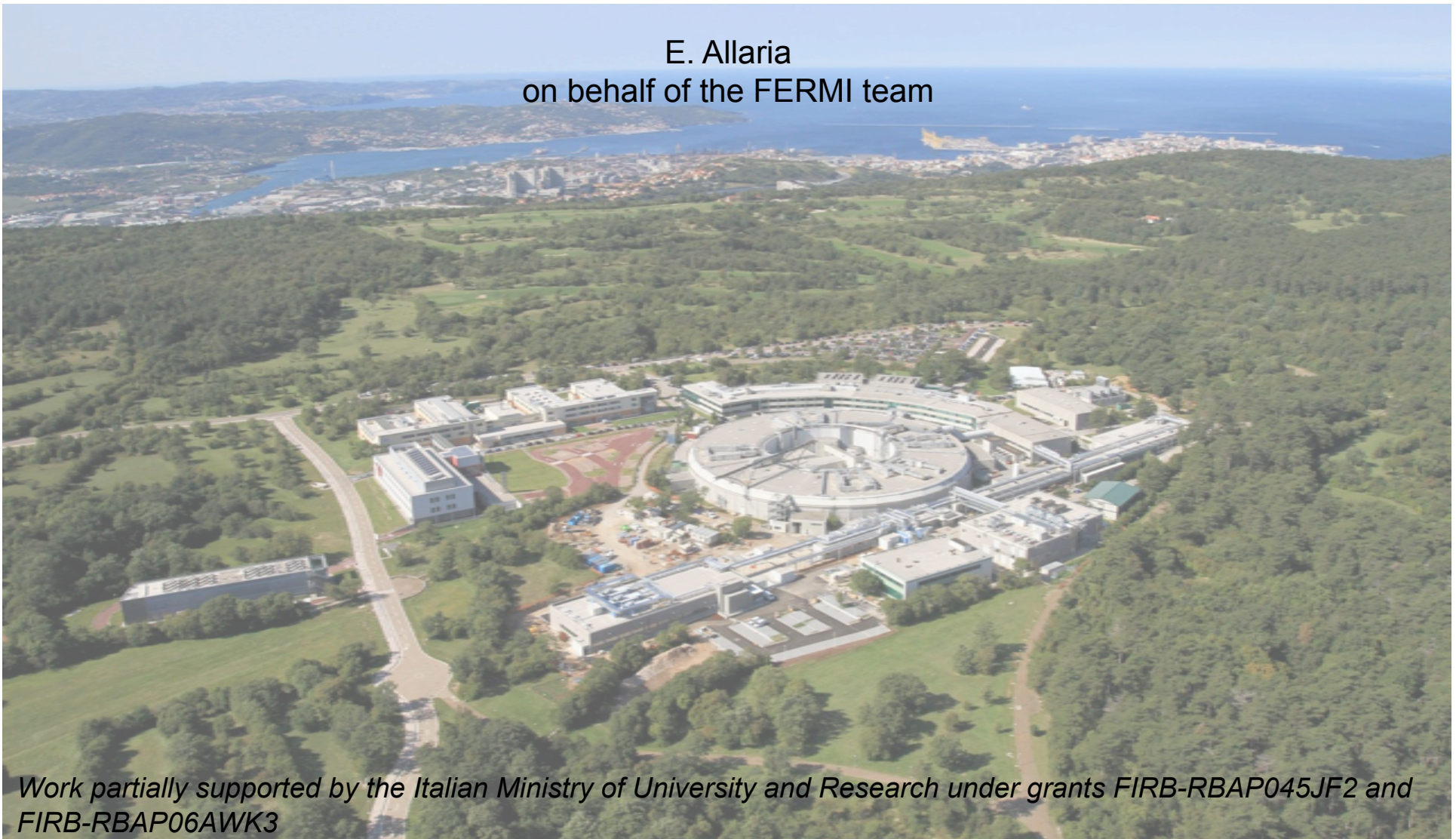


FERMI, a seeded Free Electron Laser source for experiments

E. Allaria
on behalf of the FERMI team



Work partially supported by the Italian Ministry of University and Research under grants FIRB-RBAP045JF2 and FIRB-RBAP06AWK3

- Elettra and the FERMI FEL project
 - FERMI parameters
- FEL mechanisms
 - Self Amplified Spontaneous Emission
 - High Gain Harmonic Generation
- Commissioning activities
 - FEL commissioning and results
- FEL experimental results at FERMI
 - Coherence properties
 - Spectral characterization
- Future plans

SINCROTRONE TRIESTE is a nonprofit shareholder company of national interest, established in 1987 to construct and manage synchrotron light sources as international facilities.

ELETTRA Synchrotron Light Source:
up to 2.4 GeV, top-up mode,
768 proposals from 39 countries in 2010

FERMI@Elettra FEL:

100 – 4 nm HGHG, fully funded

☐ Sponsors:

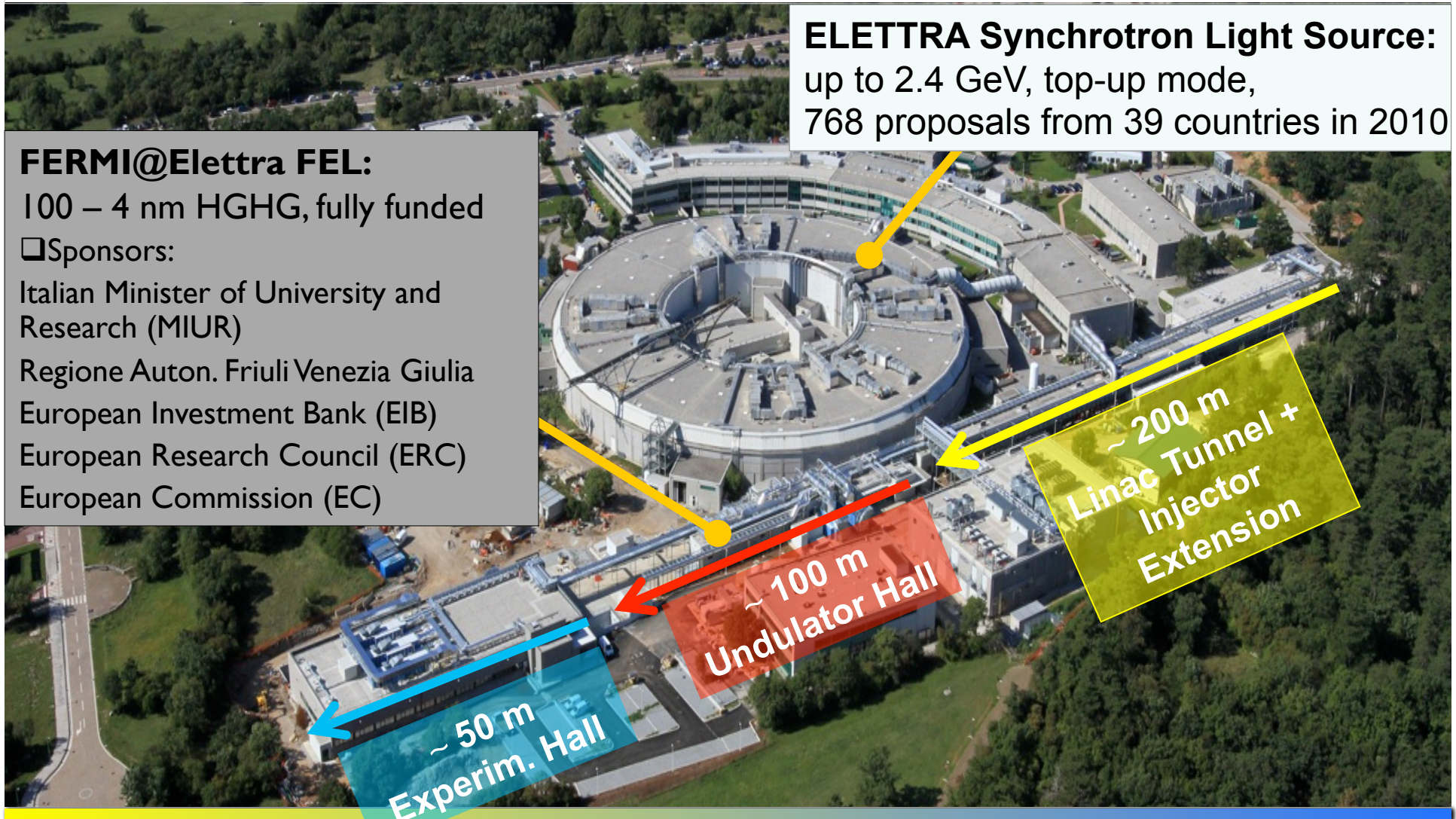
Italian Minister of University and Research (MIUR)

Regione Auton. Friuli Venezia Giulia

European Investment Bank (EIB)

European Research Council (ERC)

European Commission (EC)



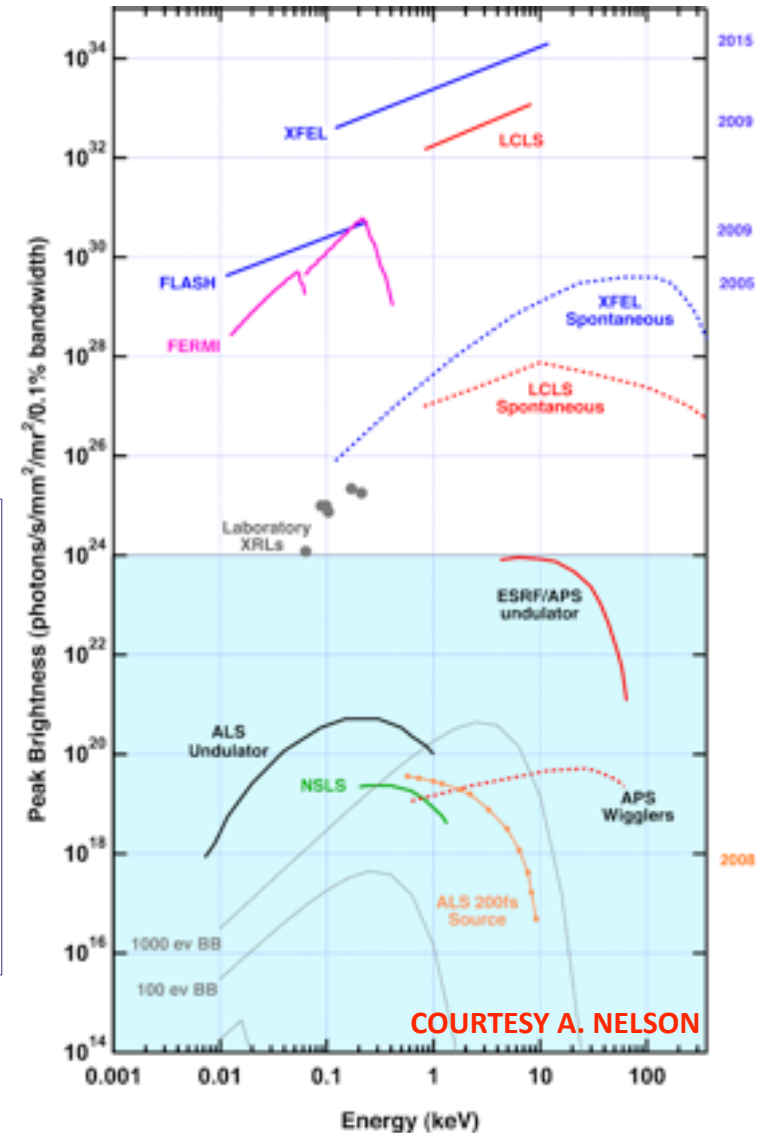
FERMI, a Seeded Free Electron Laser Source for Experiments

FERMI@Elettra single-pass FEL user-facility.

Two separate FEL amplifiers will cover the spectral range from 100 nm (12eV) to 4 nm (320 eV).

The two FEL's will provide users with ~100fs photon pulses with unique characteristics.

- | | |
|----------------------------------------------------------|---------------------------------|
| <input type="checkbox"/> <u>high peak power</u> | 0.3 – GW's range |
| <input type="checkbox"/> <u>short temporal structure</u> | sub-ps to 10 fs time scale |
| <input type="checkbox"/> <u>tunable wavelength</u> | APPLE II-type undulators |
| <input type="checkbox"/> <u>variable polarization</u> | horizontal/circular/vertical |
| <input type="checkbox"/> <u>seeded harmonic cascade</u> | longitud. and transv. coherence |



E. Parmigiani (Head of Scientific Programs)

▶ **Low Density Matter** (*coord. C. Callegari*):

- ▶ structure of nano-clusters *brightness*
- ▶ high resolution spectroscopy *narrow bw, λ -tunability*
- ▶ magnetism in nano-particles *circular polarization*
- ▶ catalysis in nano-materials *fs pulse and stability*

▶ **Elastic and Inelastic Scattering** (*coord. C. Masciovecchio*):

- ▶ Transient Grating Spectroscopy (collective dynamics at the nano-scale) *bw Fourier Transform Limit*
- ▶ Pump & Probe Spectroscopy (meta-stable states of matter) *brightness, λ -tunability*

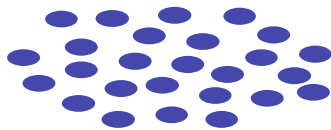
▶ **Diffraction and Projection Imaging** (*coord. M. Kiskinova*): Single-shot & Resonant Transverse Coherent Diffraction Imaging

- ▶ morphology and internal structure at the nm scale
- ▶ chemical and magnetic imaging *brightness*

High brightness light sources in VUV - X-ray spectral range defined as 4th Generation Light sources are based on the high gain Free Electron Laser mechanism.

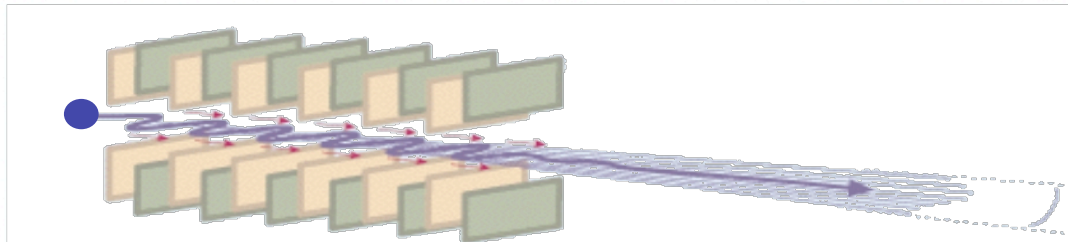
Free-Electron Lasers exploit the spontaneous and/or induced emission of a relativistic electron beam “guided” by the periodic and static magnetic field generated by an undulator

- 1) Relativistic **electron beam**
 - Energy (γ)
 - Current (I)
 - Emittance (ϵ)
 - Energy spread ($\delta\gamma$)
 - Dimensions (σ)
- 2) **Undulator**



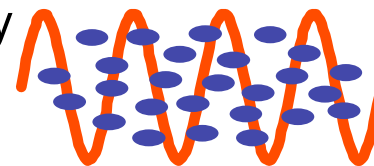
Resonance condition

$$\lambda = \frac{\lambda_w}{2\gamma^2} (1 + K^2)$$



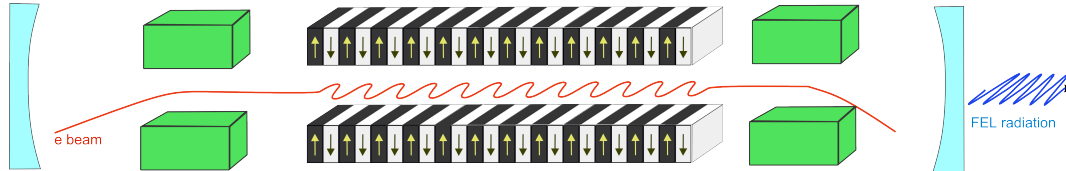
Magnetic period (λ_w)
Magnetic strength (K)
Undulator length (L)

- 3) **Electromagnetic field** co-propagating with the electron beam and **getting amplified** to the detriment of electrons' kinetic energy



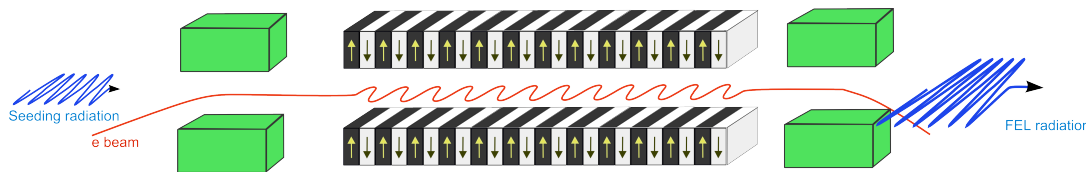
Wavelength (λ)
Power (P)

Oscillator FEL



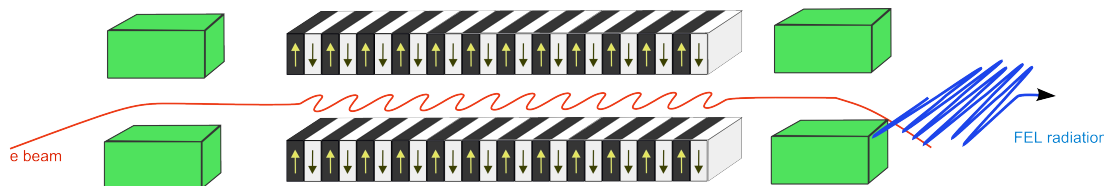
The tunability toward short wavelengths is limited by the availability of high quality mirrors.
Very good spectral quality (mirrors act as a filter).

Amplifier FEL



The tunability toward short wavelengths is limited by the availability of the seed wavelength.
Spectral quality limited by the quality of the electron beam.

SASE FEL



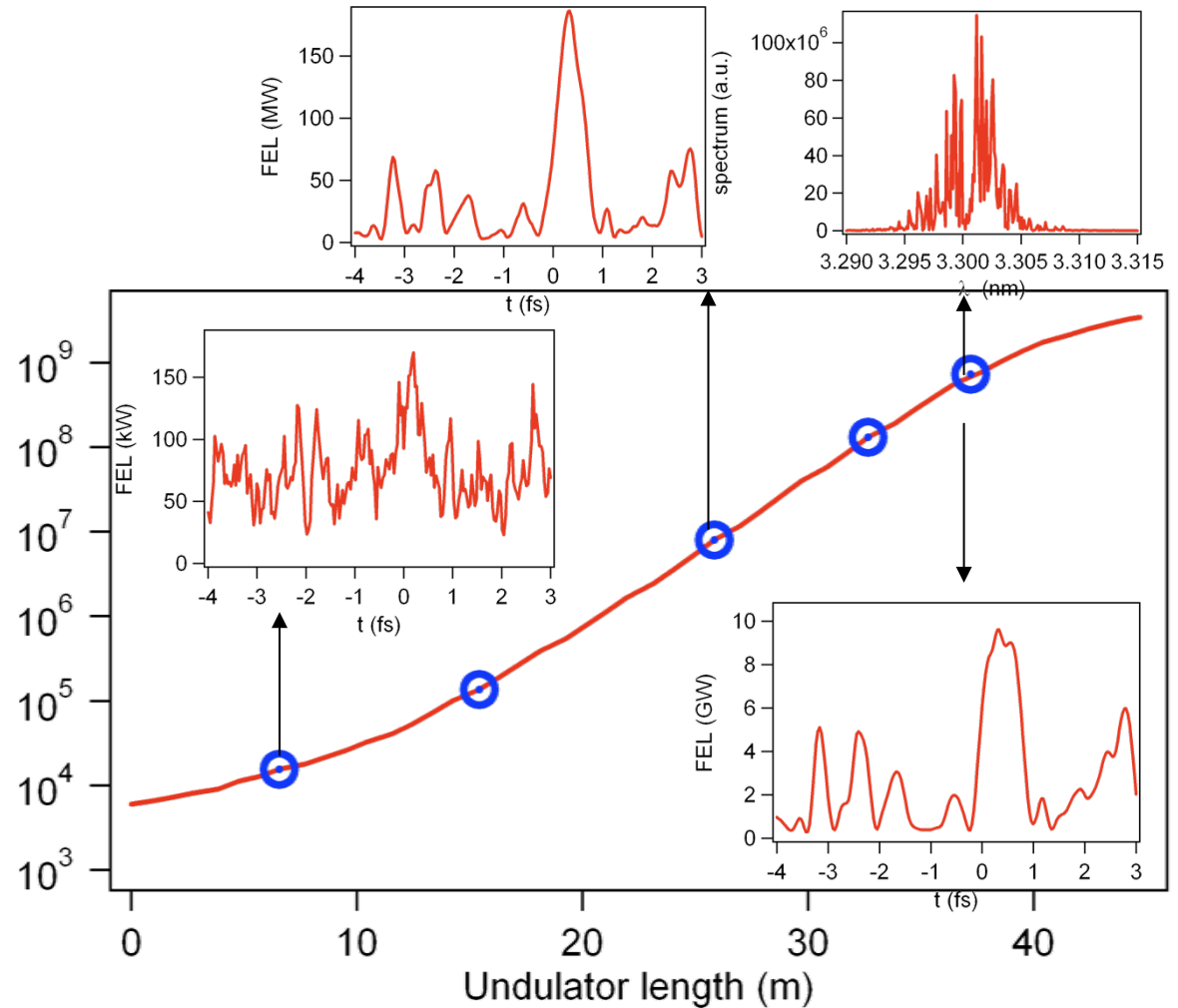
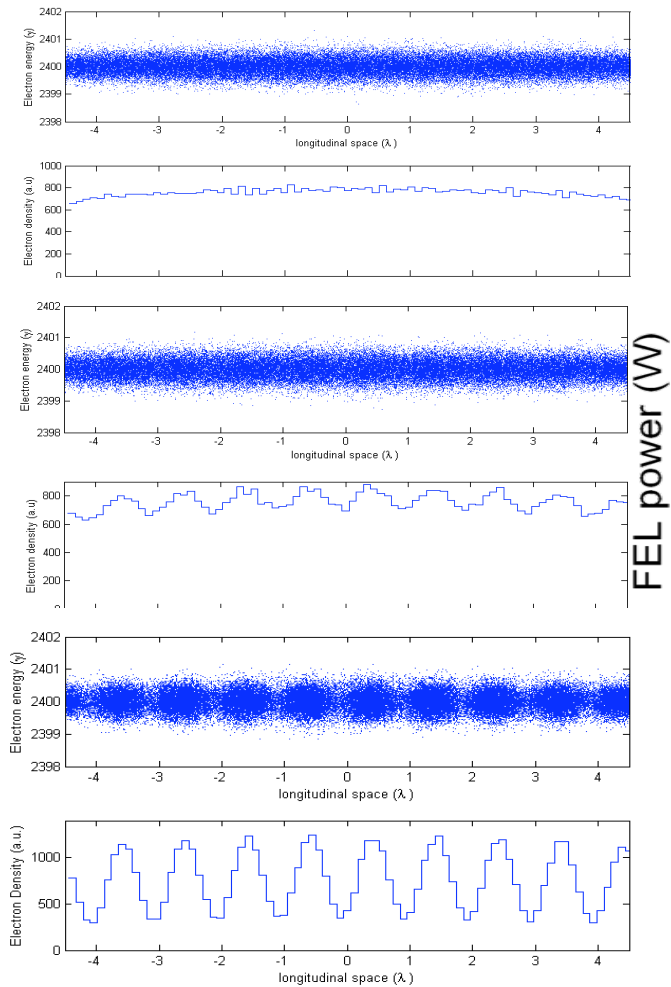
Potentially completely tunable, FEL wavelengths only depend on the resonance condition.
Tighter requirements for the electron beam parameters
Spectral properties are affected by the random startup



Self-Amplified Spontaneous Emission (SASE) FERMI @elettra



Electron beam phase space



- SASE FELs are the most suitable for X-ray FELs (FLASH, LCLS, SACLA, ...)
- As it has been shown by the LCLS and SACLA, SASE can be successfully operated at wavelength below 1nm.

However:

- Desired power is reached only with long undulators
- Spectral properties are affected by the startup from noise

For the soft X-ray spectral range other facilities are looking for a seeded configuration.

The FERMI@Elettra has considered from the beginning the seeded option in order to produce high quality FEL pulses.

The seeding is also under investigation now at FLASH and recently studies started for a seeded FEL configuration at LCLS.

- A “seed” laser controls the distribution of electrons within a bunch:
- Very high peak flux and brightness (comparable to SASE FELs)
 - Temporal coherence of the FEL output pulse
 - Control of the time duration, wavelength and bandwidth of the coherent FEL pulse
 - Close to transform-limit pulse provides excellent resolving power without monochromators
 - Perfect synchronization of the FEL pulse to the seed laser
 - Reduction in undulator length needed to achieve saturation.

The problem with seeding is that there are not sources available for direct seeding in the very short wavelength range (few nanometers).

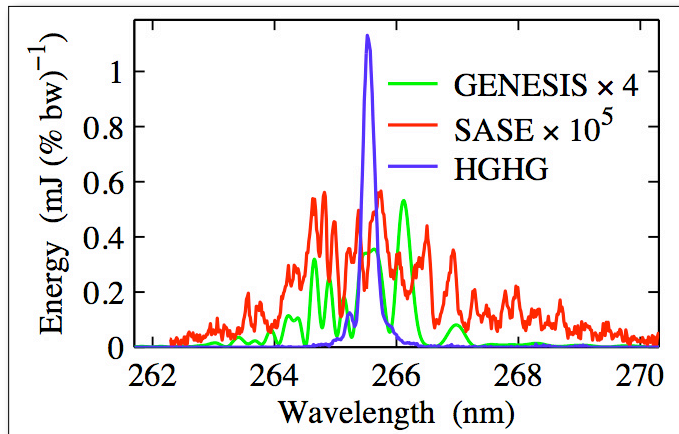
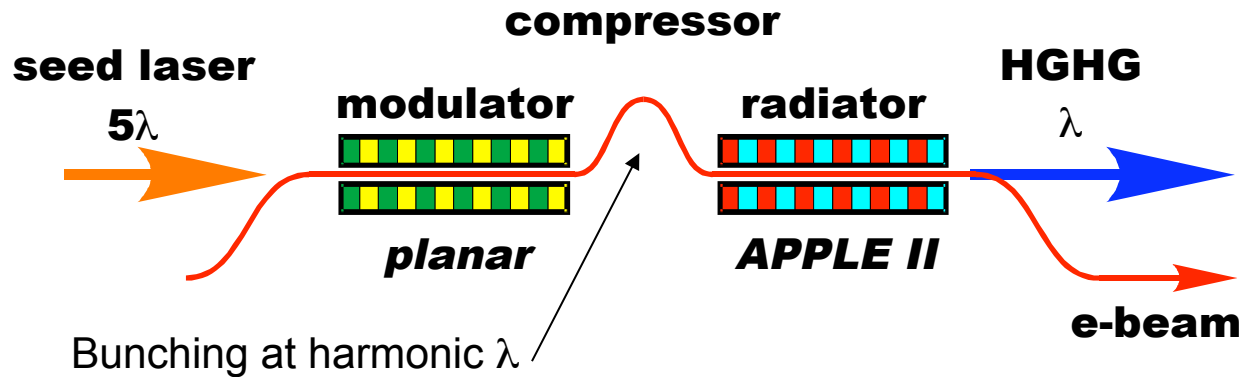
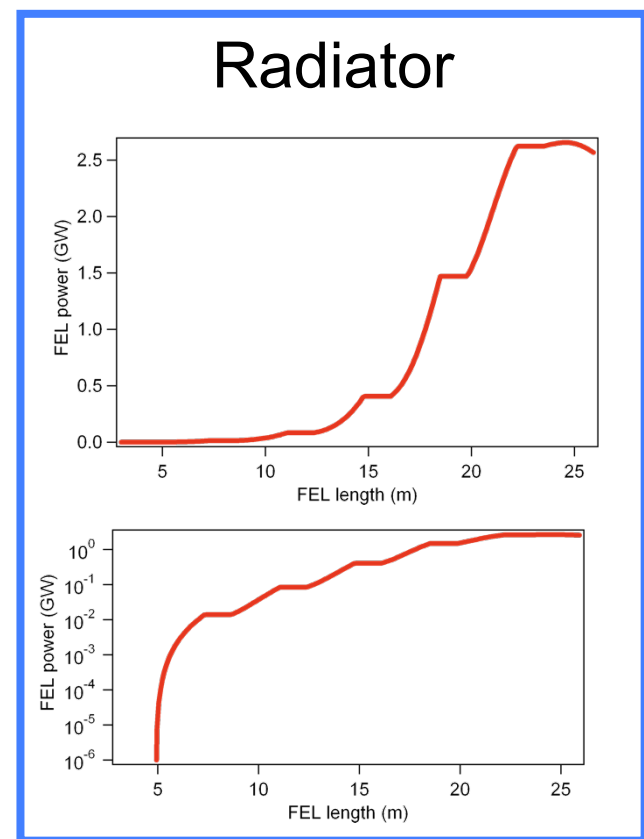
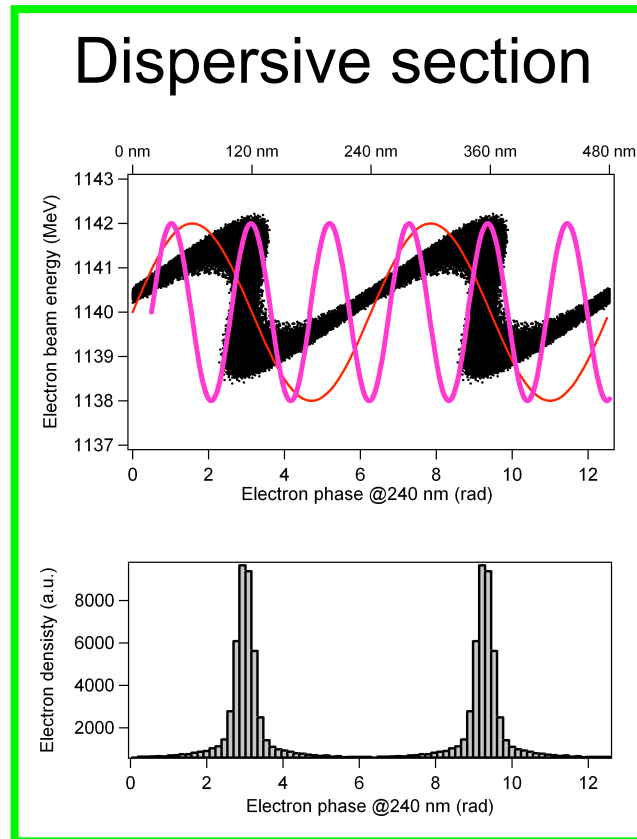
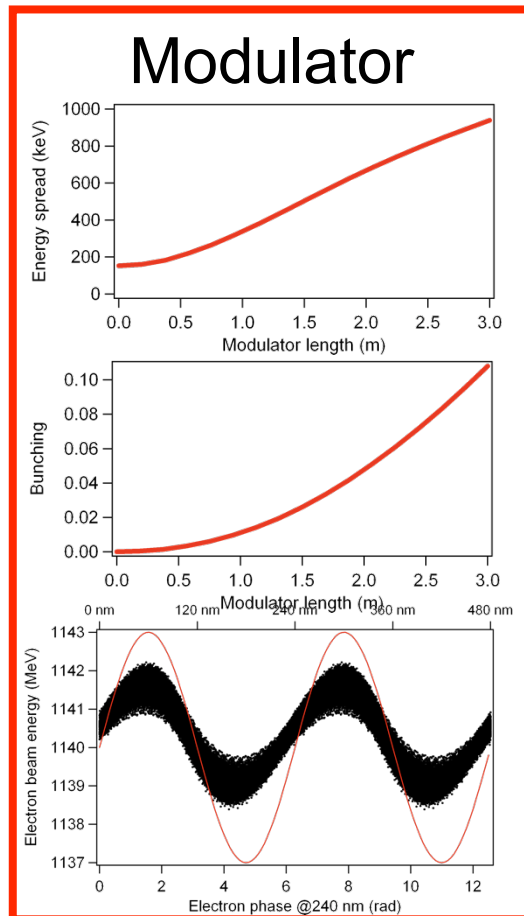
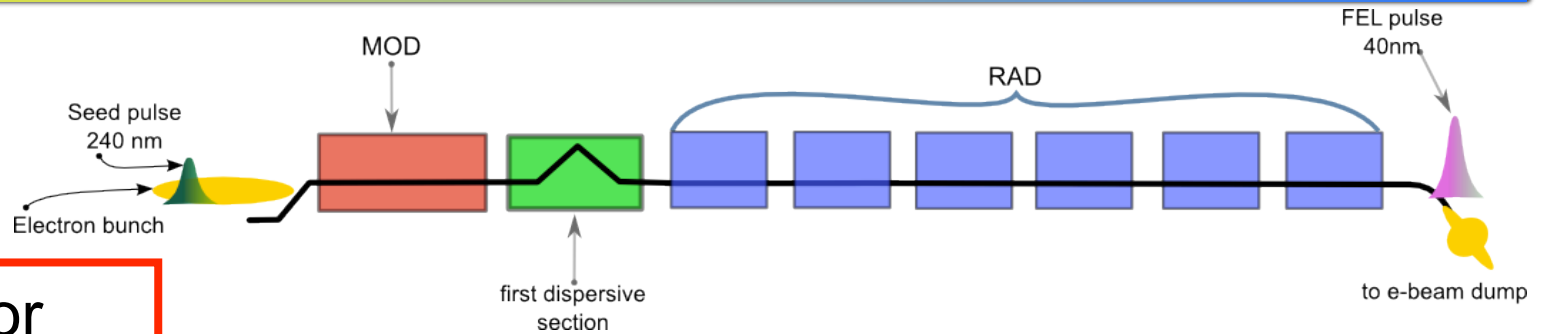


FIG. 4: Single shot HGHG spectrum for 30 MW seed (blue), single shot SASE spectrum measured by blocking the seed laser (red) and simulation the SASE spectrum after 20 m of NISUS structure (green). The average spacing between spikes in the SASE spectrum is used to estimate the pulse length.

Compared to SASE devices, generally more compact and nearly full temporally coherence output; many spectral parameters more easily controlled (e.g., pulse length, chirp).

L.H. Yu et al.
Phys. Rev. Lett. 91, 074801 (2003)

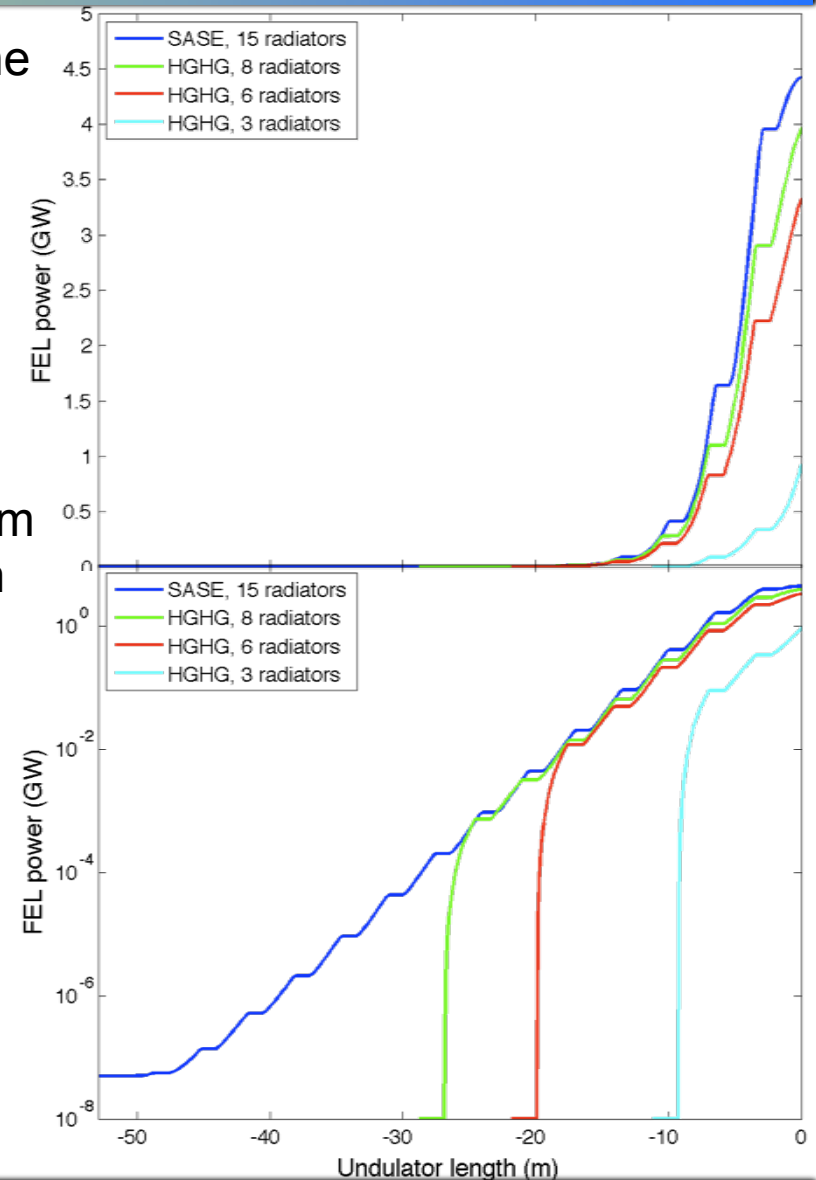


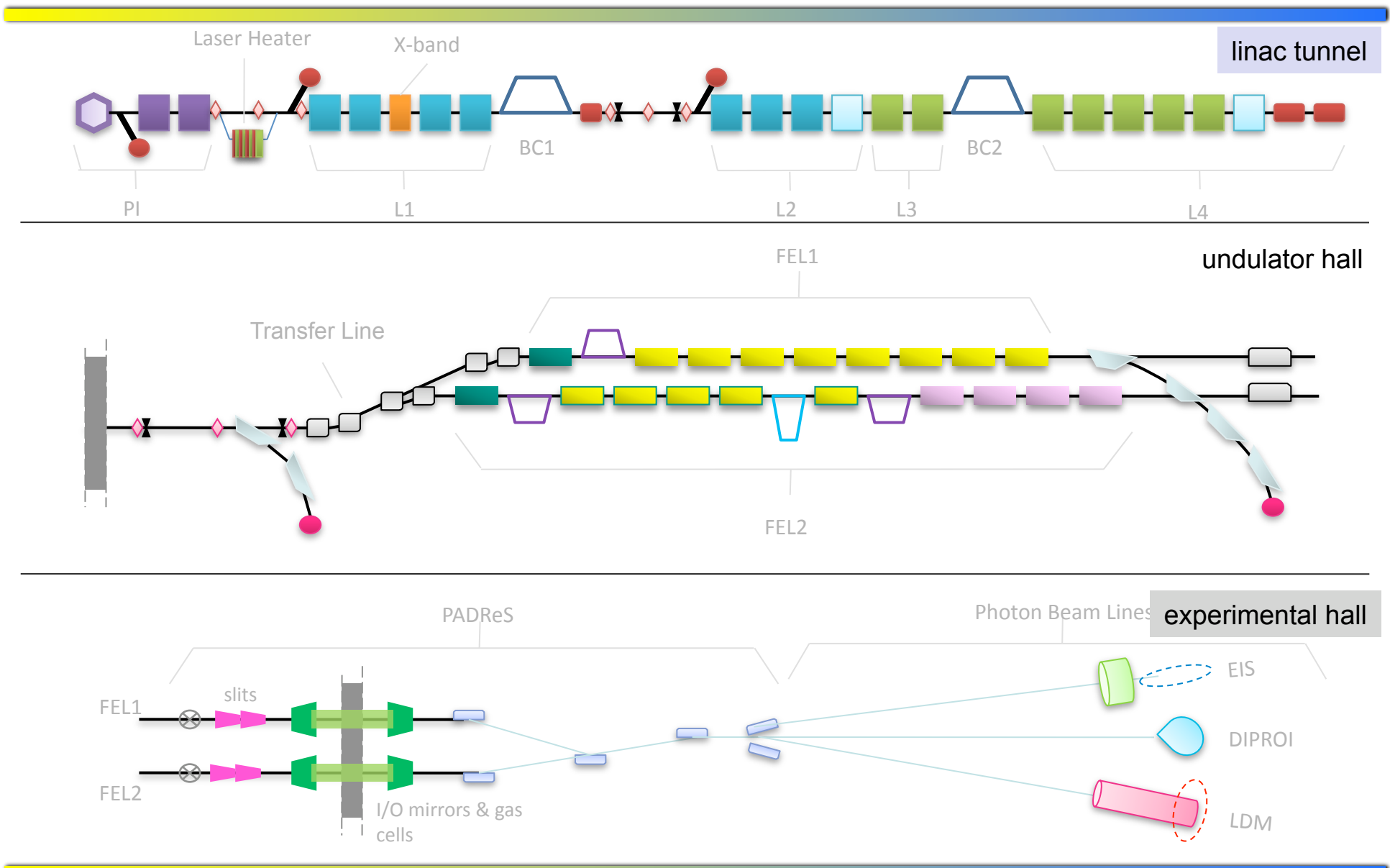
The HGHG FEL can be configured depending on the desired FEL performance and available undulator length.

An output power comparable to the one achievable from the SASE FEL can be produced using a radiator which is about half the length needed for SASE. In the case of a very short radiator only coherent emission from a prebunched electron beam is generated and no exponential gain is expected in the final radiator.

The maximum power from the FEL only slightly decreases with the undulator length.

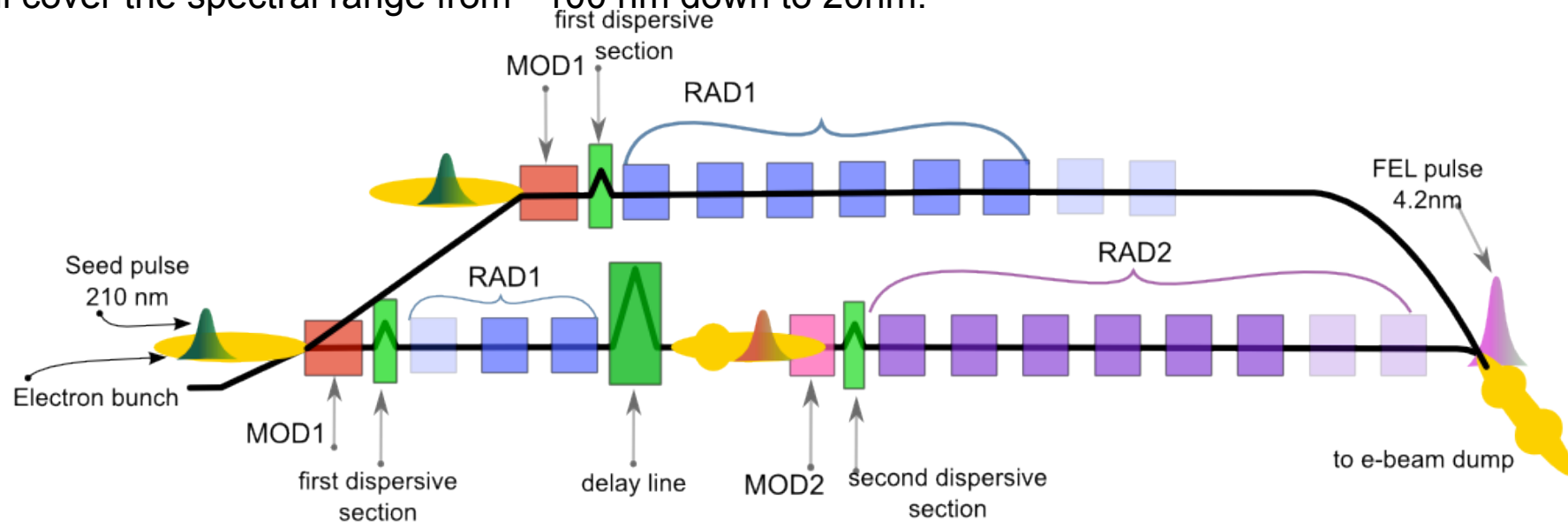
Reported simulation results are for a FERMI FEL with the parameters similar to FERMI:
 Energy: 1.2GeV, peak current: 1kA, emit: 1 mm mrad, energy spread: 150keV, wavelength: 20nm,





FERMI's two FELs will cover different spectral regions.

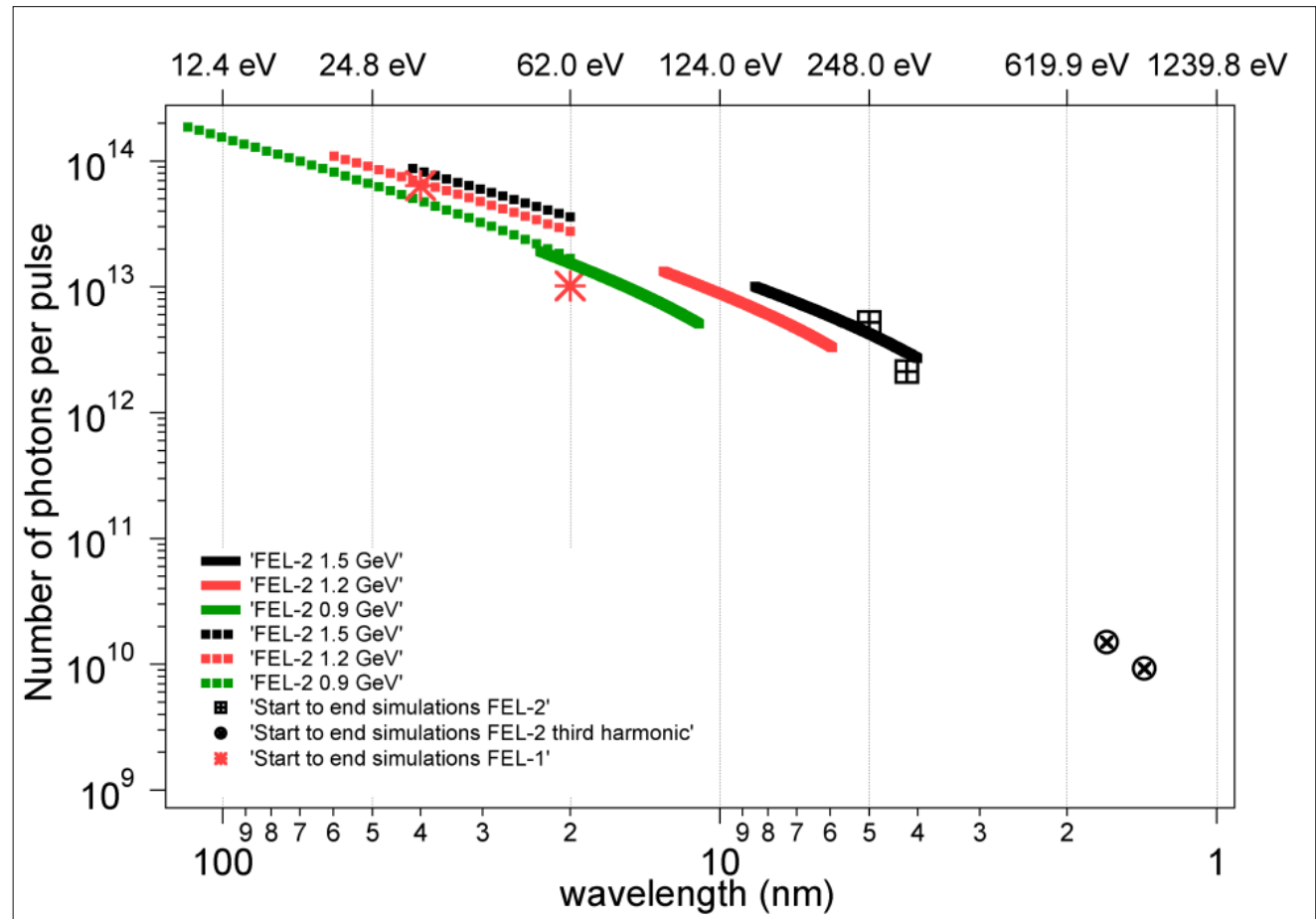
FEL-1, based on a single stage high gain harmonic generations scheme initialized by a UV laser will cover the spectral range from ~ 100 nm down to 20nm.



FEL-2, in order to be able to reach the wavelength range from 20 to ~ 4 nm starting from a seed laser in the UV, will be based on a double cascade of high gain harmonic generation. The nominal layout uses a magnetic electron delay line in order to improve the FEL performance by using the fresh bunch technique. Other FEL configurations are also possible in the future (e.g. EEHG).

Using the expected values for the electron beam parameters for FERMI we can predict the FEL performance.

| Parameter | Value | Units |
|---------------|---------|---------|
| Energy | 0.9-1.5 | GeV |
| Peak current | >750 | A |
| Emittance | <1.0 | mm mrad |
| Energy spread | <150 | keV |



Lines predicted using M.Xie formulae for expected FERMI parameters assuming 40fs pulse length
 Points Ginger and Genesis simulations for S2E files

December 2010 (250pC):

| | |
|-------------------------------------------|-----------------------------------|
| Compressed e-beam 43nm with photodiode: | ~6 nJ; ~1*10 ⁹ |
| Compressed e-beam 43nm with spectrometer: | ~3 nJ; ~5*10 ⁸ |
| Compressed e-beam down to 17 nm: | clear evidence of coherent signal |

March 2011 (250pC):

| | |
|-----------------------------------------------------------------|------------------------------------|
| Uncompressed e-beam 65 nm with the DESY gas detector: | ~0.3μJ, ~1*10 ¹¹ , ~2MW |
| Compressed e-beam 65 nm with the calibrated FERMI gas detector: | ~3 μJ, ~1*10 ¹² , ~20MW |

April 2011 (350pC):

| | |
|------------------------------------------------------|-------------------------------------|
| Compressed e-beam 65 nm with spectrometer (average): | ~2 μJ, ~6*10 ¹¹ , ~12MW |
| Compressed e-beam 43 nm with spectrometer (average): | ~5 μJ, ~1*10 ¹² , ~30MW |
| Down to ~24 nm | ~0.3 μJ, ~4*10 ¹⁰ , ~2MW |

Experimental stations:

| | |
|------------------------------------------------------------|------------------------------------|
| <i>Timex 65 nm with Al filter (March 25th):</i> | ~135nJ (peak), 70-80nJ (average). |
| <i>LDM 65 and 52 nm:</i> | estimated from PADReS measurements |

June-July 2011:

350pC-450pC

| | |
|-------------------------------------------------------|-------------------------------------------|
| Compressed e-beam 52nm with calibrated* photodiode: | ~30μJ, ~1*10 ¹³ [PRELIMINARY] |
| Compressed e-beam 43nm with calibrated* photodiode: | ~45μJ, ~1*10 ¹³ [PRELIMINARY] |
| Compressed e-beam 32.5nm with calibrated* photodiode: | ~100μJ, ~1*10 ¹³ [PRELIMINARY] |
| Compressed e-beam 20nm with calibrated* photodiode: | ~40μJ, ~1*10 ¹² [PRELIMINARY] |

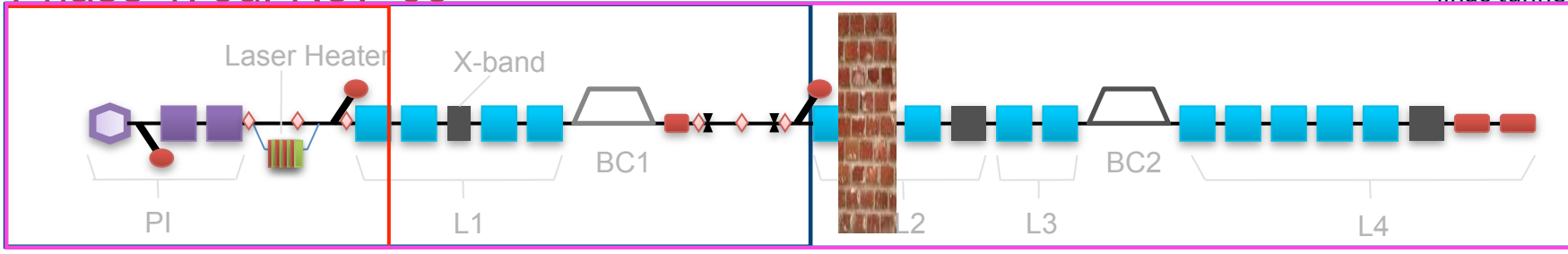
Experimental stations:

| | |
|---------------------|--------------------------------------------------|
| <i>LDM 52 nm:</i> | ~--- |
| <i>Timex 52nm:</i> | ~20μJ at the sample |
| <i>DIPROI 32nm:</i> | ~ 2μJ in the chamber with ~4μJ at the photodiode |

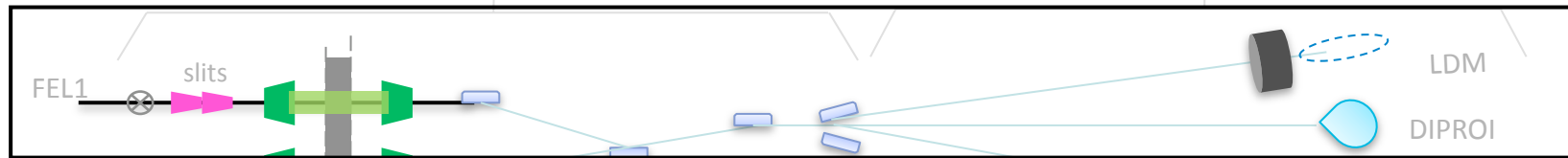
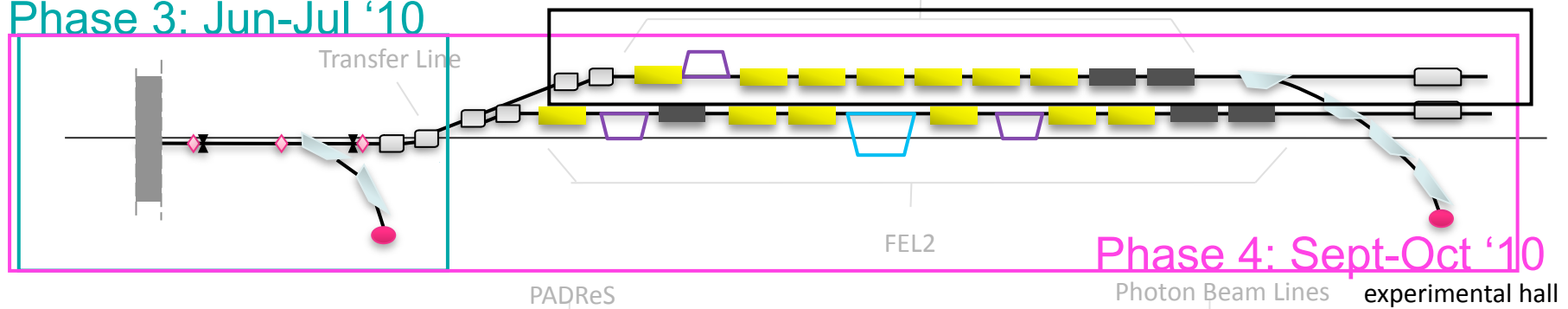
*** Photodiode calibration to be confirmed**

Phase 1: Jul-Nov '09

Phase 2: Feb-Mar '10



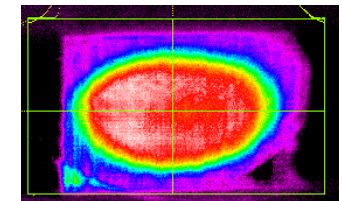
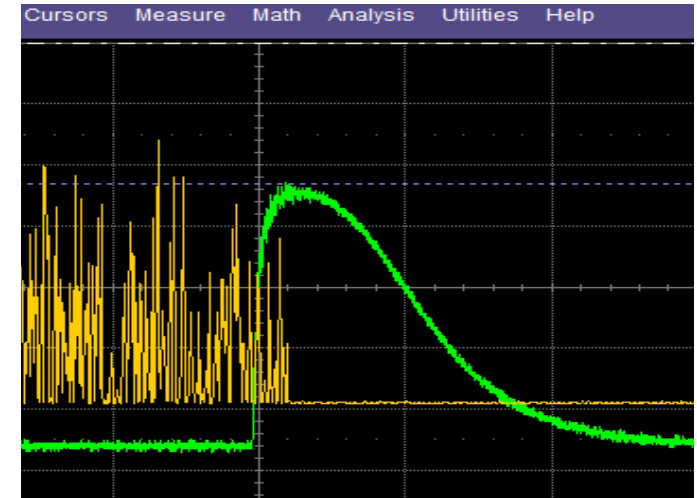
Phase 3: Jun-Jul '10



Phase 5: Nov.-Dec. '10.

Seeded FEL1: single stage HGHG scheme

- ❑ PI Laser, Gun & Injector commissioned: Sept. 2009 – **2.5 months.**
- ❑ Linac & First Bunch Length Compressor: 2010 – **3.5 months.**
- ❑ Transfer Line to Main Beam Dump: 2010 – **1.5 month.**
 - ▶ FELI Undulators installed in 3 weeks.
- ❑ **FIRST LASING** at 43 nm: 2010 – **1.5 months.** (Dec.13, first try!)
- ❑ X-ray Transport & Diagnostics: 2011 – **2 months.**
 - ▶ Tests with Beam Lines *LDM* and *TIMEX*: 3 weeks.
- ❑ Exponential Gain, Stability, Polariz. & Tunability: 2011 – **1.5 months.**
- ❑ 65 – 32.5 nm to *LDM*, *TIMEX* & *DIPROI* Lines: 2011 – **3 weeks.**

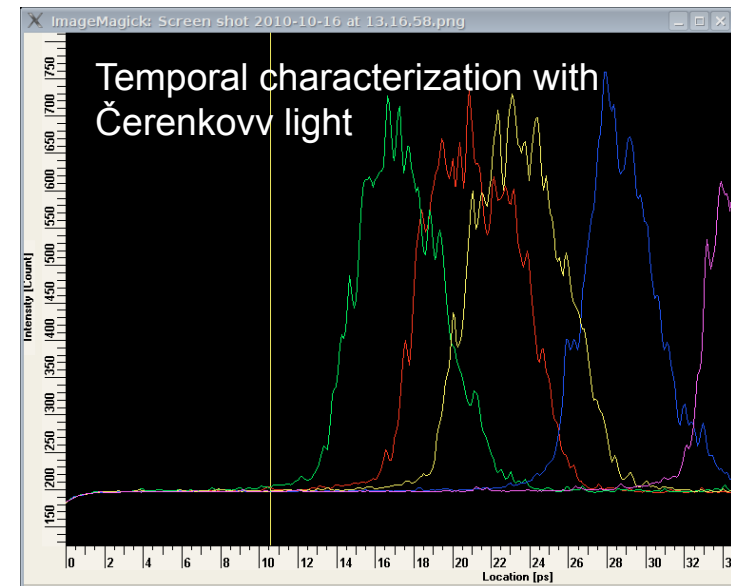
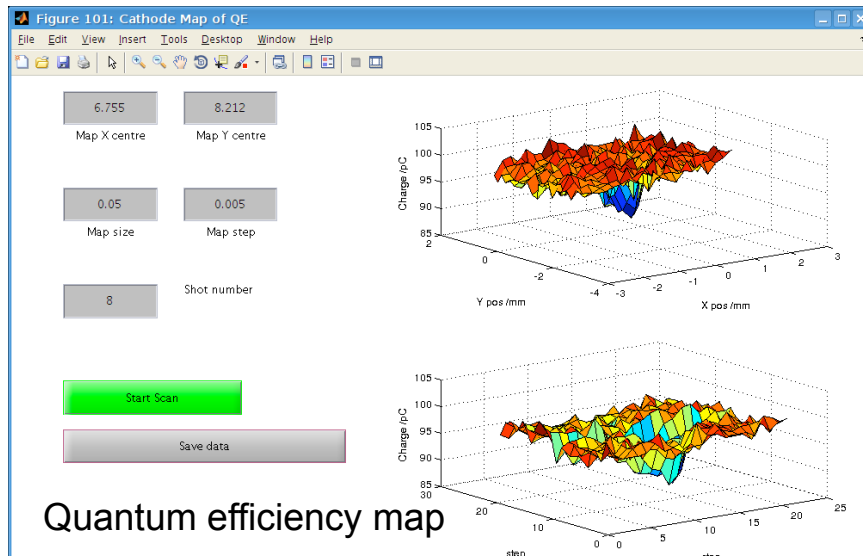
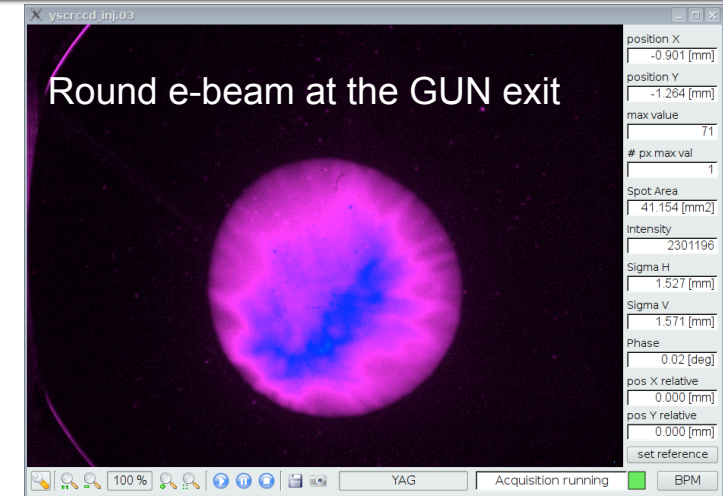


The photo-injector gun has been used for about 2 years without major problems. Beam charge of 800pC have been extracted and measured. Operations have been mainly using 350 or 200pC.

Degradation of the quantum efficiency has been experience but every run has been successfully re-establishes with an ozone cleaning procedure. The copper cathode has not yet been changed.

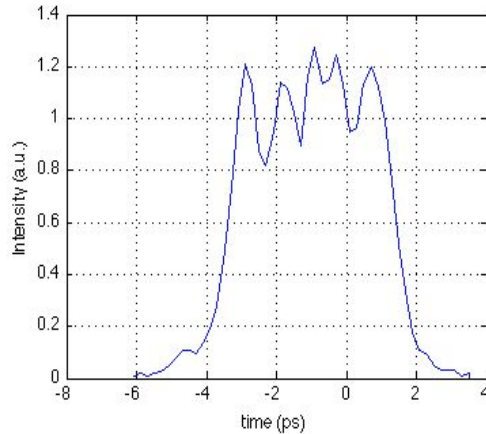
Gun is operated at 100MV/m filed with a fault rate smaller than $3e-3$.

Bright electron beams (5 ps, 50 A, 1 mm mrad) space-charge repulsion forces are balanced by acceleration and magnetic focusing trough a solenoid.



M. Trovo', L. Badano, C. Spezzani

Laser: FWHM=5ps, $\phi=1.3\text{mm}$

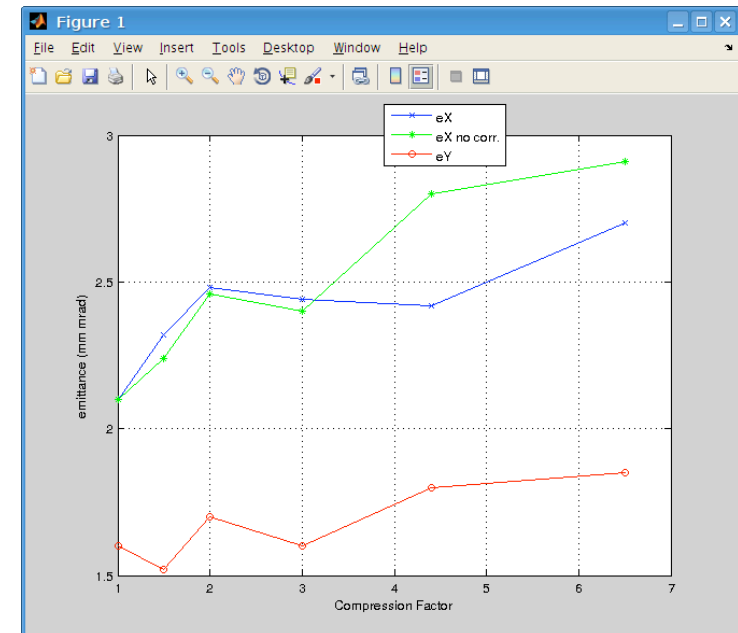


Most of the **emittance studies** have been focused on a **350pC** generated by a 5ps laser pulse. Typical values for the **projected emittance** in the laser heater region (100MeV) are below 1 mm mrad. The best measured values **0.8 mm mrad**. We experience a **critical dependence** of the emittance on the **optics and trajectory** while propagating the beam.

A good LINAC configuration allow to propagate to BC1 (350 MeV) a beam with about 50% emittance increase ($e_{x,y}=1.3\text{ mm mrad}$) before compression.

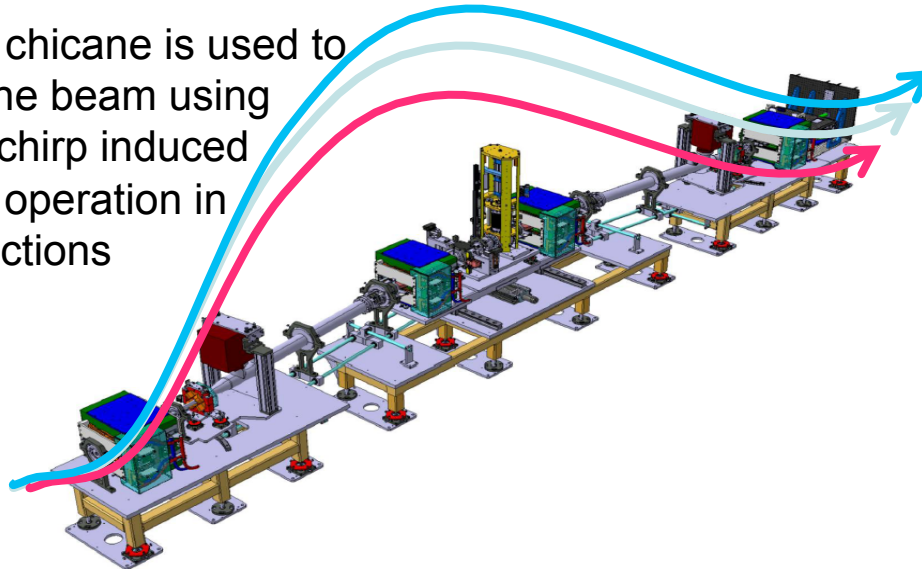
Effects of the **compression** on the emittance are not easy to avoid, generally an **increase** close to a **factor 2** is measured in the **horizontal plane**. An accurate tune of the LINAC optics can reduce the emittance growth to another 50%.

| | ϵ_x | ϵ_y |
|--------------------------------|--------------|--------------|
| BC1 Off | 1.1 | 1.0 |
| BC1 Off, with L01 at +25deg | 1.7 | 1.3 |
| BC1 On, CF=1 | 1.3-1.6 | 1.3-1.5 |
| BC1 On, CF=6.5 (L01 at +25deg) | 1.9 | 1.4 |

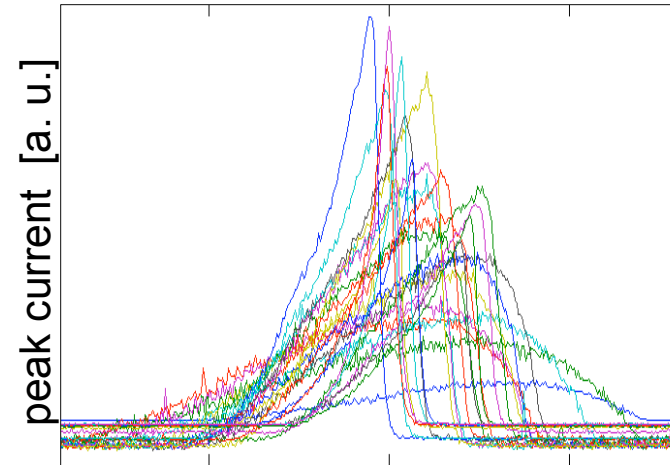


Ref. Di Mitri, Penco

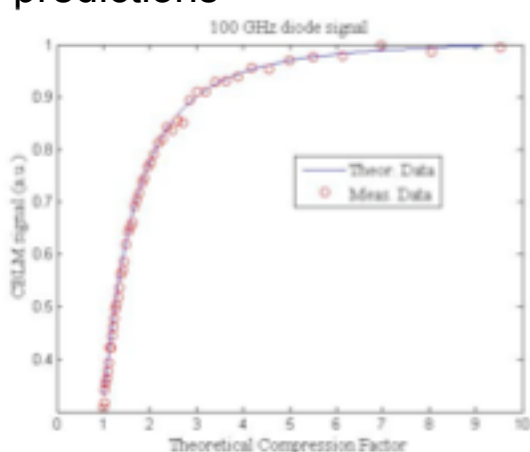
A magnetic chicane is used to compress the beam using the energy chirp induced by off-crest operation in previous sections



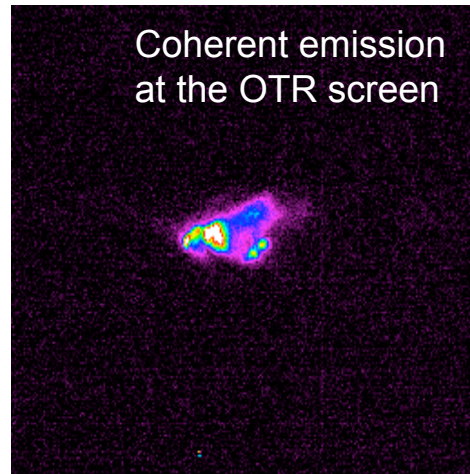
Bunch are compressed from 6 ps to about 0.7 ps



Measurements from the new diagnostic using the coherent signal of the bunch are in agreement with the theoretical predictions

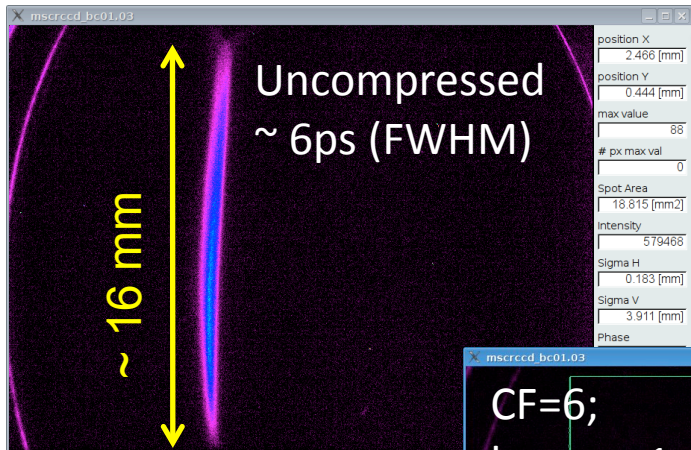
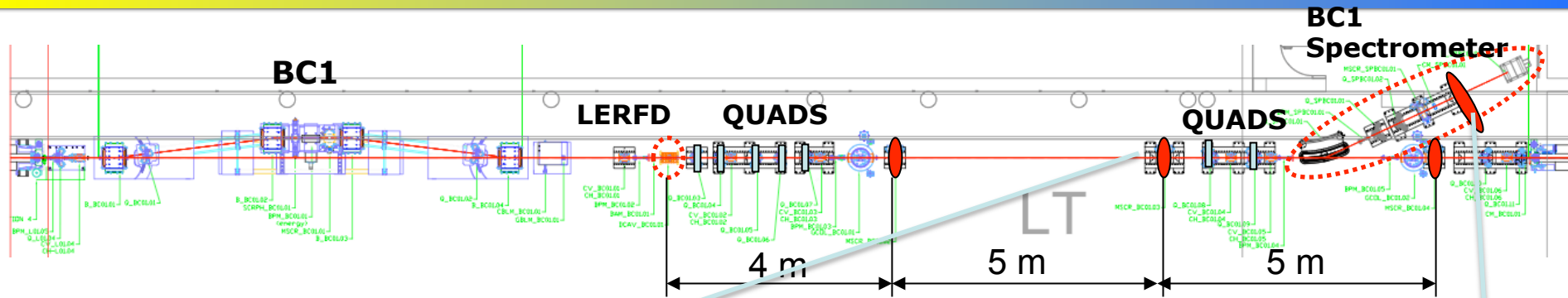


Coherent emission at the OTR screen

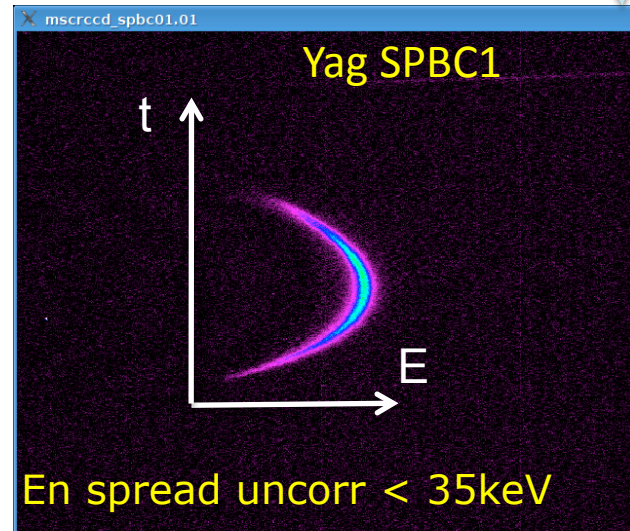
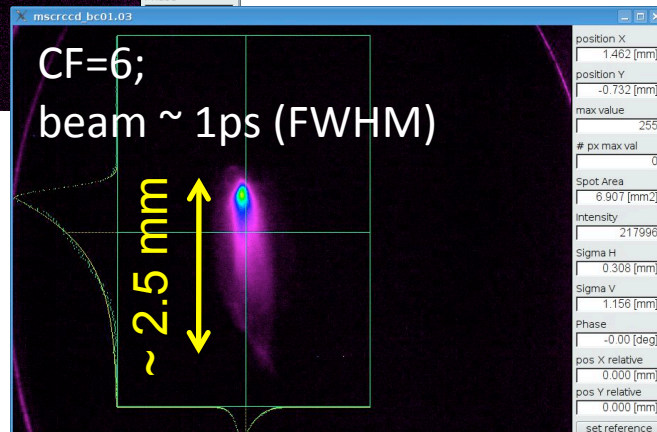


A **high peak current** (500 A over 0.7 ps FWHM) is generated by compressing the bunch length.

When the beam is compressed OTR screen after BC1 are affected by COTR, usually YAG screens can be used.



Deflecting cavity is now routinely used for bunch length measurements.

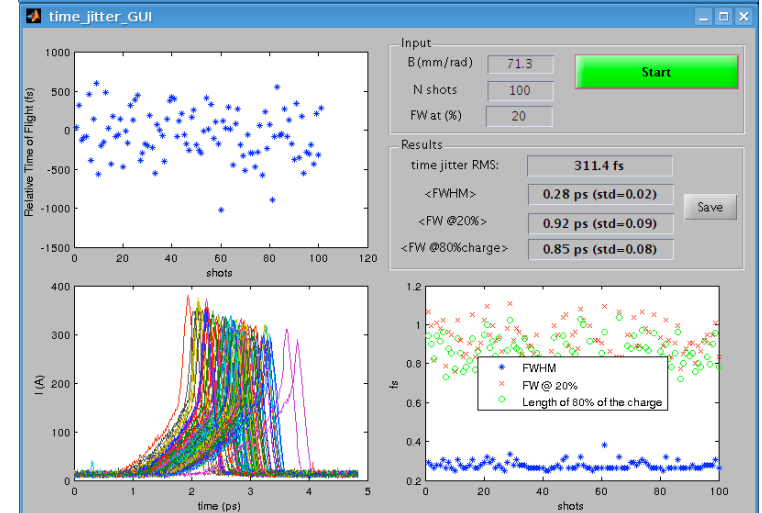
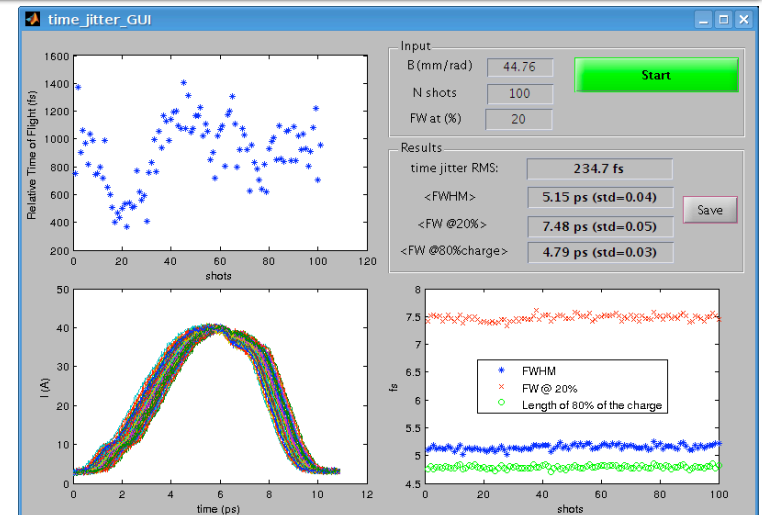
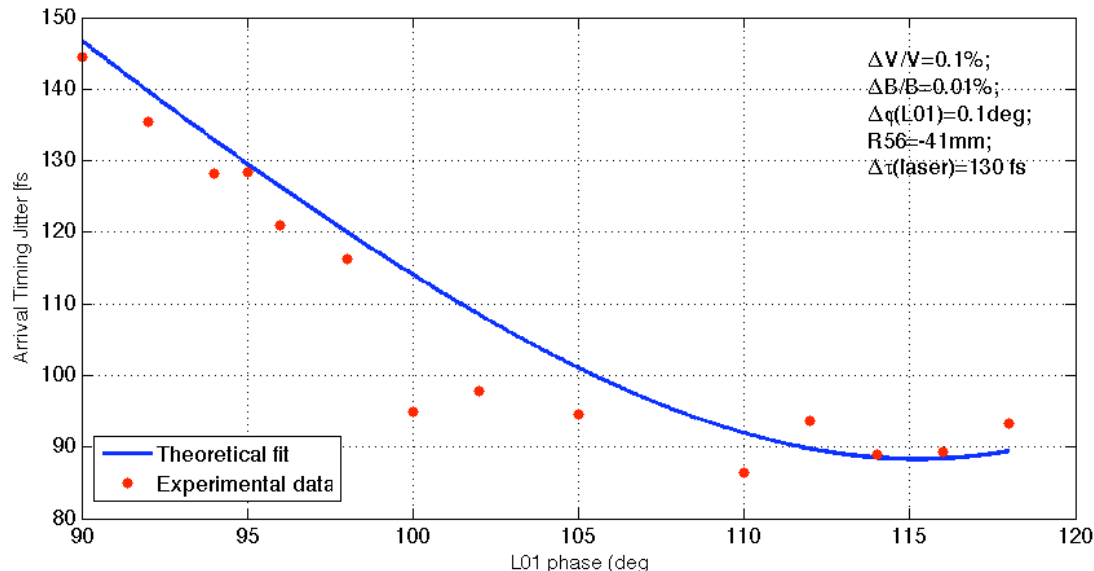


Combined with the spectrometer allow to reconstruct the phase space and can be also used for slice parameters.

A crucial parameter for the seeded FEL is the timing jitter of the electron beam.

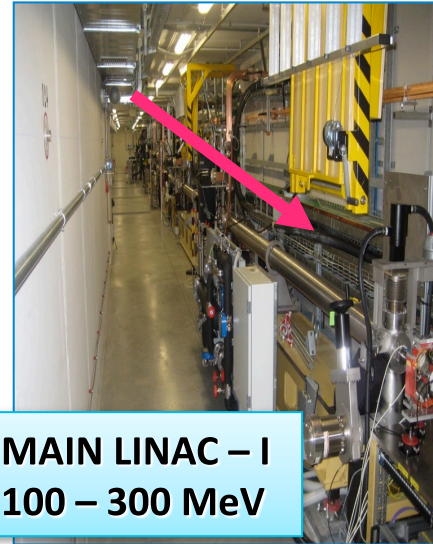
A dedicated diagnostic (Bunch Arrival Monitor) is under commissioning, BAM jitter measurements are confirmed by measurements with deflecting cavity.

A significant increase of the timing jitter as we start compressing the beam was experienced for not optimized Low Level RF system. Optimizing the LLRF system allow to decrease the timing jitter from about 300fs to about 100fs.

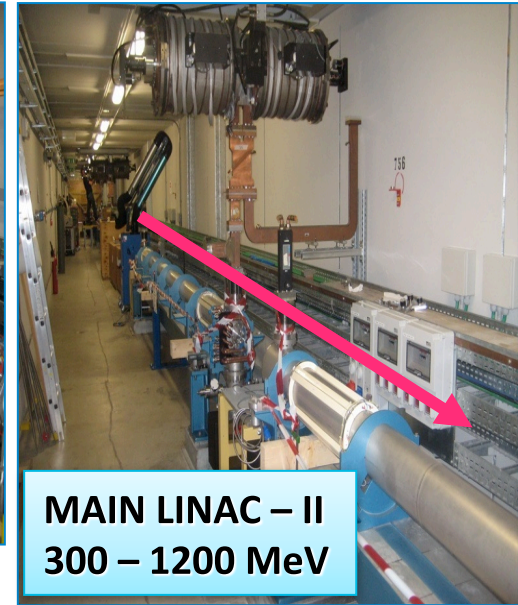


The goal energy of 1.2 GeV at the end of the linac has been achieved. Shot to shot energy stability is usually smaller than $1e-3$.

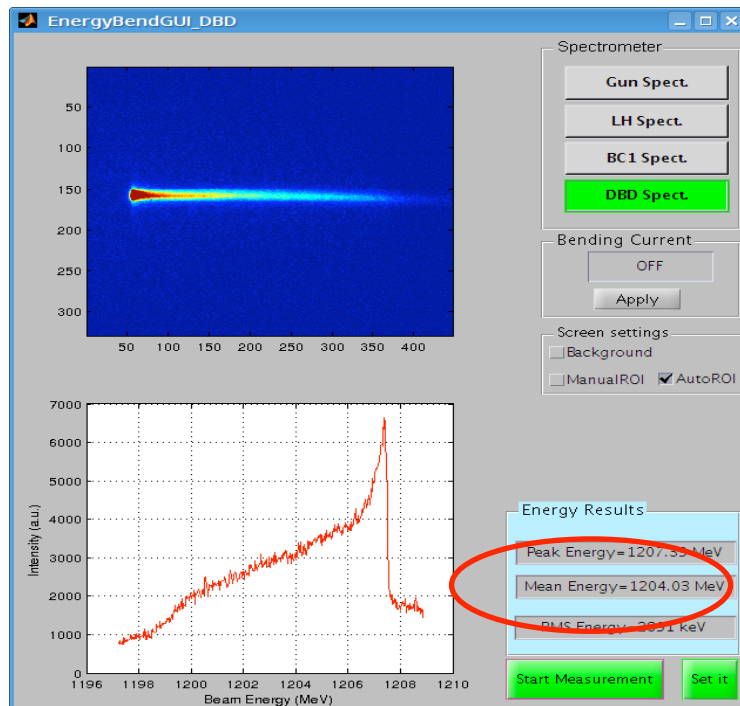
An emittance growth is measured from BC1 to LINAC end, probably due to residual dispersion in LINAC the problem is under investigation.



MAIN LINAC – I
100 – 300 MeV

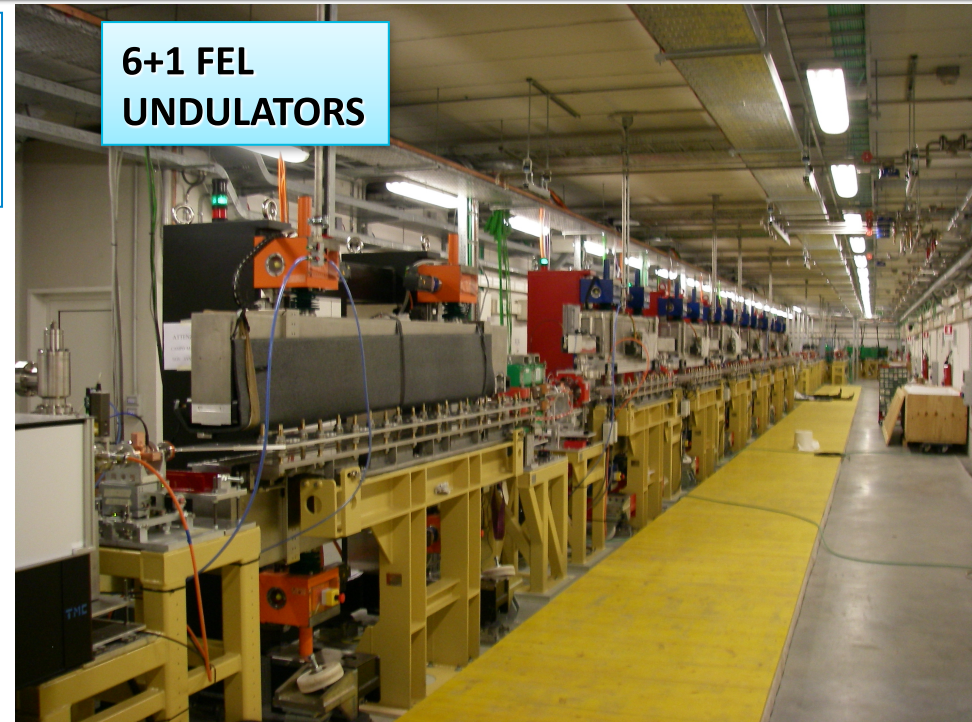
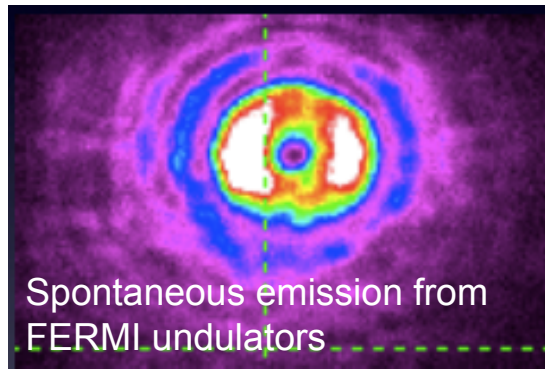


MAIN LINAC – II
300 – 1200 MeV

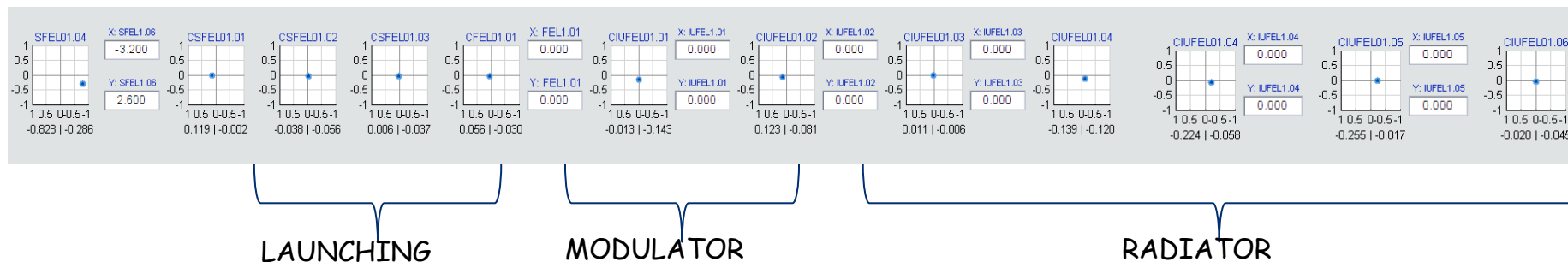


By using the phase reversal modulation on the sled cavities 1.3GeV is easily achieved with the present linac.

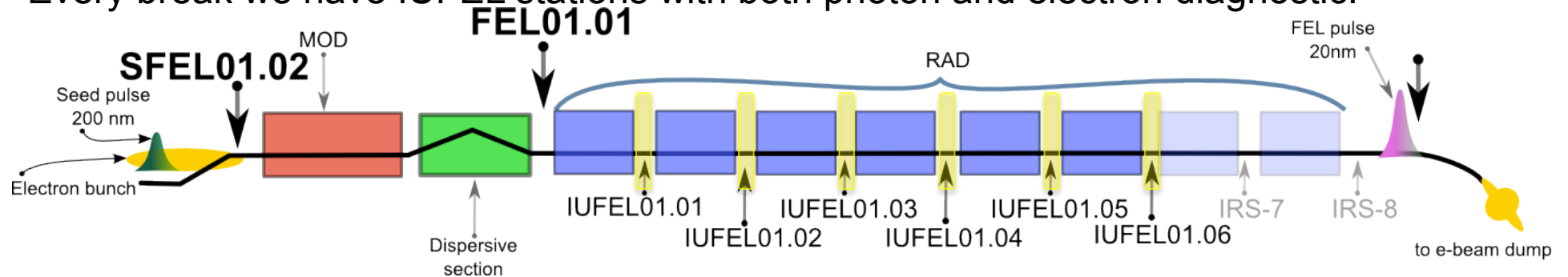
APPLE-II type undulators allow **high brilliance FEL** with **variable polarization** and **tunable wavelength**



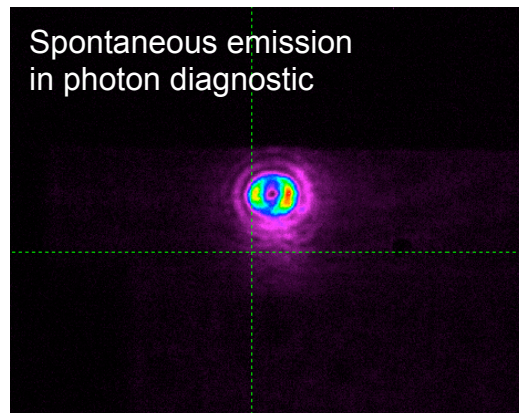
Beam-based straight line defined in the undulator: quadrupoles (x,y) and undulators (y only) moved to minimize beam steering. All C-BPMs set to zero. All correctors turned off.



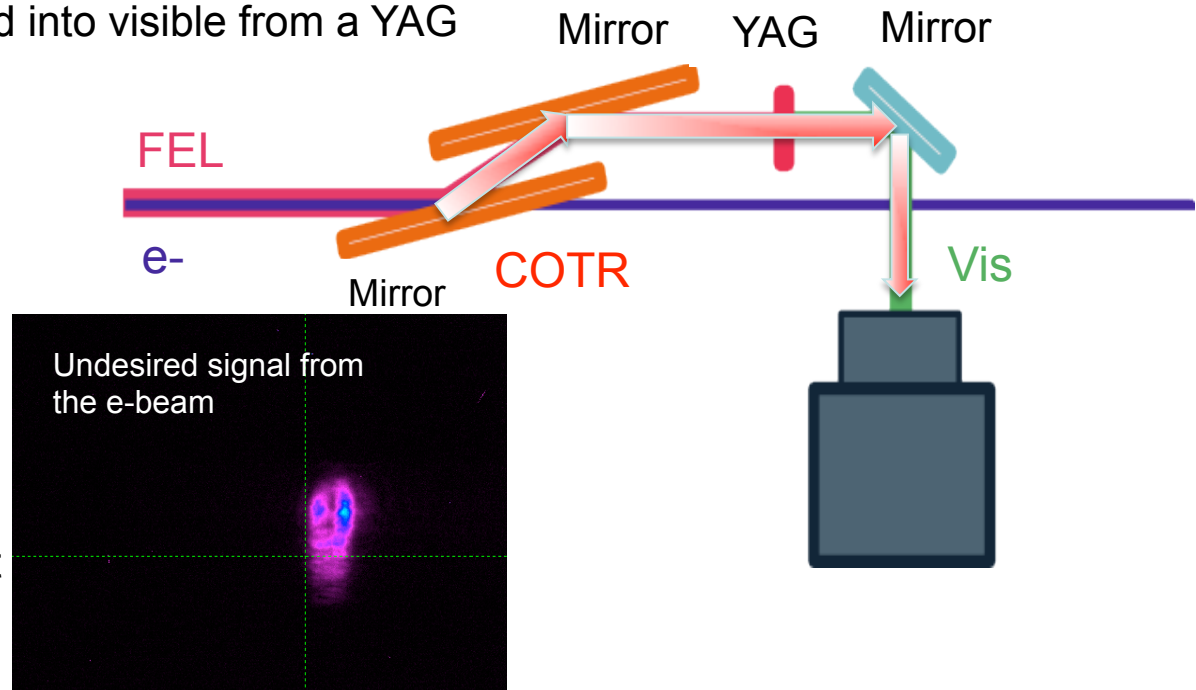
Every break we have IUFEL stations with both photon and electron diagnostic.

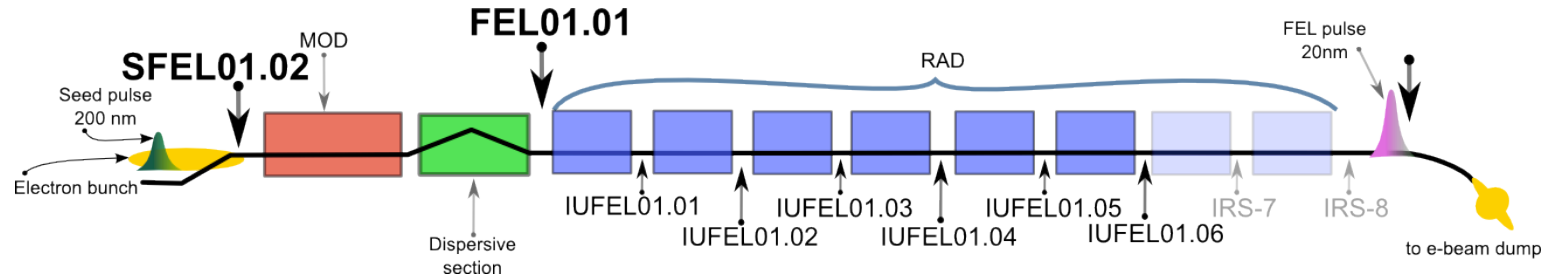


Photon diagnostic use a copper mirror to separate FEL from electrons. VUV light is then converted into visible from a YAG and detected by a CCD camera.

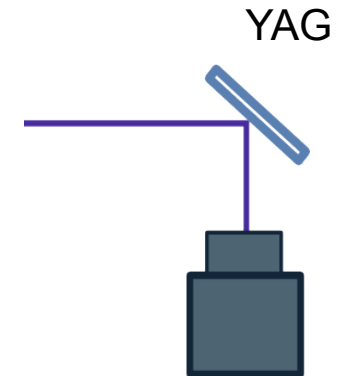


Depending on the electron alignment the measure is affected by OTR emission on the first mirror.



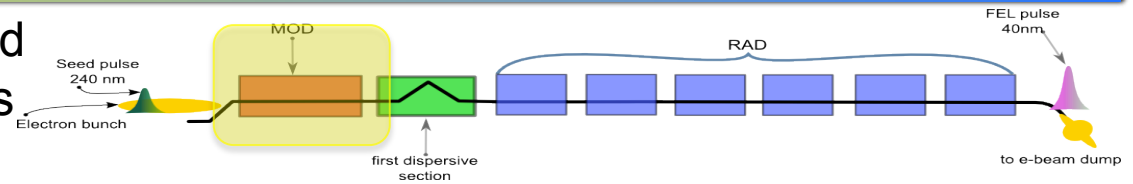


Most of IUFEL stations have a simple YAG to detect electrons. Also photons from the seed can be seen on this screen.

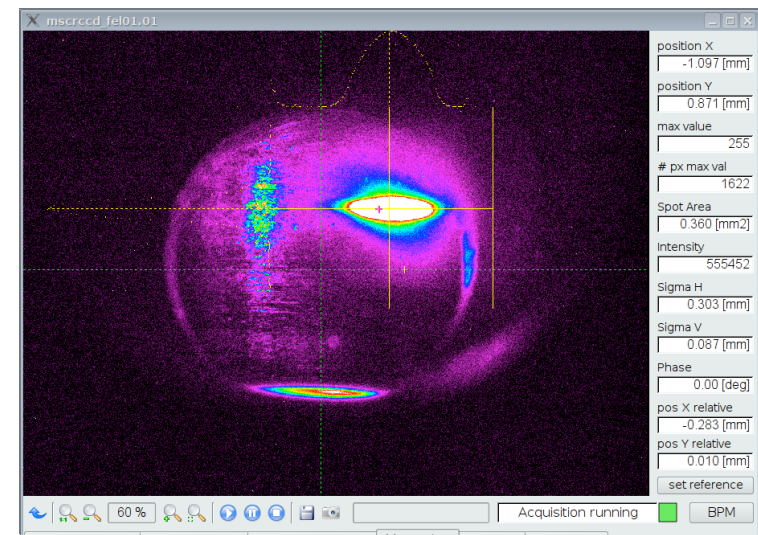
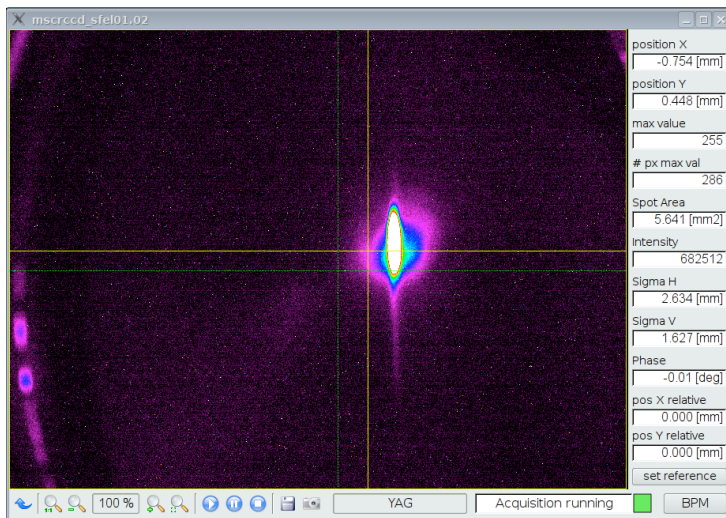
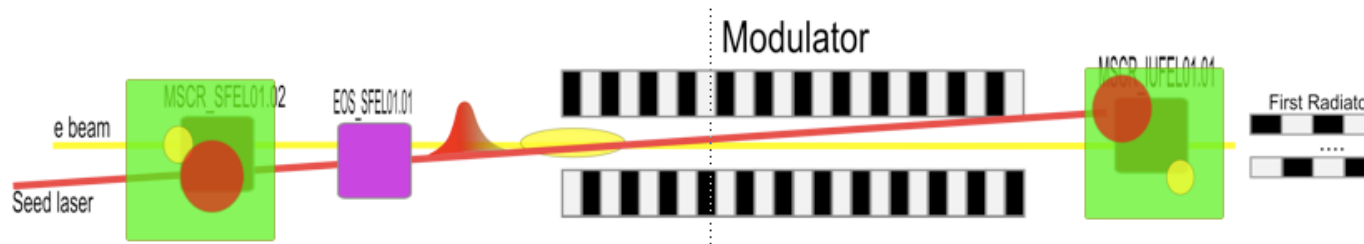


Original designed uses OTR in this position but for the moment it has not been possible to used them a we will continue using the YAG

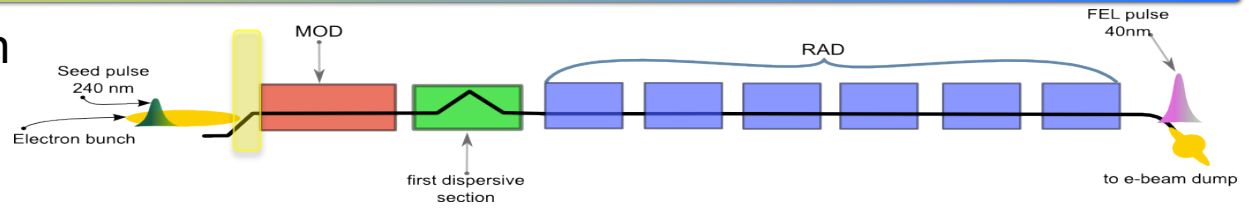
Spatial **overlap** of the **seed laser** and the **electron beam** in **modulator** has been tested using two YAG screens around the modulator.



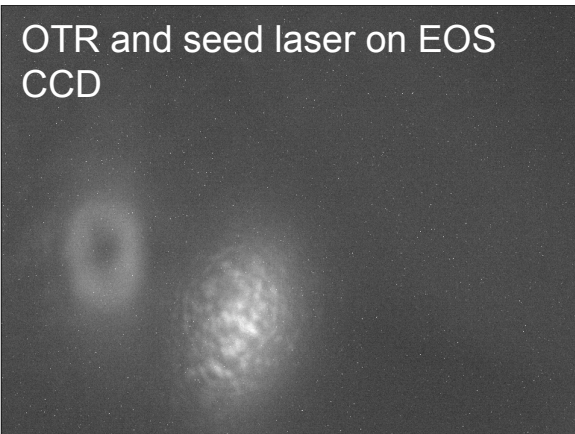
Using the seed laser steering mirror it is possible to superpose the two beams.



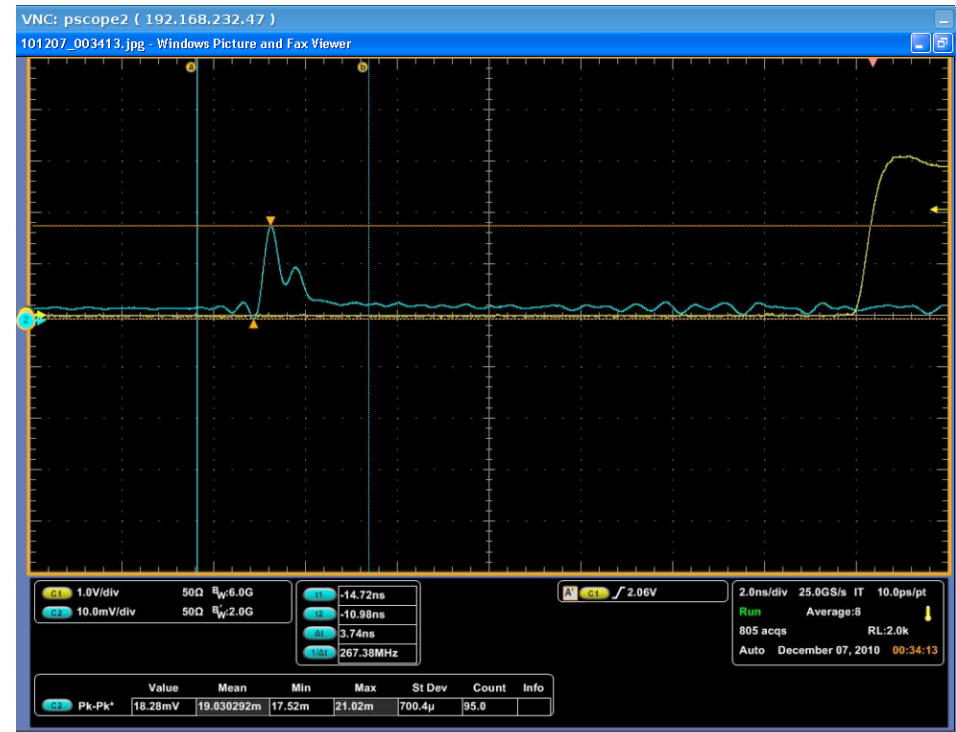
The seed laser and the electron beam are detected on a photodiode installed in the Electron Optical Sampling chamber.



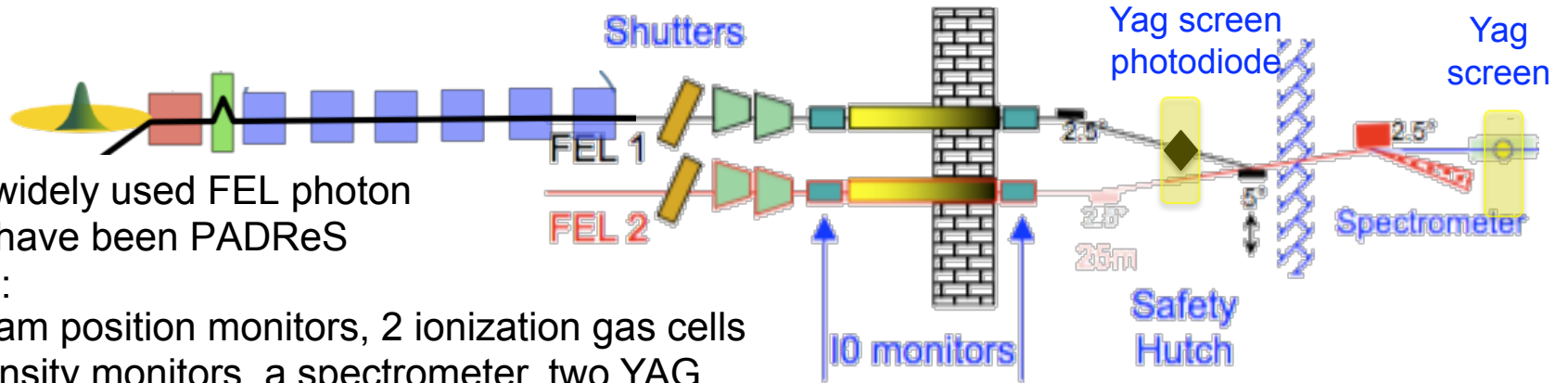
An aluminum foil is used to reflect the laser out of the chamber and to produce OTR. The two signals can be detected with a CCD and a fast photodiode



After the signal optimization using a CCD it has been possible to obtain the timing parameters that allow to overlap in time e-beam and seed laser within ± 50 ps

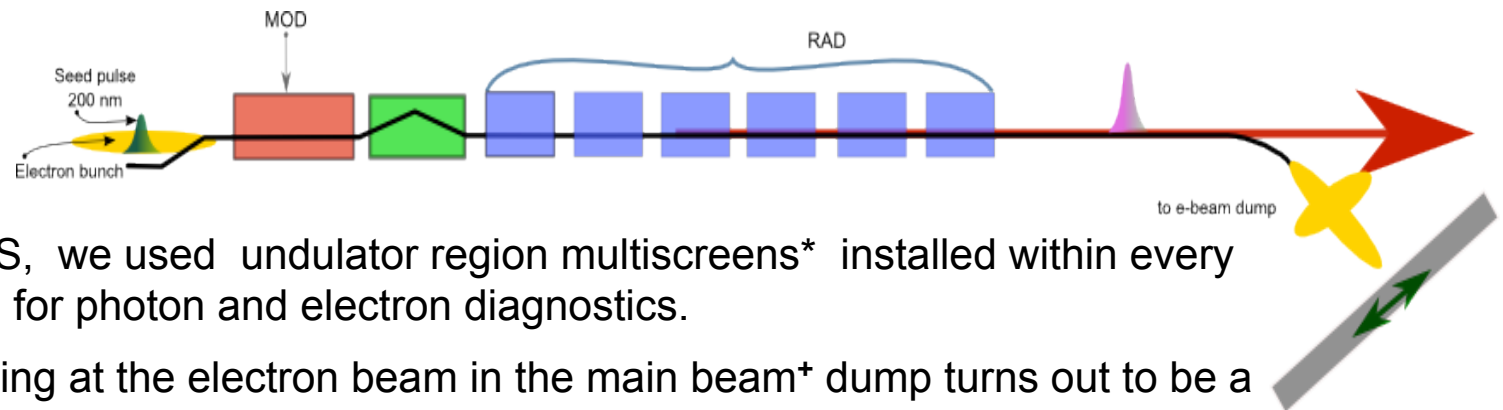


Photon diagnostic PADReS[#] (Photon Analysis Delivery and Reduction System)



The mostly widely used FEL photon diagnostics have been PADReS components:
 2 photon beam position monitors, 2 ionization gas cells used as intensity monitors, a spectrometer, two YAG screens and a calibrated photodiode.

#D. Cocco, C. Svetina, M. Zangrando



In addition to PADReS, we used undulator region multiscreens* installed within every undulator break, both for photon and electron diagnostics.

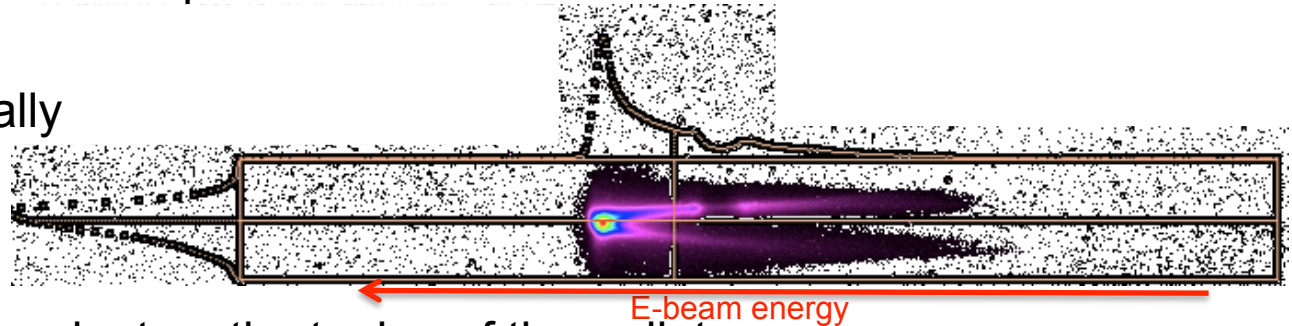
The YAG screen looking at the electron beam in the main beam+ dump turns out to be a pivotal diagnostic system for FEL optimization.

*M. Veronese

+L. Badano

A useful diagnostic for detecting the occurrence of the seeding is provided by the MBD spectrum of the electron beam.

When the seed laser temporally overlaps the electron beam, a dark “shadow” appears on the MBD image.



The effect is only slightly dependent on the tuning of the radiators, and the dispersive sections R56.

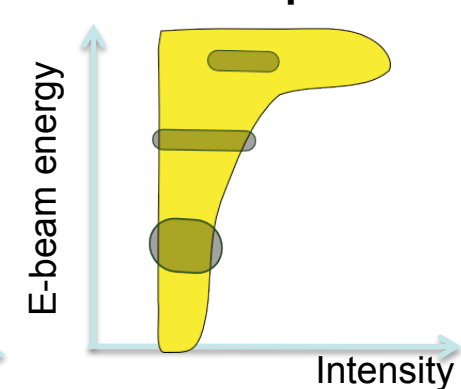
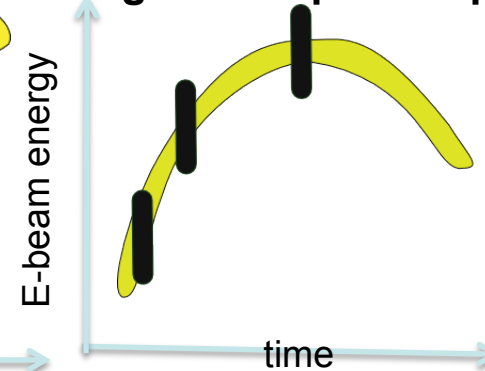
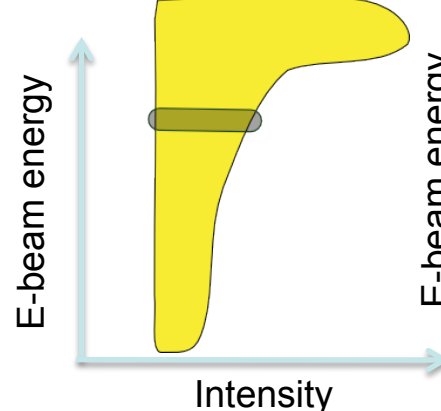
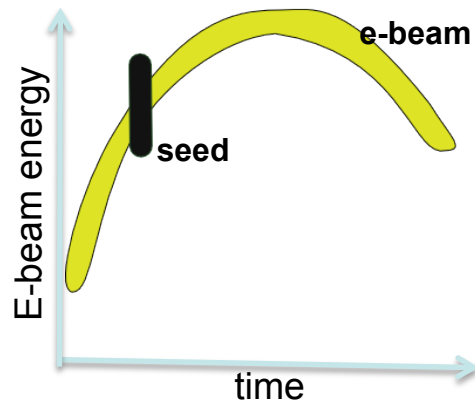
We used this effect several times to confirm proper overlap since it (often) was more reliable than FEL intensity diagnostics.

Longitudinal phase space

e-beam spectrum

Longitudinal phase space

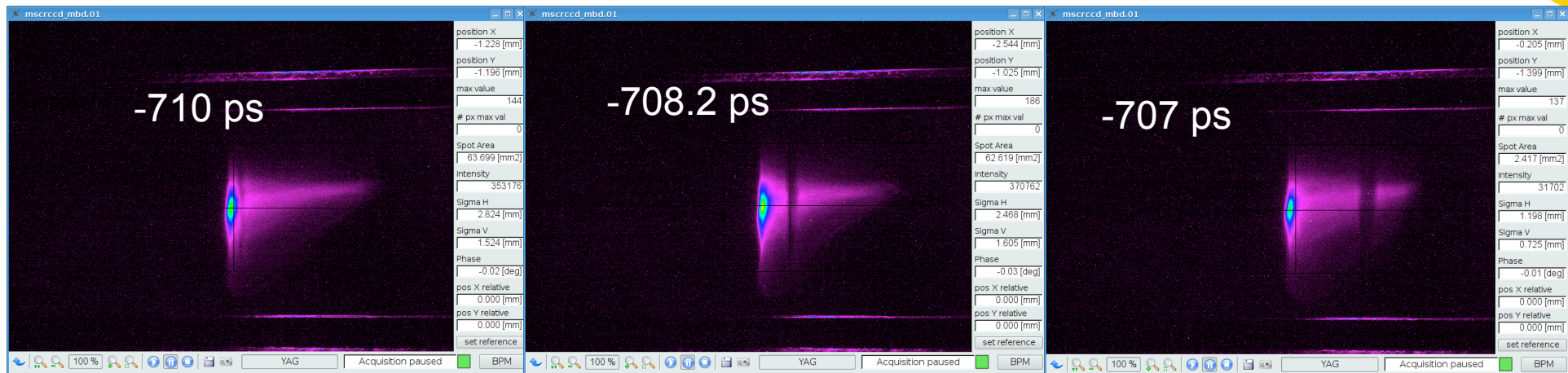
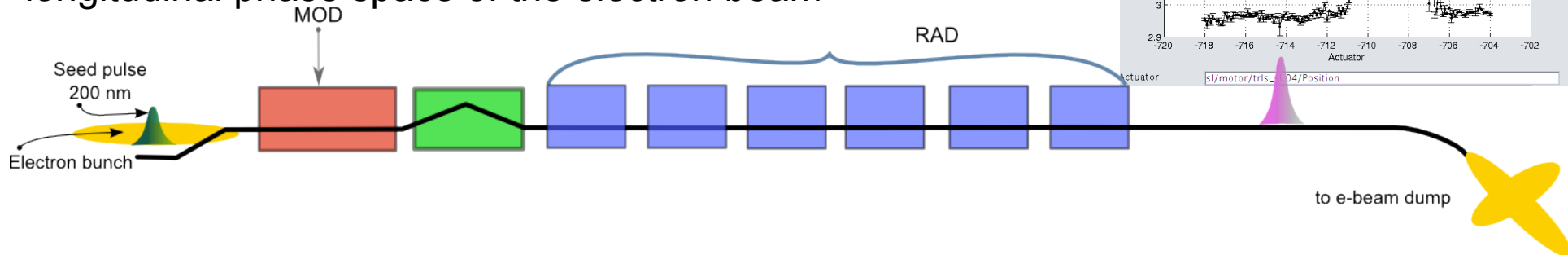
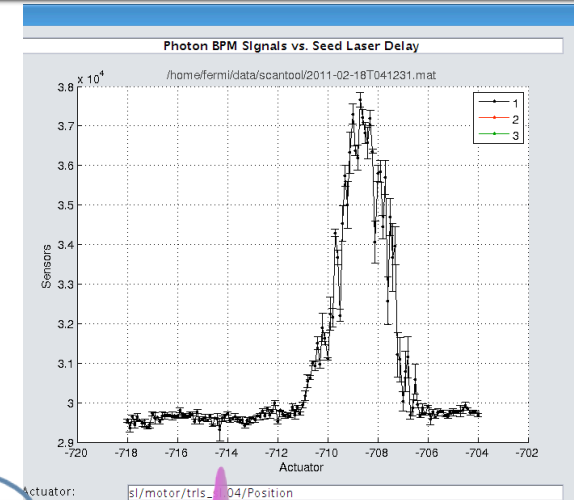
e-beam spectrum



After the seed alignment procedure the FEL easily ON.

Evidence of a band in the e-beam spectrometer in Main Bean Dump as a result of the seeding.

The seeding can provide information on the longitudinal phase space of the electron beam

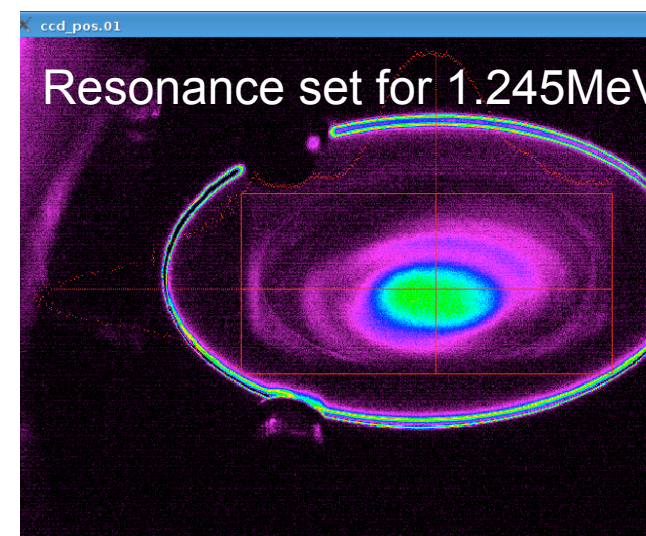
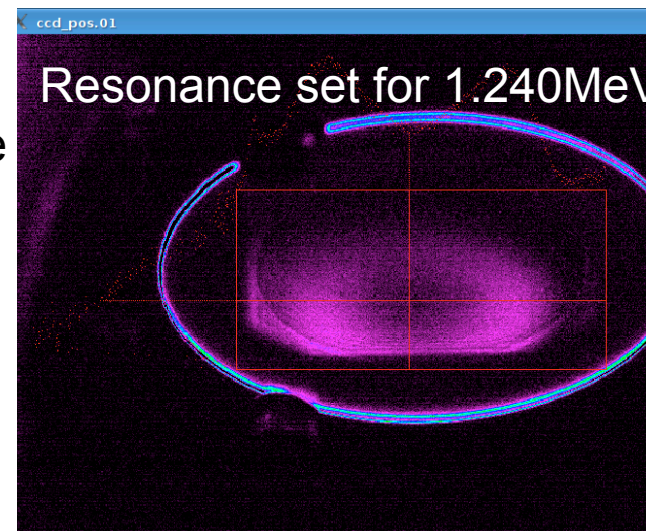


Optimizing the undulator tuning (both in K and electron beam position) is done by examining the far-field FEL spot size. This has been very critical for FERMI since the signal detected by the I0 monitor is sensitive to the FEL spatial mode.

A small undulator mismatch (of the order of $\Delta K \sim 0.1\%$) can produce a “doughnut” transverse mode with resonance moved to the outer portions of the electron beam.

The image of the FEL radiation on the PADReS YAG screen has become a critical diagnostic for FEL optimization, allowing us to increase the overall power by a factor 10 or greater.

The images to the right correspond to 43nm with 6 radiators tuned to resonance.



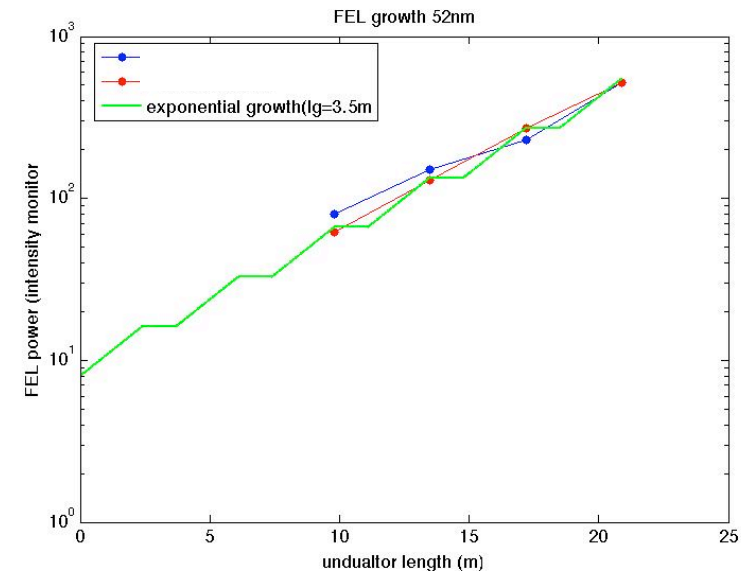
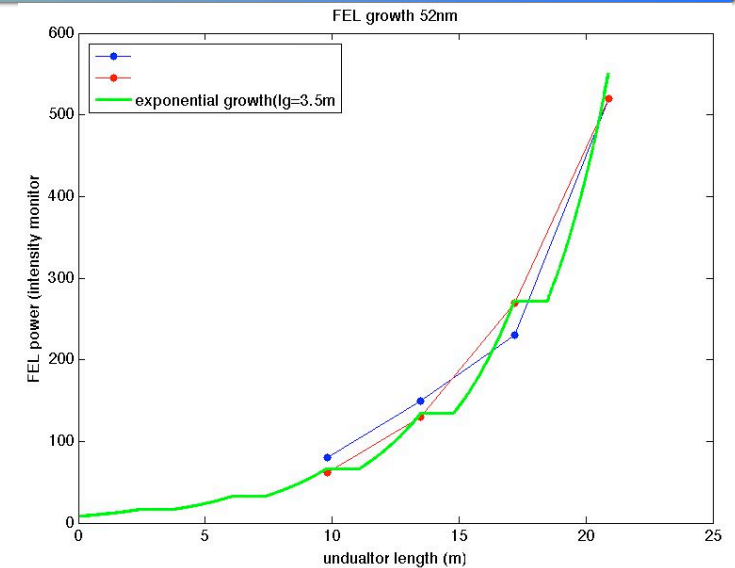
Once the FEL is optimized to produce an on-axis, TEM00 mode, there is a clear evidence of FEL gain with increasing undulator number.

Data reported here refer to the measured FEL power with the calibrated photodiode. Experimental conditions were:

350pC beam at 1.24GeV, electron beam compressed about a factor 3, seeding at 260nm, undulators tuned at 52nm ($h=5$).

Maximum measured FEL energy $\sim 20\mu\text{J}$

Measured data fit well with the exponential growth predicted for a 200A beam with 4 mm mrad emittance and 600keV energy spread in the radiator. Although 4 mm mrad is the typical emittance measured in the FEL region, the slice emittance could be better than this but also the electron beam matching could be worse. A more careful beam characterization is needed.

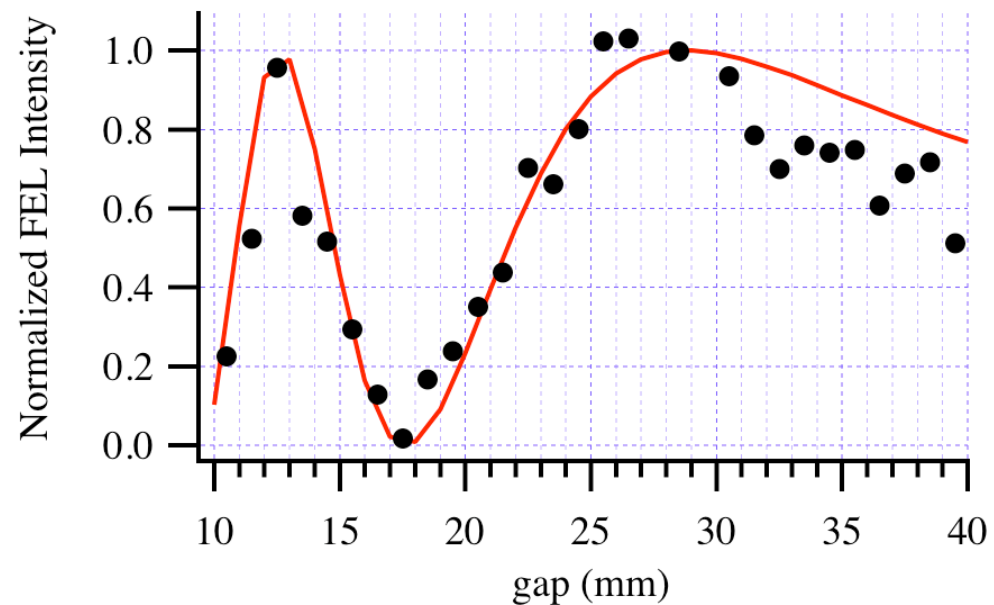


The radiation's longitudinal coherence is confirmed by the clear effect of the phase shifter.

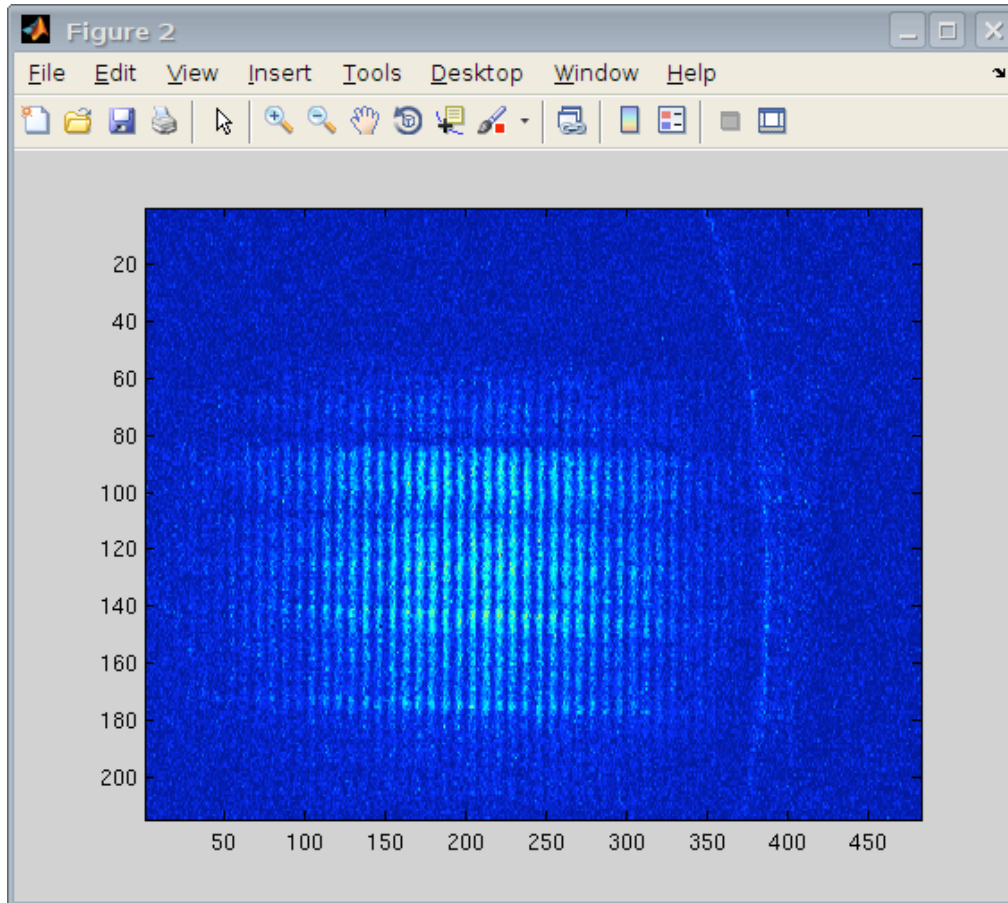
Using the phase shifter it is possible to coherently sum or subtract the emission from two consecutive undulators tuned to the same harmonic of the seed laser.

Data here reported refer to the case of a 350pC beam compressed by about a factor 3. The beam is seeded at 260nm and passes through two consecutive radiators tuned to 52 nm.

The measured FEL intensity (dots) as a function of the phase shifter gap is compared with a simple interference model of the phase shifter (solid line) based on the measured magnetic field.



Once the FEL output been optimized to a Gaussian TEM00 mode, measuring the spatial coherence(*) has been straightforward.



Several images have been acquired for different wavelength and slits conditions. Preliminary analysis indicate a very good degree of spatial coherence. Further quantitative analysis is ongoing.

Image of interference fringes of 52-nm FEL light produced by a double slit. Images taken with 9 second of acquisition (i.e., 90 pulses).

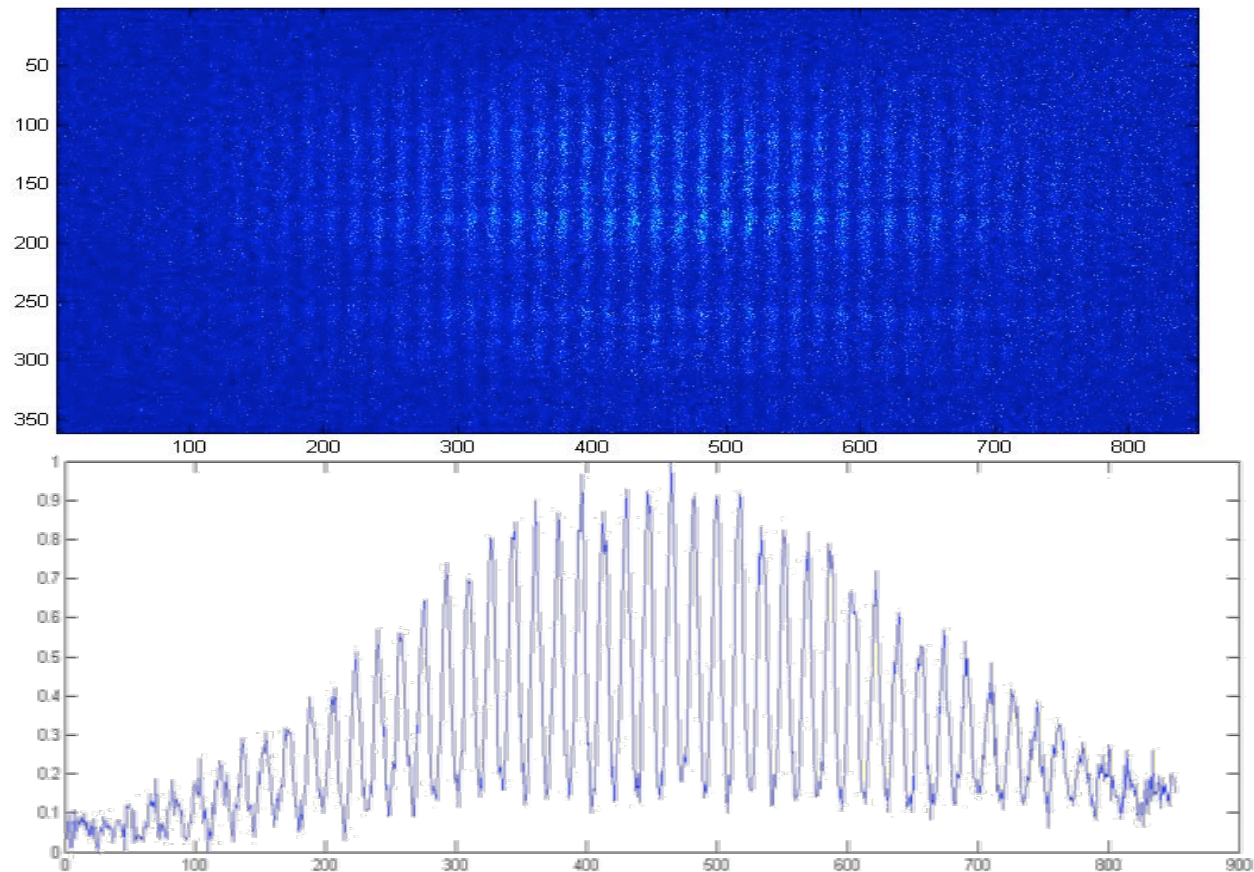
(*) Young slit experiments on FERMI proposed by F. Parmigiani and setup implemented by C. Svetina.

Double slit experiments were repeated at 32.5 nm, also showing very good transverse coherence.

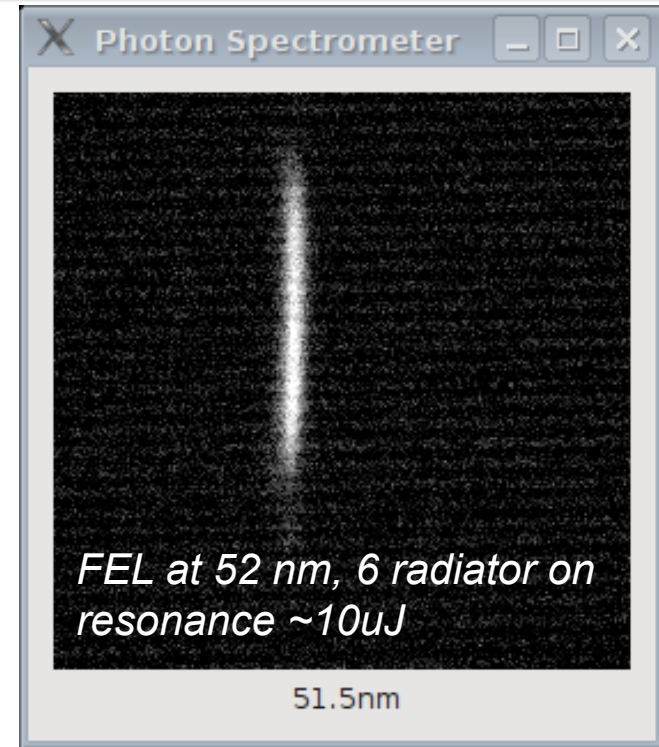
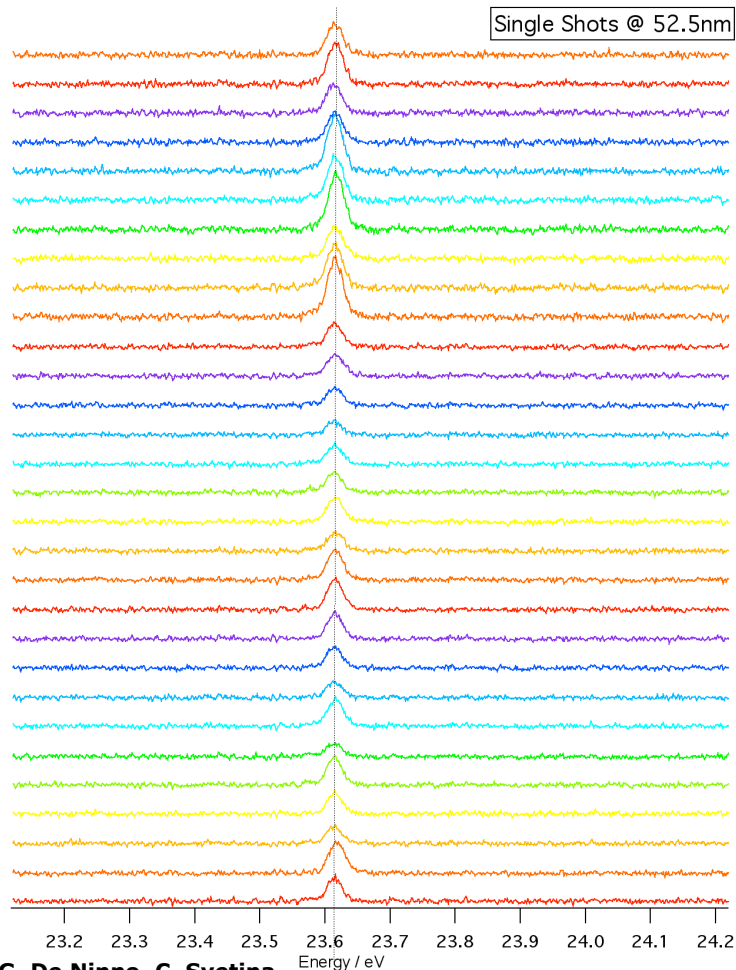
Fringe visibility is very high along the entire FEL pulse and also for relative large slits separation, indicating a very high degree of transverse coherence.

Quantitative analysis is ongoing.

FEL at 32.5 nm, 6 radiators,
450pC, compression ~3.
Slit separation = 0.8 mm,
width = 20 μm

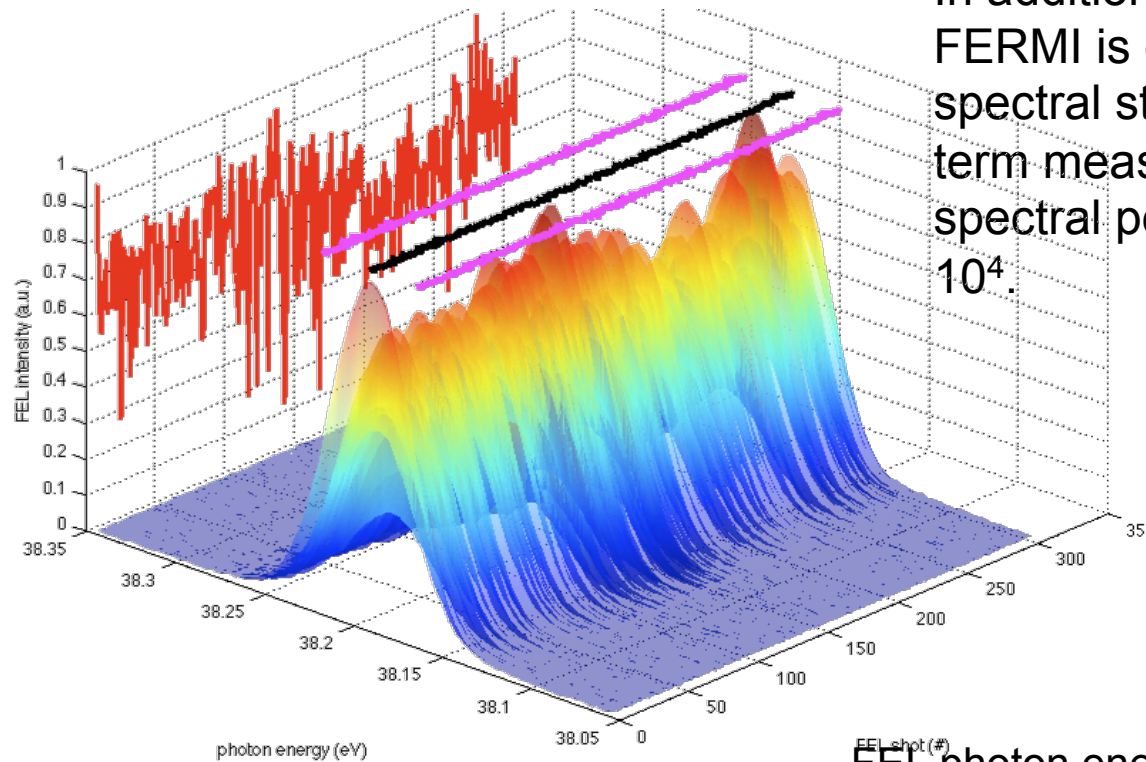


From the first day of FEL operation there has been clear benefits from using the external seeding in terms of bandwidth and photon energy stability.



Single shot measurements at 52nm show a bandwidth of 28meV (~0.1%)

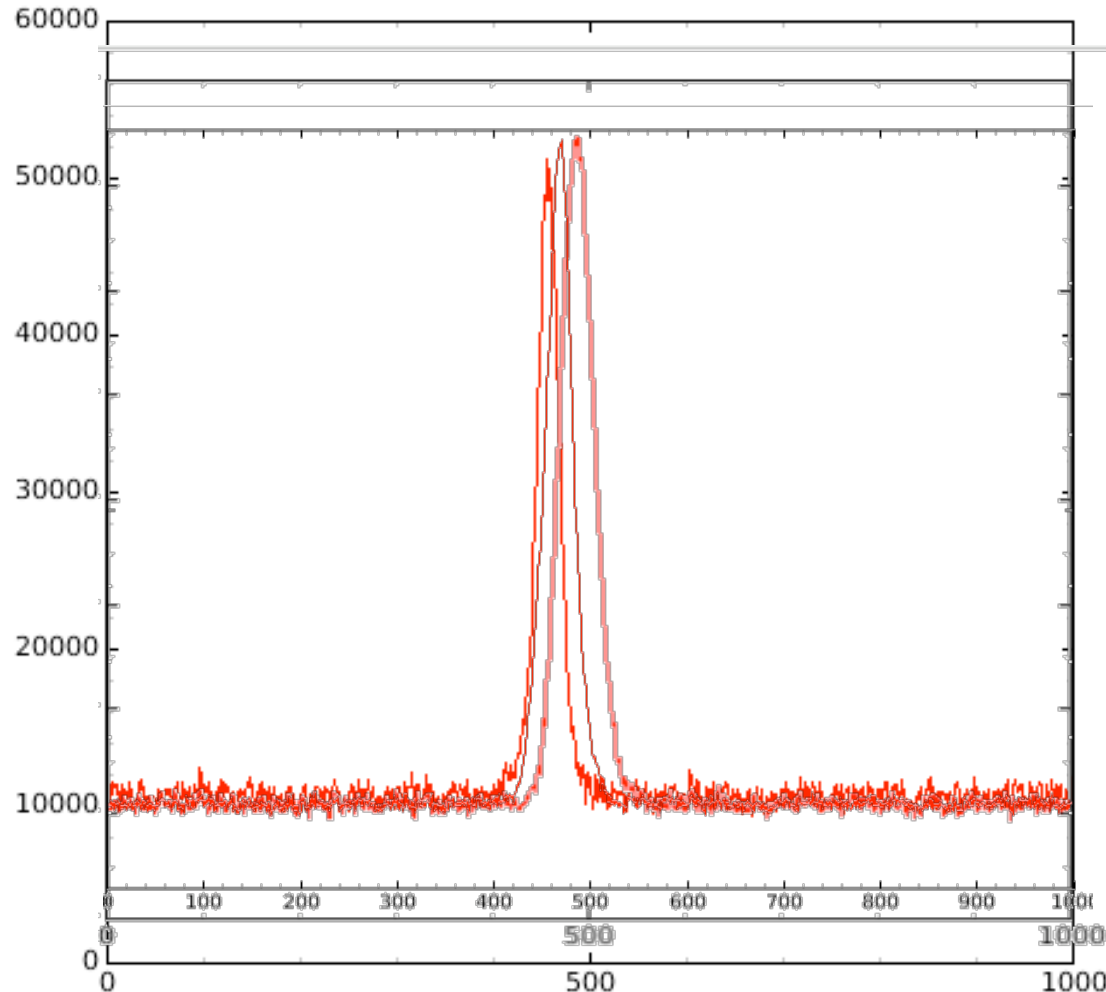
Typical seed laser parameters:
Pulse length ~150fs FWHM with a measured bandwidth at 260nm of 0.8nm (15meV)



In addition to the very narrow spectrum FERMI is characterized by excellent spectral stability. Both short and long term measurements show that the spectral peak jitter of less than 1 part in 10^4 .

Reported data refer to an electron beam of 350pC at 1.24GeV compressed about a factor 3. The 6 radiators are tuned to 32.5nm.

| | |
|-------------------|-----------------|
| FEL photon energy | ~ 38.19eV |
| fluctuations | = 1.1meV (RMS) |
| fluctuations | = 3e-5 (RMS) |
| FEL bandwidth | = 22.5meV (RMS) |
| fluctuations | = 5.9e-4 (RMS) |
| fluctuations | = 3% (RMS) |

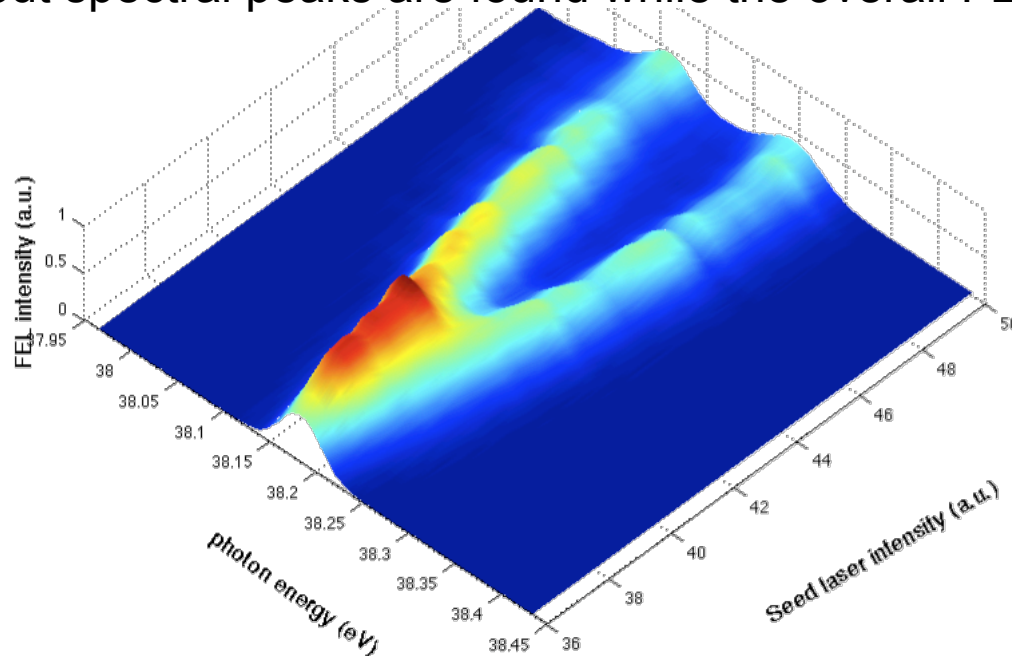


A small FEL tuning range around 52nm has been achieved by changing the seed laser wavelength of 1 nm (0.4%).

After tuning of the seed laser wavelength, the undulator resonance is changed accordingly to maximize the FEL power.

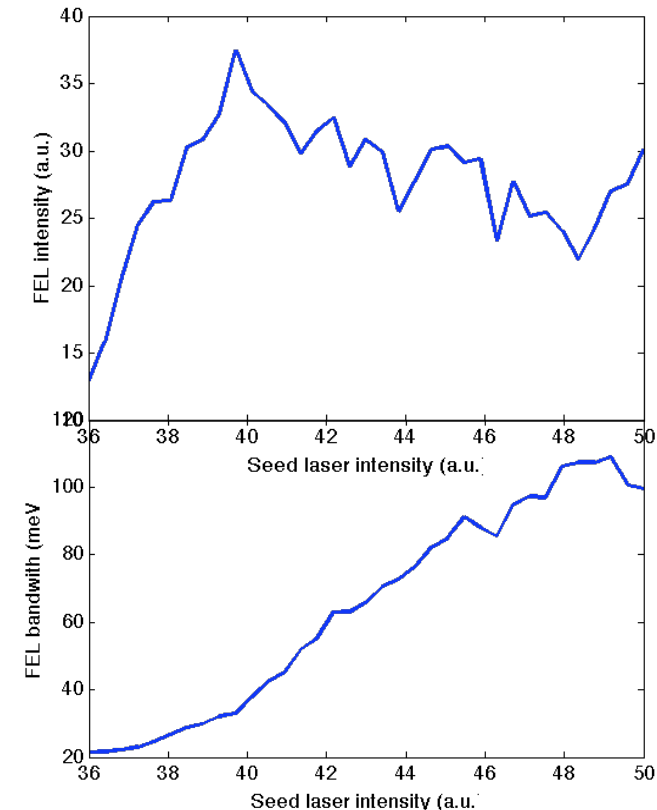
In the future, we larger tuning ranges will be possible using the tunability of the seed laser based on the Optical Parametric Amplifier.

The measured FEL bandwidth depends critically on the FEL optimization. By controlling the seeding process (seed intensity and/or R56 strength) it is possible to control the FEL bandwidth. In case of a very strong seed (overbunching regime) two output spectral peaks are found while the overall FEL intensity is only slightly affected.



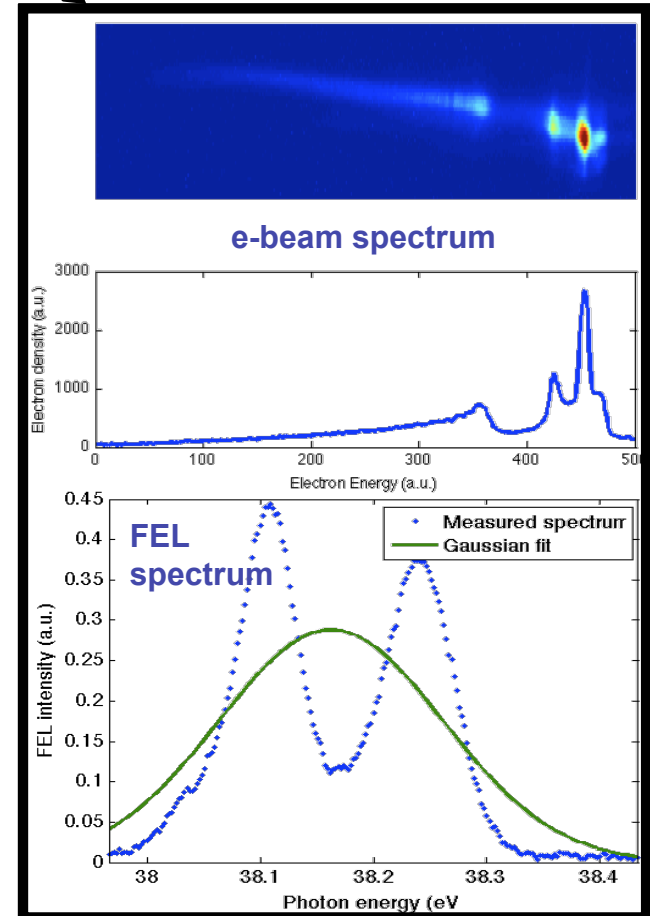
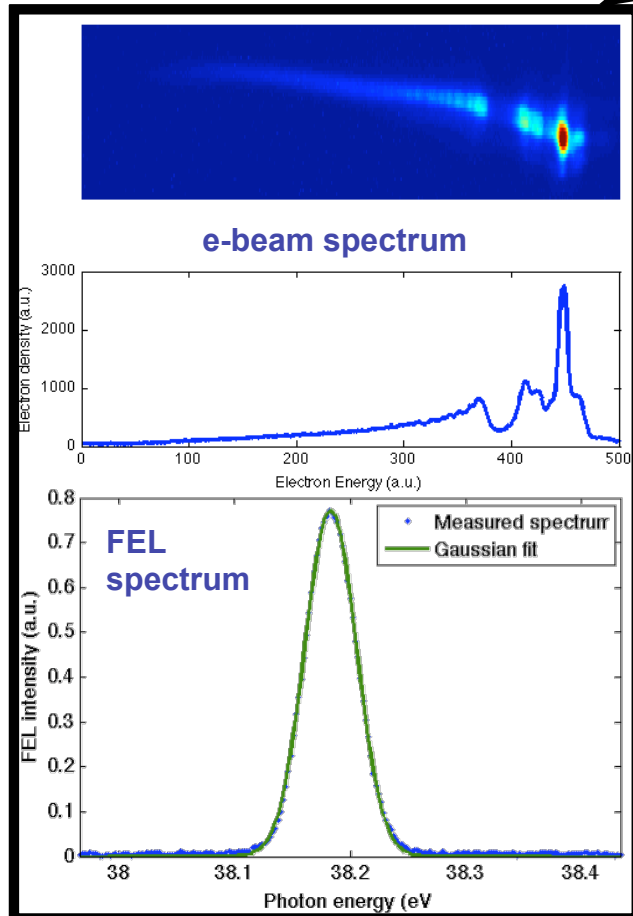
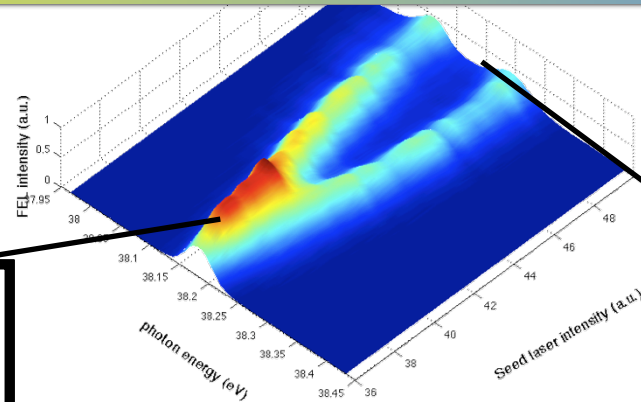
Both RMS and FWHM fits to the FEL bandwidth show an increase up to a factor 5 .

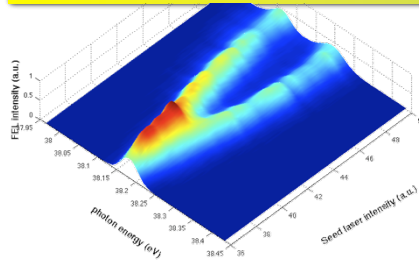
Data refers to a 350pC beam compressed ~ 3 at 1.24GeV. R56 $\sim 16\mu\text{m}$ and the 6 radiators tuned at 32.5nm in circular polarization.



The effect of a stronger seeding process is clearly visible also in the e-beam spectrometer.

The double peak in the FEL spectrum is associated with a larger “hole” in the e-beam spectrum.

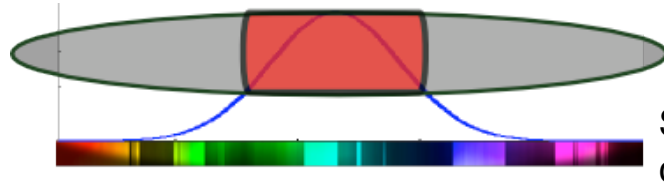




The two spectral peaks individually have a bandwidth only slightly larger than that of lower seed power, single peak FEL emission ($\sim 30\text{meV}$).

A possible interpretation of the phenomenon is that in the case of overbunching, only the head and the tail of the seed are producing an effective bunching used in the radiator and they are at different wavelength due to a residual chirp in the seed laser.

Only electrons that see **optimal seed intensity** contribute to FEL

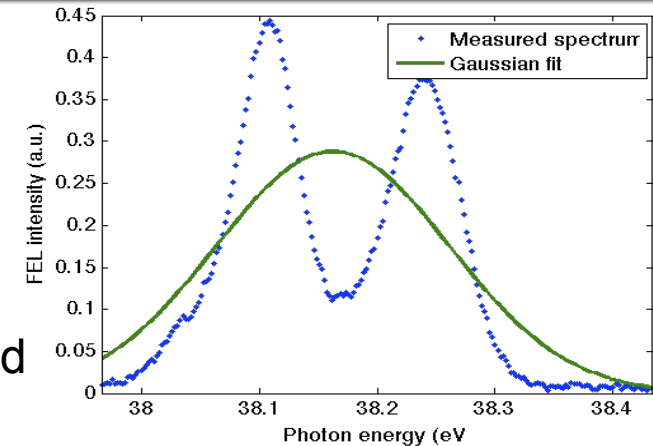


Seed laser with a residual chirp on the head and tail.

For **strong** seeding, electrons in the central region go in **overbunching** and do not contribute(*)



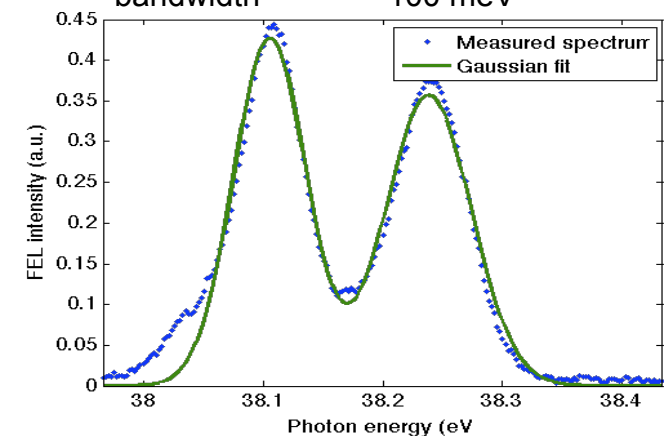
(*) M. Labat, et al. Phys. Rev. Lett. 103, 264801 (2009)



Single Gaussian fit

photon energy = 38.16

bandwidth = 100 meV



First peak

Intensity = 1
photon energy = 38.11
bandwidth = 30

Second peak

Intensity = 0.8 (a.u.)
photon energy = 38.24eV
bandwidth = 36meV

The effect of a chirped seed pulse has been studied as a possible means to increase the overall number of photons per pulse.

FEL in circular polarization has been produced

FEL has been operated also with the fully tunable, Optical Parametric Amplified seed laser.

Coherent Harmonic Generation has been extended efficiently down to the 13th harmonic (20nm) with an estimated FEL power of tens of μJ .

Clear evidence of coherent emission has been demonstrated down to 17nm (15th harmonic). Studies are ongoing to better optimize the FEL and quantify the produced intensity.

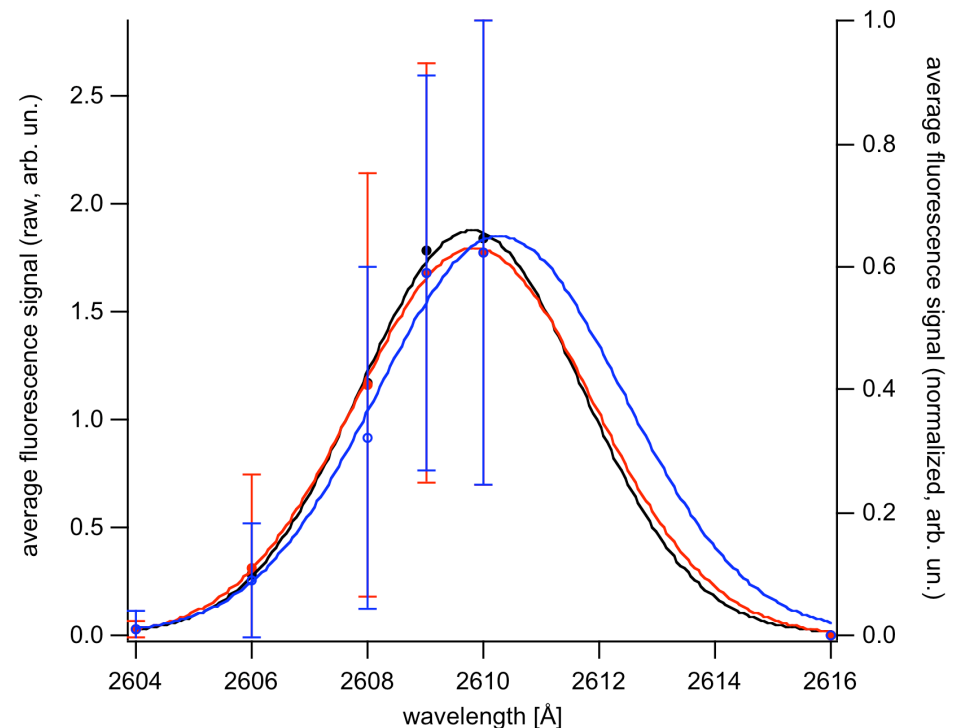
We have evidence of coherent transition radiation from the bunched electron beam at 130 nm or shorter wavelength.

During the last run, the Low Density Matter team(*) used the FEL tunability to scan the photon energy and measure the signal at an absorption resonance around the He 1s-4p .

The experiment immediately showed the dependence of the fluorescence signal on the FEL wavelength.

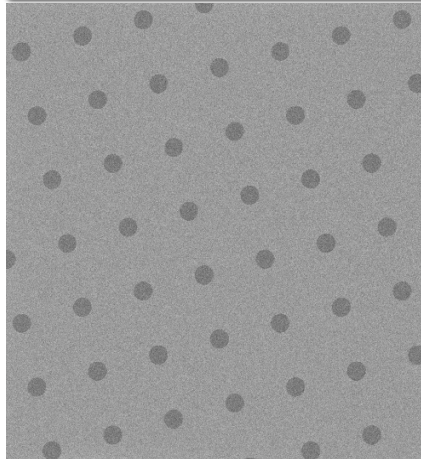
The atomic resonance is much narrower than the FEL linewidth, so measurement of the output spectrum by the LDM group allows independent determination of the input FEL spectrum.

A more detailed data analysis of other collected data is ongoing, to study the effect of FEL intensity on the ionization process.



(*) C. Callegari, V. Feyer, G. Cautero, A. Moise, K. Prince, R. Richter, R. Sergo (Sincrotrone Trieste) V. Lyamayev, M. Mudrich, F. Stienkemeier, U. Person (University of Freiburg) L. Avaldi, P. Bolognesi, M. Coreno, P. O' Keeffe (CNR-IMIP, Montelibretti) M. Alagia, M. de Simone, A. Kivimäki, (CNR-IOM, TASC, Trieste) M. Devetta, P. Milani, P. Piseri, T. Mazza (University of Milan) S. Stranges (University "La Sapienza", Rome) T. Möller, Y. Ovcharenko (TU Berlin) M. Drabbels (EPFL Lausanne).

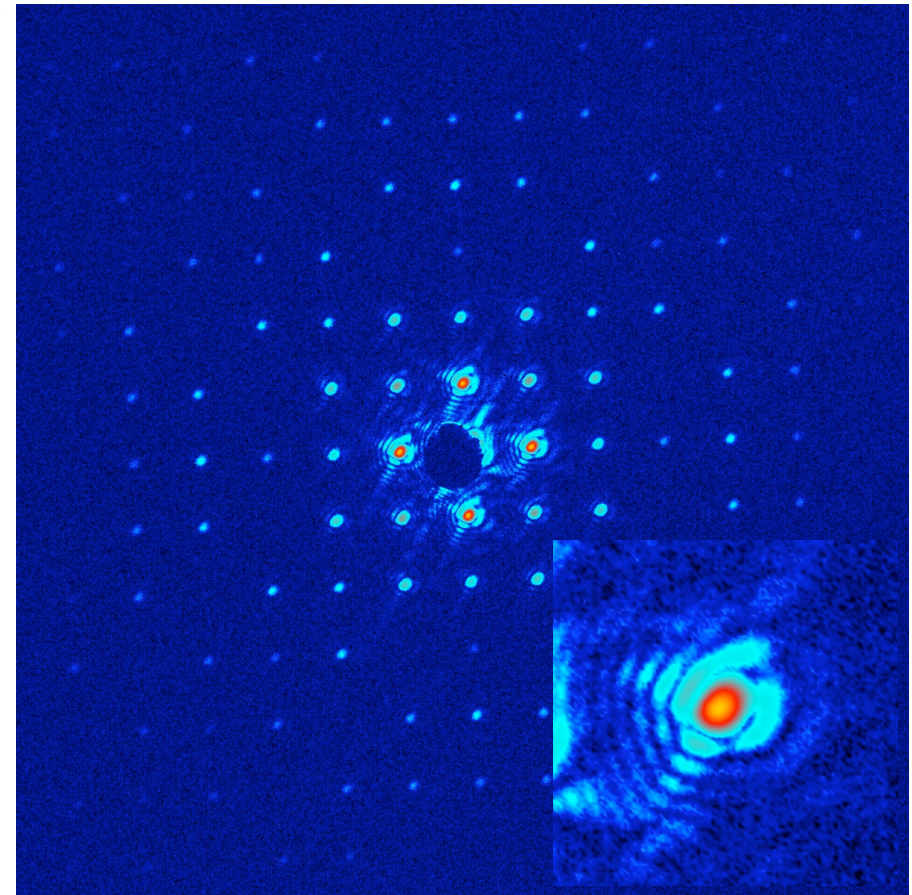
First use of FEL light in experimental chambers (2/2)



FEL radiation at 32.5 nm was used to start commissioning of DiProl experimental chamber. The FEL signal was filtered by using two Al filters (400 and 800 nm) which reduces the pulse intensity by more than an order of magnitude

Due to problems with the K-B focusing mirror the FEL spot was produced using a 20 μm pin-hole in front of the specimen.

Under these conditions coherent scattering images were recorded from a periodic array by integrating the signal for 10-100 seconds. This supposes that single shot CDI will be possible removing the filters and using the focusing K-B mirror



F. Capotondi, E. Pedersoli, R.H. Menk, M. Kiskinova and H. Chapman et al. (CFEL-DESY), J. Hajdu et al. (Uppsala), M. Bogan et al. (SLAC), M. Pivovarov, A. Nelson et al. (LLNL)

F. Capotondi, E. Pedersoli, M. Kiskinova, R.H. Menk, C. Svetina, S. Spampinati, S. Bassanese, E. Allaria

26/07/2011

- FEL-1
 - FEL-1 optimization is expected to be concluded in 2011.
 - Two projects are already underway to implement HHG sources as a possible seeds for FERMI at $\lambda \sim 30\text{nm}$ and more long term below 10 nm
 - When FEL-2 comes online FEL-1 could be temporarily configured for HHG tests (2013 and beyond).
- FEL-2
 - FEL-2 has been already shown to be very compatible with the ECHO scheme. A possible temporary modification could be done (if agreeable with users) for testing ECHO at the 50th or higher harmonic.

Success at FERMI has been the result of a concerted and unified effort by the entire FERMI team and the support staff at Sincrotrone Trieste. The physics commissioning team thanks all the people involved in the project (including consultants, guests and advisory committee members) that contributed to the design, construction and commissioning of FERMI over the the past 6 years.



My special thanks goes to the following people that contributed to FERMI's success by working on most of the commissioning shifts over the past two years:
 P. Craievich, S. Di Mitri, W. Fawley, L. Froehlich, G. De Ninno, G. Penco, S. Spampinati, C. Spezzani, M. Trovo'