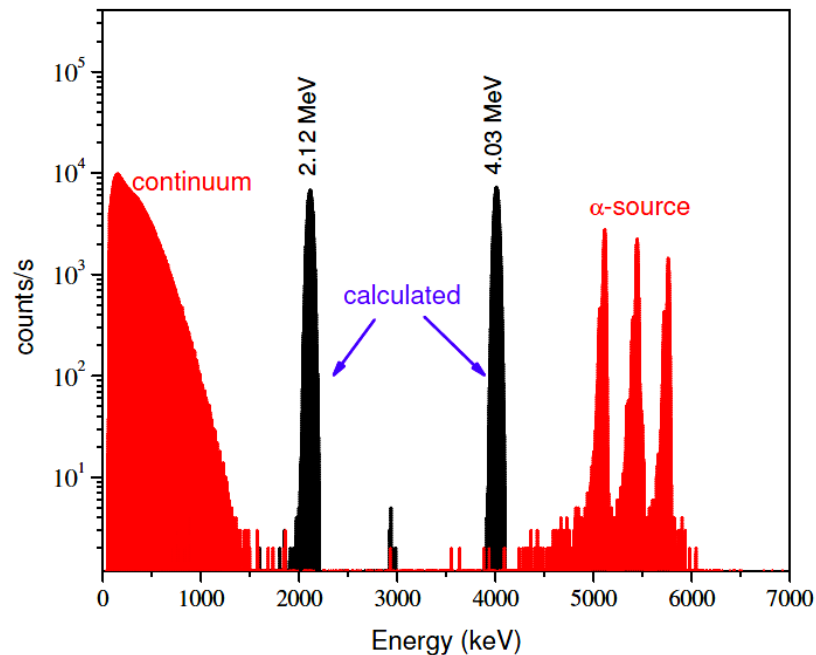
- AmCuPu-source and ^{44}Ti (83 kBq) are both observed by the 500 μm Si-detector



- A strong increase in the background continuum; 1 % to 94 % counting rate.
- Continuum ends point is at almost 1.7 MeV.
- **IS IT GOOD?**

❑ Can $^{44}\text{Ti}(\alpha, p)^{47}\text{V}$ be performed?

- Detecting the protons after the reaction.
 - Using Lise & Trim codes.
 - Starting at $E_\alpha = 4 \text{ MeV} \rightarrow \text{Gold} \rightarrow E_\alpha = 3.82 \pm 0.04 \text{ MeV}$
 - $E_p = 2.8 \text{ MeV} \rightarrow \text{Au \& Al} \rightarrow E_p = 2.12 \pm 0.08 \text{ MeV}$.
- Similarly for $E_\alpha = 6 \text{ MeV}$,
- $E_p = 4.49 \text{ MeV} \rightarrow \text{Au \& Al} \rightarrow E_p = 4.03 \pm 0.08 \text{ MeV}$.



- Protons peaks do not overlap with the continuum.
- PIPS detectors can be used.
- The height of the peaks is in random units.

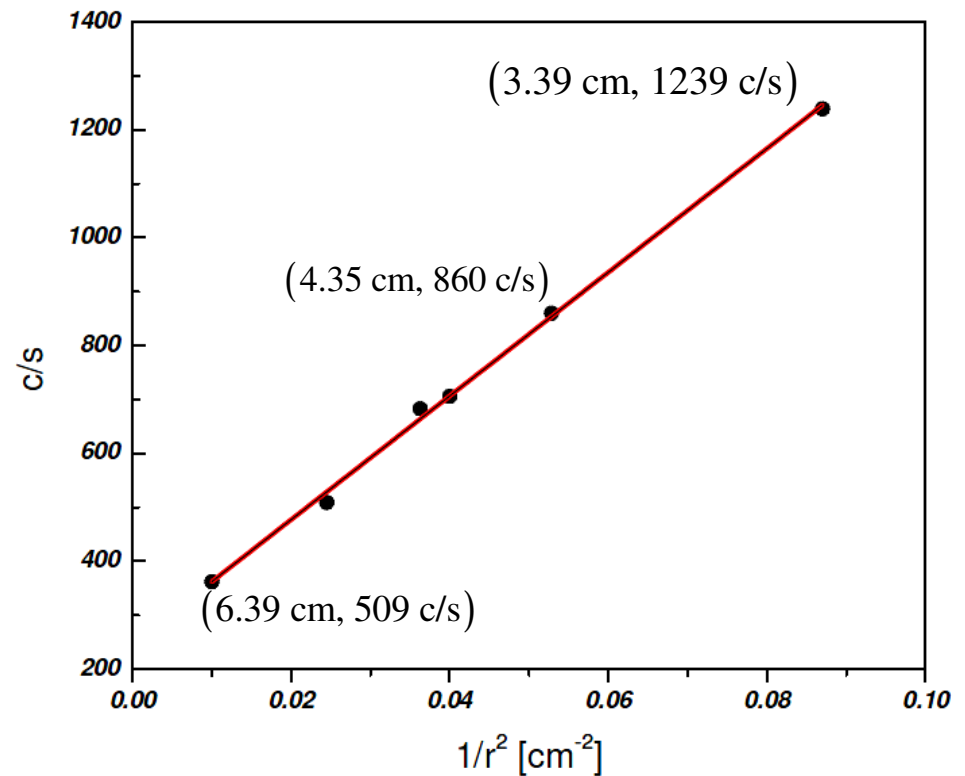
❑ Problems with radioactive ^{44}Ti target.

- Using the 83-kBq ^{44}Ti source.
- The relation between the activity and the distance for the continuum background is:

$$A = 11473 \times \frac{1}{r^2} + 247 \quad [\text{c/s}].$$

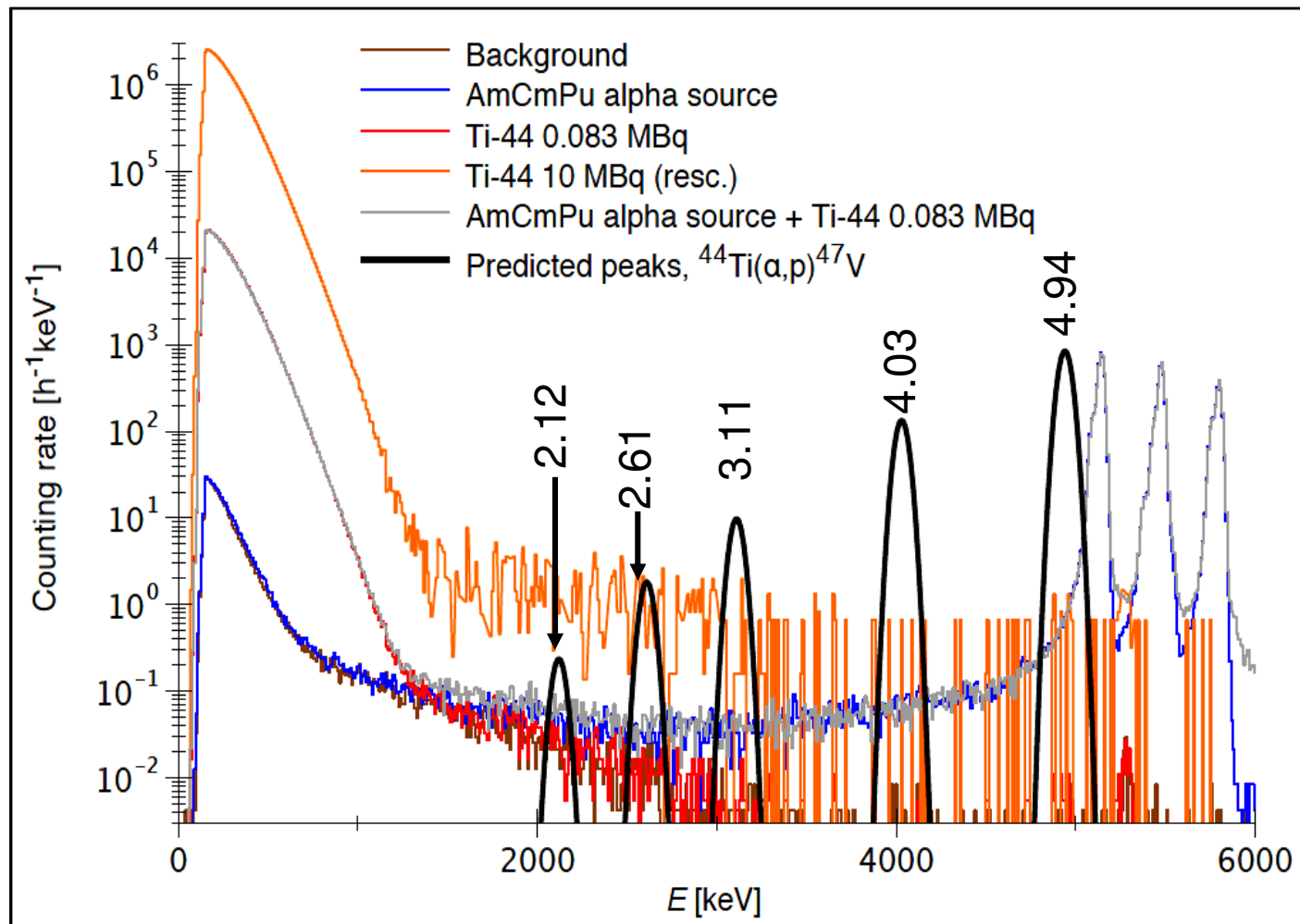
- The estimated continuum background for 10MBq-target is:

Dist. [cm]	Rate [c/s]	10 MBq [c/s]
5	706	5.5E+4
10	362	1.4E+4



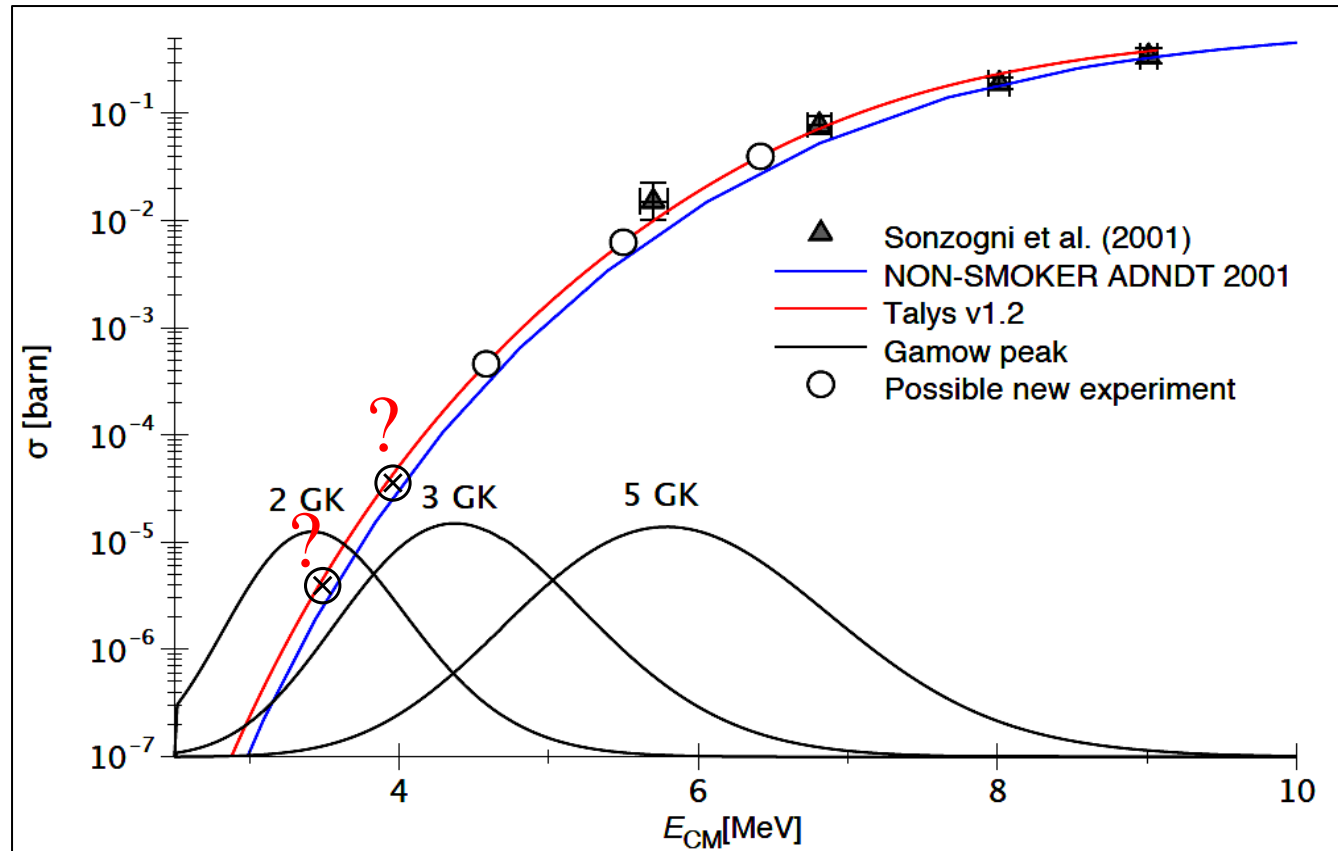
❑ Problems with radioactive ^{44}Ti target.

- Several runs with **Am** and ^{44}Ti sources are collected at the same geometry.
- Estimating the proton peaks in **300 μm Si-detector**. $E_\alpha = 4, 4.5, 5, 6, 7 \text{ MeV}$.
- The background from the cosmic rays hide peaks for $E_\alpha \leq 4.5 \text{ MeV}$.



❑ The Feasibility of the $^{44}\text{Ti}(\alpha, p)^{47}\text{V}$!

- Adopting Rauscher calculations, PRC 81, 045807 (2010).
- Gamow Windows are re-calculated for $T_9 = 2, 3, 5$.
- Talys code is used. [A. J. Koning, AIP 769, 1154 (2005)]
- Three new points are very probable to measure. Two more are question marks.



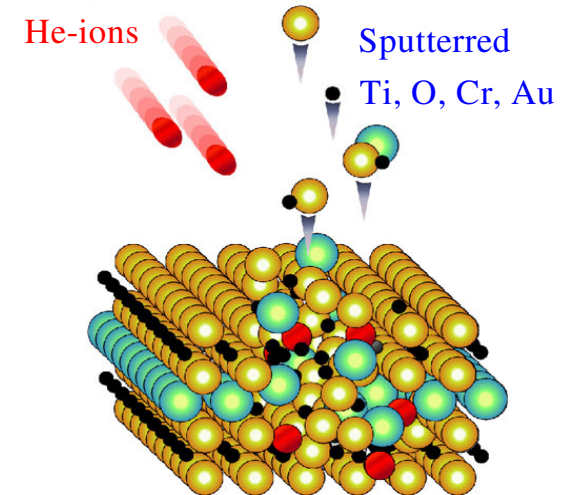
❑ Can $^{44}\text{Ti}(\alpha, p)^{47}\text{V}$ be performed?

- Estimating the yield at **5 & 10 cm** between the target and the detector:.
- Efficiency: **5 cm** $\rightarrow$ 9.6E-3, **10 cm** $\rightarrow$ 2.4E-3.
- The reaction is feasible for **$E_{lab} \geq 4$ MeV**.

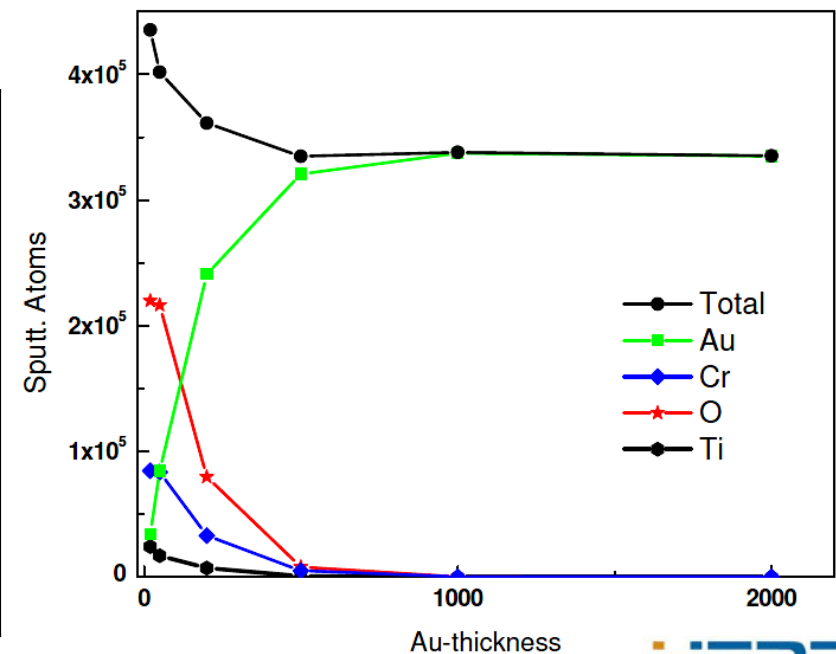
E_{lab} [MeV]	σ [mb]	Dist. 5-cm [Count/hr]	Dist. 10-cm [Count/hr]
3.5	8.96E-4	1.6	0.4
4	1.11E-2	20.3	5.8
5	4.59E-1	84.1	21.0
6	6.26E+0	1.15E+4	2.87E+3
7	3.79E+1	7.27E+4	1.82E+4

❑ The Safety Concerns of the Experiment.

- The sputtering of the radioactive atoms from the target.
- Several calculations are made using TRIDYN_FZD code*. [*Dr. M. Posselt, HZDR*] * [W. Möller & W. Eckstein, NIM B 2, 814 (1984)].
- Parameters: Fluence = 10^{19} atoms/cm² [1 μ A-10 days], NH = 10^8 [quality], E_α = 6.0 MeV @ 55°, TiO₂ [450 °A], Layers [Au varying, Cr = 50 °A]

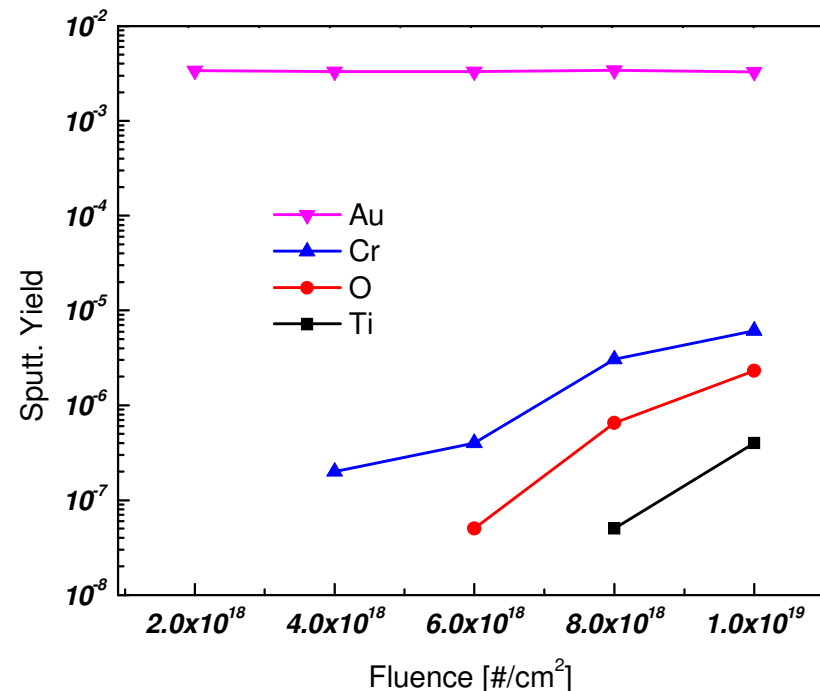
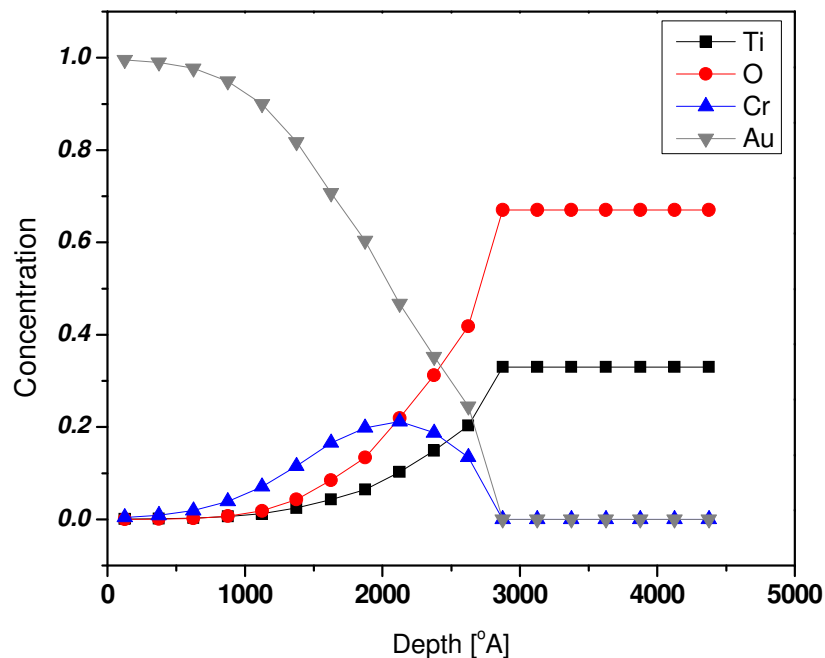


Au- Thick	20	50	200	500	1000	2000
Depth	-105.7	-91.2	-69.6	-58.7	-55.8	-38.8
Ti	24359	17058	7284	920	34	9
O	219922	216518	79827	8160	271	16
Cr	84656	83628	33066	5106	389	196
Au	34417	84578	241035	320695	337291	334982
Total	435354	401782	361212	334881	337985	335248



❑ The Safety Concerns of the Experiment.

- For our case: when the thickness of the Gold is $0.2\ \mu\text{m}$.
- The depth profiles of the components for $10^{19}\ \text{ions/cm}^2$ fluence shows a strong mixing between the atoms at all depths.
- Decreasing the fluence will rapidly decrease the sputtered ^{44}Ti atoms.



In worst case, Au-20 $^{\circ}\text{A}$ thick FL = $10^{19}/\text{cm}^2$,
the sputtered ^{44}Ti is 25 k $\rightarrow$ 10 μBq .

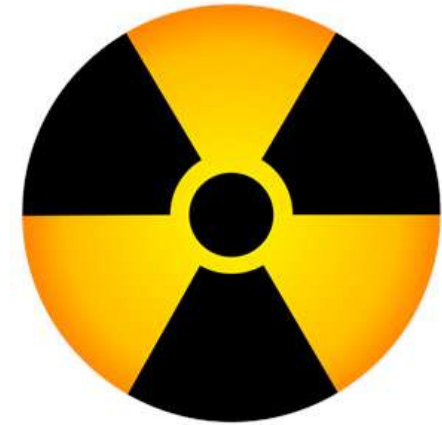
❑ The Safety Concerns of the Experiment.

Estimating the amount of the absorbed dose [Sv/h].

For 10 MBq $^{44}\text{Ti}/^{44}\text{Sc}$ source

Dose rate at 1 m distance from the source

$$4 \frac{\mu\text{Sv}}{\text{h}} \Leftrightarrow 8 \frac{\text{mSv}}{\text{yr}} \quad \text{Controlled Area}$$



In case of swallowing the source

10 MBq $^{44}\text{Ti}/^{44}\text{Sc}$ source

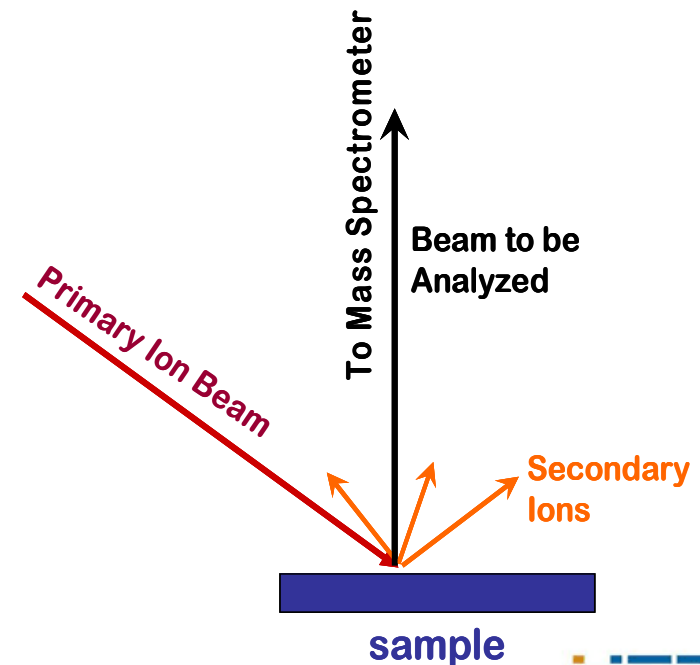
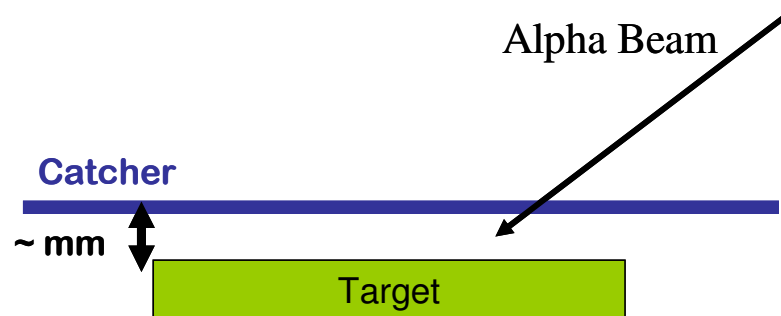
→ 66 mSv $\longleftrightarrow$ total dose

→ If it is not digested, the hazard is small.



❑ Suggestions and Future Plans.

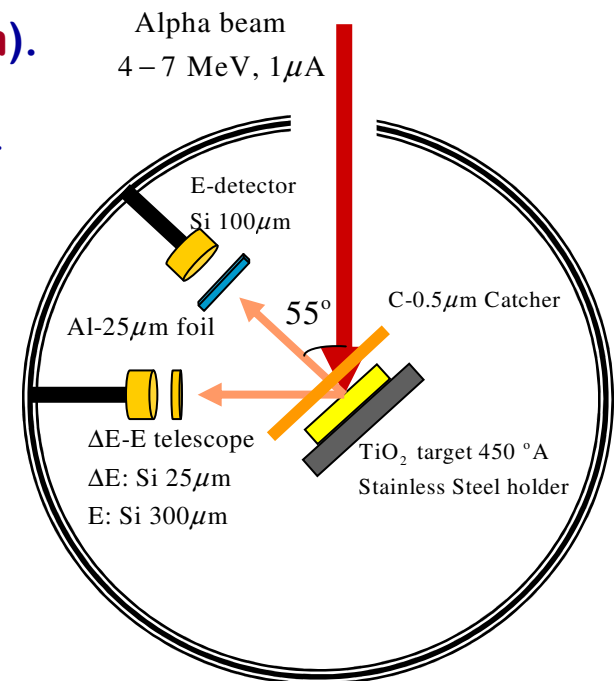
- Measuring the sputtering yield experimentally with an inert target.
 - An inert target is prepared and available, [D. Schumann, PSI].
 - The target has same properties as the radioactive one.
 - Sputtered atoms will be collected on a catcher, [^{12}C , 0.5-1 μm].
 - A Secondary Ion Mass Spectroscopy (SIMS) technique.
- Analyze the composition of the target surface and the catcher.



❑ Suggestions and Future Plans.

- The continuum background from the cosmic rays can be reduced with using a telescope-detection system.
- Two particle detectors PIPS; ΔE (25 μm) & E (300 μm).
- More PIPS detectors inside the chamber at different reaction angles will be good.

- Monte-Carlo simulation; GEANT 4.
- Calculate the efficiency of the detectors.
- Estimate the height of the peaks.
- Reduce the background.



❑ Legitimacy!

- The experiment can not be performed until:
 - Getting the 10-50 MBq ^{44}Ti -target.
 - ~MBq is high → several permissions and paper works.
 - Quite a few tools need to be available.
 - More than a few test runs must be made.

□ Summary & Conclusion

- The $^{44}\text{Ti}(\alpha, p)^{47}\text{V}$ reaction is possible with:



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- Look for promising results that are better than those from radioactive beam facilities.

Thank you