

# **The Fundamental Physics Beamline** @ **ESS**

**C. Theroine<sup>1,2</sup>**, A. Hiess<sup>1</sup>, C. Klauser<sup>2,3</sup> and T. Soldner<sup>2</sup>.

<sup>1</sup>European Spallation Source ESS, Tunavägen 24, 223 63 Lund, Sweden. <sup>2</sup>Institut Laue-Langevin, 6 Rue Jules Horowitz, 38000 Grenoble, France. <sup>3</sup>Atominstitut Technische Universität, 1020 Wien, Austria.

### The European Spallation Source : ESS

The European Spallation Source (ESS) is a facility that will provide to the scientific community new opportunities of research using neutrons. Indeed, the unique capabilities of the ESS, namely the high peak flux of the neutron beam and the inherent time structure of the beam, will permit to push the frontiers of neutron sciences. These properties will be very useful for many fundamental physics experiments.

Parameters	Value
Average beam power	5 MW
Number of target stations	1
Number of instruments in construction budget	22
Number of beam ports	48
Number of moderators	2
Separation of ports	5°
Proton kinetic energy	2.5 GeV
Average macro-pulse current	50 mA
Macro-pulse length	2.86 ms
Pulse repetition rate	14 Hz
Annual operating period	5000 h
Reliability	95%



# Advantages of long-pulsed spallation sources for Fundamental Physics

The ESS will be the first pulsed source that provides the same averaged flux as the ILL.

Advantages	Example
Time structure of the beam → separation of neutron velocity dependent systematic effects	Neutron spin rotation in magnetic fields
Time structure of the beam → Wavelength-resolved polarization for free, beam-related background, pulsed measurements to investigate spatial dependence of spectrometer response	Asymmetry measurements with polarized neutron
Pulse structure → increased signal/ background during pulse, measurement of spectrometer background between pulses	Experiments with intrinsic spectrometer background

#### The Fundamental Physics Beamline project

Instrument Parameters	Value	
Moderator	Cold	
Distance to moderator	to be defined	
Wavelength Range	~ 3- 30 Å	
Energy Resolution	to be defined	
Beam size at the sample	~ 100*120 mm2	
Divergence at the sample	~ 2°	
Size of the experimental area	Up to 30 m X 6 m X 6 m	
Detectors	3He, Semiconductor, Scintillator	
Infrastructure	High accuracy polarization	

The pulsed time structure of the ESS requires a careful optimization of the beamline parameters

#### Ultra Cold Neutron case

For He sources: Improve production rate by efficient monochromatization. Optimized extraction of cold neutrons from primary moderator.

## How to define the best configuration ?

- Collect state-of-the-art fundamental physics experiments at cold neutrons beams.
- Interact with the fundamental physics community to clarify their scientific priorities.
- Define different candidate experiments for ESS -
- Perform simulations/calculations



maximum flexibility for new scientific opportunities

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**PROPOSAL FOR THE FUNDAMENTAL PHYSICS BEAMLINE ~ END 2014** 

### An example of an experiment @ ESS : PERC

As example, we are studying the possibility to perform a state-of-the-art high precision neutron beta decay experiment which tests the Standard Model with cold neutrons : PERC, at ESS to determine the gain factor compared to the ILL research reactor.

The Proton Electron Radiation Channel PERC is a new instrument designed to measure correlation coefficients in neutron decay. For that, a strong longitudinal magnetic field is used along a neutron guide to extract charged neutron decay products. With this new experimental set-up, all systematic effects could be controlled at least on the 10<sup>-4</sup> level [1].



Scheme of PERC instrument.

Klauser et al. show analytically in [2] that it is possible to obtain up to a gain factor of 7 for a distance of 19 m between the moderator and the chopper for PERC at ESS compared to the ILL. The main hypotheses used here are :

Currently, the analytical calculations and claims are tested with simulations from Monte-Carlo code McStas [4]. Preliminary simulations are shown.



Simulation of PERC at the ILL with McStas.



Simulation of the capture flux in PERC at the ILL with McStas. Left : at 2m from the source, Middle : after the 5 Å selector, Right : after the chopper.

- The peak flux of ESS is about 30x the continuous flux of the ILL.
- The ESS pulse length is 2.86 ms [3].
- The ESS pulse frequency is 14 Hz [3].
- The reference wavelength is 5 Å.





McStas simulation results for the mean brilliance between 0.5 and 20Å : for ESS ~ 2.45e+10 n/cm^2/ster/Å/s for ILL ~ 2.60e+10 n/cm^2/ster/Å/s

> After checking the simulated set-up with McStas at the ILL and ESS, the next step of this work will be the determination of the gain factor

References :

[1]: D. Dubbers et al., Nucl. Instr. Meth. A 596, 238 (2008) or arXiv:0709.4440. [2] : C. Klauser et al., submitted in Physics Procedia (NPPatLPS 2013). [3]: S. Peggs, ESS Technical Design Report (2013), http://esss.se/documents/tdr/TDR\_final.pdf [4] : P. Willendrup et al., McStas manual (2012), http://mcstas.org/.

Gain for PERC at the ESS compared to the ILL calculated by [2].