

The Pool



B. Ohayon MuX for CKM

F. Wauters μ^- and APV

A. Knecht μ^- conversion

L. Gerchow μ^- elem. analysis

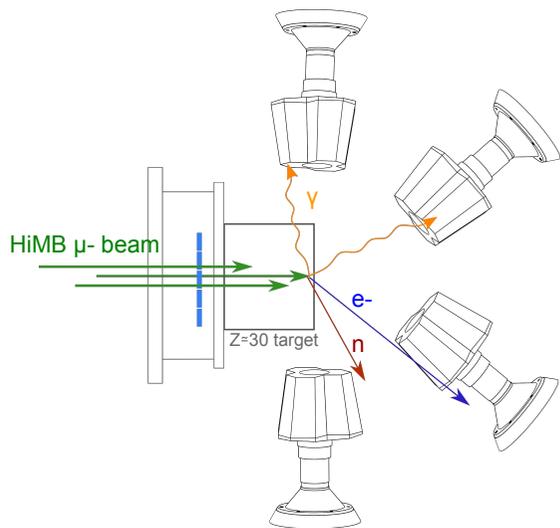
A. Antognini e^+ beam

A. Papa dark matter with e^+

A. Papa μ^+ and ALPS

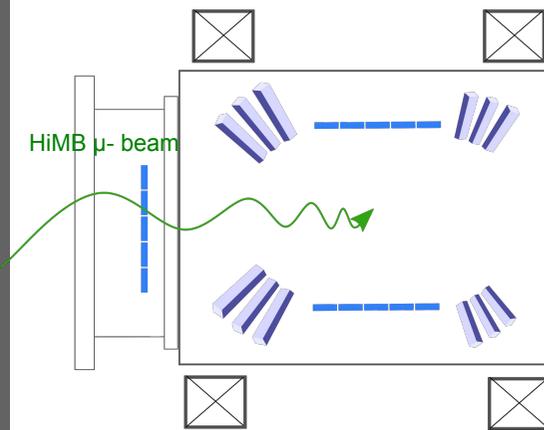
U. Langenegger $M \leftrightarrow Mbar$

S. Davidson μ^+ decays



- ❑ HPGe detectors have both high resolution and efficiency
- ❑ n-dose and rate limited
- ❑ $> \pi E I$ intensities are beneficial

With a HPGe-array, one can **not** fully exploit HiMB intensities.

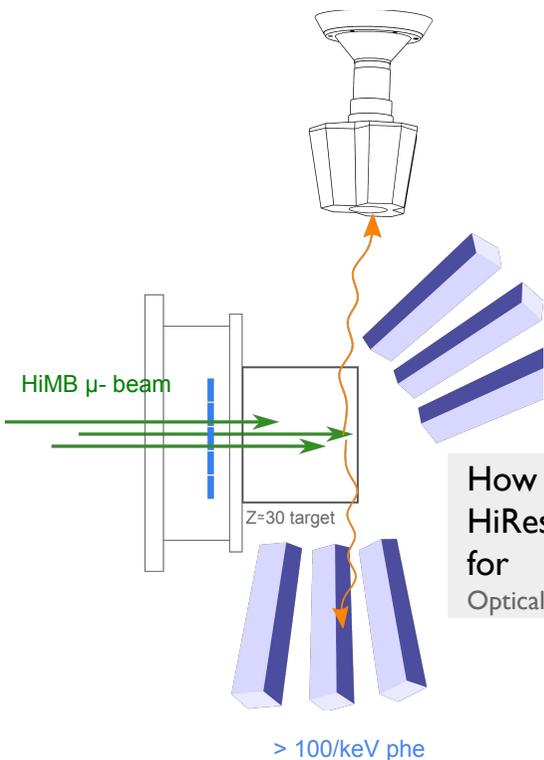


- ❑ Z = 5-10 APV experiments have a potential $< SM$ reach
- ❑ Building on work of L. Simons and Co
- ❑ Need to suppress Auger transitions
- ❑ Need to develop additional techniques to differentiate 2s1s from 2p1s X-rays, such as K-edge absorption

Can we stop a HiMB beam in a low pressure gas target?

μ APV with HiMB

and other precision muonic atom experiments

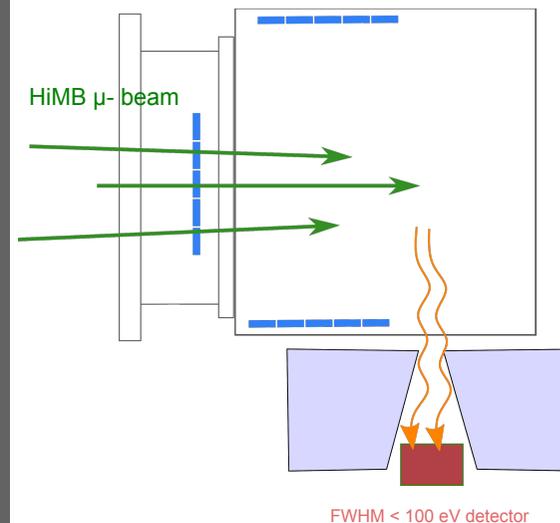


- ❑ A coincidence experiment with a close scintillator and a HPGe far detector has a 10^{-3} reach
- ❑ A pure high-resolution has a potential 10^{-4} reach
- ❑ SM μ APV $\leq 10^{-4}$

How well can current and future HiRes scintillator detectors perform for muonic X-ray measurements?
Optical Materials: X I (2019) 100021

Solid $\geq O(SM)$ amplitude APV experiment motivation?

$> 100/\text{keV}$ phe



- ❑ < 100 eV resolution can separate in energy most transitions of interest
- ❑ For Z=5, 10^{-3} precision is worth-while (1 cm² at 1m for 100 days)
- ❑ Other high-resolution measurements such as BSQED? arXiv:2011.09715
- ❑ Can we efficiently operate new detector technologies such as cryogenic microcalorimeter NIMA 770, 203-210 (2015)

FWHM < 100 eV detector

Should we bring high-resolution (< 100 eV) X-ray spectroscopy techniques to muonic X-ray studies @ PSI?

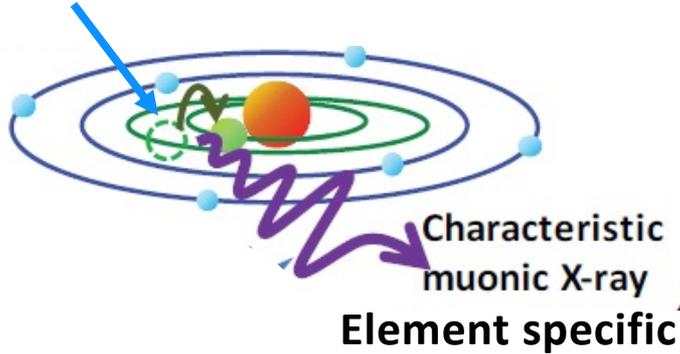
Conclusions muon conversion at HIMB

- ▶ Mu2e and COMET will perform their measurements with expected muon rates of around $1e10 \mu^-/s$
- ▶ They will improve the sensitivity over SINDRUM II by a factor 10^4 compared to muon conversion at HIMB that could improve the sensitivity at most by a factor 100
- ▶ Changes in expected branching ratio as a function of Z are on the level of <4

- ▶ Conclusion:
 - ▶ It is probably not worthwhile to pursue a muon conversion experiment at HIMB, if Mu2e and COMET reach their targeted sensitivity
 - ▶ However, if a signal is observed by either Mu2e or COMET at a relatively high branching ratio such a measurement at PSI might become again attractive

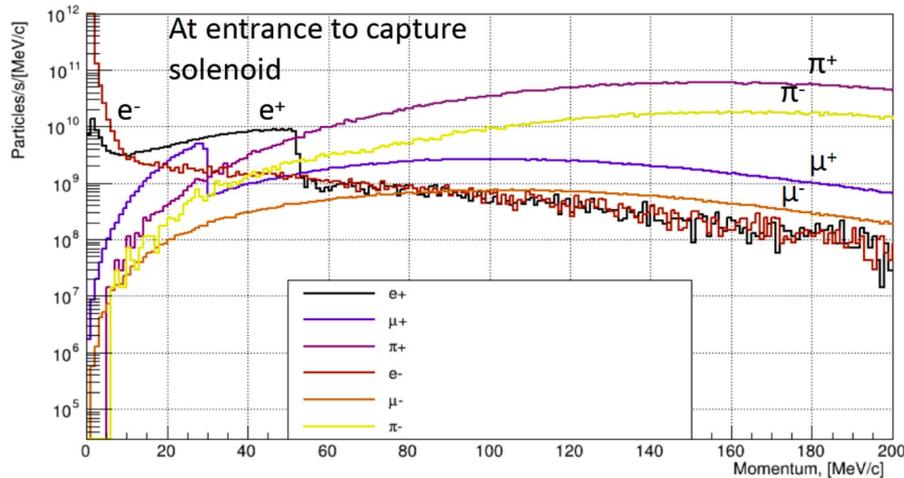
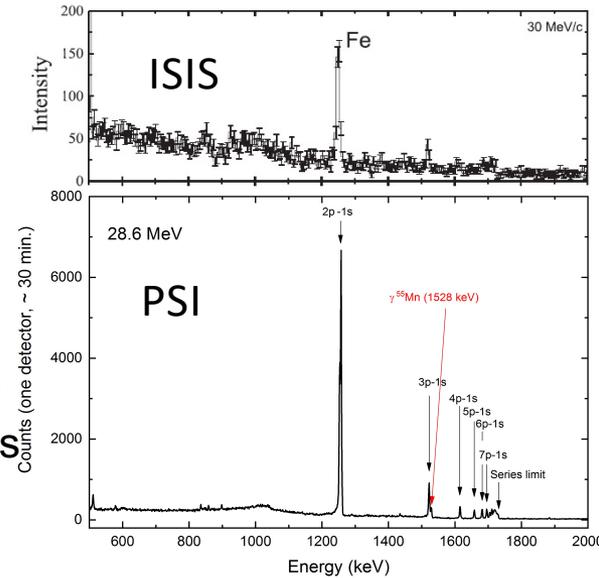
Negative Muons Elemental Analysis

Negative muon



PSI best place in the world for MIXE
CW-beam, high rates
ISIS (pulsed) vs PSI (cw)

Ideal technique to measure very rare / valuable samples



HIMBs offers

- Higher μ^- rates for reduced DAQ times
- Moreover, the capture solenoid could allow a **parasitic additional μ^- beamline**



Parasitic μ -beam

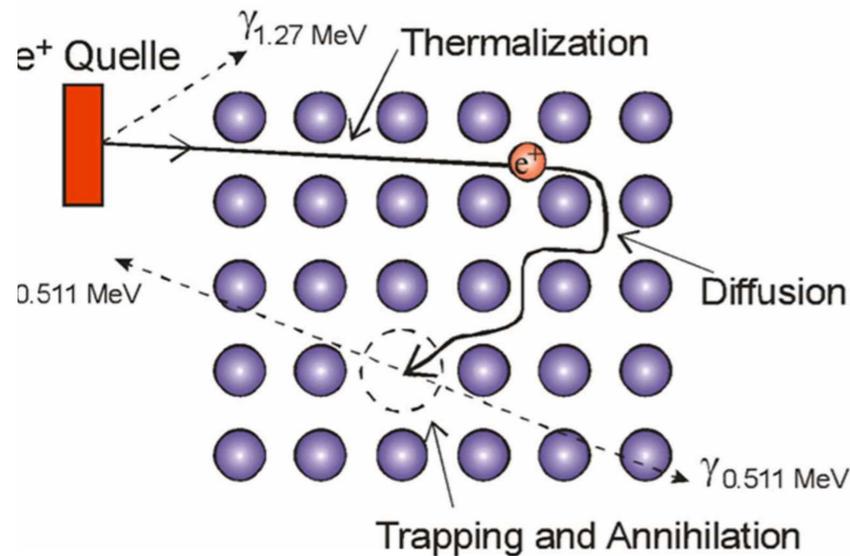
4	1	Stk	RT81.21.0020	ZUS. STRAHLINIE MUH3 BERG
3	1	Stk	RT81.21.0010	ZUS. STRAHLINIE MUH2 BERG
2	1	Stk	06.0.679	3D STRAHLINIE HIPA LAYOUT
1	1	Stk	01.0.1668	3D BODEN / ABSCHIRMUNG

Pos.	Menü	Einheit	Sachnummer	Benennung / Merkmale
Freidnummer:				
a		e		Gezeichnet 29.10.2020 RAUBER
b		f		Geprüft
c		g		
d		h		

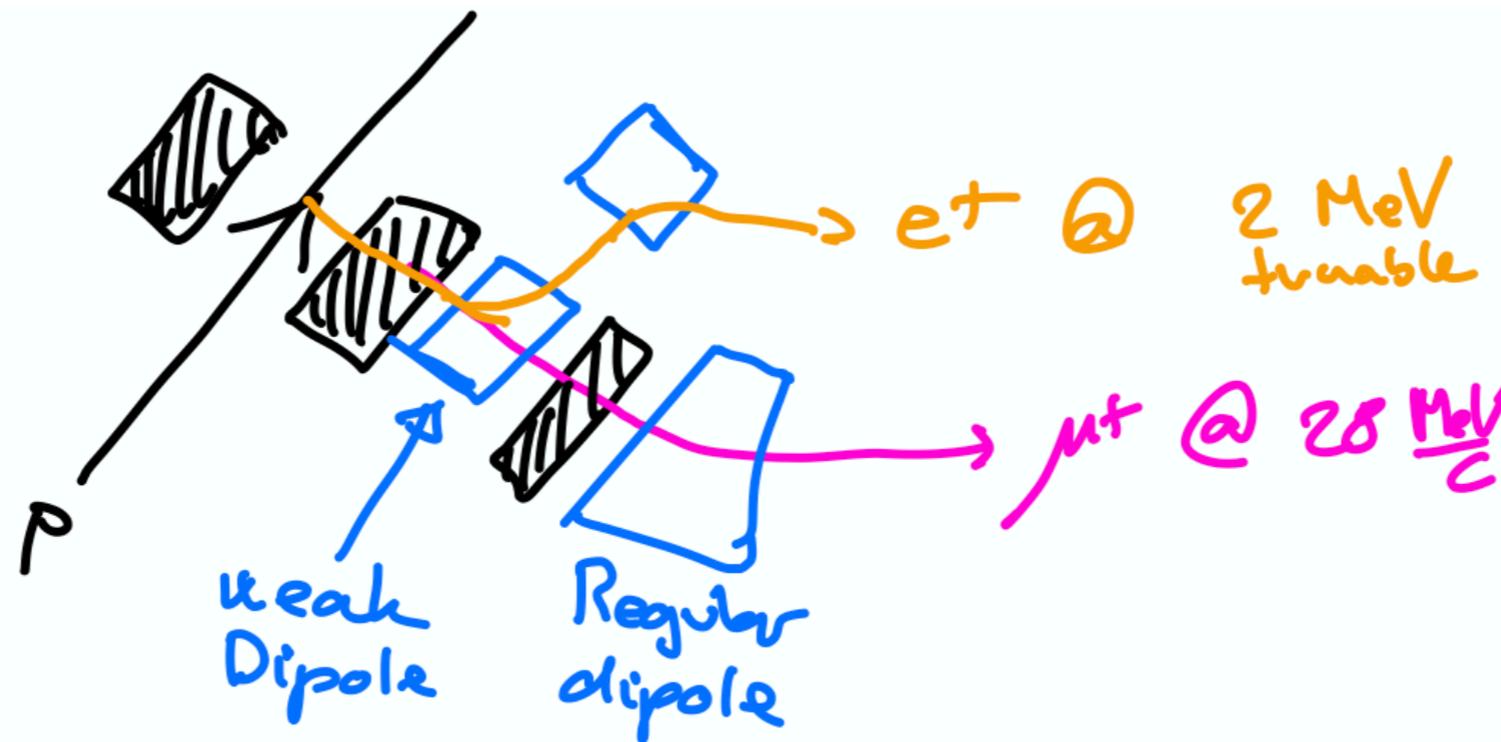
Anlage MUH2+MUH3 BERG	Ersetzt durch	
Baugruppe	Ersetzt fuer	
HIMB KONZEPT BERG	Stueckl. Nr.	
	Zusammenst. Nr.	

LAY. HIMB TARGET VERTEX 42331	0-RT81.0.0001
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Positrons for material science?

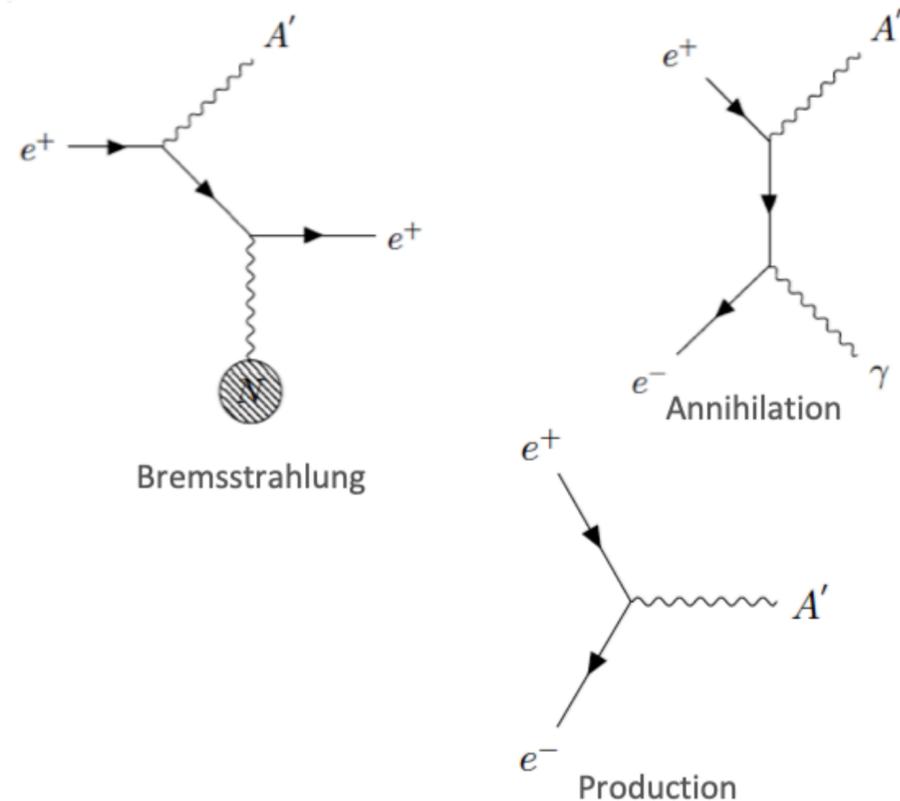


- ▶ The positron is a very sensitive probe of nm-sized vacancies
- ▶ Can PSI have a competitive MeV-energy beam that can operate parasitically to the muon beamline?
- ▶ Very very very probably NOT
- ▶ But would be interesting to keep in mind when performing the HiMB simulations

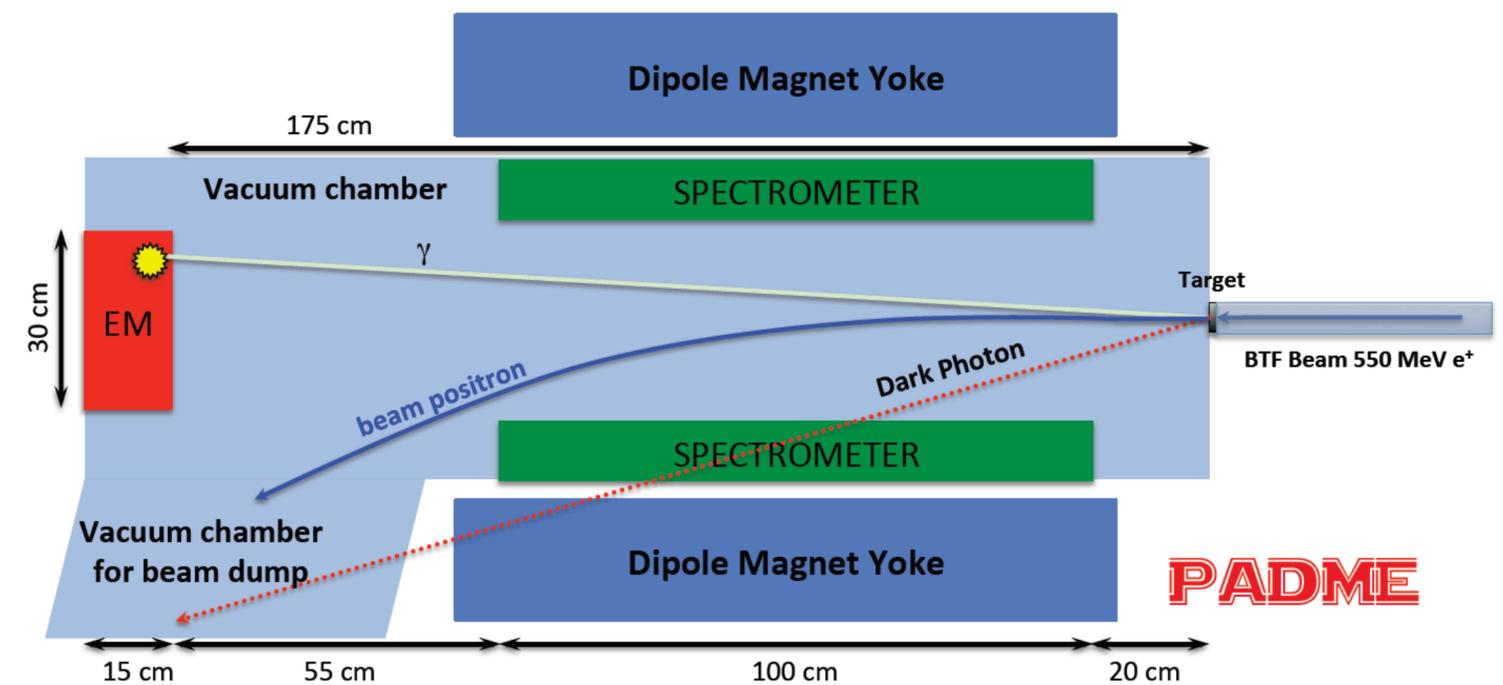


Exploiting positron beam@HiMB for Dark Matter searches

- Low energy and intense positron beams could be attractive for Dark Matter searches to complement the low mass dark matter region (mainly studied via astrophysical searches) to address both invisible ($A' \rightarrow X X$) and visible (in this case $A' \rightarrow e^+/e^-$ pair) Dark Matter decay signatures
- Experimental setup based on established technologies looks able to address the experimental requirements
- Dedicated MC simulations (with emphasis also on the beam side) must be performed to assess the final achievable sensitivities with dedicated/complemented experimental setup

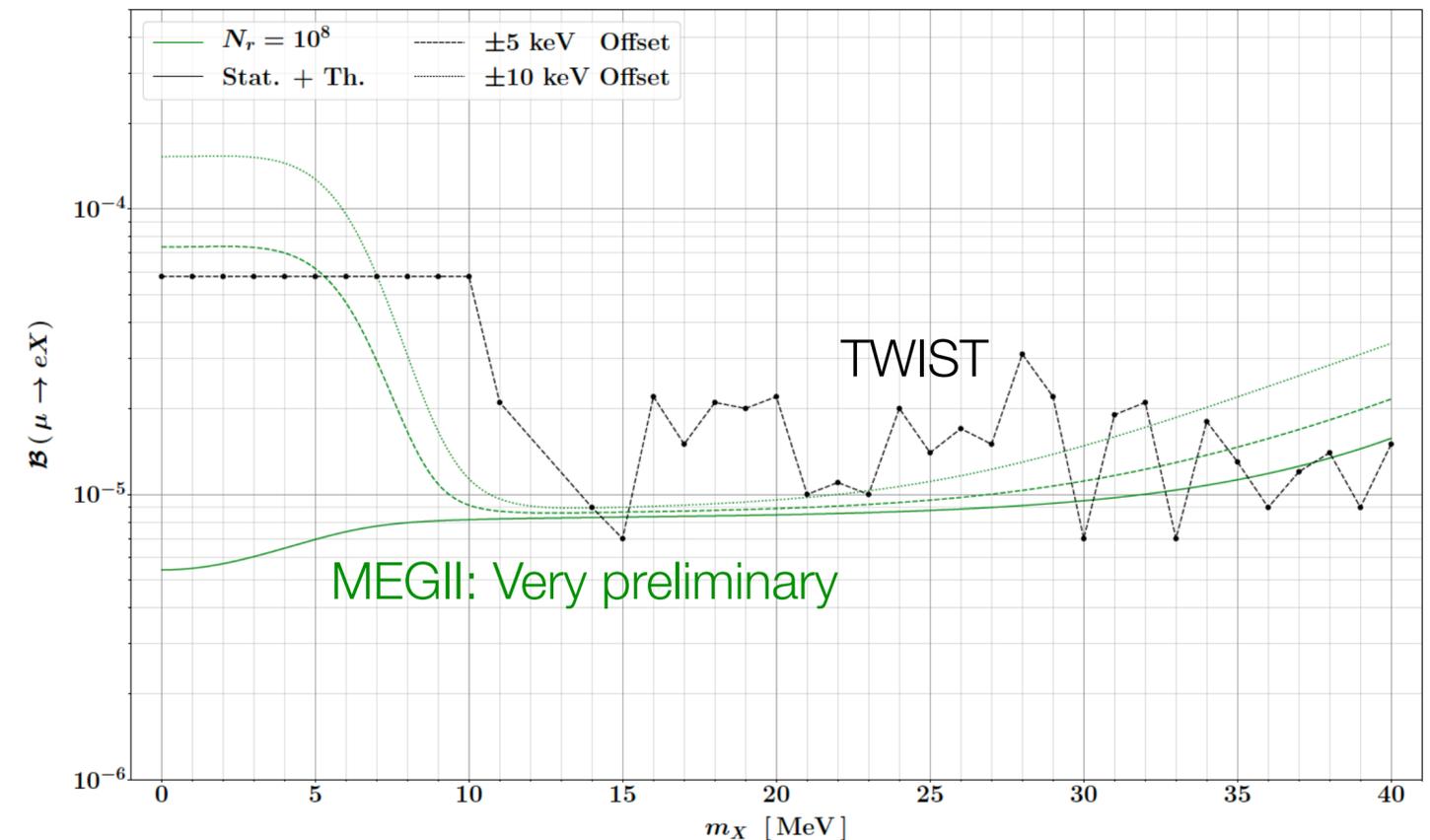
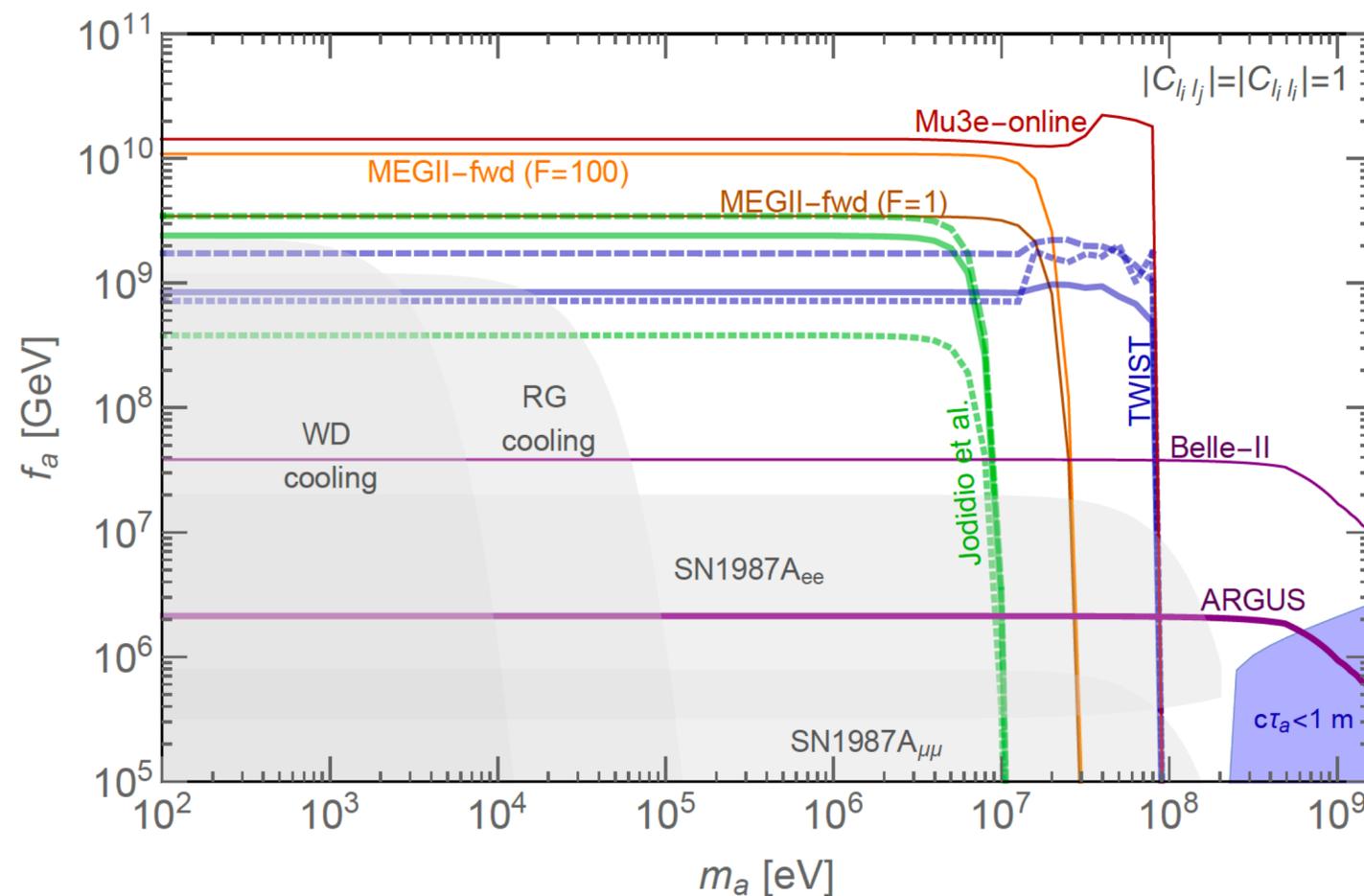


Inspired by the PADME experiment



ALP searches @HiMB

- The search for ALP clearly improves with statistics (and then will benefit from HiMB)
- In order to really take advantage for the high muon beam rates not only more performing detectors are requested (and in principle these will come with the future generation of MEG/Mu3e experiments) but also the systematics must be kept under control increasing the difficulties of such a kind of searches and pointing towards “dedicated” experimental setup
- Precise theoretical inputs of the “SM” backgrounds (i.e. standard muon decay at NNLO+LL) are crucial for the incoming sensitivities
- Detailed MC simulations are foreseen with the incoming MEGII and Mu3e apparatuses as well as with some dedicated/complementary setup



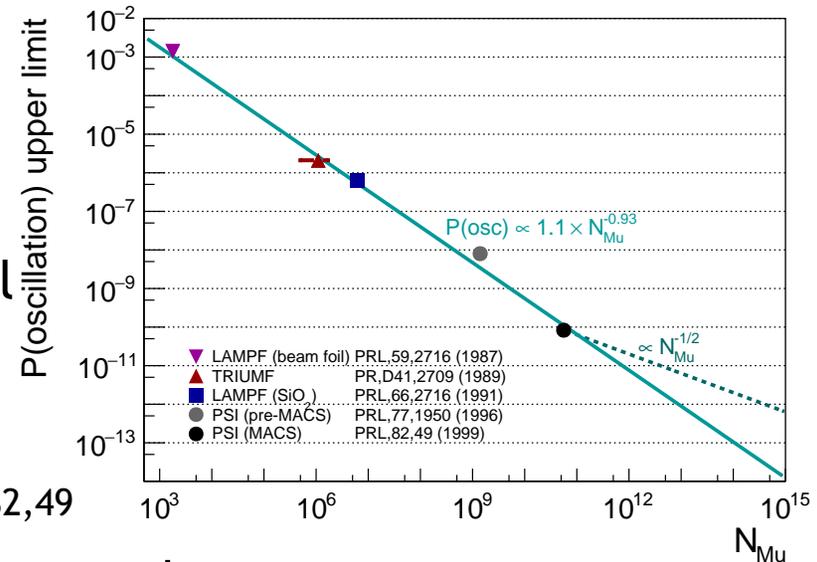
MuMu oscillations?

- Best result: MACS/PSI (1998)

- ▷ $\pi E5$ with $8 \times 10^6 \mu^+ / s$ at $p = 26 \text{ MeV}$
- ▷ $280 \mu\text{m}$ scintillator + $270 \mu\text{m}$ Mylar degr.
- ▷ SiO_2 powder target 8 mg/cm^2 on $25 \mu\text{m}$ Al replaced twice/week

- ▷ $N_{\text{Mu}} = 5.6 \times 10^{10}$ in vacuo

→ $\mathcal{P}_{\text{Mu}\overline{\text{Mu}}}(0.1 \text{ T}) < 8.3 \times 10^{-11}$ (90% CL) PRL,82,49
with 1 event observed/1.7 background expected



- Which HiMB setup?

- ▷ what minimum efficiency \times acceptance ($\equiv \epsilon_{\text{total}}$) to improve MACS result by one order of magnitude in a week of beam-time?

- HiMB-3/5: $10^{10} \mu^+ / s$ at $p_{\mu^+} = 28 \text{ MeV}$

$$\epsilon_{\text{total}} > 3 \times 10^{-4}$$

- HiMB-cool: $10^7 \mu^+ / s$ at $p_{\mu^+} = 10 \text{ keV}$

$$\epsilon_{\text{total}} > 3 \times 10^{-1} \text{ (unrealistic!)}$$

→ decay length/volume (enough time for oscillations)?

→ background level (Bhabha/rare decays/punch-through)?

⇒ To be studied further with GEANT4 setup

Summary: with $10^{10} \mu/\text{sec}$, look for NP in “SM” μ decays?

$$\mu \rightarrow e(+invisible) \sim 10^{10}/\text{sec}, \mu \rightarrow e\gamma(+inv.) \sim 600/\text{sec}, \mu \rightarrow 3e(+inv) \sim 3 \cdot 10^{-5}/\text{sec}$$

theorist: opportunity! improve Michel, +look for feebly cpled ALPs, Z' , ν_s , etc !!

exptalist: how to cut all that background...

the theory wishlist

- $\mu \rightarrow eX$ at MEG forward (BR $\lesssim 2 - 60 \times 10^{-6}$ TRIUMF)
- $\mu \rightarrow e\nu\bar{\nu}$ improved Michel parameters (heavy NP, $\mu \rightarrow e\nu\bar{\nu} + X$ for cosmo?)
- $\mu \rightarrow 3e +$ invisible (new boson cpled to e) (see eg, Phys Beyond Coll. 1901.09966)
- $\mu \rightarrow e\gamma +$ invisible : SVP look her too ... (BR $\lesssim 10^{-9}$ CrystalBox)
- ? everything :)

maybe theorists could do some work too:

- where to measure what, to have best sensitivity to which scenario?

(generalised Michel-like parameters for all SM-like decays?)

- (Invisibles who talk to ν_s (for H_0 tension): what interactions with e , μ ?)

random thought: $\mu A \rightarrow eA$ will measure the endpt spectrum of their electron; maybe look for massive NP in μ^- decays, and “massless” in μ^+ decays?