

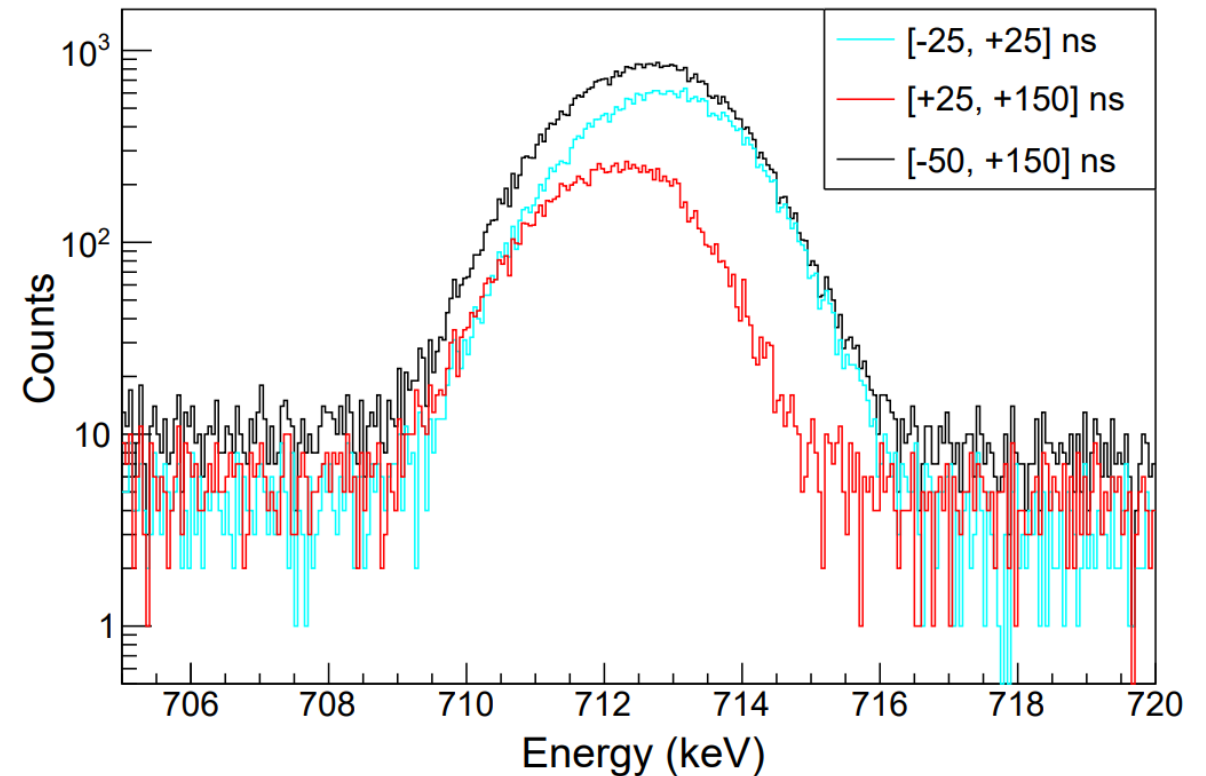
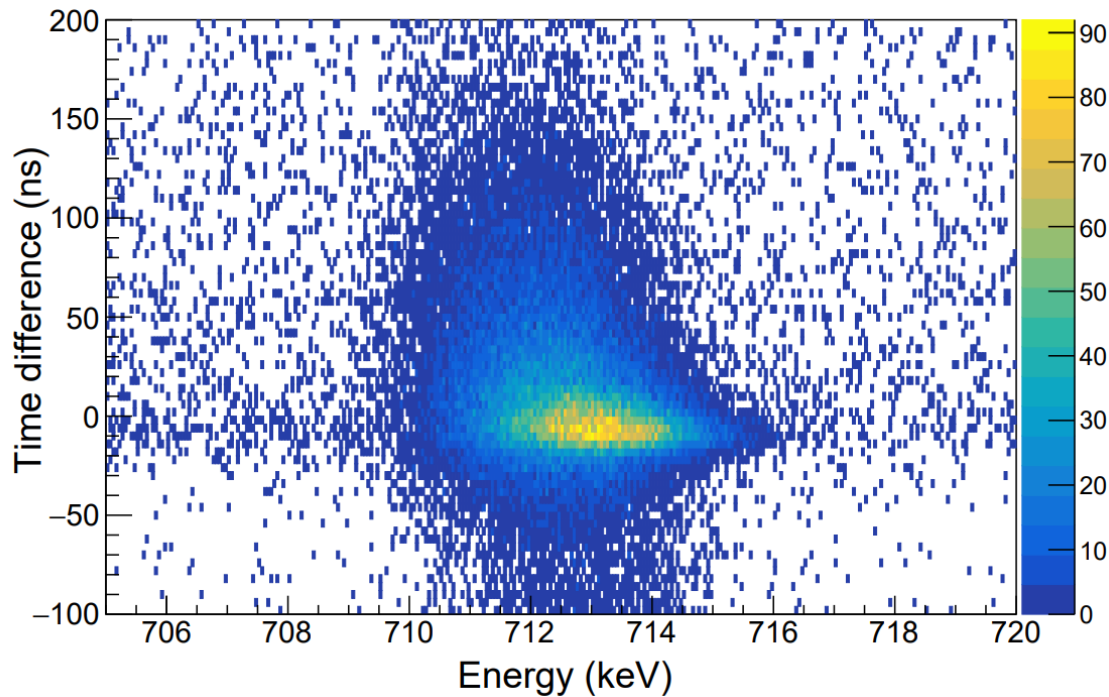
Update muX meeting 17/01

Michael Heines

Time cut systematic

Issue with bias time cuts

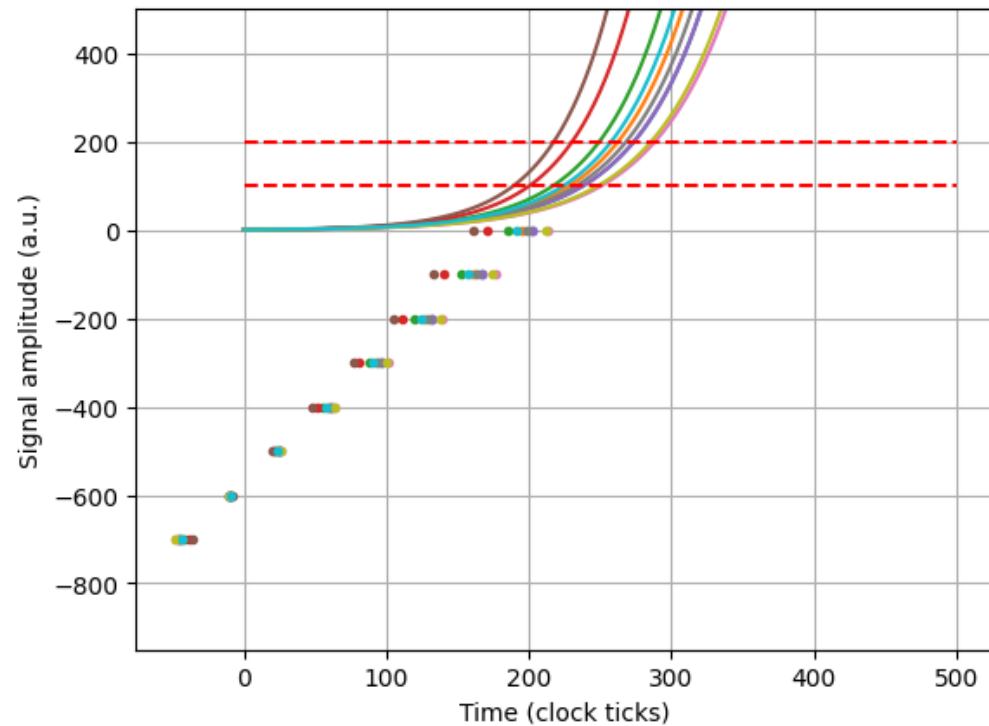
Centroid position is affected by choice of time cut!



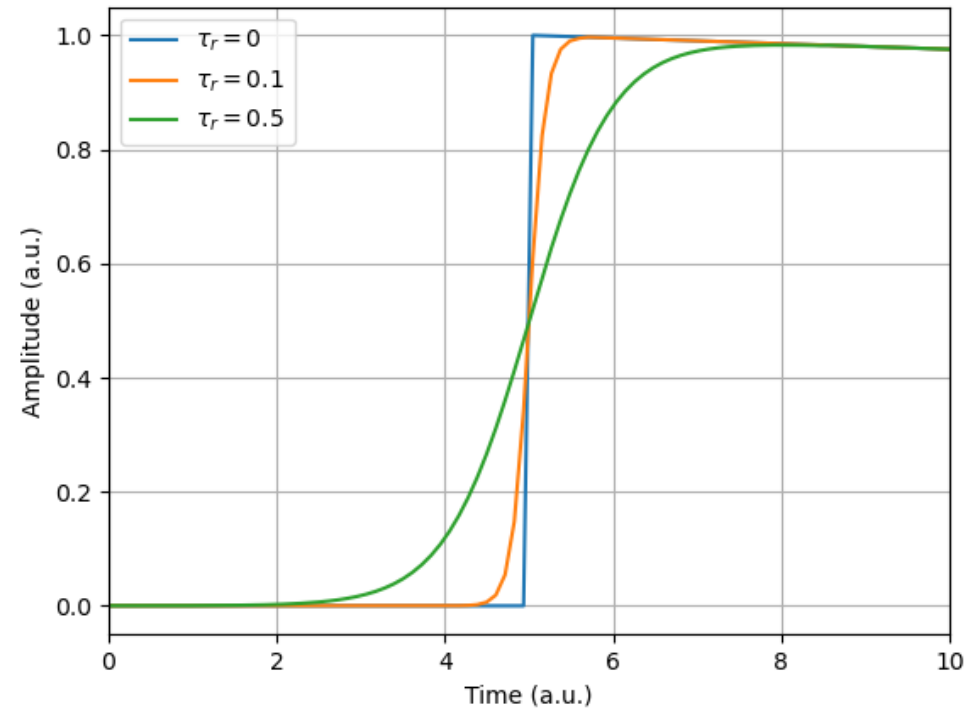
Ge09: Worst detector for this effect

Understanding the issue

- Slow rise time pulses are registered as “later”



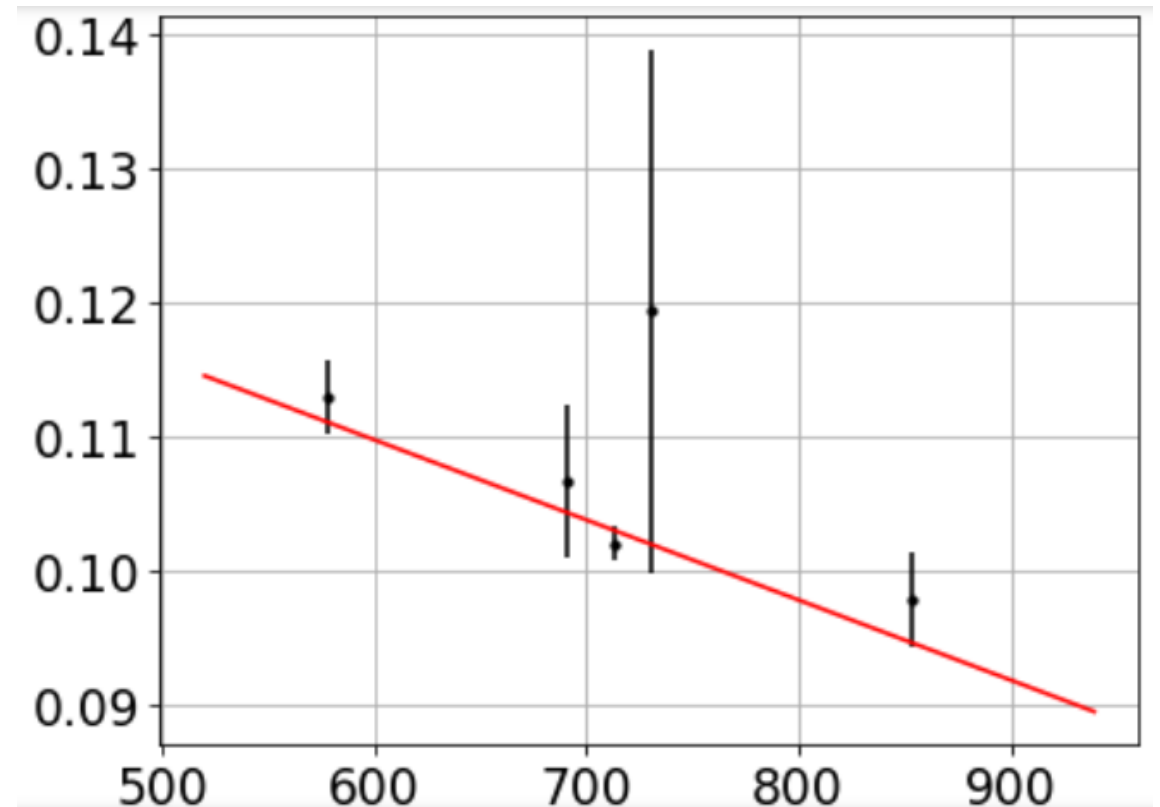
- Slow rise time pulses have lower energy



Correcting for it

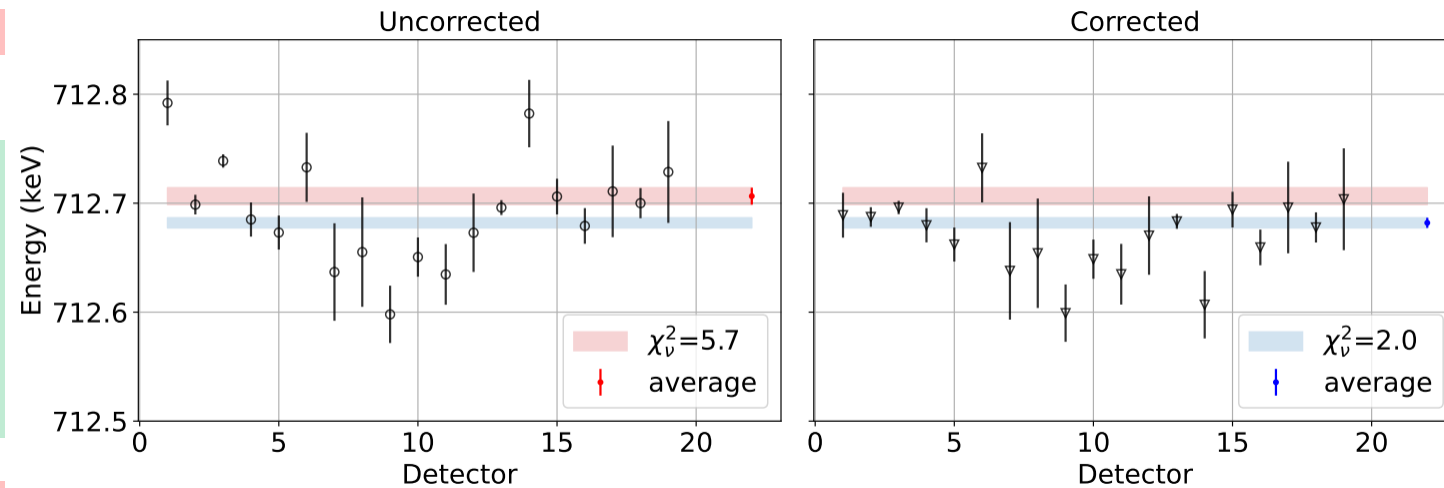
- Idea of correction:
 - Narrow time cut shifts output
 - Consider a broad time cut which takes all pulses
 - Probe the centroid difference in broad and narrow cut
 - Estimate shift at different energies → Linear fit

$$\delta = a_0 + a_1 E$$



Quite big effect

Detector	a_0 (eV)	a_1 (10^{-5})	shift in ^{39}K 2p1s (eV)
Ge01	145.6(111)	-5.97(157)	103.0(11)
Ge02	56.6(111)	-6.35(156)	11.3(6)
Ge03	86.1(108)	-6.08(154)	42.7(8)
Ge04	15.6(52)	-1.45(75)	5.3(4)
Ge05	40.4(31)	-4.12(47)	11.0(5)
Ge06A	14.9(118)	-2.02(168)	0.5(9)
Ge06B	-0.5(105)	-0.08(162)	-1.1(17)
Ge06C	16.3(155)	-2.1(22)	1.1(10)
Ge06D	-12.7(122)	1.63(177)	-1.1(12)
MB07A	1.0(85)	0.13(126)	1.9(10)
MB07B	14.8(67)	-2.09(95)	-0.2(4)
MB07C	21(24)	-2.5(34)	2.6(12)
Ge08	53.6(43)	-5.74(68)	12.7(8)
Ge09	236(24)	-8.5(32)	175(3)
Ge10	43.4(71)	-4.42(104)	12.0(8)
Ge11	45.6(44)	-3.64(66)	19.7(6)
Ge12	27.7(49)	-1.82(73)	14.8(6)
Ge13	52.0(40)	-4.18(62)	22.3(7)
Ge14	63.4(83)	-5.39(118)	25.0(6)

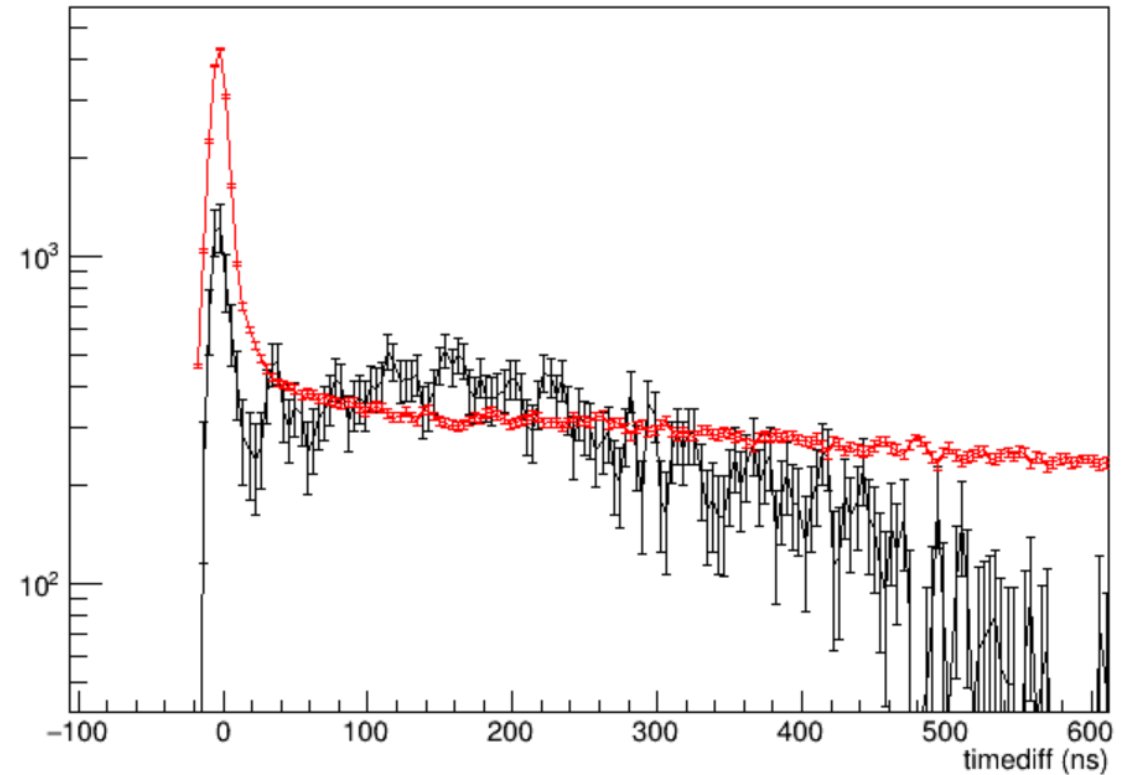
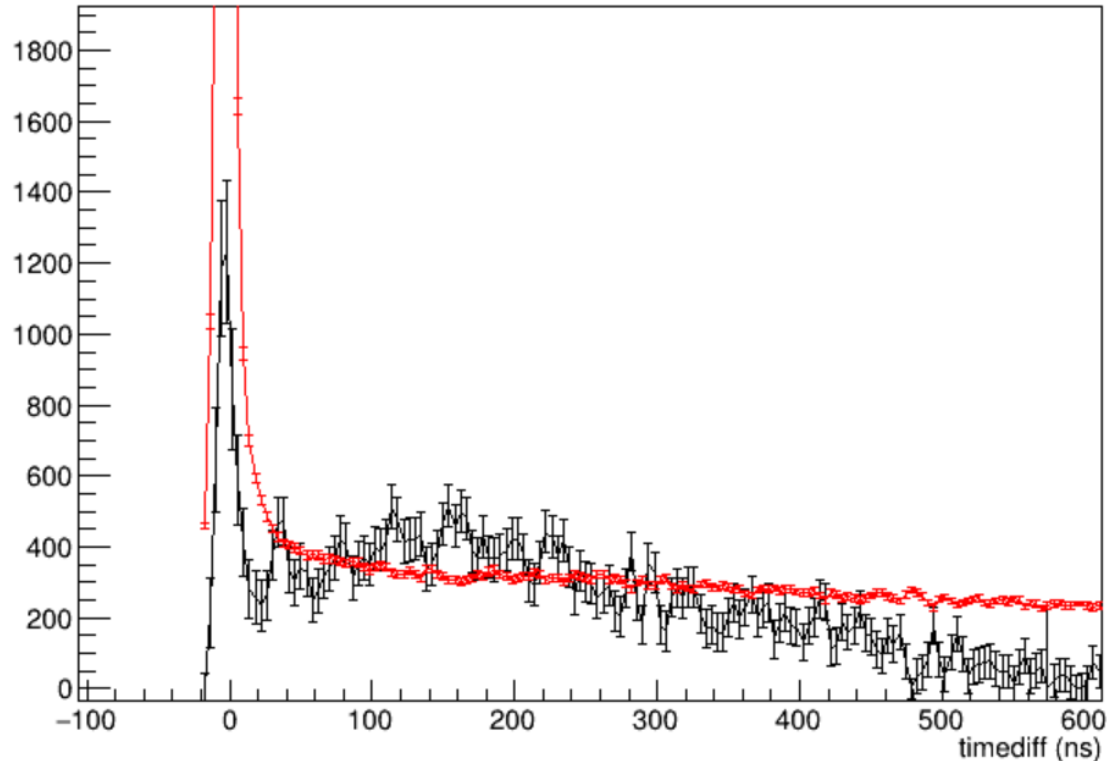


Ideal ^{40}K time cut

Time behaviour in ^{40}K data

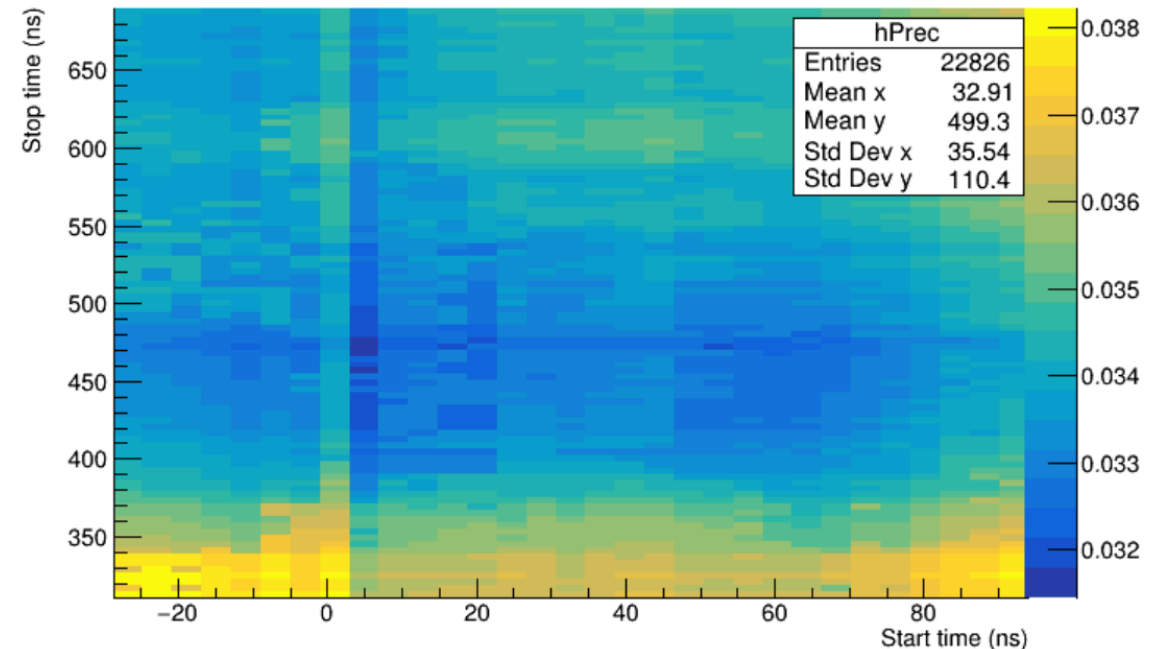
Red: Background per keV
Black: Signal integral

It makes sense to start after $t=0$ → Cut out Titanium
2p1s Compton continuum and edge



Best time cut for ^{40}K data

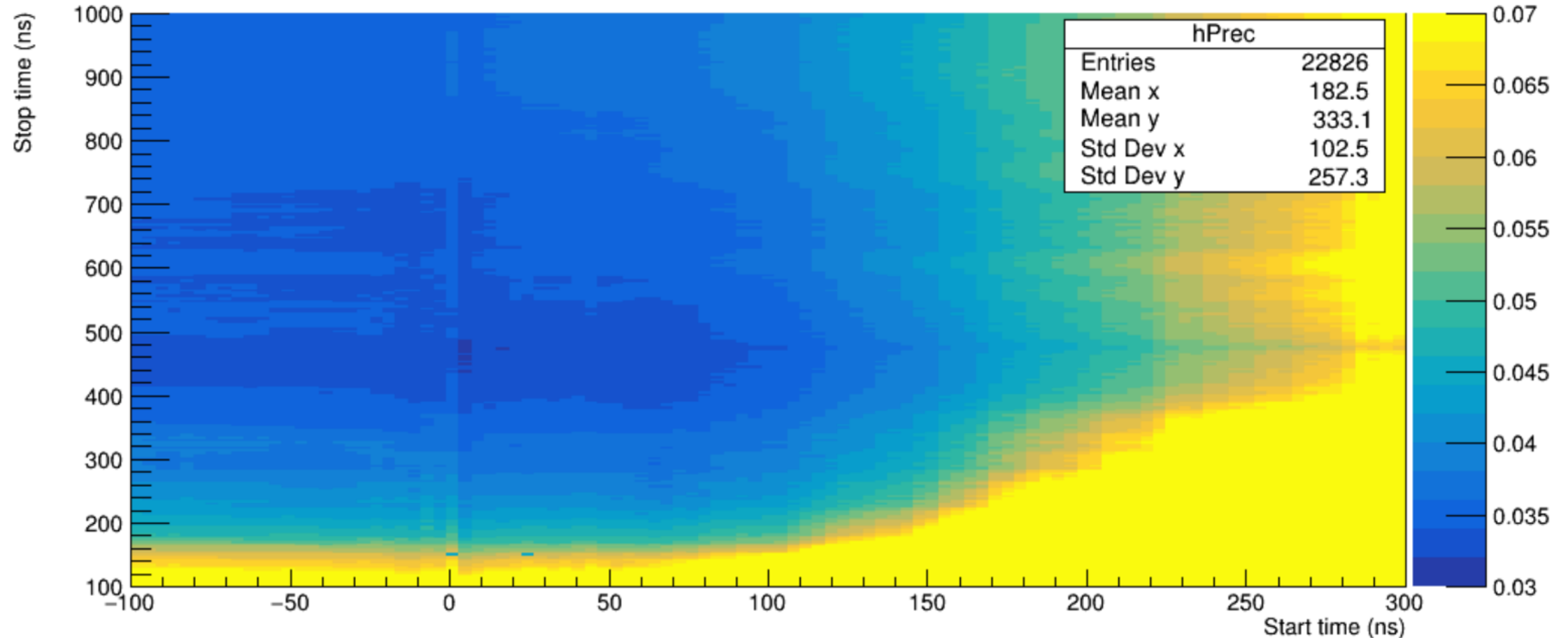
- Ideal start window: 2 local minima around 4ns and 60 ns
 - Minimal reduction in first local minimum
 - Opting for second to work with cleaner spectrum
- Ideal stop window: Consistent around 480 ns



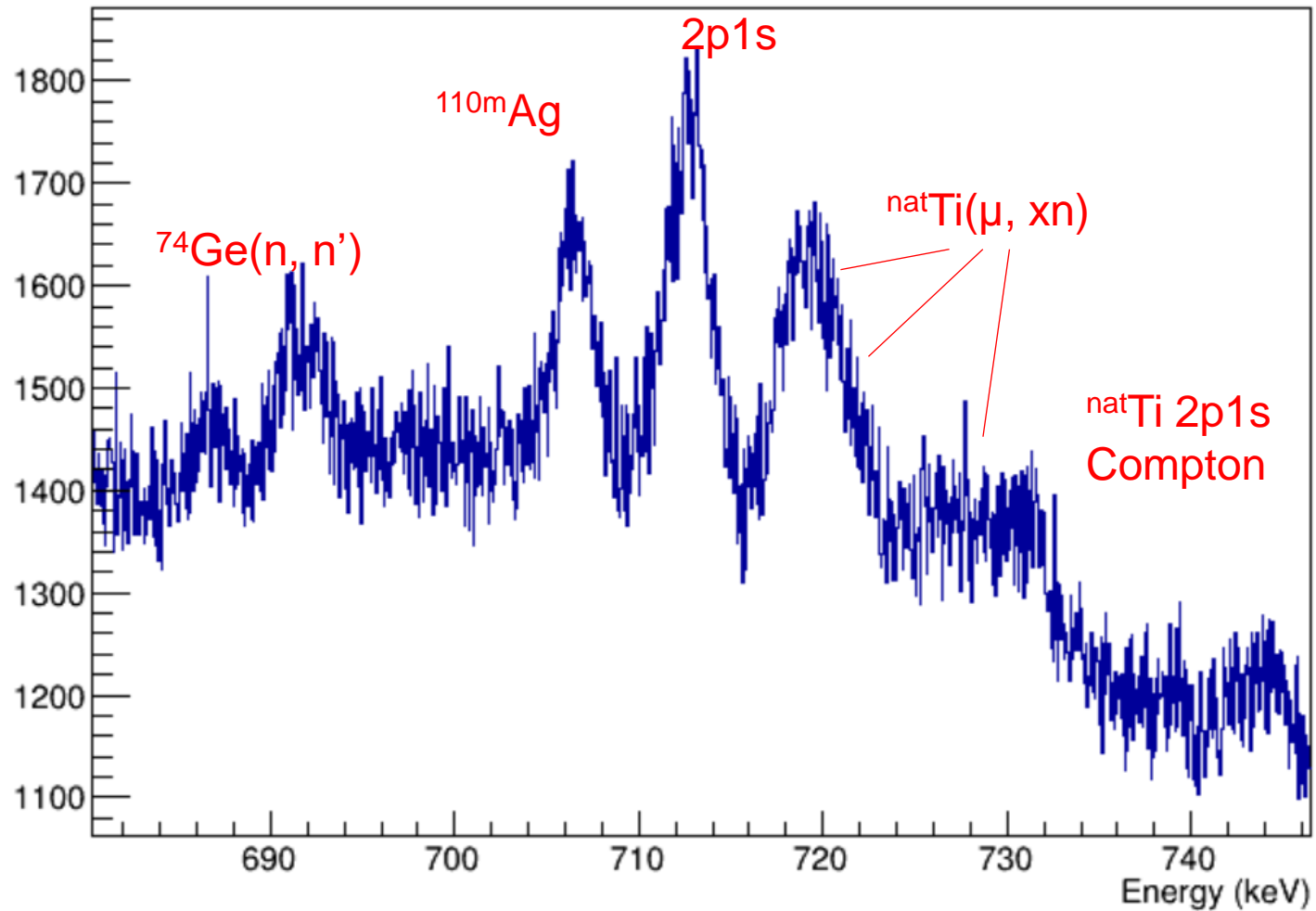
Best time cut for ^{40}K

Graph does not tell the fully story:

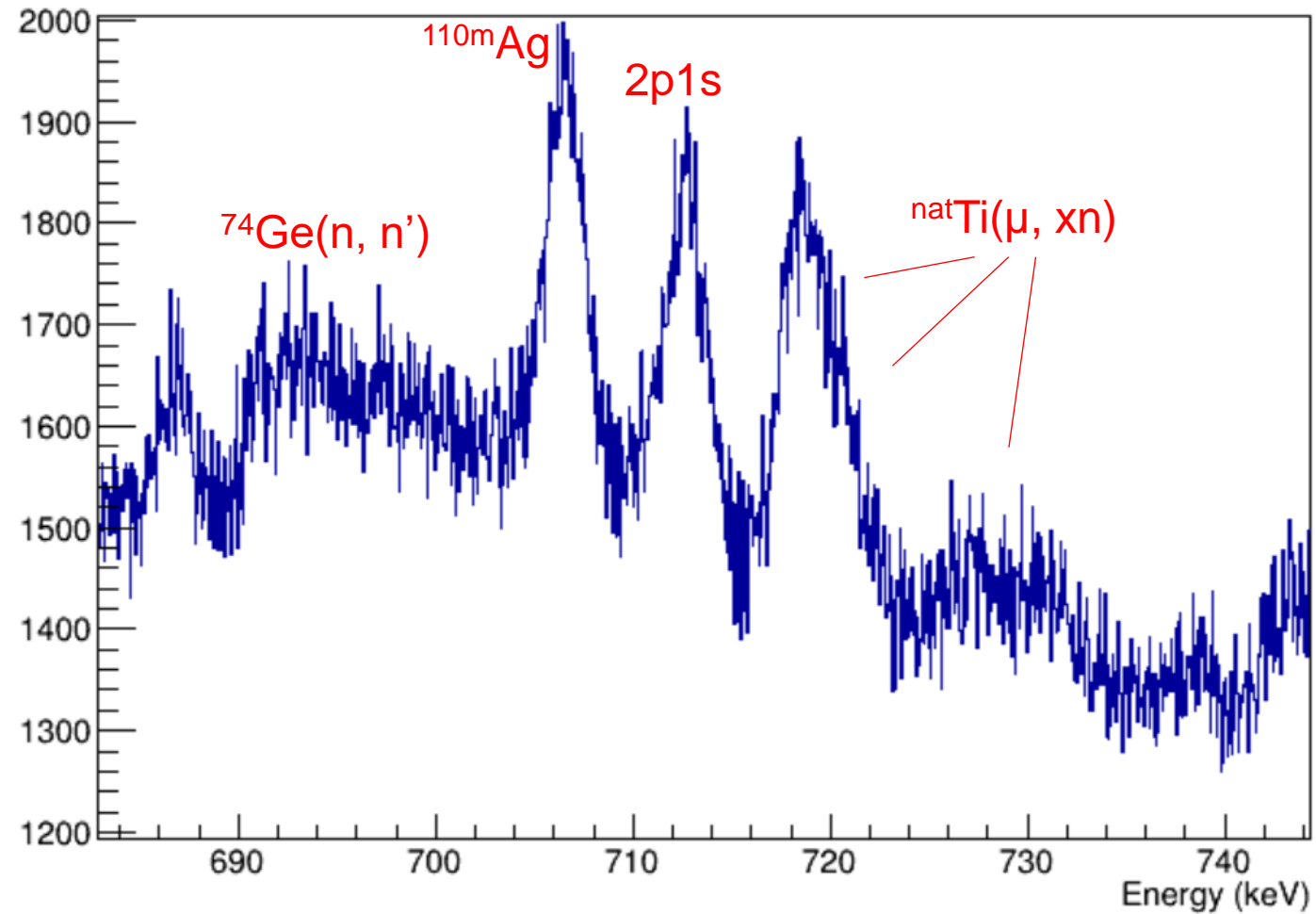
- Ti 2p1s Compton edge \rightarrow Almost fully suppressed delayed
- Capture lines \rightarrow Suppressed with delayed + shorter interval
- $^{74}\text{Ge}(n, n')$ Scattering structure \rightarrow Suppressed with delayed + shorter interval
- $^{110\text{m}}\text{Ag}$ calibration line \rightarrow Suppressed with shorter interval



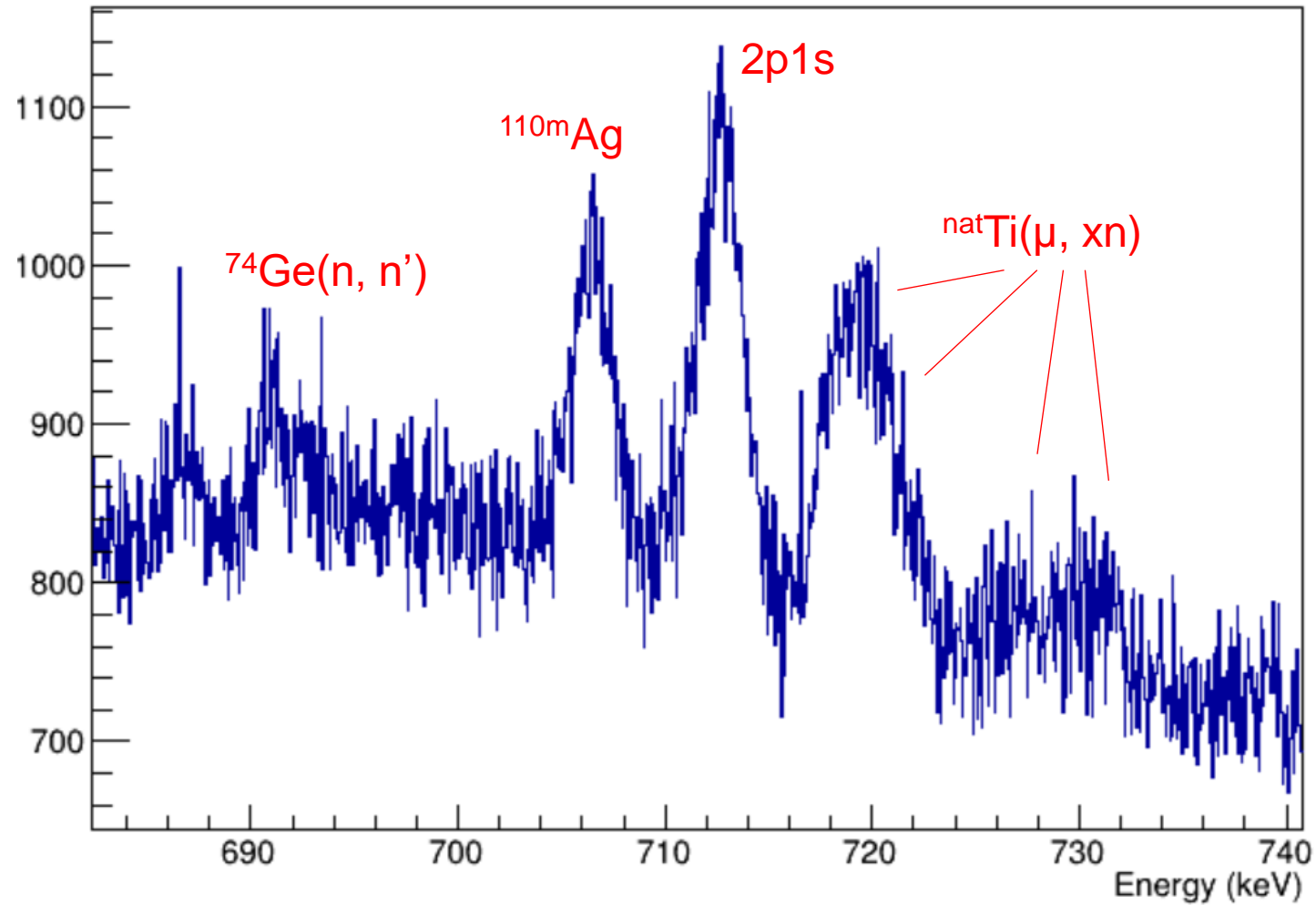
The standard [-50, 500] ns cut



Long delayed [50, 1000]ns cut



The finalized spectrum



Radius extraction

Nuclear polarization

- Contributions:
 - Nuclear part:
 - Low lying states (mostly uncorrelated between isotopes)
 - Giant resonance (highly correlated between isotopes)
 - Nucleon part:
 - nucleon polarization (highly correlated between isotopes)

NP uncertainties

- Error on nuclear part based on spread based off ^{40}Ca \rightarrow $\sim 20\%$
- Error on nucleon part based of Misha's estimates \rightarrow $\sim 10\%$
- Adding uncertainties linearly as a conservative estimate

- ^{35}Cl : NP = $103.9(208) + 14.8(15) = 118.7(223)$ eV
- ^{37}Cl : NP = $100.1(201) + 16.0(16) = 116.1(217)$ eV

- Relative uncertainty of $\sim 19\%$

NP differences

- Assume:
 - 100% correlation in giant resonance and nucleon parts
 - 0% correlation in low lying states part
- Will get exact partitioning soon, but based of rough estimates, the total correlation becomes ~91%
- $CI \Delta NP \approx 2.6(94) \text{ eV} \rightarrow \sim 8\%$ of largest NP value compared to Fricke's assumption of 10%

Abusing correlation for $\delta \langle r^2 \rangle$

$$\delta \langle r^2 \rangle = R_2^2 - R_1^2 = (R_2 - R_1)(R_2 + R_1)$$
$$\delta \langle r^2 \rangle = 2 R_1 \Delta R + (\Delta R)^2 = \frac{2R_{k\alpha,1}\Delta R}{V_{2,1}} + (\Delta R)^2$$

After doing some math:

$$\Delta R = \frac{1}{V_{2,2}} \left[\Delta R_{k\alpha} + R_{k\alpha,1} \frac{V_{2,1} - V_{2,2}}{V_{2,1}} \right]$$

Correlation matrices can be used to get best $\delta \langle r^2 \rangle$, likely limited to the relative precision of $\Delta R_{k\alpha}$ due to high level of correlation

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$$\delta \langle r^2 \rangle = R_2^2 - R_1^2 = (R_2 - R_1)(R_2 + R_1)$$

$$\delta \langle r^2 \rangle = \underbrace{2 R_1 \Delta R}_{\sim 1e-3} + \underbrace{(\Delta R)^2}_{\ll 2 R_1 \Delta R} = \frac{2 R_{k\alpha,1} \Delta R}{V_{2,1}} + (\Delta R)^2$$

After doing some math:

$$\Delta R = \frac{1}{V_{2,2}} \left[\underbrace{\Delta R_{k\alpha}}_{O(e-2)} + R_{k\alpha,1} \frac{V_{2,1} - V_{2,2}}{V_{2,1}} \right]$$

$\sim 1e-3$ ~ 4 $\sim 5(4)e-4$
 $O(e-2)$ $O(e-3)$

Based on estimates, left term about 10x bigger than right term

Correlation matrices can be used to get best $\delta \langle r^2 \rangle$, likely limited to the relative precision of $\Delta R_{k\alpha}$ due to high level of correlation

