

Particle physics experiments with slow neutrons

- Motivation
- Some experiments (mainly beta decay)
- Cold or ultracold?

Torsten Soldner
Institut Laue Langevin
Grenoble, France

Strong

- Nuclear force (n,γ), (n,p)...
- Neutron optical potential
- Neutron polarisability
→ *Potential of quark confinement*

Weak

- *Neutron beta decay*
- Weak hadronic processes
→ p(n_{pol},γ)d, ⁶Li(n_{pol},t)α ...
→ neutron weak spin rotation

Electromagnetic

- Magnetic moment
↔ quark magnetic moment
- Neutron neutrality
→ *Charge quantisation* (SM++, GUTs, magnetic monopoles?)

Gravitation

- Newton: $V = -G \frac{m_n M}{r}$

but: Falling quantum object

GUTs?

- + favourable symmetry breaking:
- observable $n\bar{n}$

Additional heavy bosons?

- Right-handed, leptoquark, exotics
- Modification of V—A weak interaction
- New CP violation

SUSY? Additional higgs fields?...

- New CP violating phases
- EDM
- CP violation in decay

Short range forces?

- Light bosons (axions...)?
- Extra dimensions?
- $V \propto \frac{\alpha e^{-r/\lambda}}{r}$
- Spin-dependent interactions

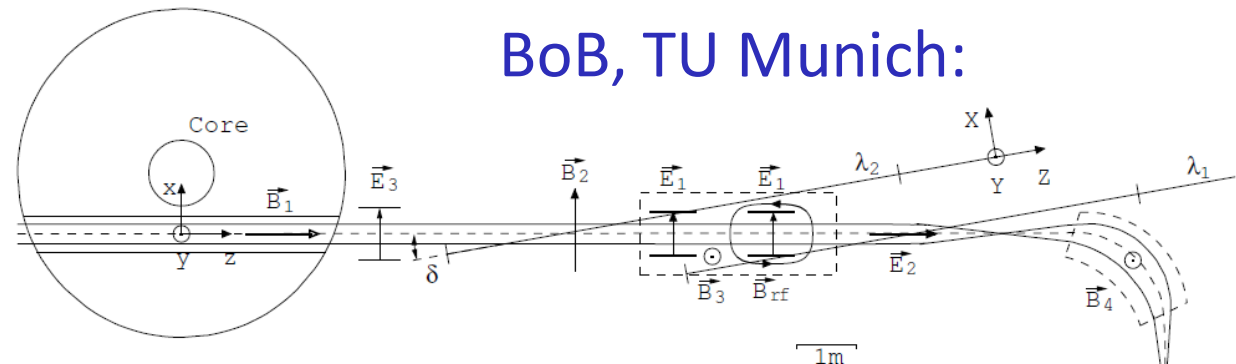
$$n \rightarrow p e \bar{\nu}_e \quad 99.7\%$$

$$n \rightarrow p e \bar{\nu}_e \gamma \quad 3.13(35) \cdot 10^{-3}$$

$$n \rightarrow H \bar{\nu}_e \quad 4 \cdot 10^{-6} \text{ (predicted)}$$

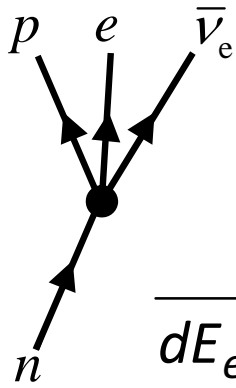
$\bar{\nu}_e$	n	p	e	$ m_s m_l\rangle$	W [%]
←	←	←	→	$ +-\rangle$	44.14(5)
←	←	→	←	$ -\!+\rangle$	55.24(4)
←	→	→	→	$ ++\rangle$	0.622(11)
→	←	←	←	$ --\rangle$	0
→	→	→	←	$ -\!+\rangle$	0
→	→	←	→	$ +-\rangle$	0

- Next slide
- Precise branching ratio?
- Photon spectrum?
- Angular correlations?
- **Talk S. Gardner, Tuesday**
- Signature: $T_H = 326 \text{ eV}$
- First detection?
- Occupation of hyperfine states?



W. Schott et al, EPJA 30 (2006) 603, **Poster** CsI(Tl)

Neutron beta decay and weak interaction



$$H = \frac{G_w}{2\sqrt{2}} \sum_{i=V,A,S,T} \{ L_i(\bar{p}\mathbf{O}_i n) (\bar{e}\mathbf{O}_i(1-\gamma_5)\nu) + R_i(\bar{p}\mathbf{O}_i n) (\bar{e}\mathbf{O}_i(1+\gamma_5)\nu) \}$$

$$\frac{dW}{dE_e d\Omega_e d\Omega_\nu} \propto G_e(E_e) \left\{ 1 + a \frac{\mathbf{p}_e \mathbf{p}_\nu}{E_e E_\nu} + b \frac{m_e}{E_e} + \mathbf{P} \left(A \frac{\mathbf{p}_e}{E_e} + B \frac{\mathbf{p}_\nu}{E_\nu} + D \frac{\mathbf{p}_e \times \mathbf{p}_\nu}{E_e E_\nu} \right) \right\}$$

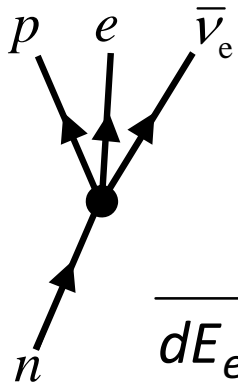
$$a = \frac{N_{\uparrow\uparrow} - N_{\uparrow\downarrow}}{N_{\uparrow\uparrow} + N_{\uparrow\downarrow}} \quad \uparrow\uparrow \quad = \frac{|L_V|^2 - |L_A|^2 - |L_S|^2 + |L_T|^2 + |R_V|^2 - |R_A|^2 - |R_S|^2 + |R_T|^2}{|L_V|^2 + 3|L_A|^2 + |L_S|^2 + 3|L_T|^2 + |R_V|^2 + 3|R_A|^2 + |R_S|^2 + 3|R_T|^2}$$

$$A = \frac{N_{\uparrow\uparrow\uparrow} - N_{\uparrow\uparrow\downarrow}}{N_{\uparrow\uparrow\uparrow} + N_{\uparrow\uparrow\downarrow}} \quad \uparrow\uparrow\uparrow \quad = A(L_V, L_A, L_S, L_T, R_V, R_A, R_S, R_T)$$

$$B \quad \uparrow\uparrow \quad = B(L_V, L_A, L_S, L_T, R_V, R_A, R_S, R_T)$$

$$D \quad \uparrow\uparrow \quad = D(L_V, L_A, L_S, L_T, R_V, R_A, R_S, R_T)$$

Neutron beta decay and V—A theory



$$H = \frac{G_W}{2\sqrt{2}} \sum_{i=V,A} L_i (\bar{p} \mathbf{O}_i n) (\bar{e} \mathbf{O}_i (1 - \gamma_5) \nu)$$

$$\lambda = \frac{L_A}{L_V}$$

$$\frac{dW}{dE_e d\Omega_e d\Omega_\nu} \propto G_e(E_e) \left\{ 1 + a \frac{\mathbf{p}_e \mathbf{p}_\nu}{E_e E_\nu} + b \frac{m_e}{E_e} + \mathbf{P} \left(A \frac{\mathbf{p}_e}{E_e} + B \frac{\mathbf{p}_\nu}{E_\nu} + D \frac{\mathbf{p}_e \times \mathbf{p}_\nu}{E_e E_\nu} \right) \right\}$$

$$a = \frac{N_{\uparrow\uparrow} - N_{\uparrow\downarrow}}{N_{\uparrow\uparrow} + N_{\uparrow\downarrow}} \quad \uparrow\uparrow \quad = a(L_V, L_A, L_S, L_T, R_V, R_A, R_S, R_T)$$

$$a = \frac{1 - |\lambda|^2}{1 + 3|\lambda|^2}$$

$$A = \frac{N_{\uparrow\uparrow} - N_{\uparrow\downarrow}}{N_{\uparrow\uparrow} + N_{\uparrow\downarrow}} \quad \uparrow\uparrow \quad = A(L_V, L_A, L_S, L_T, R_V, R_A, R_S, R_T)$$

$$A = -2 \frac{|\lambda|^2 + \text{Re}(\lambda)}{1 + 3|\lambda|^2}$$

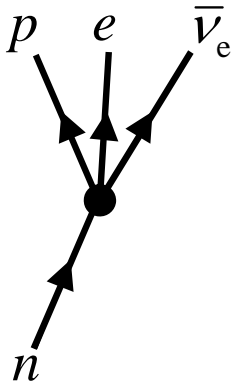
$$B = \quad \uparrow\uparrow \quad = B(L_V, L_A, L_S, L_T, R_V, R_A, R_S, R_T)$$

$$B = 2 \frac{|\lambda|^2 - \text{Re}(\lambda)}{1 + 3|\lambda|^2}$$

$$D = \quad \uparrow \quad = D(L_V, L_A, L_S, L_T, R_V, R_A, R_S, R_T)$$

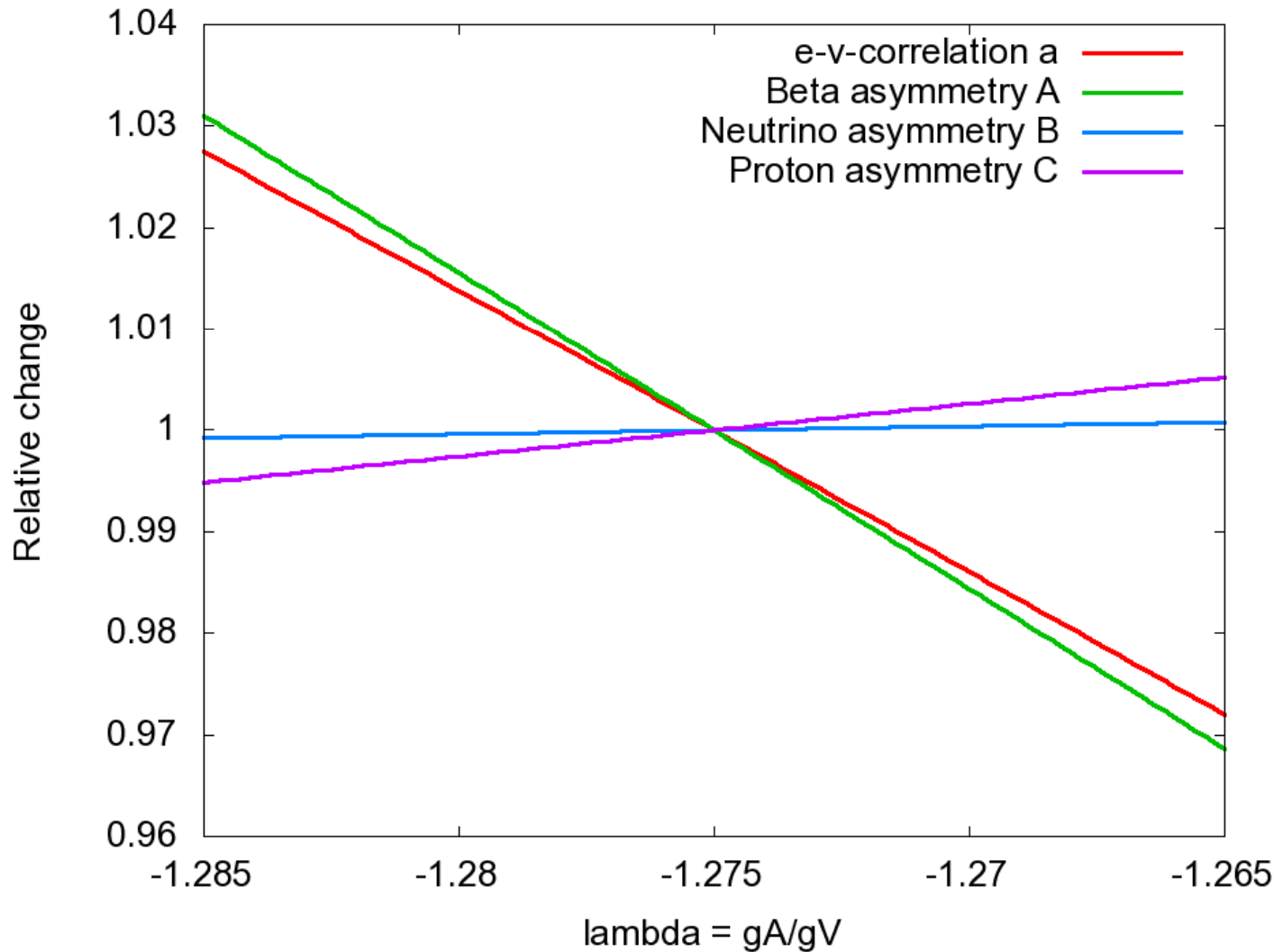
$$D = 2 \frac{\text{Im}(\lambda)}{1 + 3|\lambda|^2}$$

Neutron beta decay and V–A theory



$$H = \frac{G_W}{2\sqrt{2}} \sum_{i=V,A} L_i (\bar{p} \mathbf{O}_i n) (\bar{e} \mathbf{O}_i (1 - \gamma_5) \nu)$$

$$\lambda = \frac{L_A}{L_V}$$



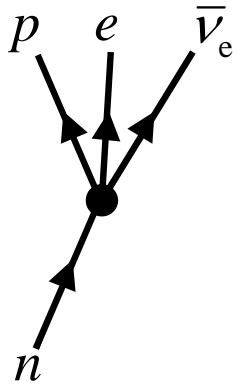
$$a = \frac{1 - |\lambda|^2}{1 + 3|\lambda|^2}$$

$$A = -2 \frac{|\lambda|^2 + \text{Re}(\lambda)}{1 + 3|\lambda|^2}$$

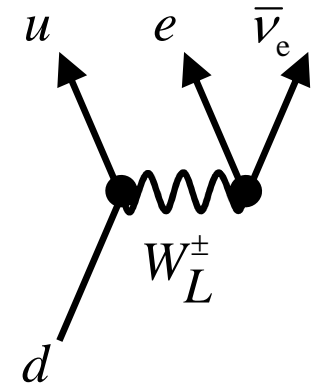
$$B = 2 \frac{|\lambda|^2 - \text{Re}(\lambda)}{1 + 3|\lambda|^2}$$

$$C = -x_C \frac{4\text{Re}(\lambda)}{1 + 3|\lambda|^2}$$

Neutron beta decay and the Standard Model



$$H = \frac{G_F}{\sqrt{2}} V_{ud} \bar{p} \gamma^\mu (1 + \lambda \gamma_5) n \bar{e} \gamma_\mu (1 - \gamma_5) \nu$$



$$H = \frac{G_F}{\sqrt{2}} V_{ud} \bar{u} \gamma^\mu (1 - \gamma_5) d \bar{e} \gamma_\mu (1 - \gamma_5) \nu$$

$$\begin{pmatrix} u_L \\ d_L \end{pmatrix} \quad \begin{pmatrix} e_L \\ \nu_{eL} \end{pmatrix}$$

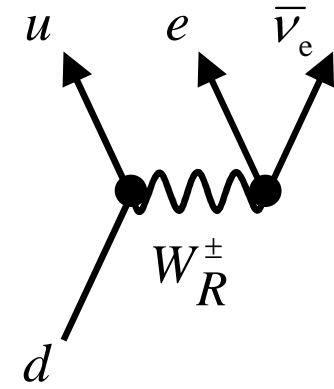
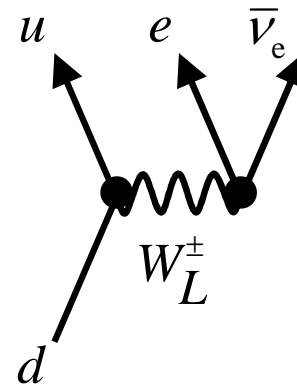
$$\tau = \frac{4908.7(1.9)s}{|V_{ud}|^2 (1 + 3\lambda^2)}$$

$$A = -2 \frac{|\lambda|^2 + \text{Re}(\lambda)}{1 + 3|\lambda|^2}$$

Right-handed interactions (V+A)

$$\begin{pmatrix} W_L \\ W_R \end{pmatrix} = \begin{pmatrix} \cos \zeta & -\sin \zeta \\ \sin \zeta & \cos \zeta \end{pmatrix} \begin{pmatrix} W_1 \\ W_2 \end{pmatrix}$$

$$(a, A, B, \tau) \rightarrow \left(\lambda', \zeta, \delta = \frac{m_1^2}{m_2^2} \right)$$



(broken) $SU(2)_L \times SU(2)_R$

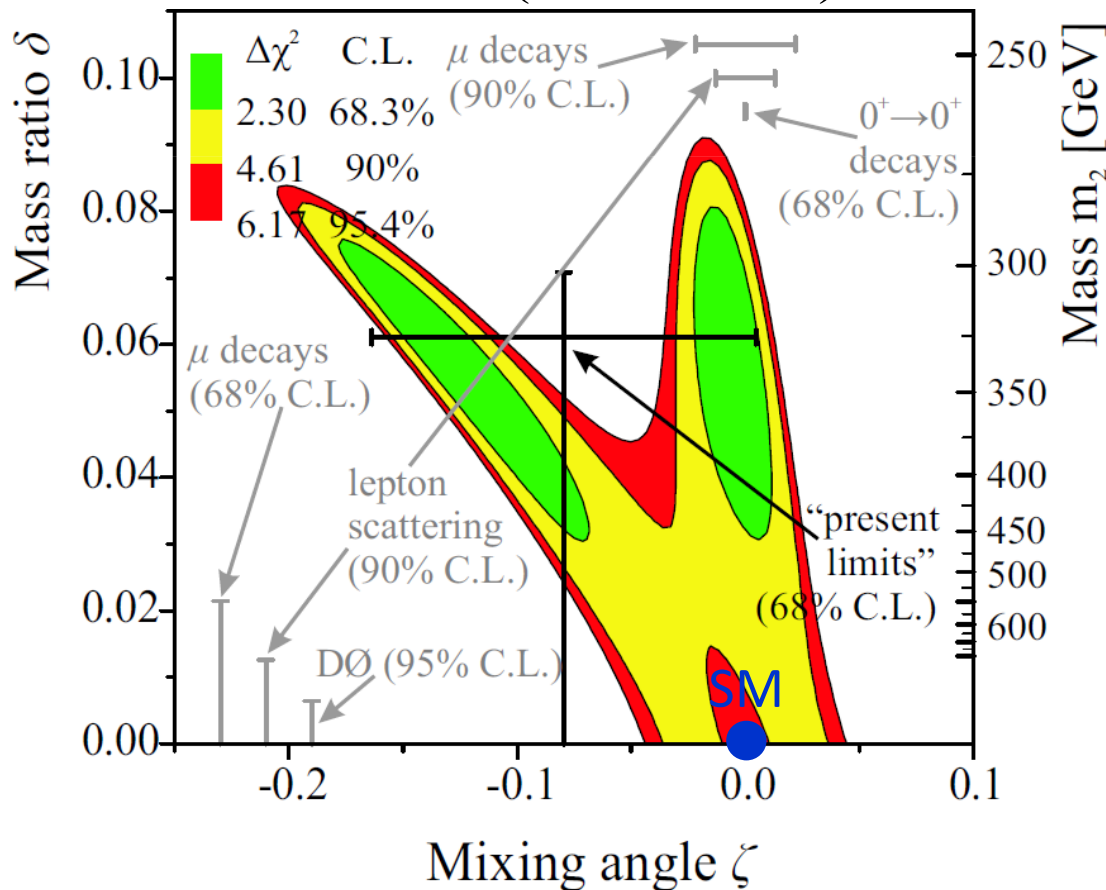
→ deviation from maximal parity violation (V+A)

→ also: additional phases for CP violation

Other analyses:

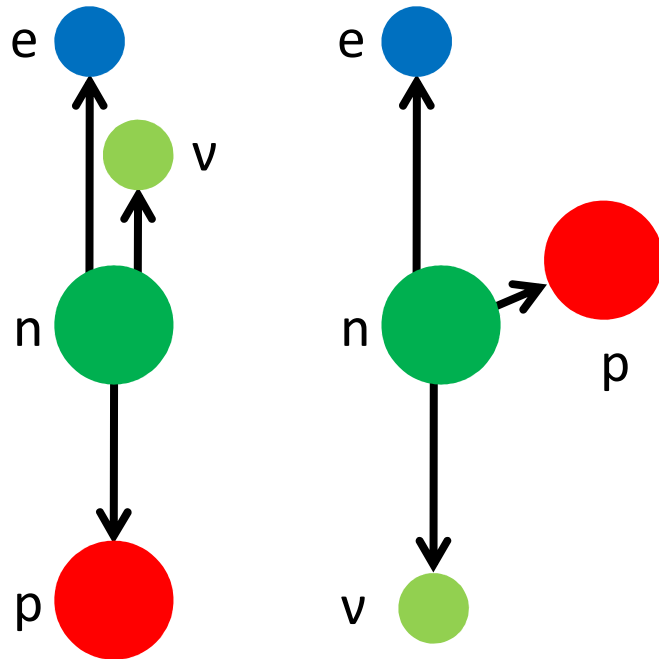
- Left-handed scalar and tensor
- Right-handed scalar and tensor

G. Konrad et al, arxiv:1007.3027



$$dW \propto 1 + a \frac{\mathbf{p}_e \mathbf{p}_\nu}{E_e E_\nu}$$

$$a = \frac{N_{\uparrow\uparrow} - N_{\uparrow\downarrow}}{N_{\uparrow\uparrow} + N_{\uparrow\downarrow}}$$

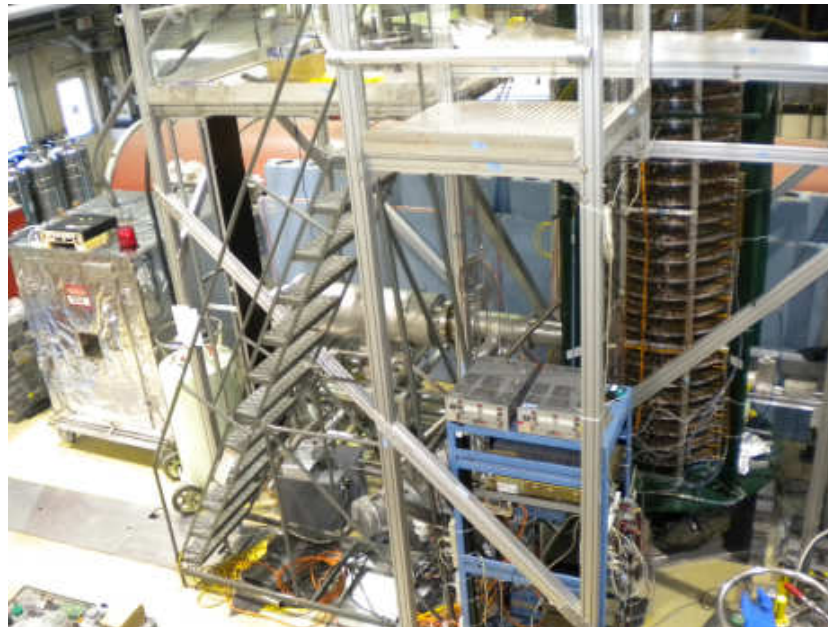


[PDG10] $a = -0.103 \pm 0.004$
(4%)

Methods

Goal: $\Delta a/a = O(10^{-3})$

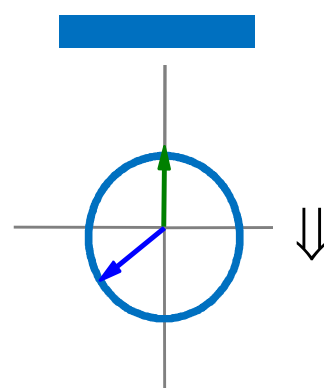
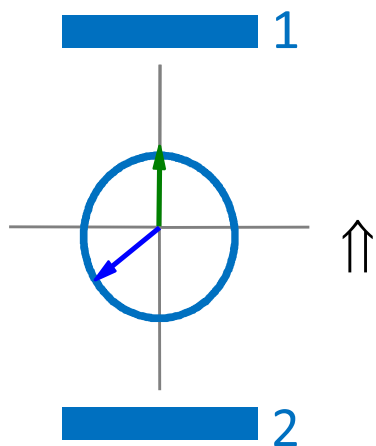
- **Proton spectrum 0...751 eV**
aSPECT (Mainz, Munich, ILL)
Talk G. Konrad, Tuesday
- **Correlation e-p, proton TOF**
aCORN, running at NIST
Nab, in progress



aCORN@NG6 (NIST)
Courtesy F.E. Wietfeldt

$$dW \propto 1 + AP \frac{p_e}{E_e}$$

$$A = \frac{N_{\uparrow\uparrow} - N_{\uparrow\downarrow}}{N_{\uparrow\uparrow} + N_{\uparrow\downarrow}}$$



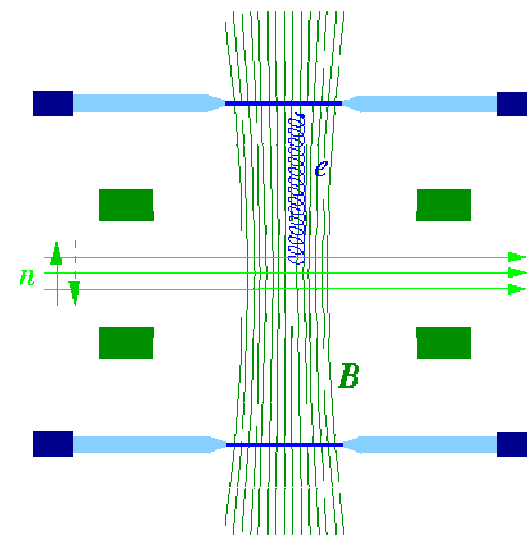
$$Ak\beta = \frac{N_{\uparrow} - N_{\downarrow}}{N_{\uparrow} + N_{\downarrow}}$$

$$kA\beta = \frac{N_{\uparrow} - N_{\downarrow}}{N_{\uparrow} + N_{\downarrow}}$$

$$k = \left\langle FP \frac{p_e}{E_e} \right\rangle$$

Requirements

- Solid angle
- Statistics
- Polarisation / Flipping
- Background
- Energy calibration



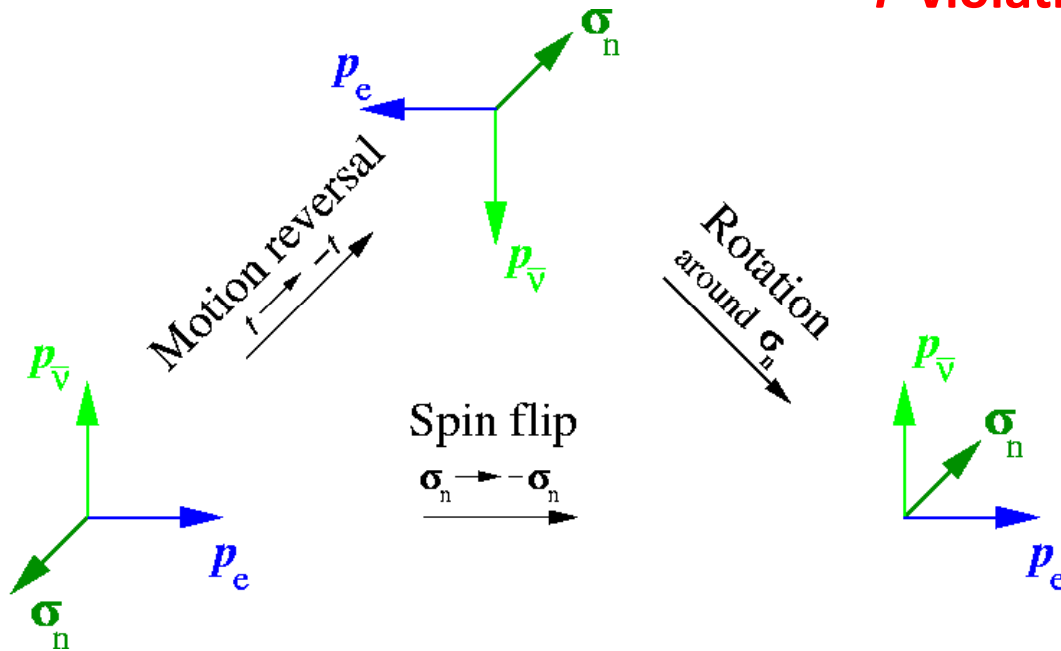
2x2π solid angle with
Magnetic field

Time reversal violation correlation D

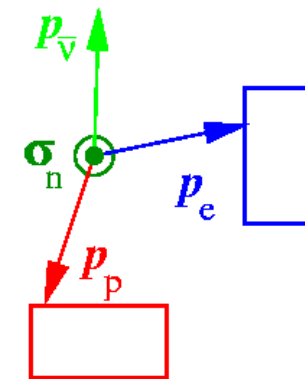
$D = 0$ in SM

$$\frac{dW}{dE_e d\Omega_e d\Omega_{\bar{\nu}}} \propto G_e(E_e) \cdot \left\{ 1 + a \frac{\mathbf{p}_e \mathbf{p}_{\bar{\nu}}}{E_e E_{\bar{\nu}}} + b \frac{m_e}{E_e} + \frac{\boldsymbol{\sigma}_n}{\sigma_n} \left(A \frac{\mathbf{p}_e}{E_e} + B \frac{\mathbf{p}_{\bar{\nu}}}{E_{\bar{\nu}}} \right) + D \frac{\boldsymbol{\sigma}_n}{\sigma_n} \cdot \frac{\mathbf{p}_e \times \mathbf{p}_{\bar{\nu}}}{E_e E_{\bar{\nu}}} \right\}$$

P violation \Leftrightarrow Asymmetry with spin-flip



Principle Set-Up



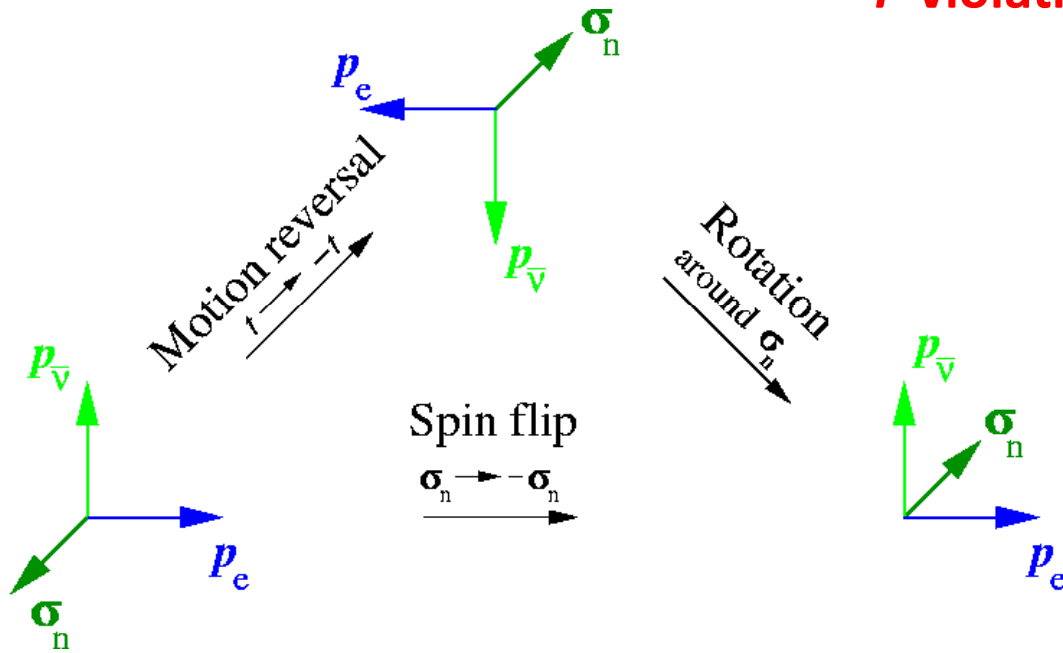
$$\alpha = \frac{N_{ep}^{\uparrow} - N_{ep}^{\downarrow}}{N_{ep}^{\uparrow} + N_{ep}^{\downarrow}} = DP\kappa_D + AP\kappa_A + BP\kappa_B$$

Time reversal violation correlation D

$D = 0$ in SM

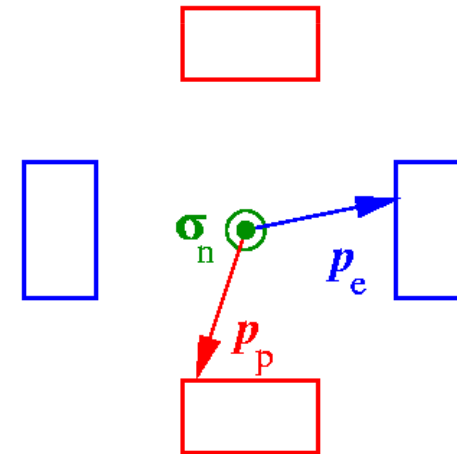
$$\frac{dW}{dE_e d\Omega_e d\Omega_{\bar{\nu}}} \propto G_e(E_e) \cdot \left\{ 1 + a \frac{\mathbf{p}_e \mathbf{p}_{\bar{\nu}}}{E_e E_{\bar{\nu}}} + b \frac{m_e}{E_e} + \frac{\sigma_n}{\sigma_n} \left(A \frac{\mathbf{p}_e}{E_e} + B \frac{\mathbf{p}_{\bar{\nu}}}{E_{\bar{\nu}}} \right) + D \frac{\sigma_n}{\sigma_n} \cdot \frac{\mathbf{p}_e \times \mathbf{p}_{\bar{\nu}}}{E_e E_{\bar{\nu}}} \right\}$$

P violation \Leftrightarrow Asymmetry with spin-flip



**Breaking of detector symmetry
 \Rightarrow Systematic effects**

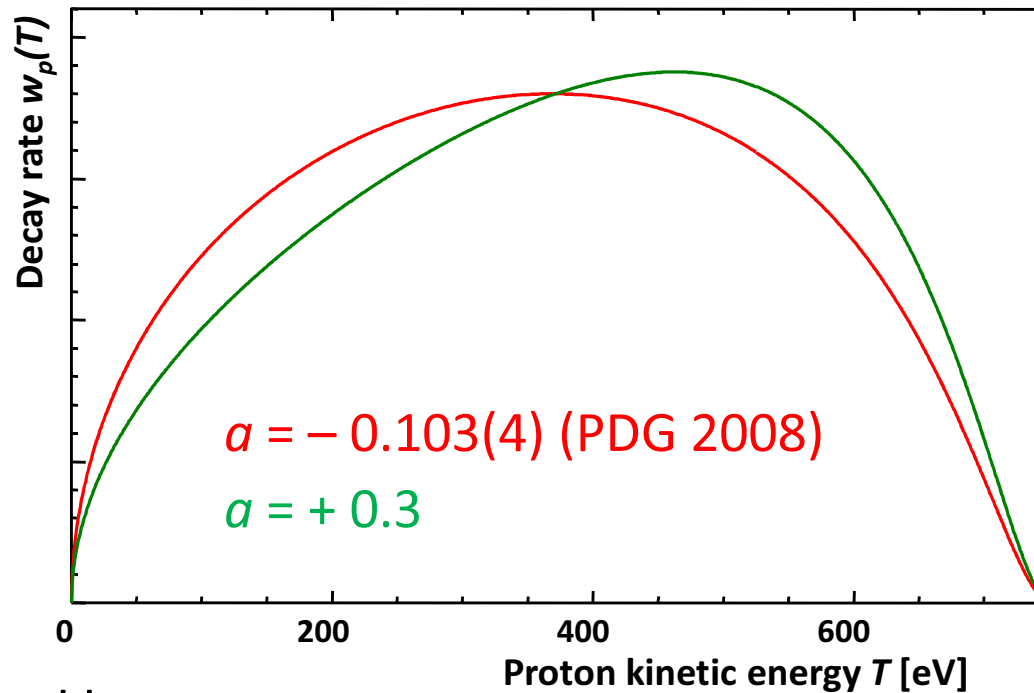
Principle Set-Up



$$D = \frac{\alpha^{00} - \alpha^{01} - \alpha^{10} + \alpha^{11}}{4PK_D^{00}}$$

Cold or ultracold? Velocity!

	Cold		Ultracold	
Velocity		800 m/s		5 m/s
Doppler eff.	$E_p = 751 \text{ eV}$	+ 4 eV	$E_p = 751 \text{ eV}$	+ 0.02 eV



Compensated by geometry
and 4π proton detection

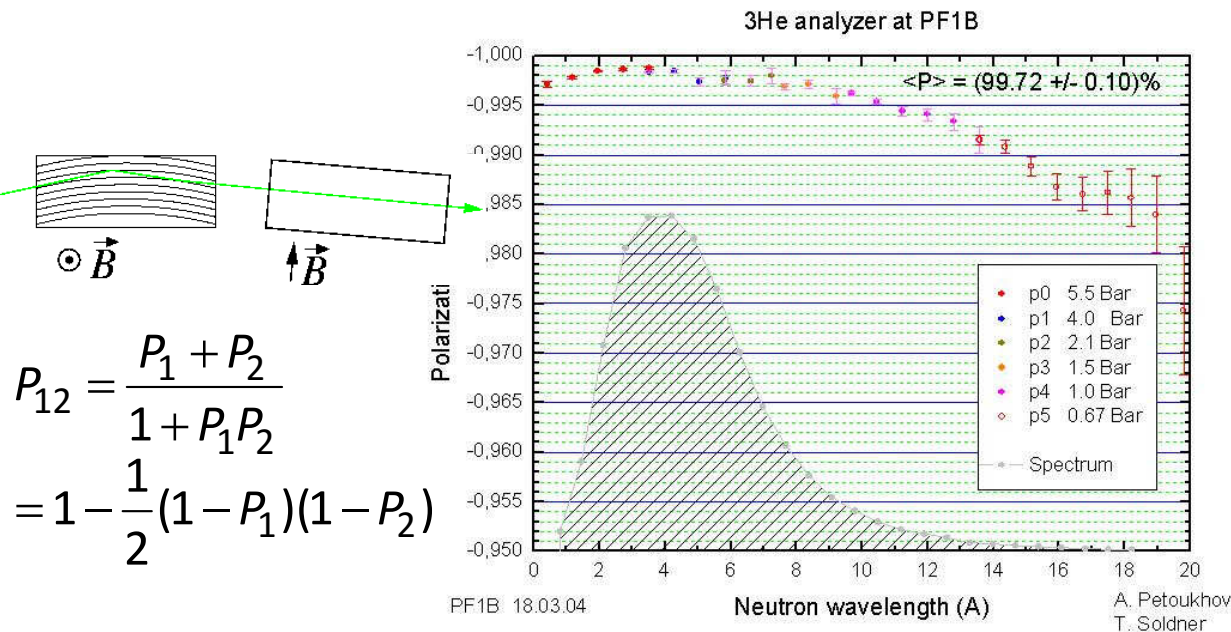
$v \times E$

large

small

Cold or ultracold? Statistics! Polarisation!

	Cold	Ultracold
Statistics – Density	In guide: $10^5/\text{cm}^3$ Polarised 98%: $10^4/\text{cm}^3$ Polarised 99.7%, background optimised: $3 \cdot 10^2/\text{cm}^3$	Demonstrated: $100/\text{cm}^3$ Promised: $10^3 \dots 10^5/\text{cm}^3$ Polarised 90% PF2 EDM: $1/\text{cm}^3$
Polarisation	99.7(1)%	Believed: 100% Demonstrated (UCNA): >99.5%



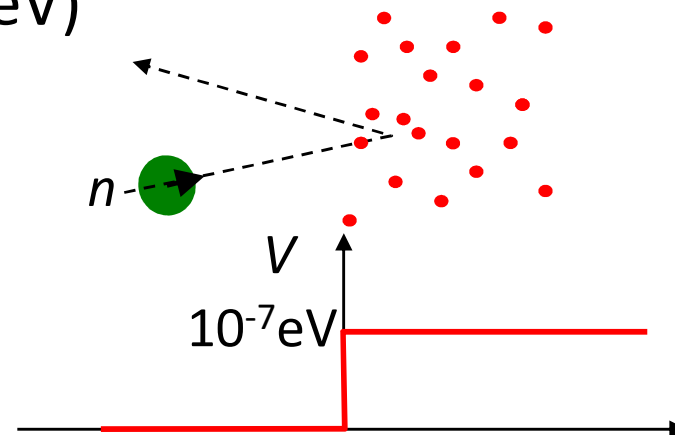
PSI EDM magnet
Photo stolen from P. Fierlinger

Kinetic energy

- 100 neV for 4 m/s

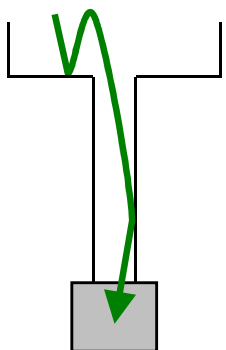
Neutron optical potential

- $O(100 \text{ neV})$



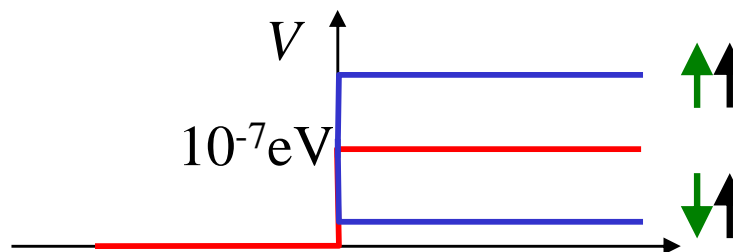
Gravitation

- 100 neV / m



Magnetic interaction

- 100 neV / 1.7 T



Cold or ultracold? Symmetry!

	Cold (3 meV, 800 m/s)	Ultracold (100 neV, 5 m/s)
Magnetic field 1 T	Spin-dependent focussing or defocussing (Perkeo II): 0.2%	Spin-dependent loading efficiencies (UCNA): 30%*
Gravitation	Vertical gradient: small	Vertical gradient: few%/10 cm

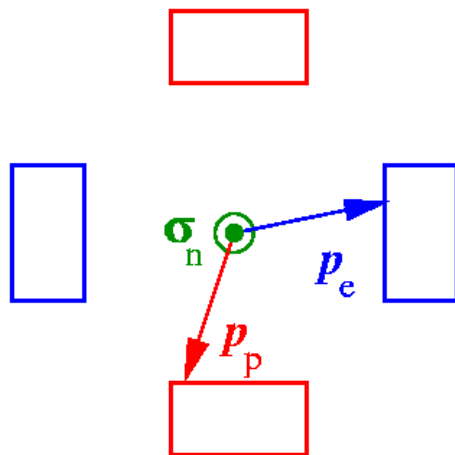
Example: height difference for ^{199}Hg and UCN (EDM)

Require storage

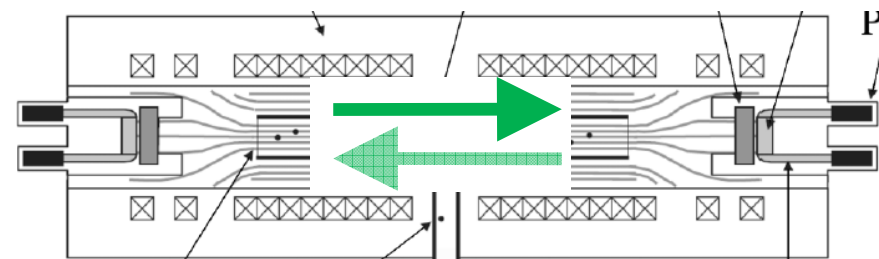
→ energy loss electrons

→ **proton detection??**

Spatial Symmetry



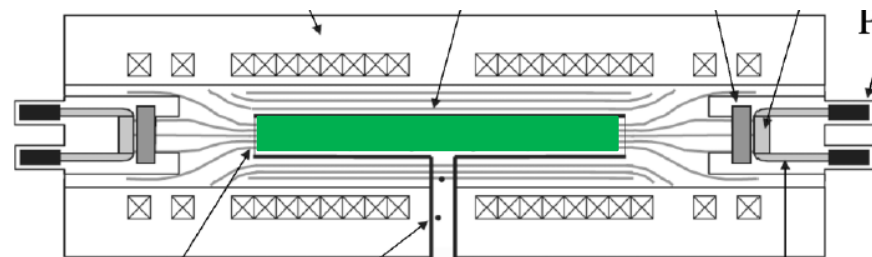
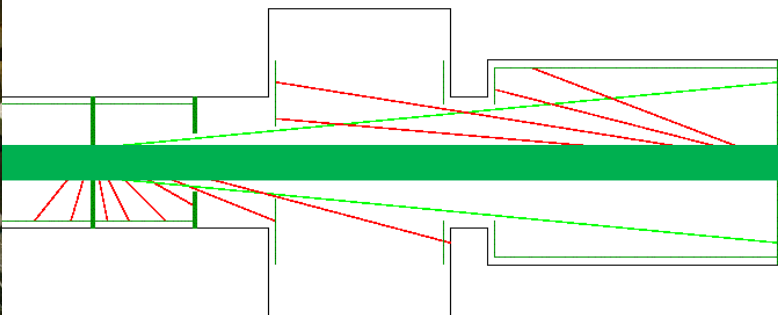
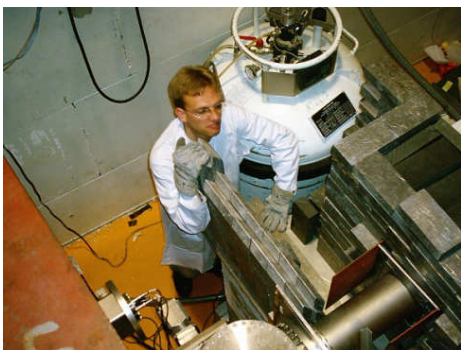
Spin symmetry



* Thanks to A. Young for info

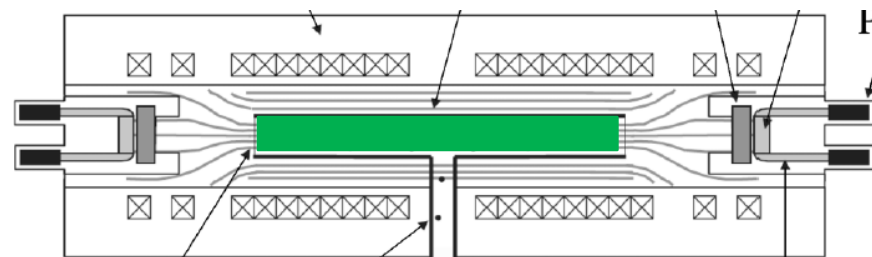
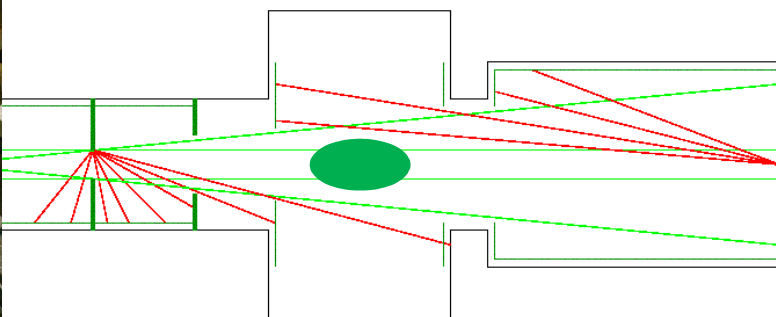
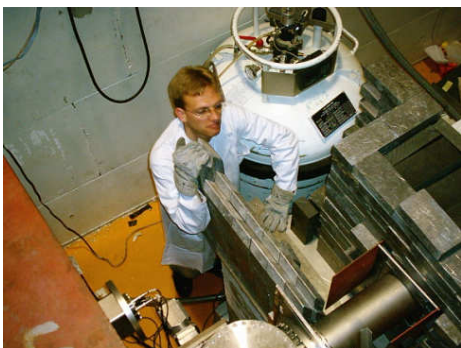
Cold or ultracold? Background!

	Cold (800 m/s)	Ultracold (5 m/s)
Obs.time	10^{-3}s/m	0.2s/m
Sg/beamBG Decay: $10^{-3}/\text{s}$	Observation time (1m): 10^{-3}s Decay/absorption: 10^{-6} <i>Decay & absorption separate!</i>	Collisions: 30/s Loss/collision: $3 \cdot 10^{-5}$ Decay/absorption: 1 <i>Decay & absorption close!</i>
	Sg/BG Perkeo II: (SCI only) 10^3	Sg/BG UCNA: (MWPC+SCI) $5 \cdot 10^3$



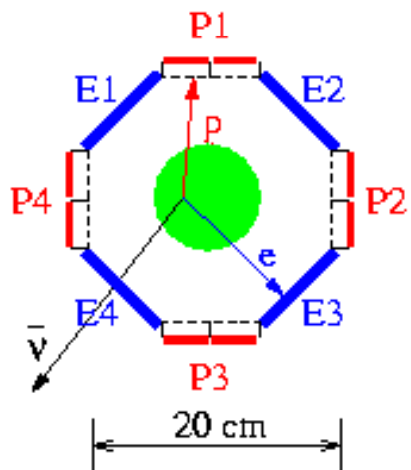
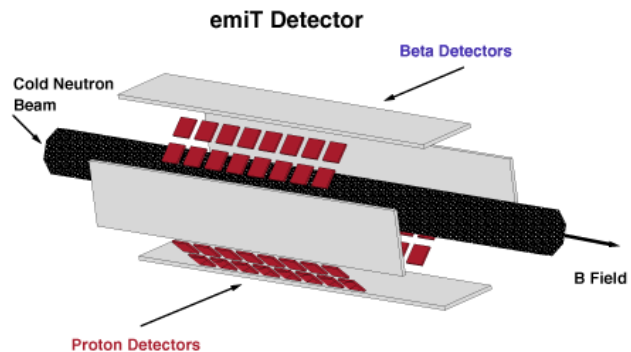
Cold or ultracold? Background!

	Cold (800 m/s)	Ultracold (5 m/s)
Obs.time	10^{-3}s/m	0.2s/m
Sg/beamBG Decay: $10^{-3}/\text{s}$	Observation time (1m): 10^{-3}s Decay/absorption: 10^{-6} <i>Decay & absorption separate!</i>	Collisions: 30/s Loss/collision: $3 \cdot 10^{-5}$ Decay/absorption: 1 <i>Decay & absorption close!</i>
	Sg/BG Perkeo II: (SCI only) 10^3	Sg/BG UCNA: (MWPC+SCI) $5 \cdot 10^3$



Or use pulsed beam!

emiT

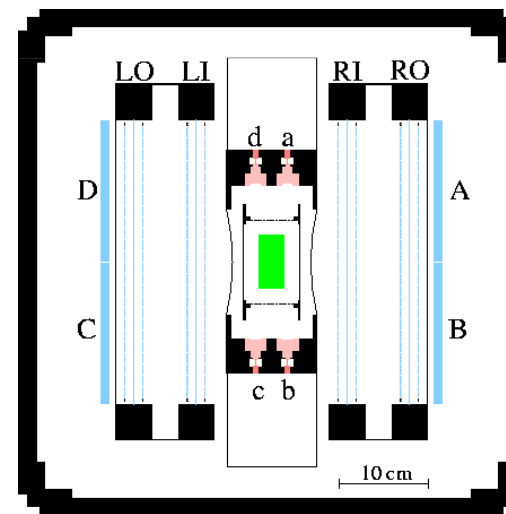
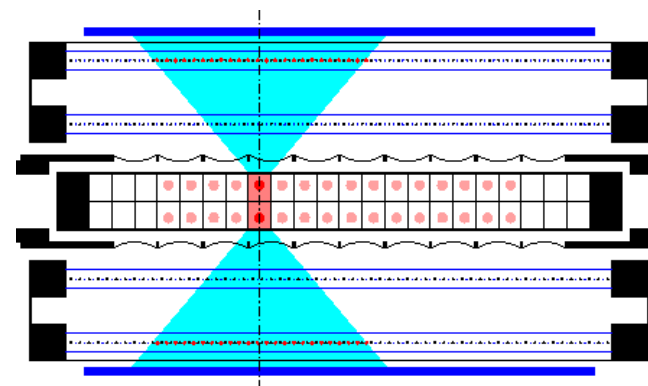


$$D = (-6 \pm 12^{\text{stat}} \pm 5^{\text{syst}}) \cdot 10^{-4}$$

L.J. Lising et al, PRC 62 (2000) 055501.

and ongoing analysis

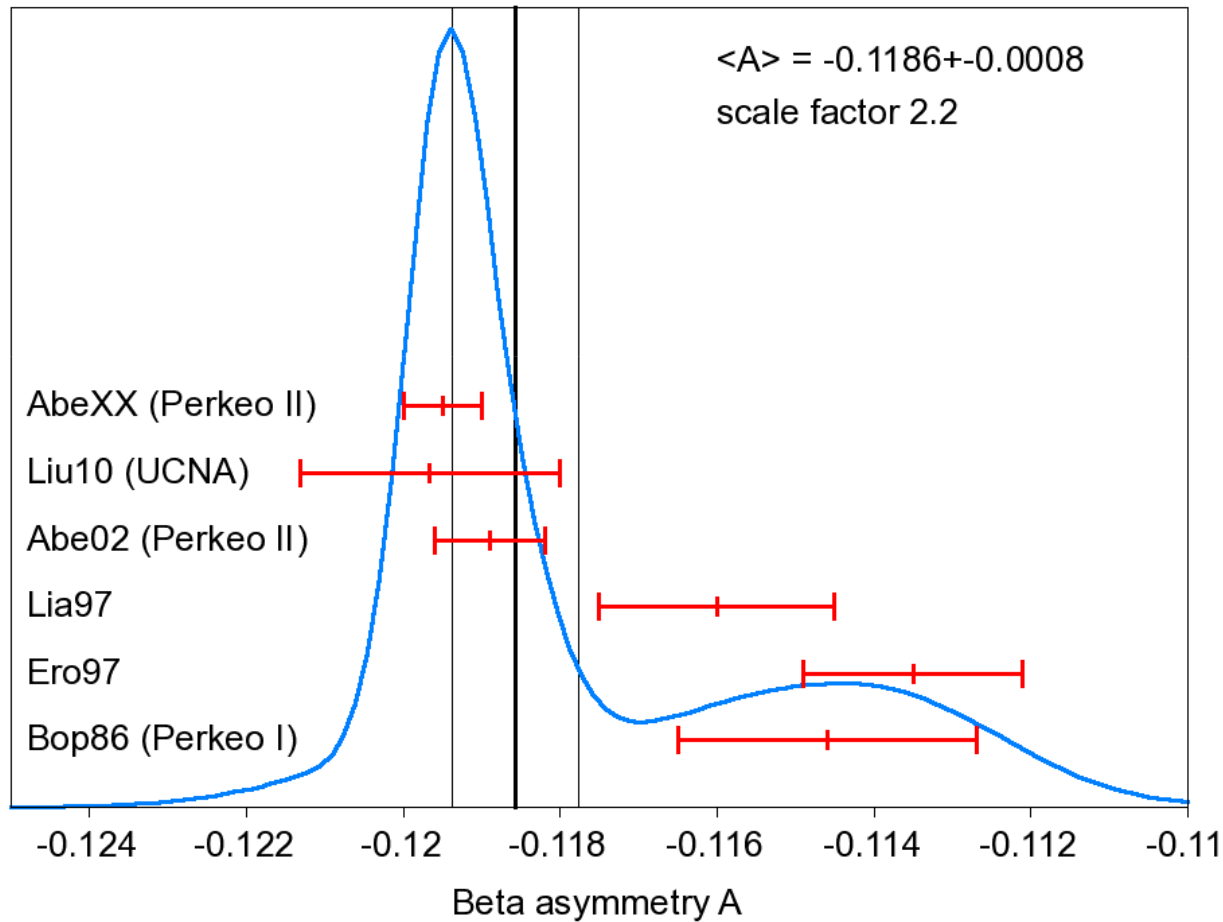
Trine



$$D = (-2.8 \pm 6.4^{\text{stat}} \pm 3.0^{\text{syst}}) \cdot 10^{-4}$$

T. S. et al, Phys. Lett. B 581 (2004) 49.

Talk on R coefficient
K. Bodek,
Tuesday



Experiments

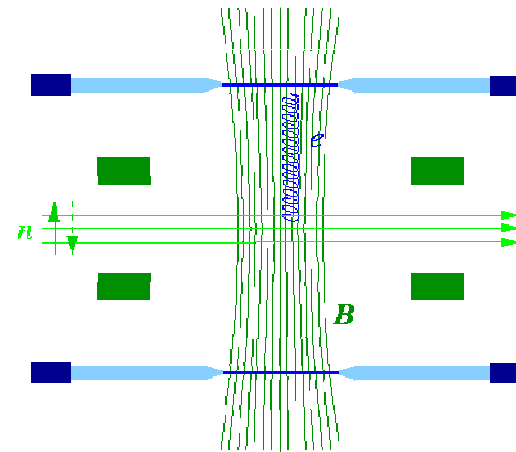
- Perkeo II
Final publication soon
- Perkeo III
Analysis in progress
Talk B. Märkisch, Tuesday
- UCNA
First results (2009, 10)
Talk A. Young, Tuesday
- PERC
funded
- abBA
farer future

Limitations in cold neutron beam experiments



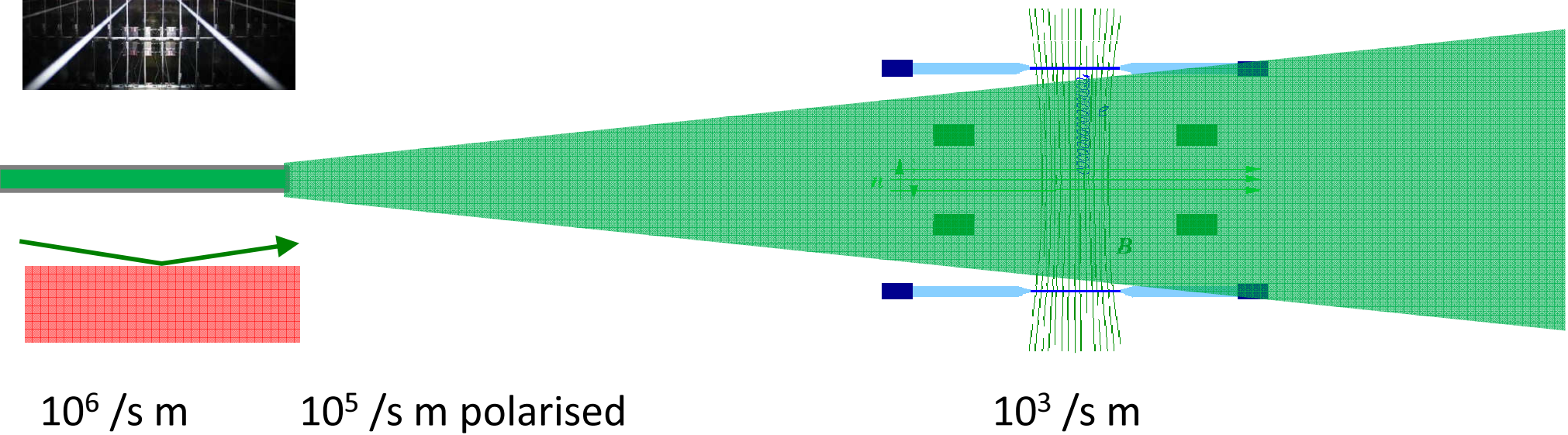
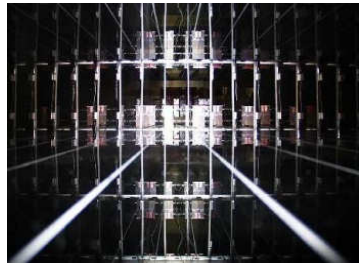
10^6 /s m

10^5 /s m polarised

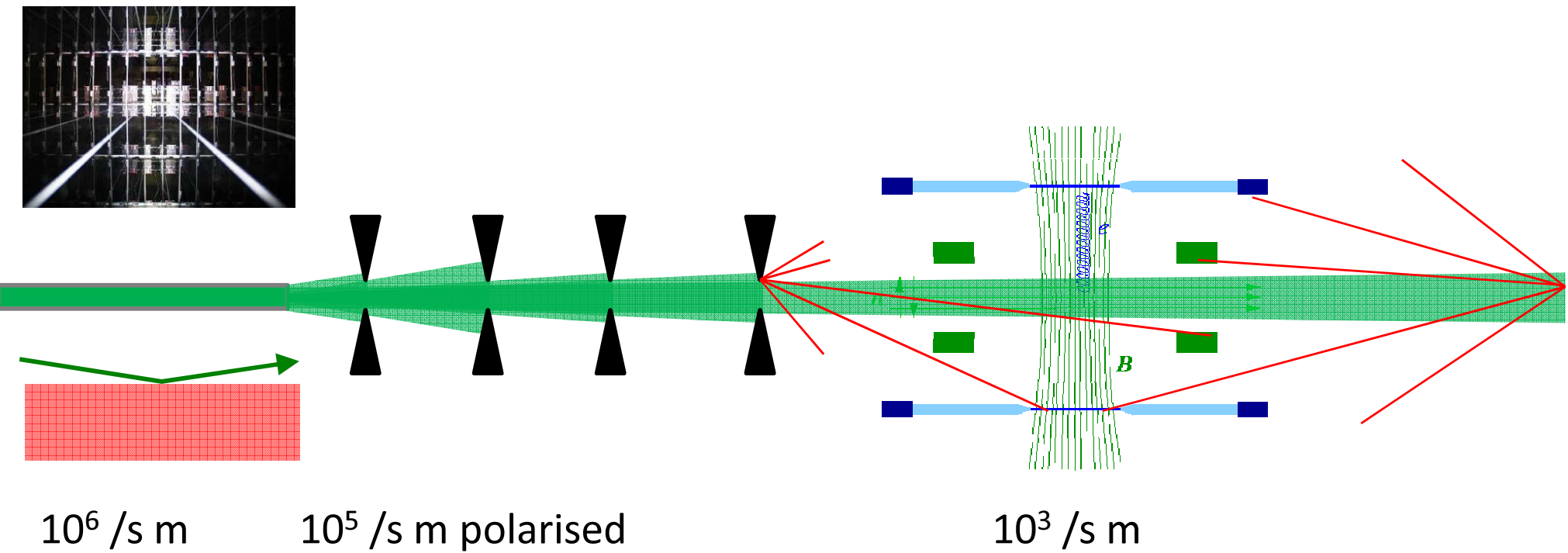


10^3 /s m

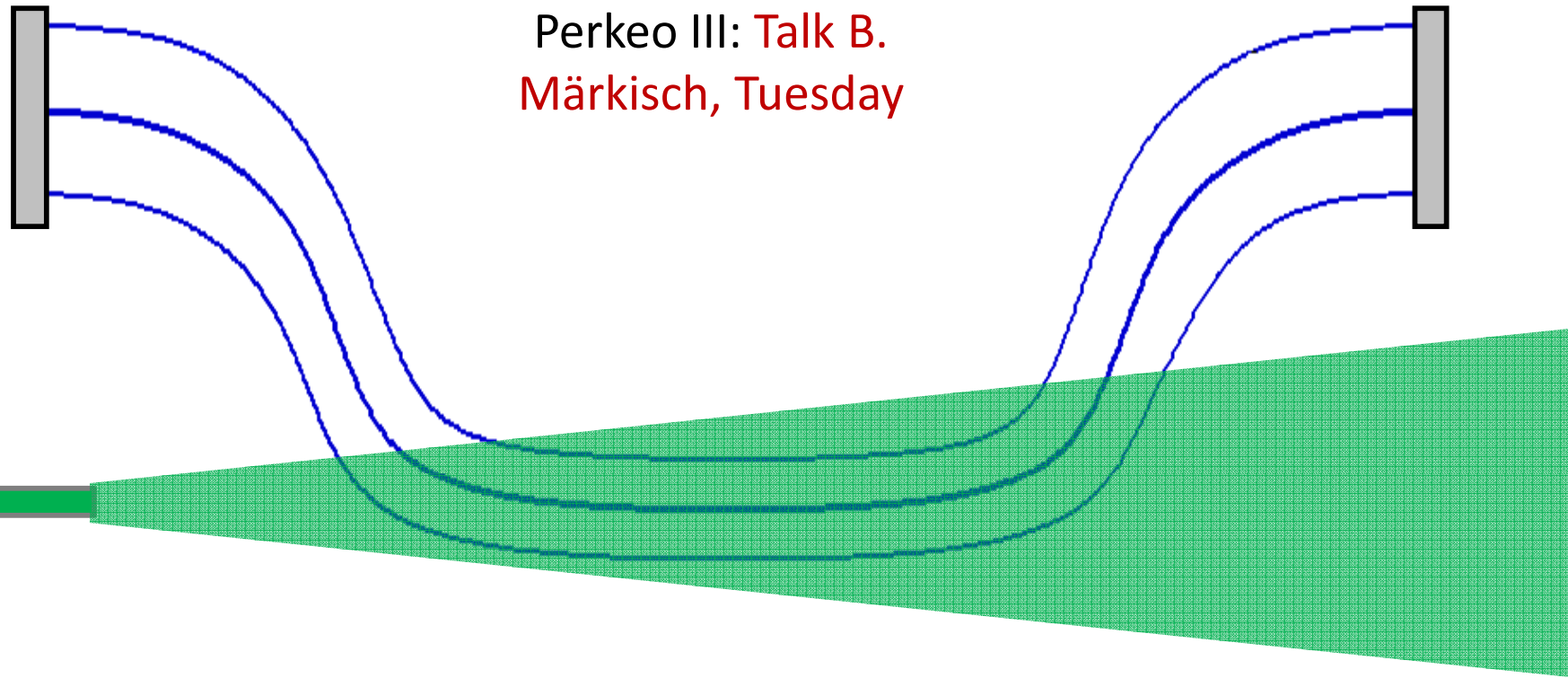
Limitations in cold neutron beam experiments



Limitations in cold neutron beam experiments



Limitations in cold neutron beam experiments



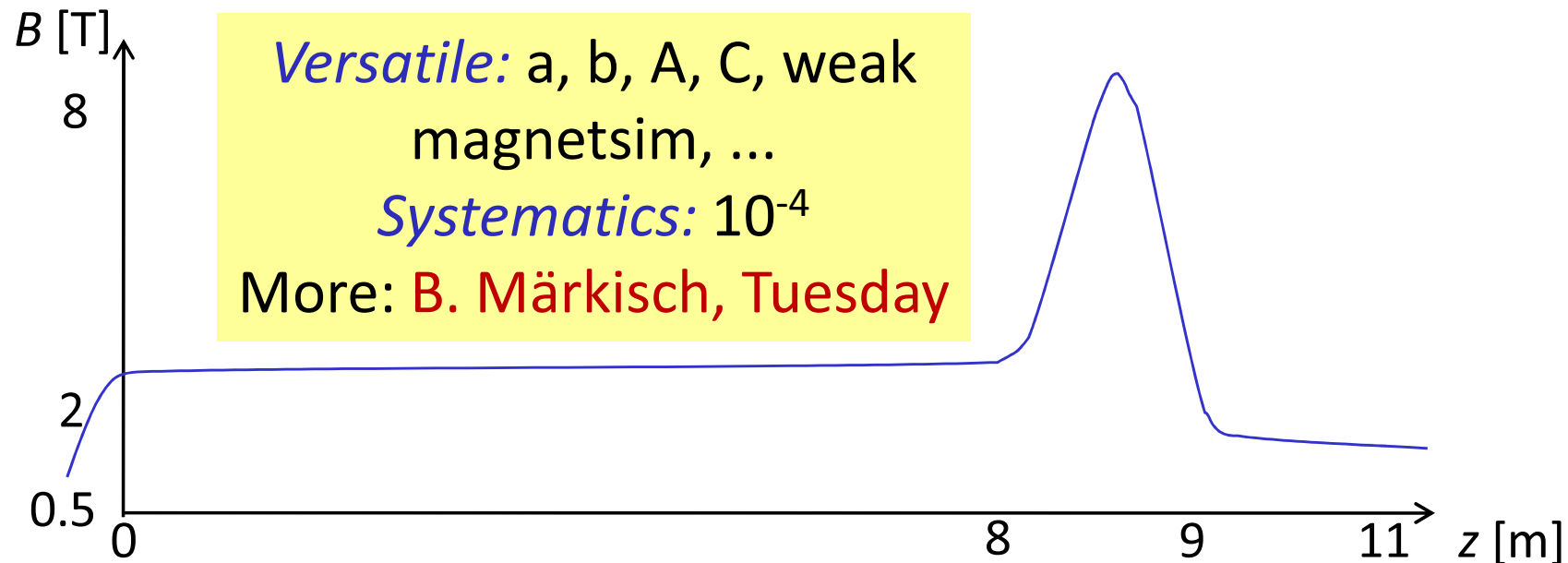
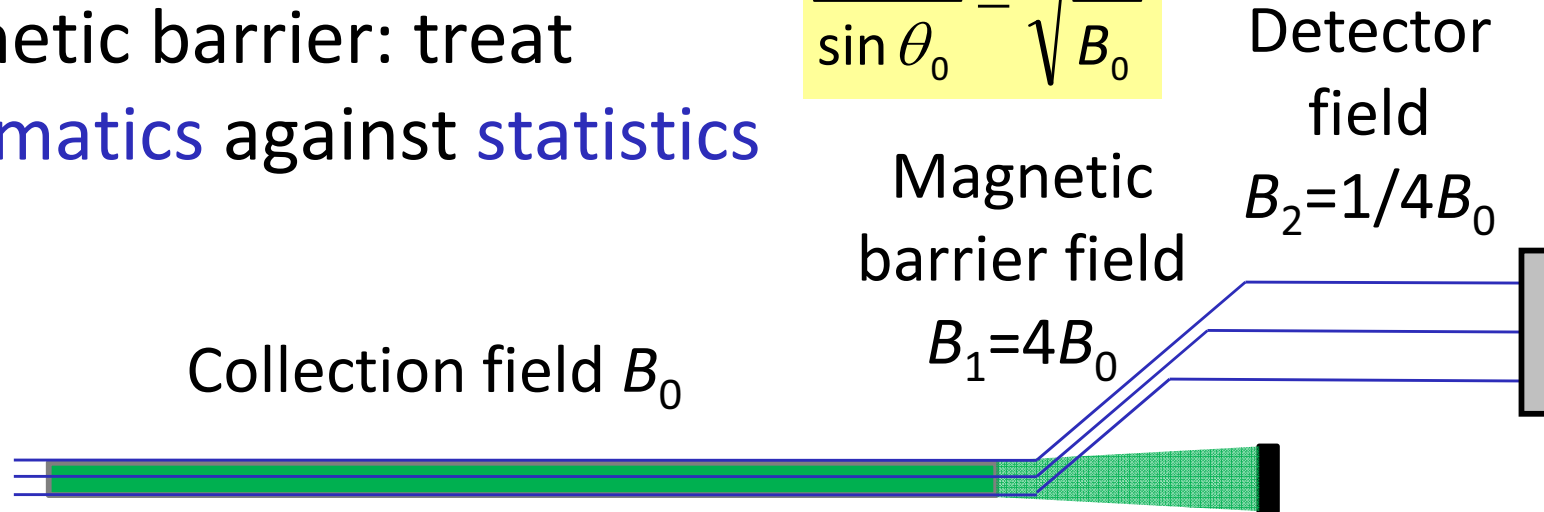
Perkeo III: Talk B.
Märkisch, Tuesday

10^6 /s m

10^5 /s m polarised

Magnetic barrier: treat
systematics against statistics

$$\frac{\sin \theta}{\sin \theta_0} = \sqrt{\frac{B}{B_0}}$$



D. Dubbers et al., NIMA 596 (2008) 238 and arXiv:0709.4440

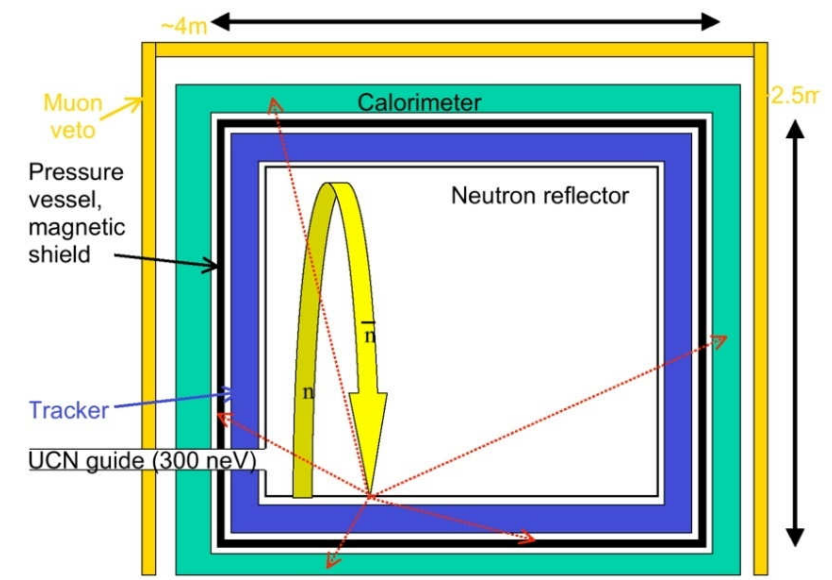
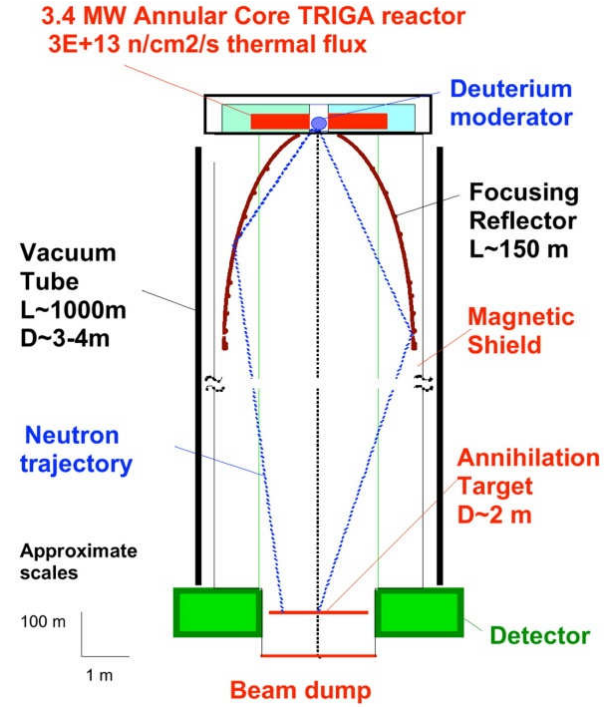
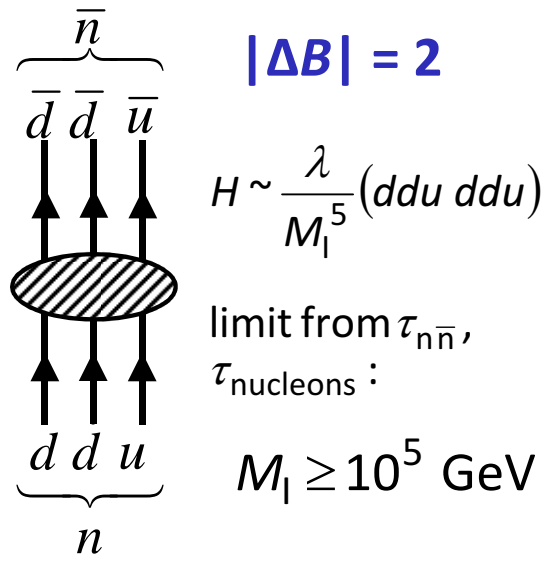
Cold or ultracold? $n\bar{n}$?

	Cold (800 m/s)	Ultracold (5 m/s)
Free flight	0.1 s	1 s

$\tau_{n\bar{n}} \geq 0.86 \cdot 10^8 \text{ s}$ (90% c.l.)
 M. Baldo-Ceolin et al: Z. Phys. C 63 (1994) 409

$$P_{\bar{n}}(t) \approx \left(\frac{t}{\tau_{n\bar{n}}} \right)^2$$

$$P_{\bar{n}}(t) = \left(\frac{t_{\text{free}}}{\tau_{n\bar{n}}} \right)^2 N = \frac{t_{\text{free}} t_{\text{storage}}}{\tau_{n\bar{n}}^2}$$



Intermediate scales
 (GUT breaking, SUSY seesaw, extradim)

Factor 30

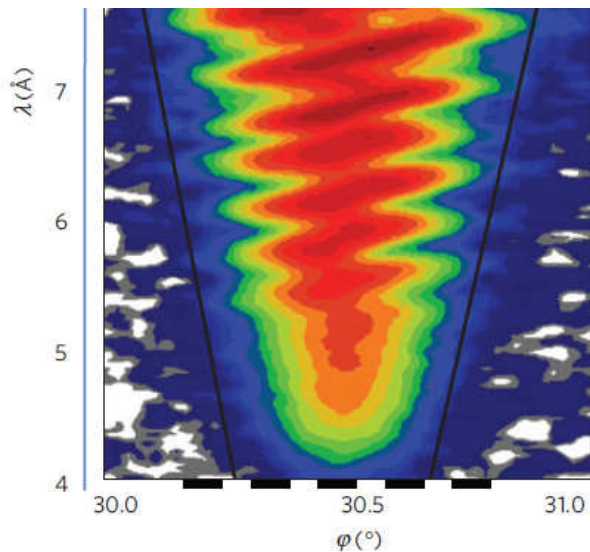
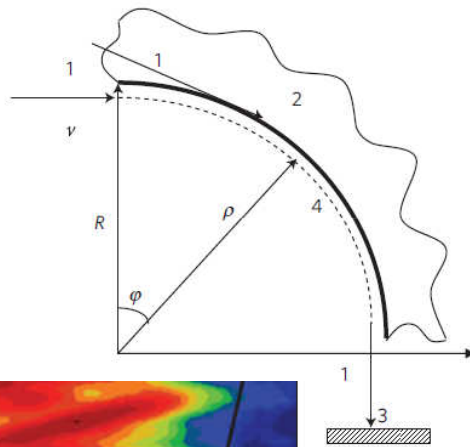
W. Snow, NIMA 611 (2009) 144

Factor 7

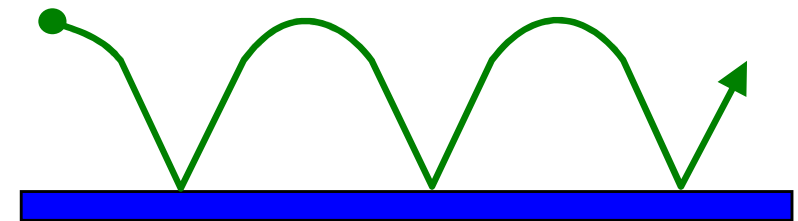
Cold or ultracold? Gravitational quantum states!

	Cold (800 m/s)	Ultracold (5 m/s)
Observ. time	10^{-3}s/m	0.2s/m

Zentrifugal quantum states



V. Nesvizhevsky et al,
Nature Physics 6
(2010) 114



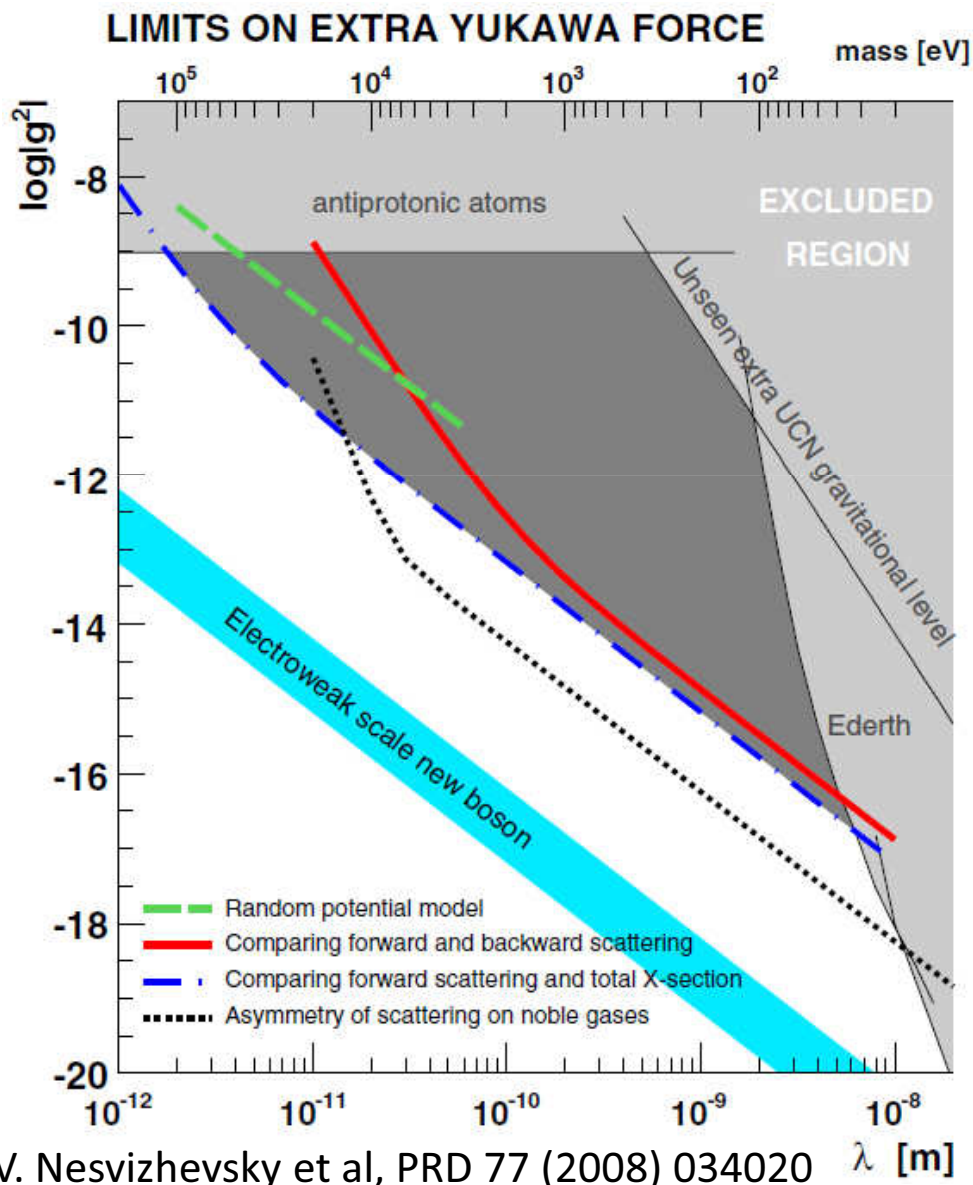
n	E_n [peV]	z_n [μm]
1	1.41	13.7
2	2.46	24.0
3	3.32	32.4

$$\Delta E \Delta t > \hbar \quad \Delta E = 0.12\text{peV} \Rightarrow \Delta t > 5\text{ms}$$

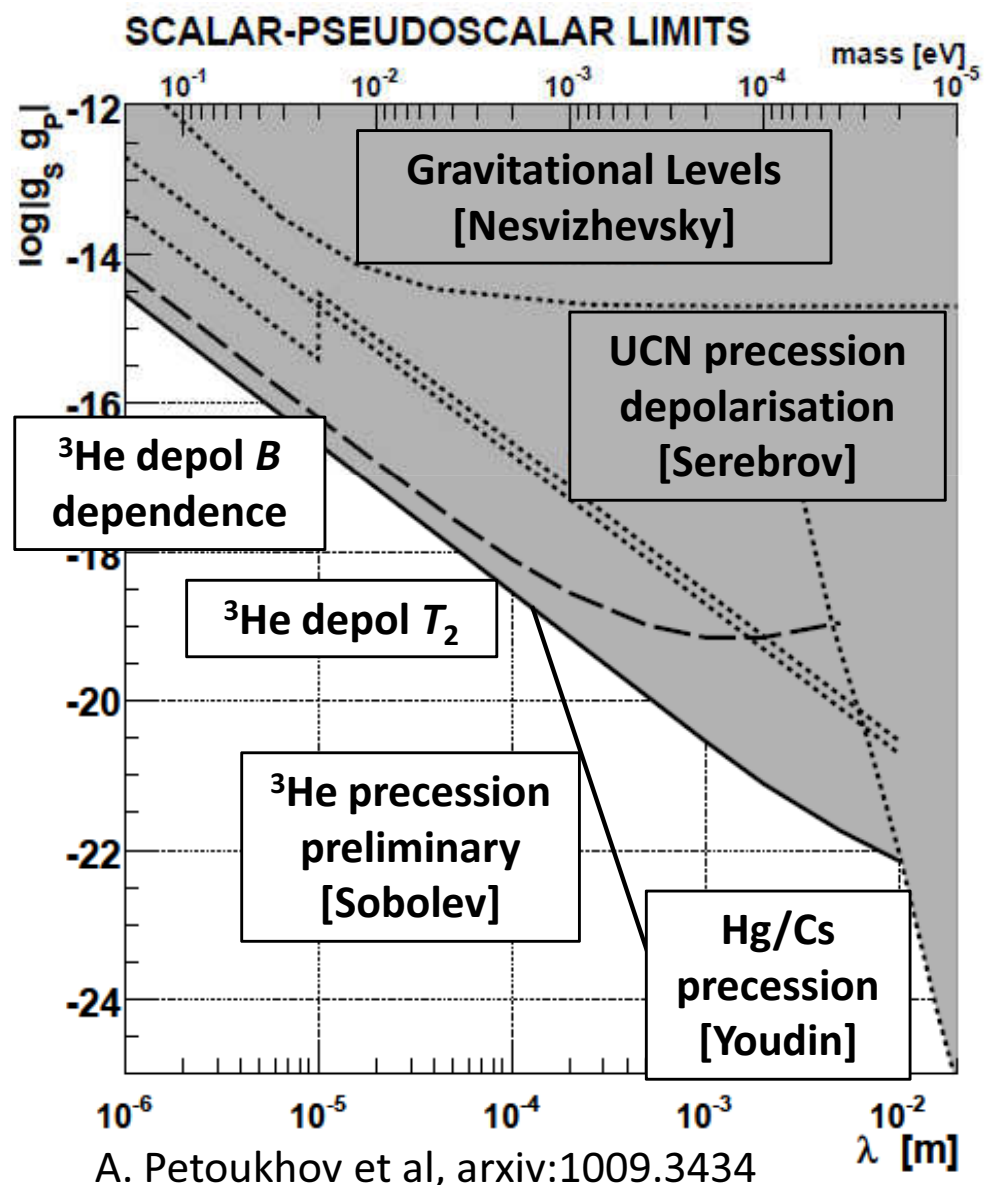
Manageable mirror size: $\sim 10\text{cm}$

$$v_x < 20\text{m/s} \quad \rightarrow \text{VCN or UCN}$$

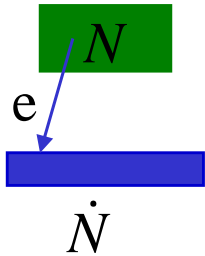
Spin-indepenent



Spin-dependent



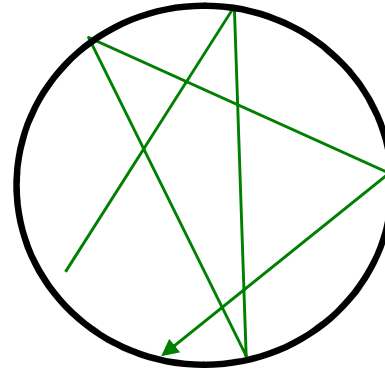
Beam experiments



$$\tau_n = \frac{N}{-\dot{N}}$$

- Two absolute measurements:
 - ? Solid angle
 - ? Thresholds
 - ? Backscattering
 - ...
- Stated precision $4 \cdot 10^{-3}$

Storage experiments



$$\tau_n = \frac{t}{\ln \frac{N(t)}{N_0}}$$

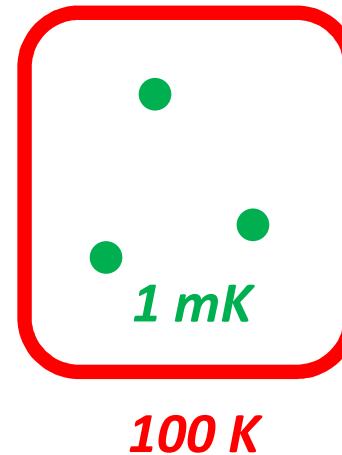
- Counting survivors after storage
- Relative measurement (?)
- Losses \rightarrow extrapolation techniques
- Stated precision $9 \cdot 10^{-4}$

Obvious

- Faster neutrons
 - More collisions
 - Faster losses

Also obvious

- Thermal non-equilibrium
 - Heating



→ Spectrum changes somehow predictable

Talk A. Fomin

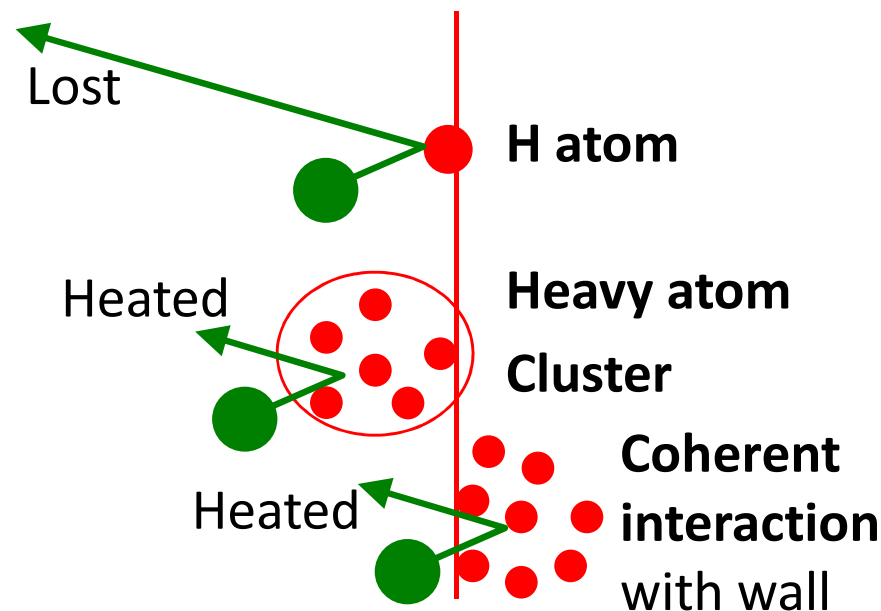
Obvious

- Faster neutrons
 - More collisions
 - Faster losses

→ Spectrum changes somehow predictable

Also obvious

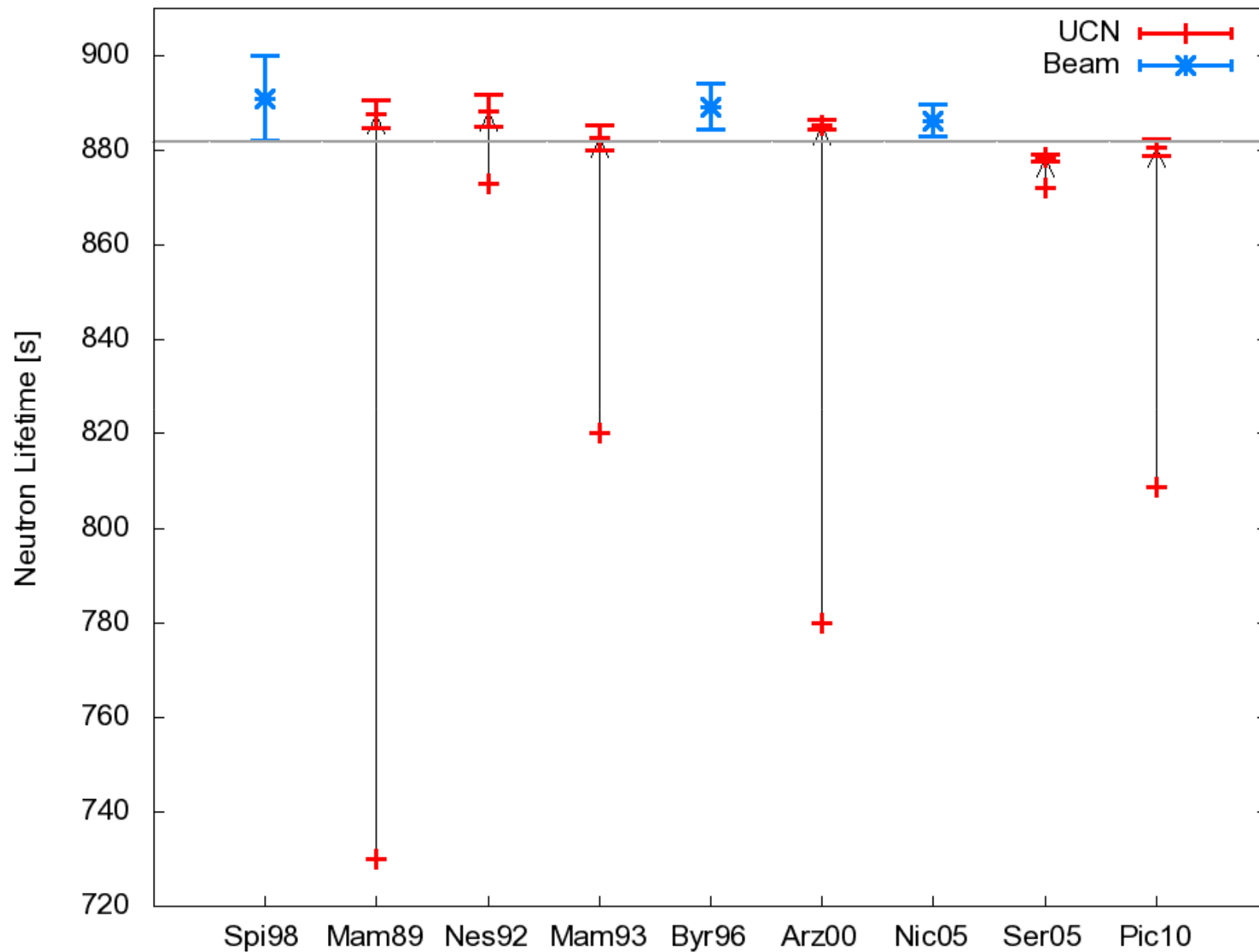
- Thermal non-equilibrium
 - Heating



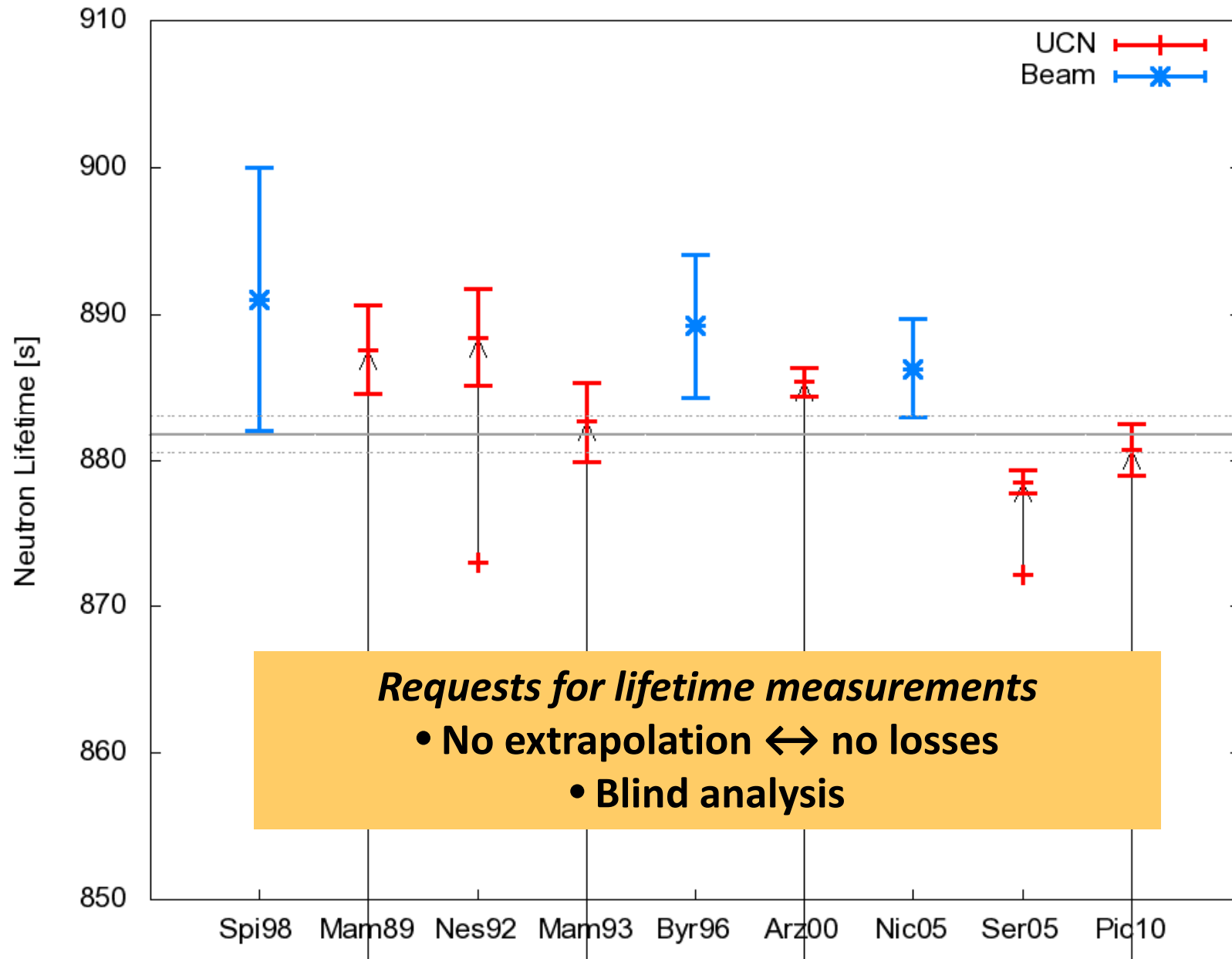
→ Unpredictable spectrum change

→ ***Concept of storage time constant invalid***

Neutron lifetime – Status

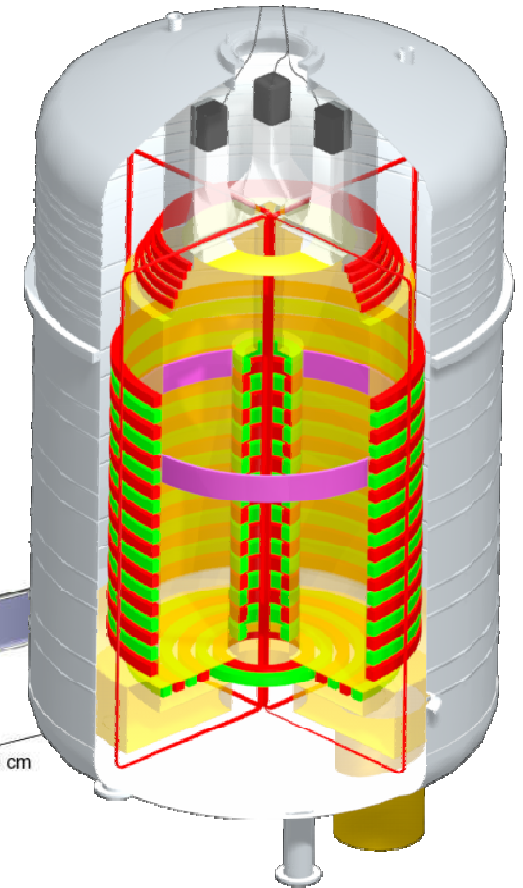
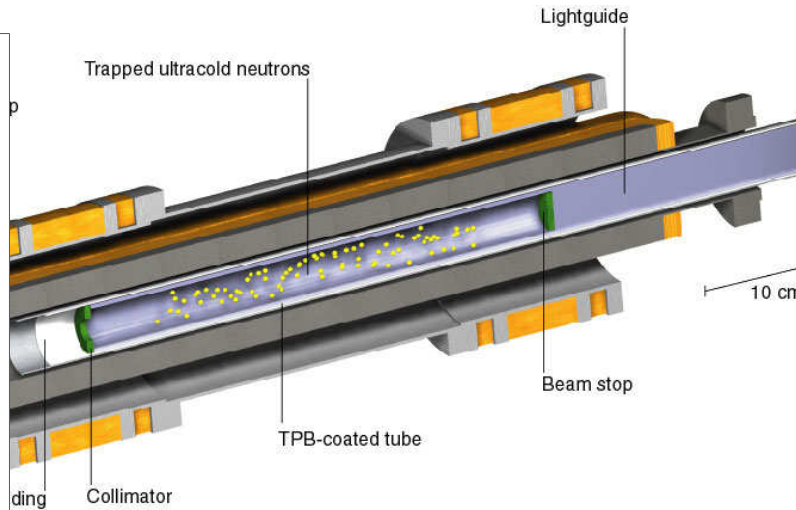
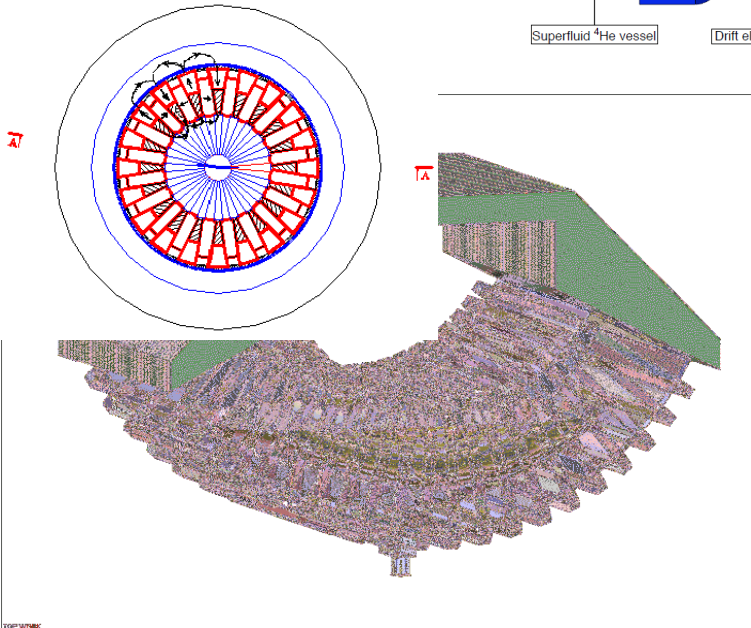
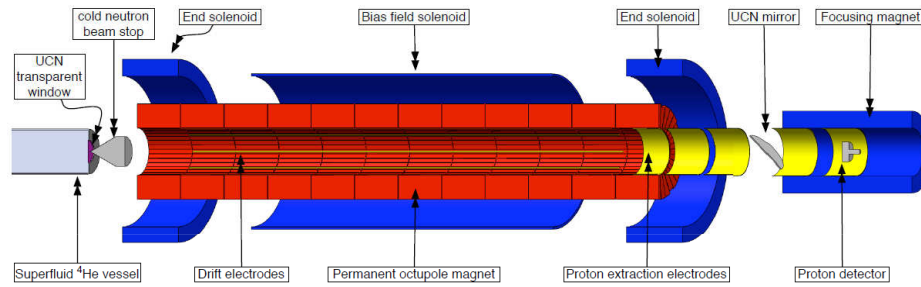
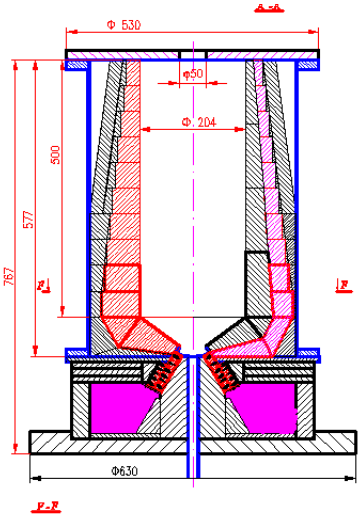


Neutron lifetime – Status



A possible solution? Magnetic trapping

Talk V. Ezhov



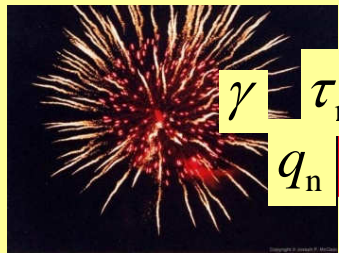
Talk R. Picker

- Neutron particle physics:
 - Precision measurements & theory (A, B, C, τ , gravitational states, occupation of bob states)
 - Search deviations from 0 ($D, d_n, q, | \rightarrow$ state)
 - Search for the forbidden ($n\bar{n}$)
 - *Very active field*
- Cold or ultracold neutrons?
 - Doppler effect, $v \times E$, statistics, polarisation, symmetry (space, spin), background, observation time
 - Choose the best for **you!**
- Neutron lifetime: **No losses, please!**

The Neutron Guide to the Universe

Unknown physics

Known physics



Planck

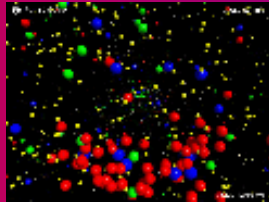
GUT

γ $\tau_{n\bar{n}}$ m_{WR}, ζ
 q_n g d_n φ

Inflation

λ
 a_{WM}
 $\sigma_{N\nu}$

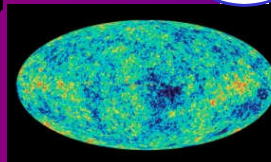
Electro-weak



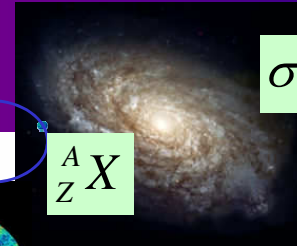
Quarks, hadrons

τ_n
 α_n
 V_{ud}
 N_ν

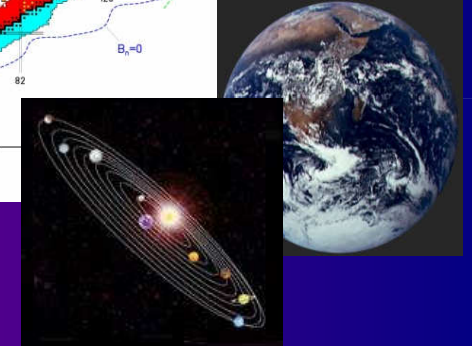
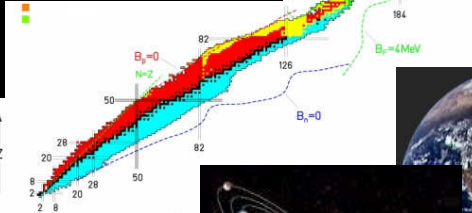
Nucleo-synthesis



Atom formation



Stars
Galaxies



Solar system



Today

$10^{19}m_n$ $10^{15}m_n$

$10^{-43}s$ $10^{-35}s$

10^3m_n $1m_n$

$10^{-11}s$ $10^{-6}s$

3min 300 000 a

3 000 K

$8 \cdot 10^9a$ $13.7 \cdot 10^9a$