

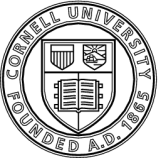
# Low-flux Measurements with Cornell's LCLS Integrating Pixel Array Detector

Hugh T. Philipp<sup>a</sup>, Mark W. Tate<sup>a</sup>, Sol M. Gruner<sup>a,c</sup>

<sup>a</sup> Lab of Atomic and Solid state Physics, Department of Physics, Cornell University.

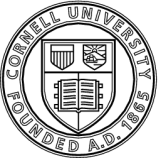
<sup>c</sup> Cornell High Energy Synchrotron Source, Cornell University.





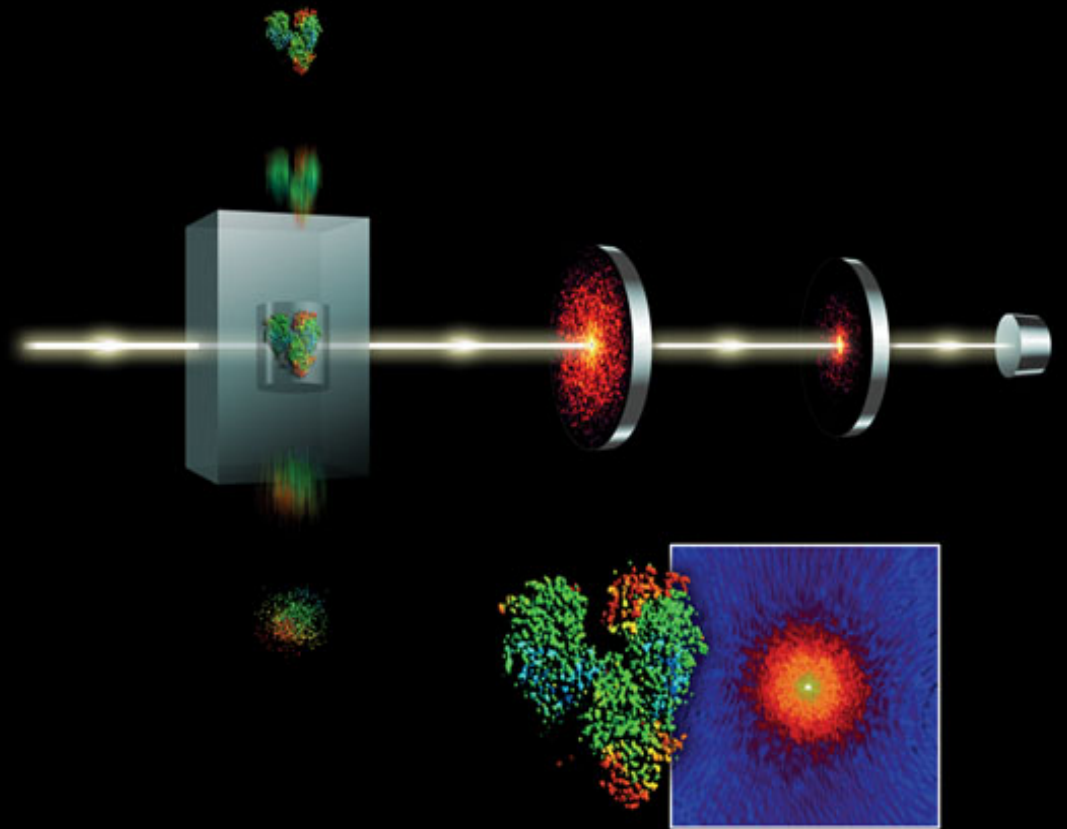
# *Outline*

- ♦ Coherent x-ray imaging and detector need.
- ♦ Original specifications.
- ♦ Brief overview of PAD design.
- ♦ Single photon measurements.
- ♦ Thoughts on low flux measurement with integrating PAD.
- ♦ Measurements and data reduction with Cornell's single module detector.



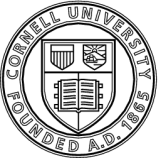
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# *CXI Instrument*



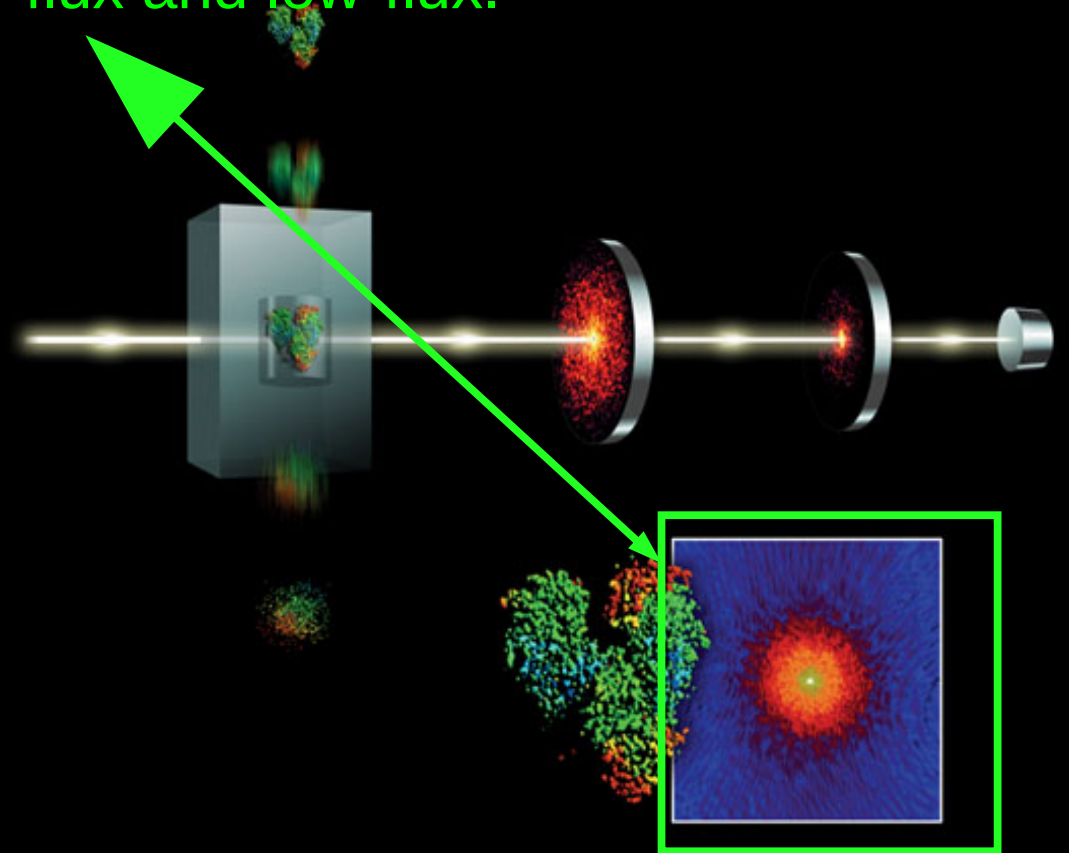
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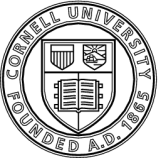


# *CXI Instrument*

- Combination of extremely high-flux and low-flux.

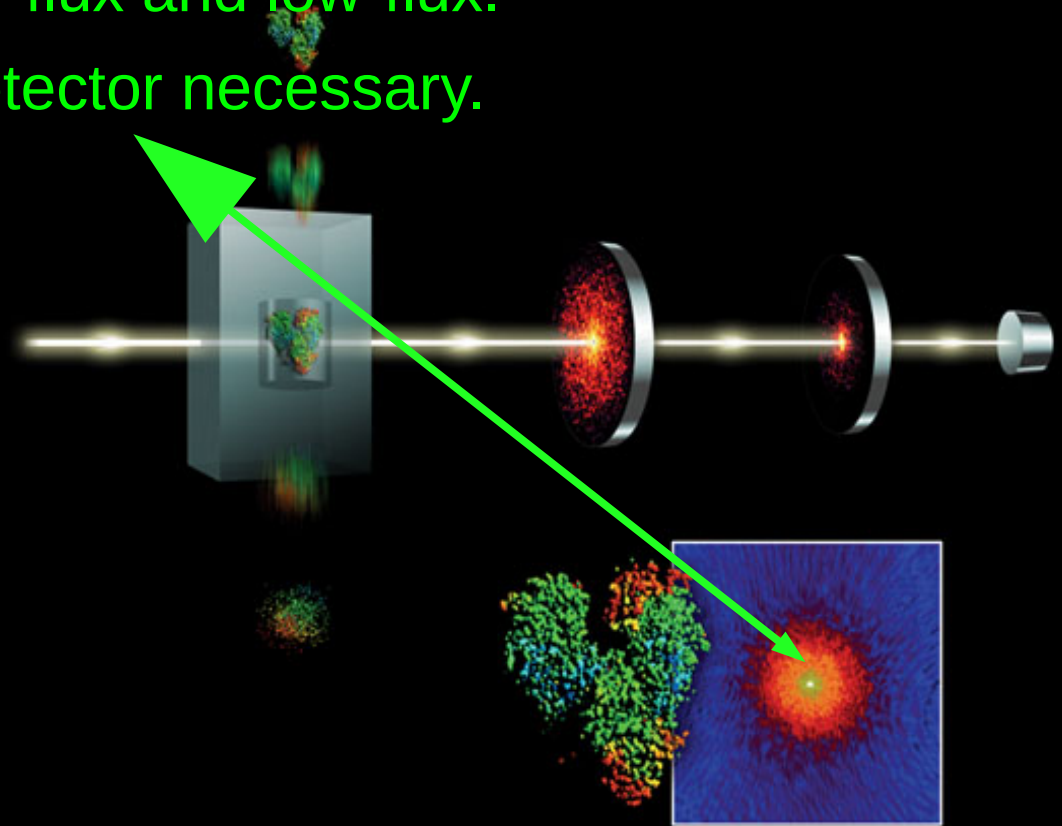


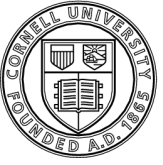




# *CXI Instrument*

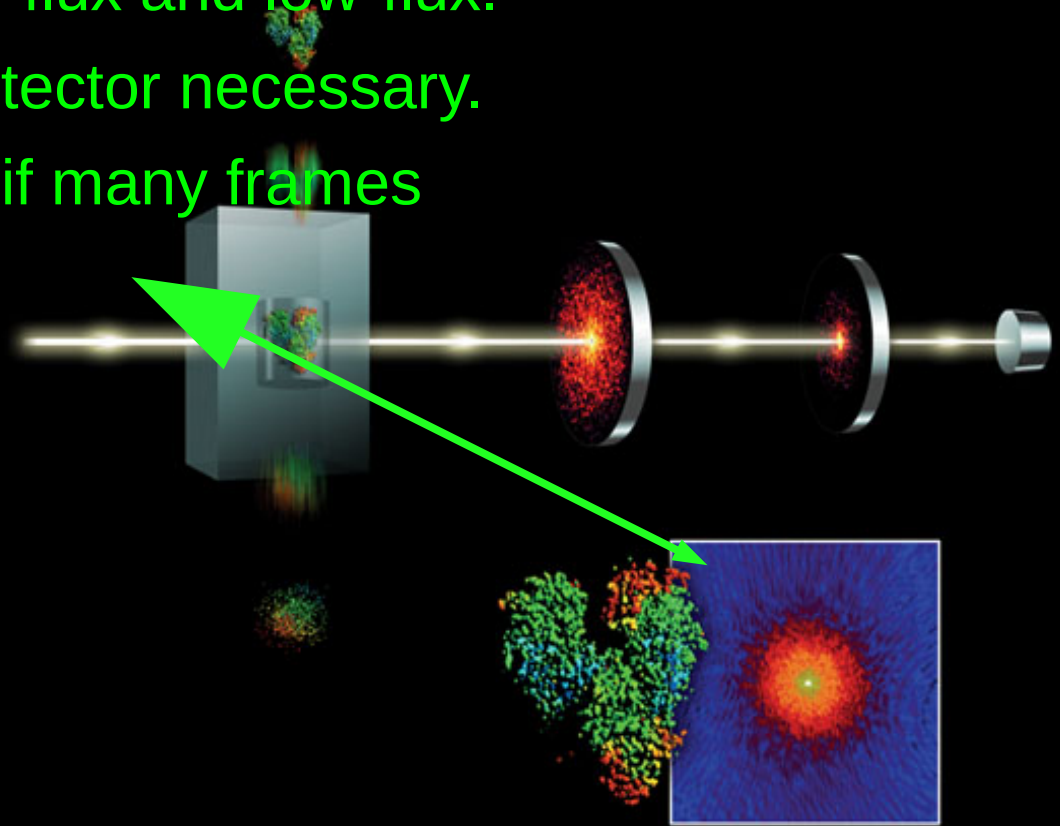
- Combination of extremely high-flux and low-flux.
- High-flux makes integrating detector necessary.



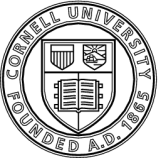


# *CXI Instrument*

- Combination of extremely high-flux and low-flux.
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- Low-flux data only meaningful if many frames added.

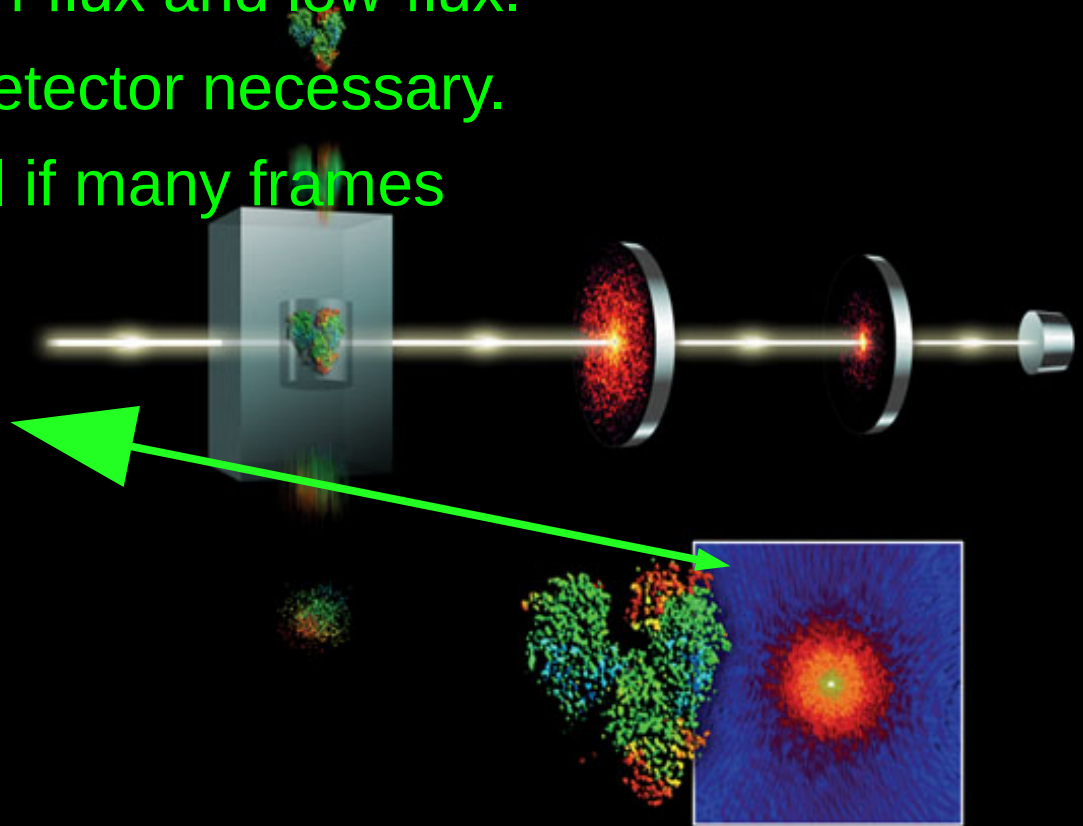


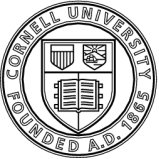




# *CXI Instrument*

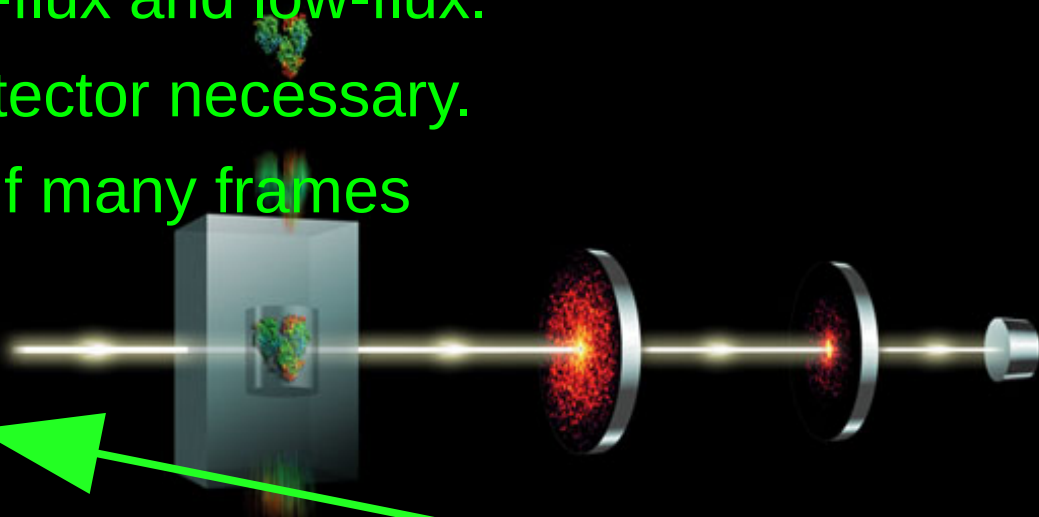
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- Low-flux data defined as  $\ll 1$  photon/pixel/frame.



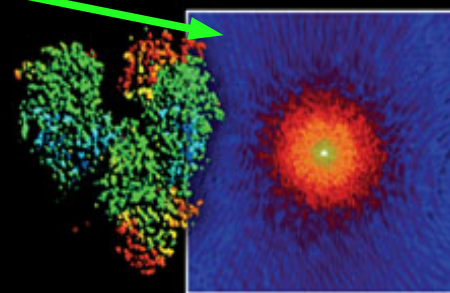
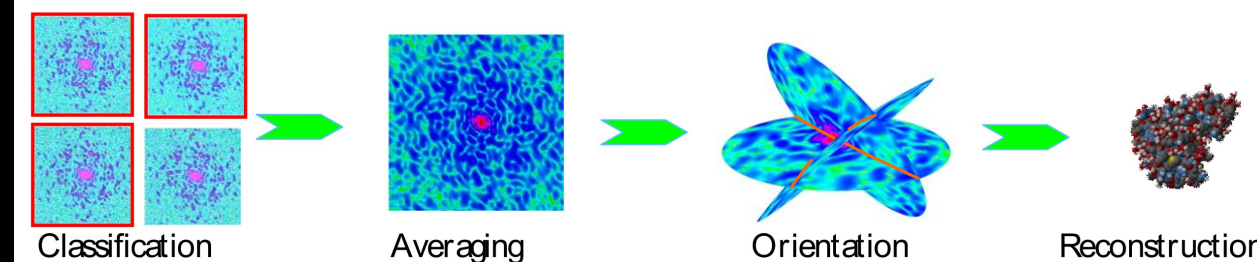


# *CXI Instrument*

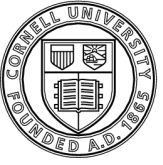
- Combination of extremely high-flux and low-flux.
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Combine  $10^5$ - $10^7$  measurements

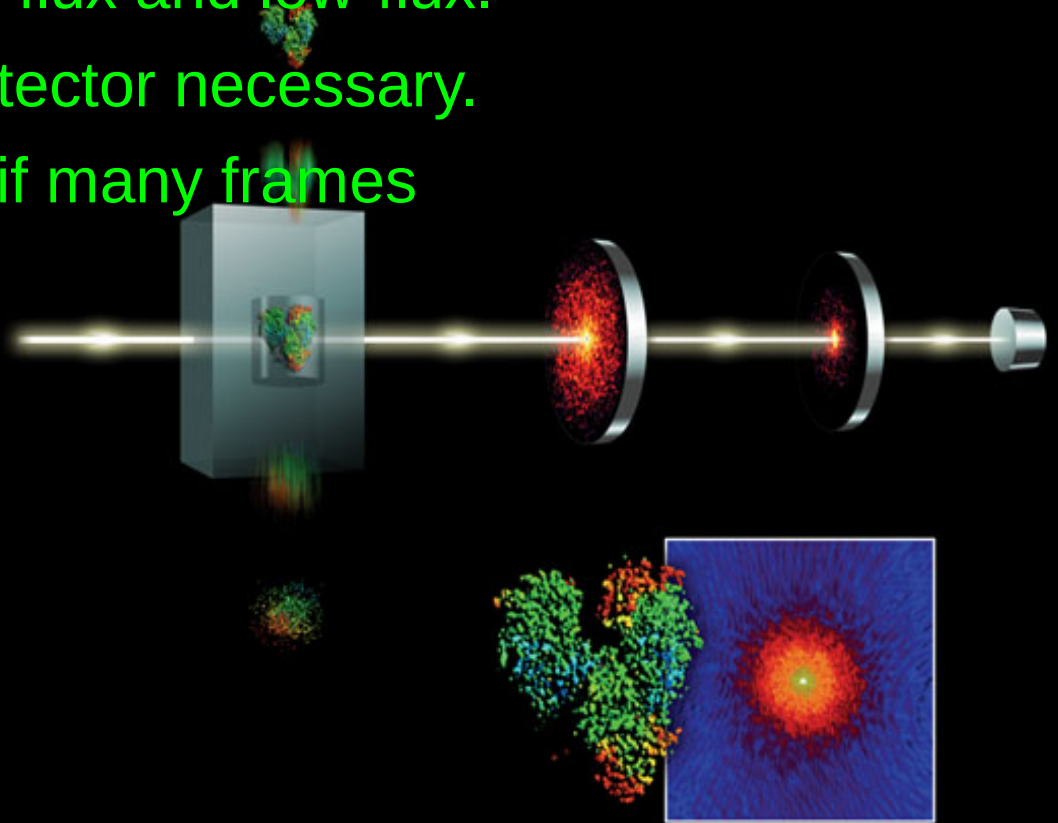




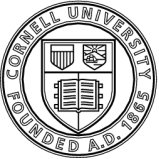


# *CXI Instrument*

- Combination of extremely high-flux and low-flux.
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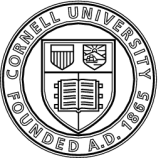
How do you deal with this when your detector noise is (say)  $1/7$  photon?



# *Original Specs Given for Detector*

Parameter	Requirement
Energy Range	4-8 keV
Well-depth/pixel	$10^3$
Readout frame rate	120 Hz
Signal/Noise	>3 for single 8 keV photon
DQE	> 90% at 8 keV
Pixel size	100-200 $\mu\text{m}$
Detector area	> 500x500 pixels

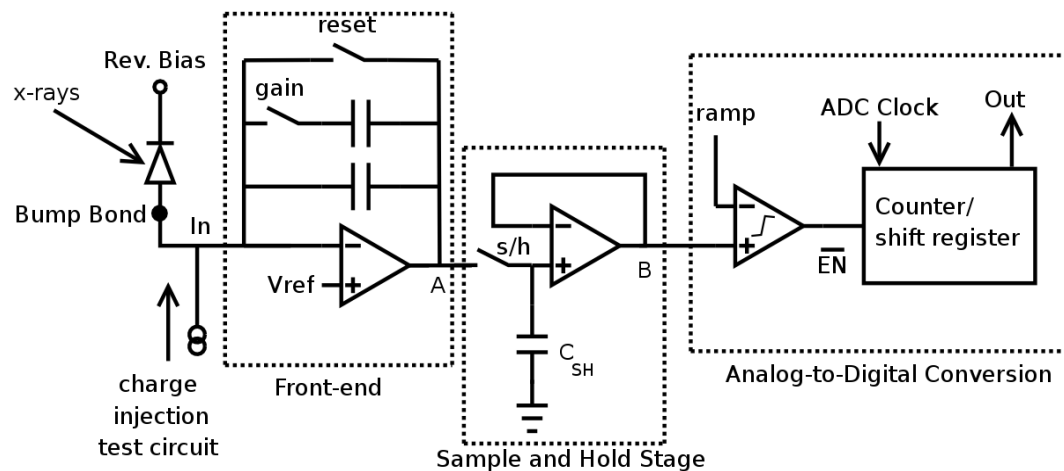




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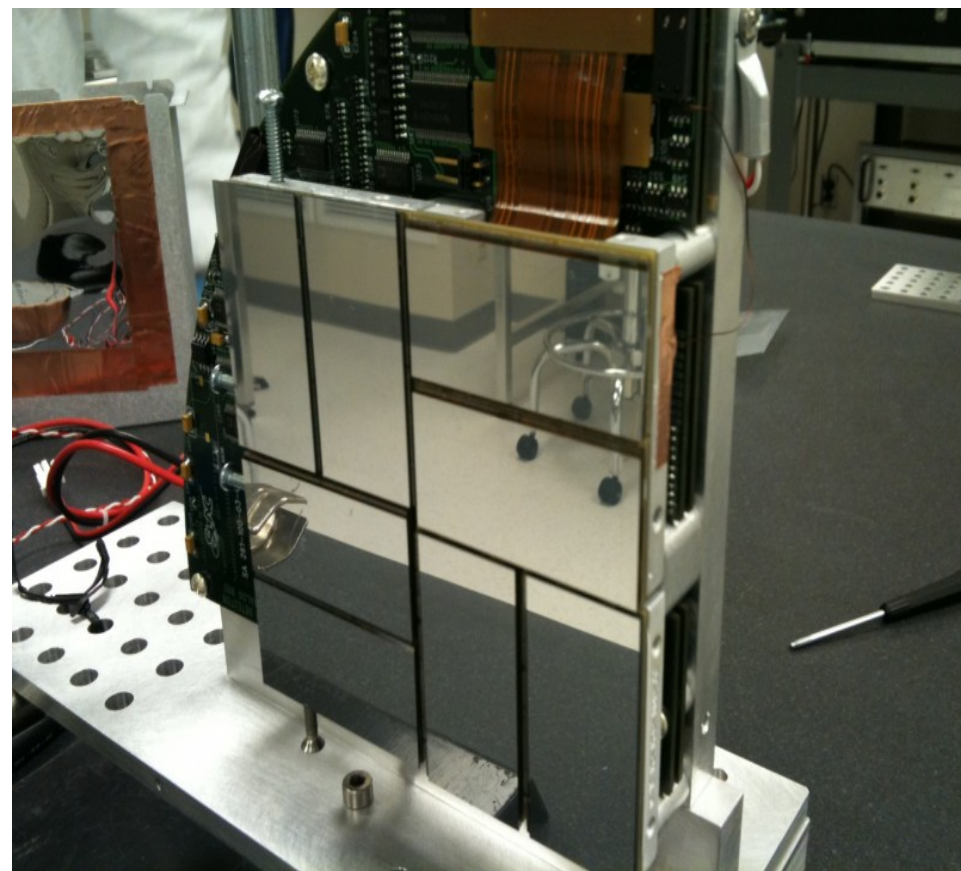
# Cornell's LCLS PAD

## Basic pixel design.



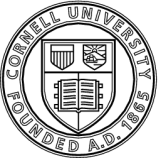
Single ASIC: 192x185 pixels  
ASIC: TSMC 0.25 micron.

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1 quadrant of SLAC's detector based  
off Cornell's ASIC.

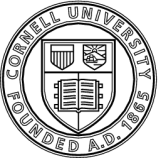
Picture from:  
<http://www.amaroq.com/ryan/2010/04/02/pictures-of-cxi-detector>



# *Specs for Detector*

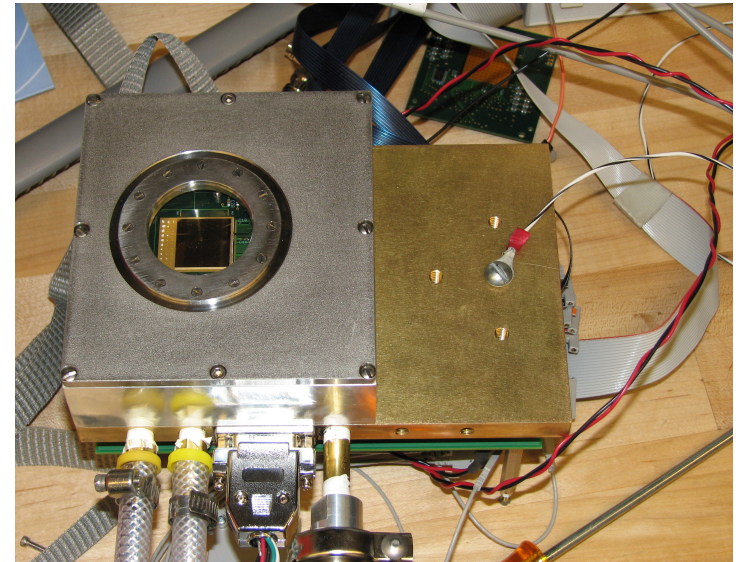
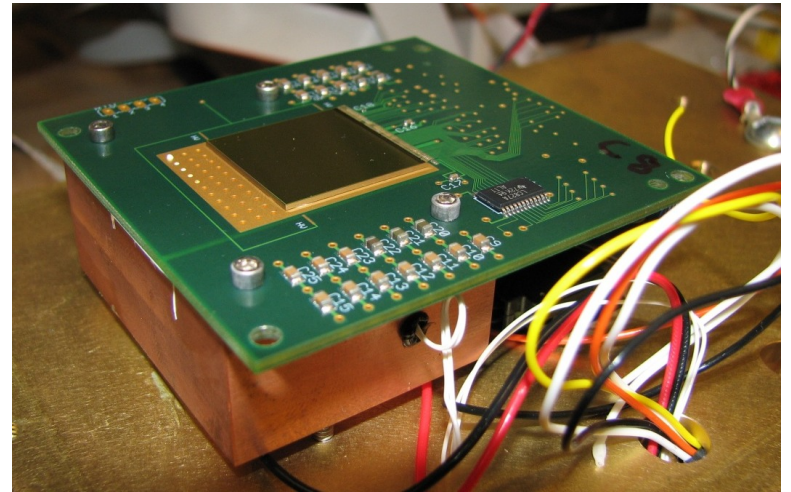
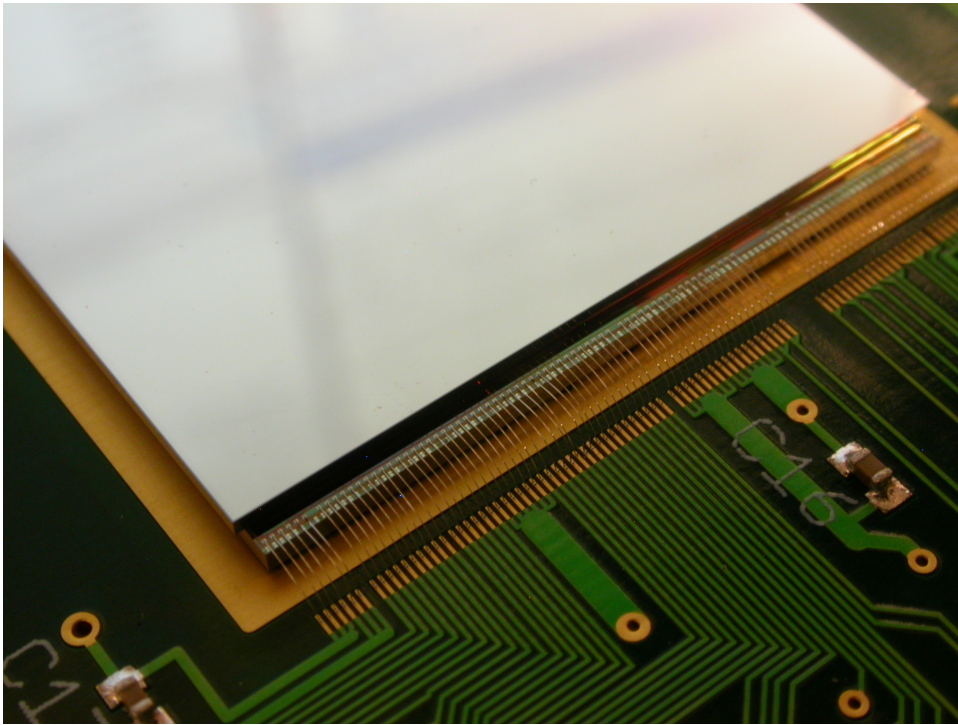
Parameter	Requirement	
Energy Range	4-8 keV	
Well-depth/pixel	$10^3$	~2500
Readout frame rate	120 Hz	
Signal/Noise	>3 for single 8 keV photon	7
DQE	> 90% at 8 keV	
Pixel size	100-200 $\mu\text{m}$	110x110 $\mu\text{m}$
Detector area	> 500x500 pixels	1500x1500





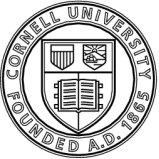
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# *Single Module*

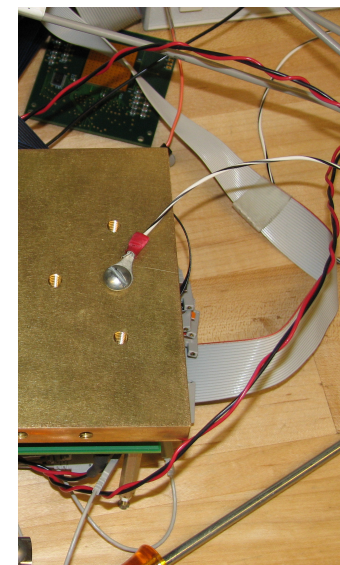
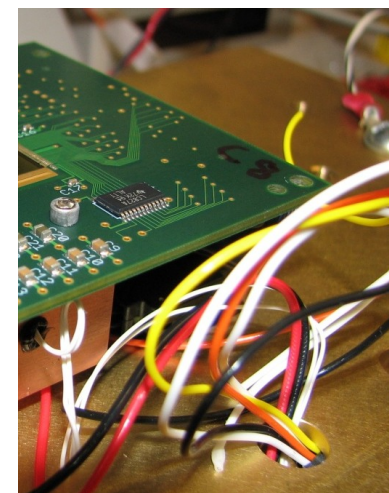
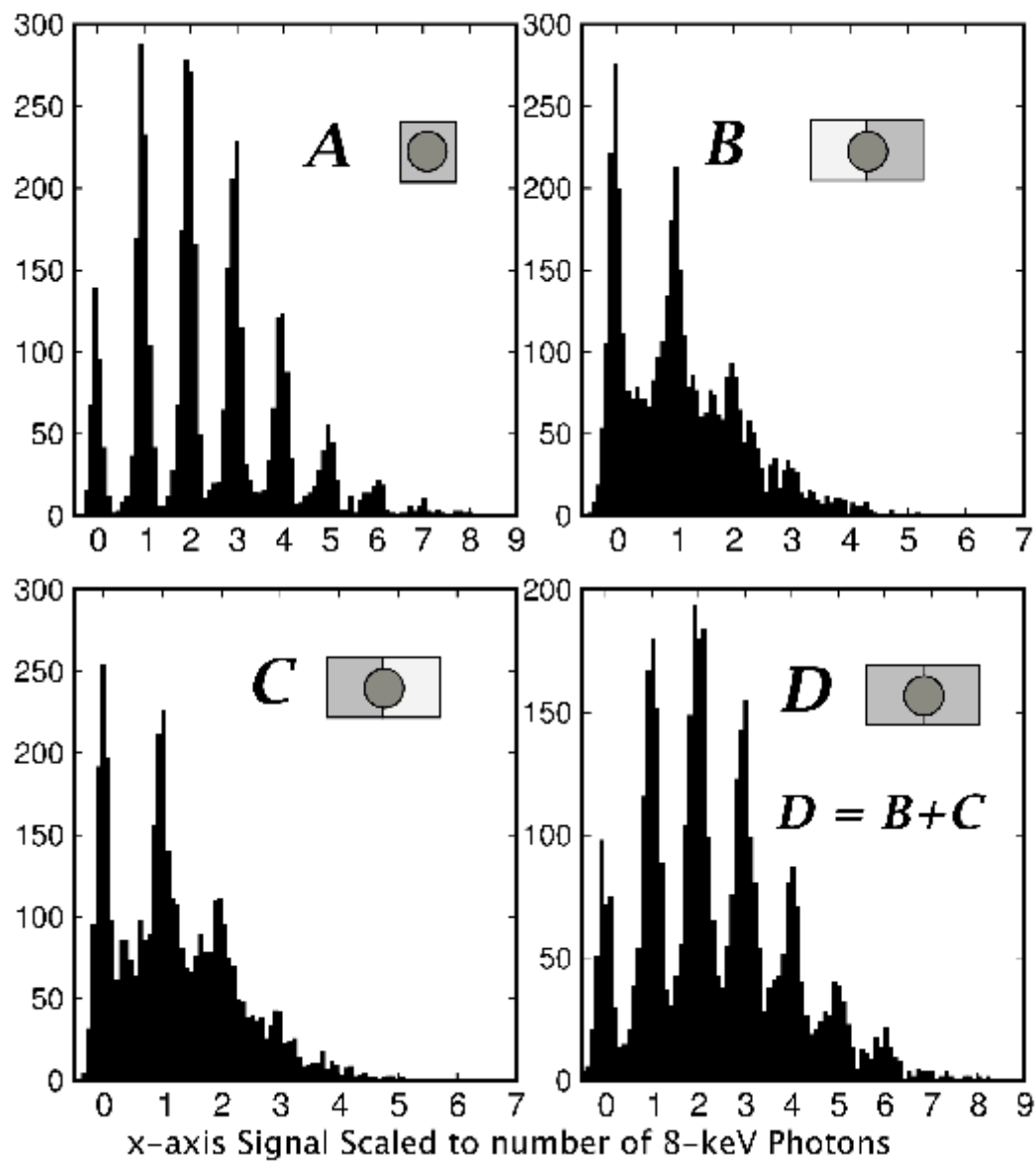


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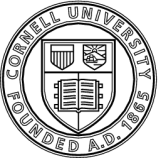
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# Single Photon Detection

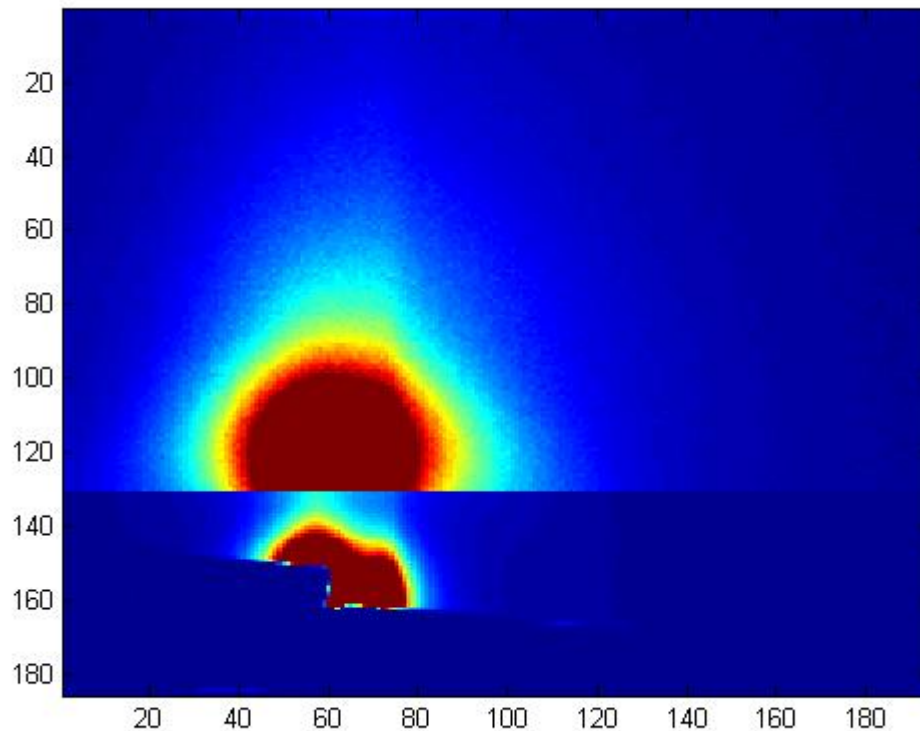






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# *Tested in Experiments*

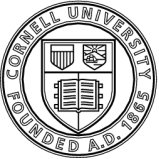


External movie

In-situ CNT forest growth.  
Collaboration with  
John Hart group,  
University of Michigan.

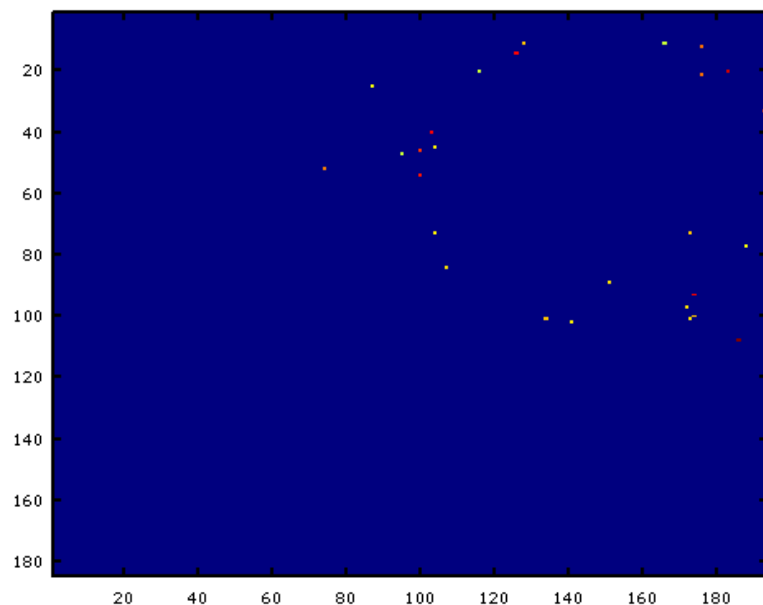
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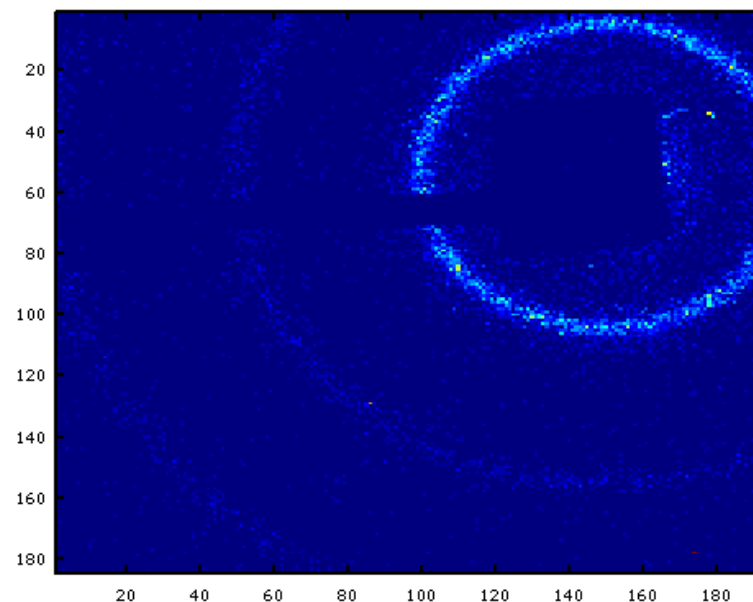


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# *Adding Many Frames*



Single Frames



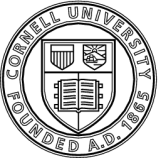
Added, thresholded frames  
Up to ~350 frames

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Have external movie.

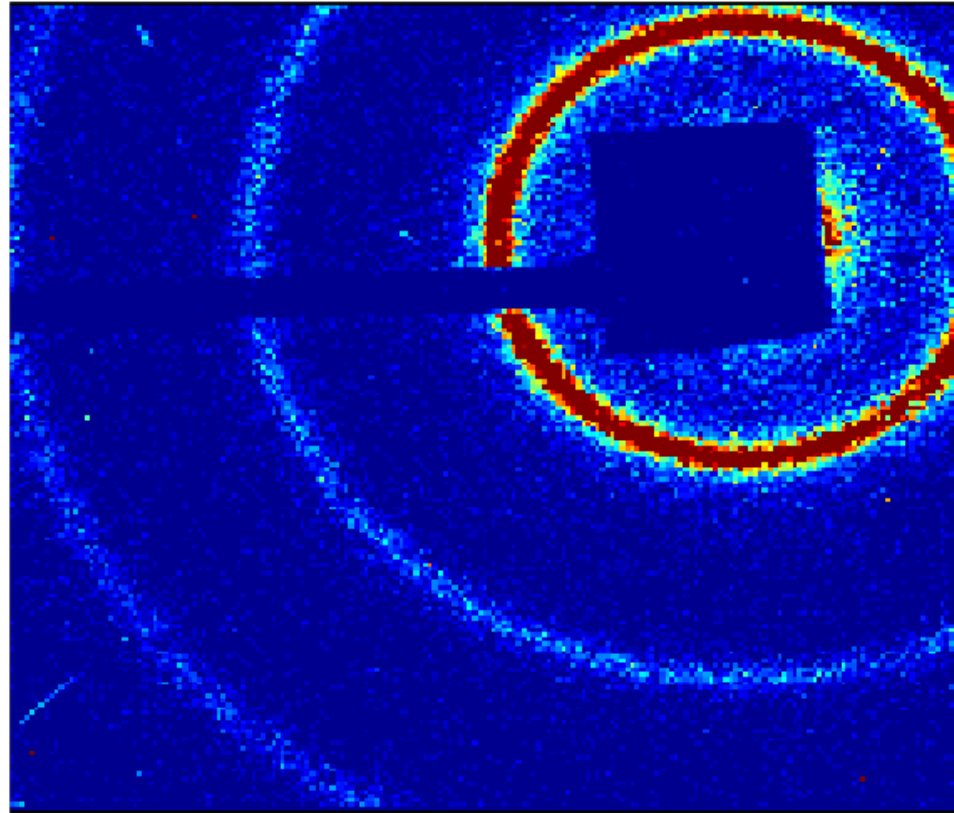
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# *Adding Many Frames*

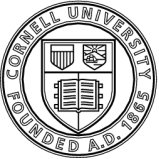


3000 frames, low flux data

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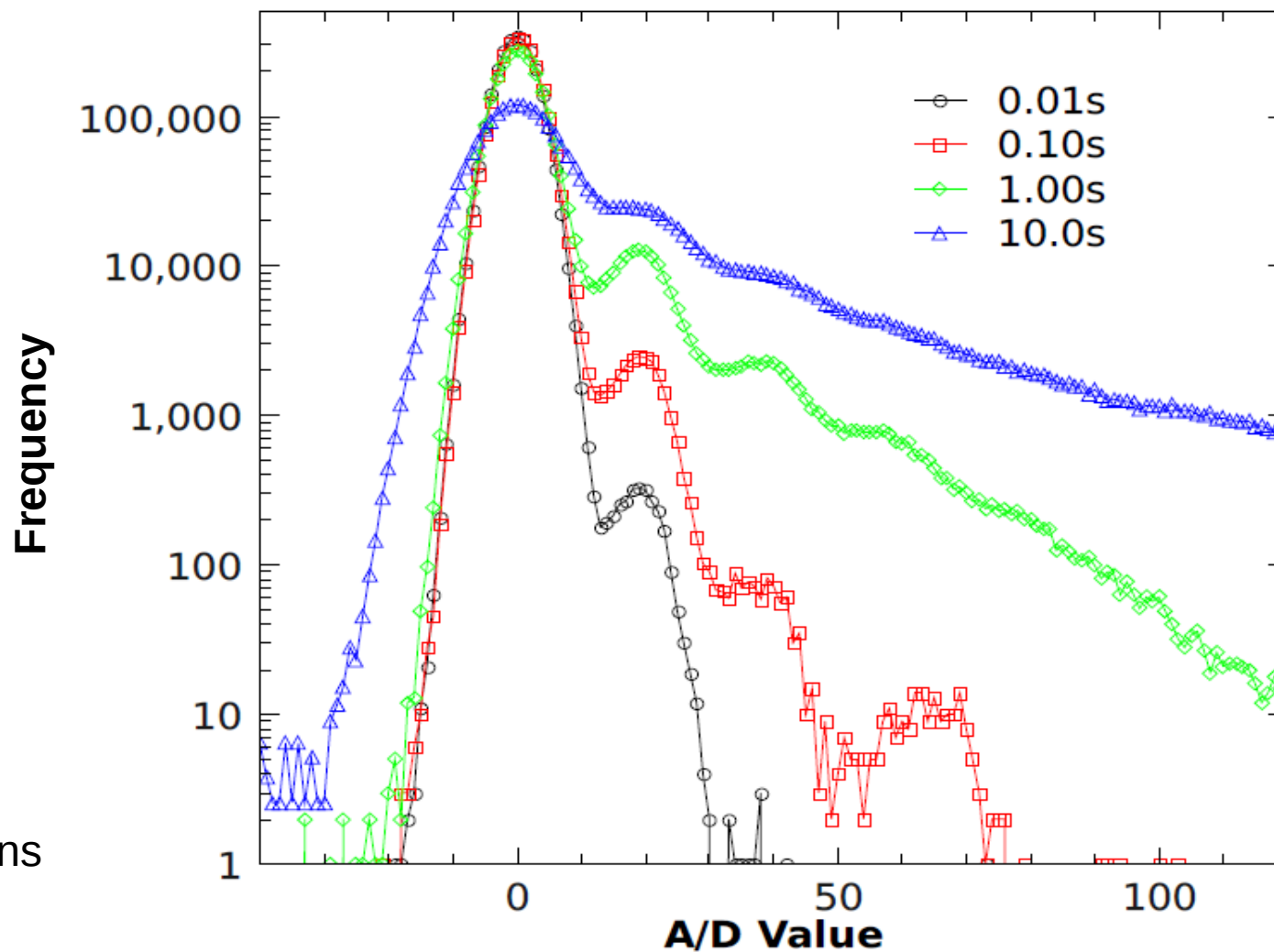
Have video file.

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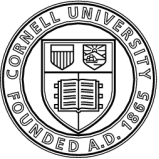


# *Whole Array Peak Histogram*

## Photon Peak Histogram

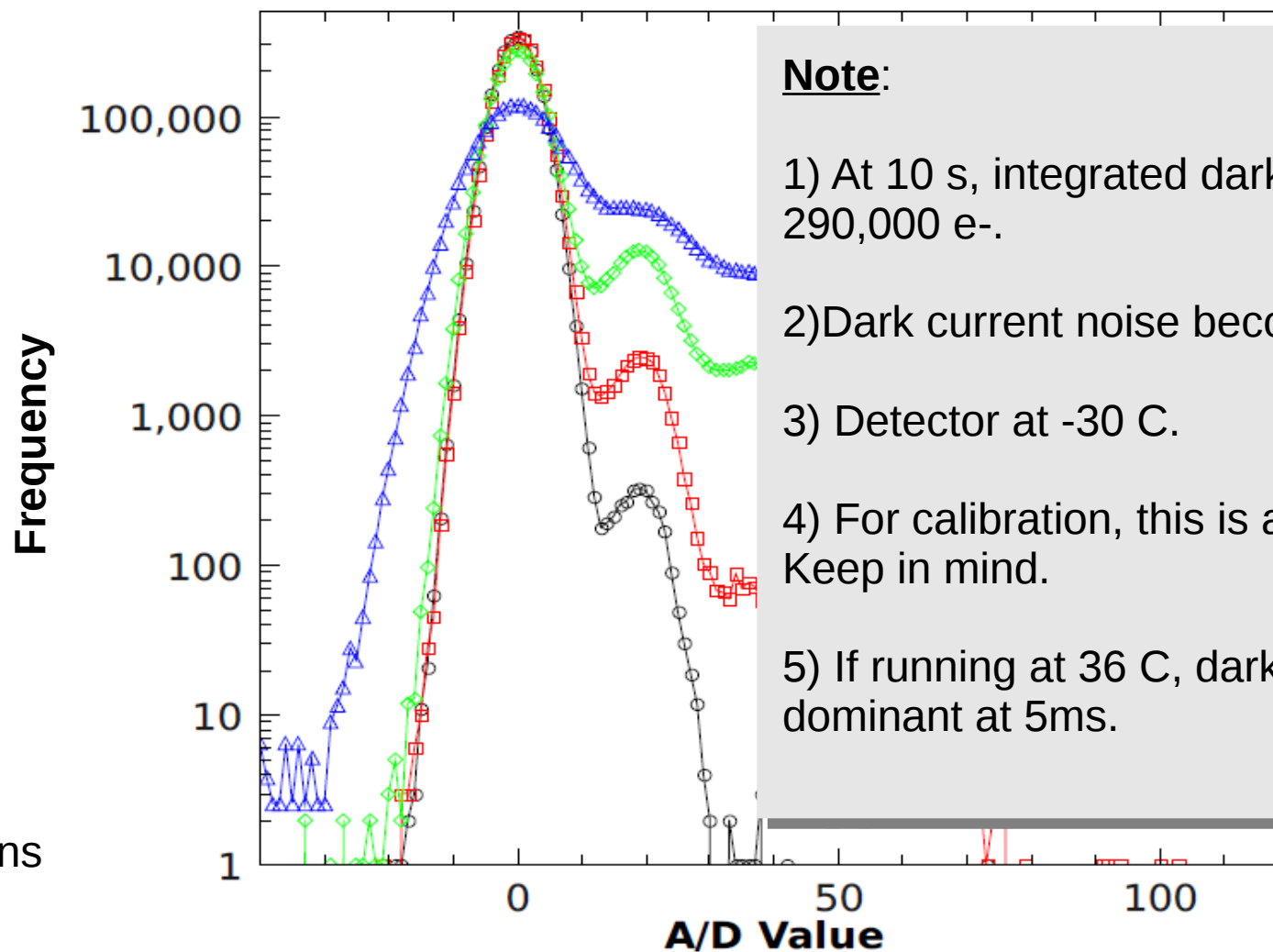


8 keV photons



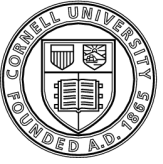
# *Whole Array Peak Histogram*

## Photon Peak Histogram



### Note:

- 1) At 10 s, integrated dark current is 290,000 e<sup>-</sup>.
- 2) Dark current noise becomes dominant.
- 3) Detector at -30 C.
- 4) For calibration, this is a good thing to Keep in mind.
- 5) If running at 36 C, dark current noise dominant at 5ms.



# *Adding many frames*

- Basic problems

- 1) How well do you know the background.

Under best circumstances you have:  $\sigma_{back} = \frac{\sigma_{pix}}{\sqrt{N_{back\ frames}}}$

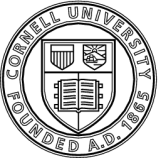
- So, if you want to have single photon sensitivity:

$$N_{Sigframes} * \sigma_{back} = \frac{N_{Sigframes} \sigma_{pix}}{\sqrt{N_{backframes}}} \ll 1$$



(hand wave)  $N_{backframes} \gg \sigma_{pix}^2 N_{sigframes}^2$





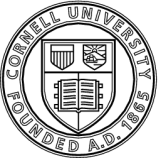
# *Adding many frames*

- ◆ Basic problems

1) How well do you know the background.

- ◆  $N_{back\ frames} > \sigma_{pix}^2 N_{sig\ frames}^2$  (very optimistic)

- ◆ Are there any small scale drifts, frame to frame variations?



# *Adding many frames*

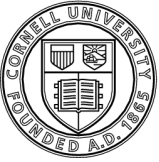
## ♦ Basic problems

1) How well do you know the background.

♦  $N_{back\ frames} > \sigma_{pix}^2 N_{sig\ frames}^2$  (at very least)

♦ Are there any small scale drifts, frame to frame variations?

♦ Errors in background add when adding frames.



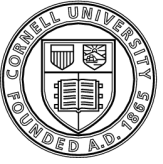
# *Adding many frames*

## ♦ Basic problems

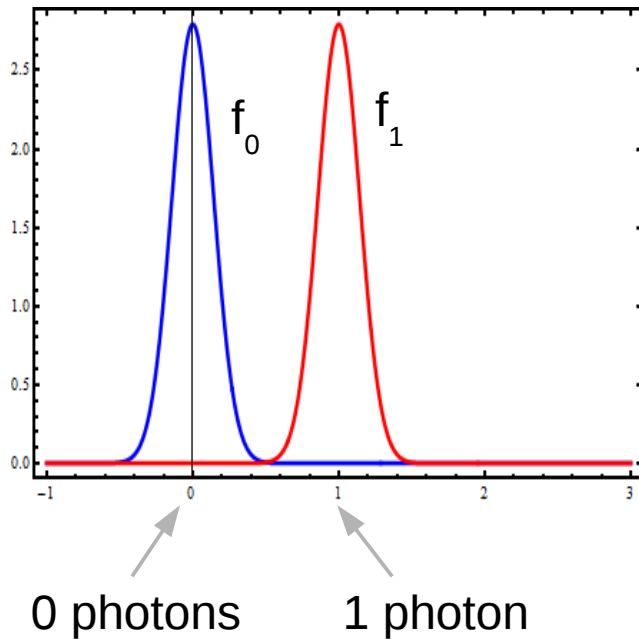
1) How well do you know the background.

♦  $N_{back\ frames} > \sigma_{pix}^2 N_{sig\ frames}^2$  (at very least)

- ♦ Are there any small scale drifts, frame to frame variations?
- ♦ Errors in background add when adding frames.
- ♦ Applying a threshold allows you to avoid such stringent requirements.

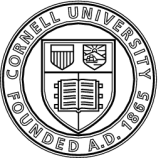


# Thresholding

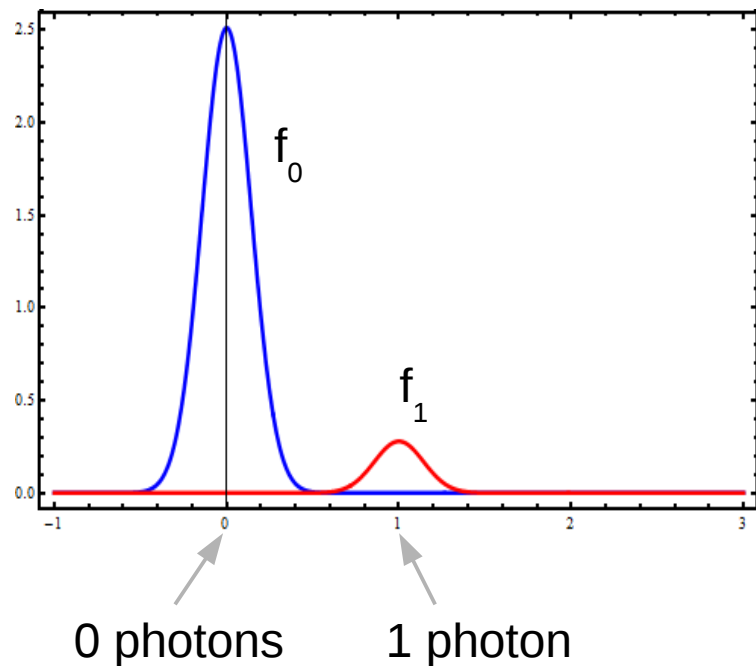


- Two distributions: 0 and 1 photons.
- If frequency (photon/pixel/frame) for 1 photon is  $A$ .
- Question – for some  $A$ , at what level does the threshold need  $f_0$  'noise' is low compared to  $f_1$  'signal'.





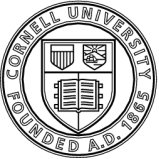
# Thresholding



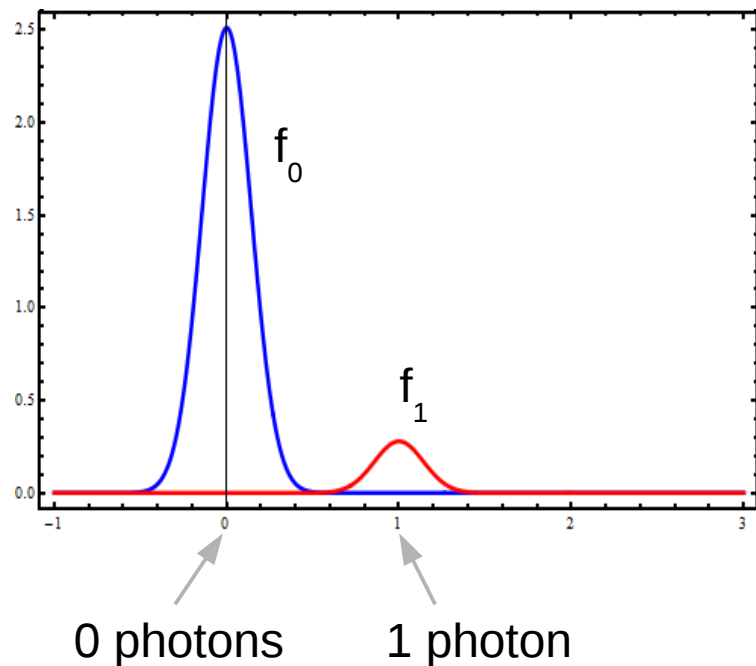
- Two distributions: 0 and 1 photons.
- If frequency (photon/pixel/frame) for 1 photon is A.
- Question – for an A, at what level does the threshold need f<sub>0</sub> 'noise' is low compared to f<sub>1</sub> 'signal'.

For half signal from each:

$$\int_{thr}^{\infty} f_0 = \int_{thr}^{\infty} f_1$$



# Thresholding

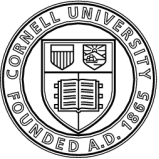


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- Question – for an  $A$ , at what level does the threshold need  $f_0$  'noise' is low compared to  $f_1$  'signal'.

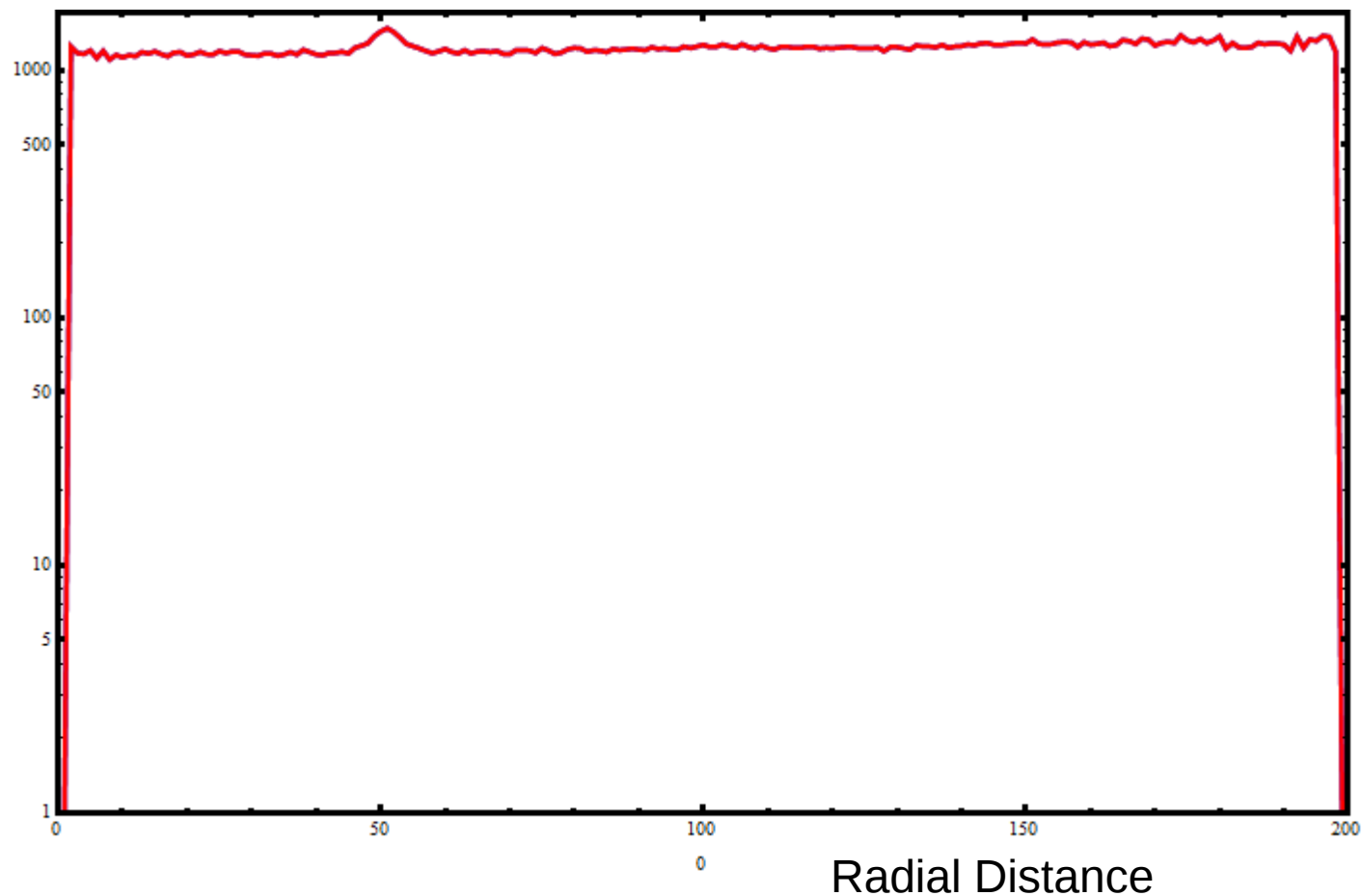
For half signal from each:

$$\int_{thr}^{\infty} f_0 = \int_{thr}^{\infty} f_1$$

For  $A = 1/1000$ , and  $\sigma = 1/7$  photon, at threshold of 0.44 photons, half the signal is from  $f_0$



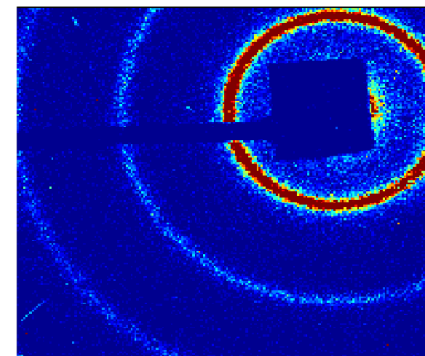
# *Varying Threshold*

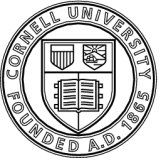


(Azimuthal integration around detector)

~ 20 ADU/x-ray

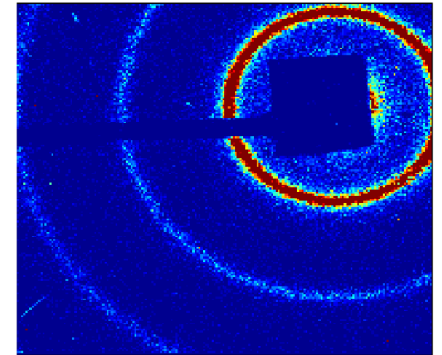
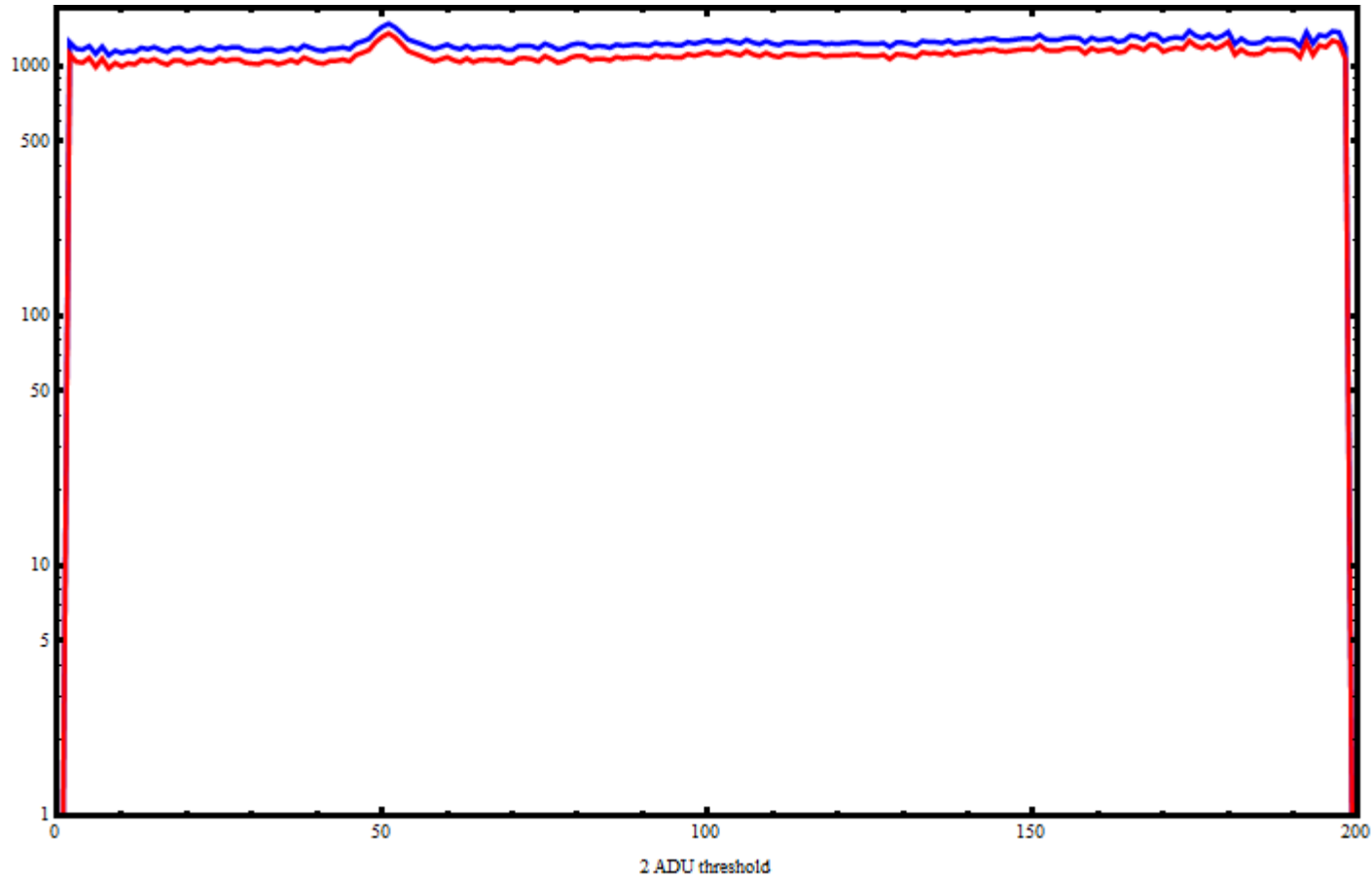
1000 frame sum





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# *Varying Threshold*

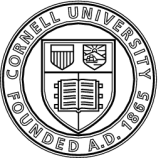


(Azimuthal integration around detector)

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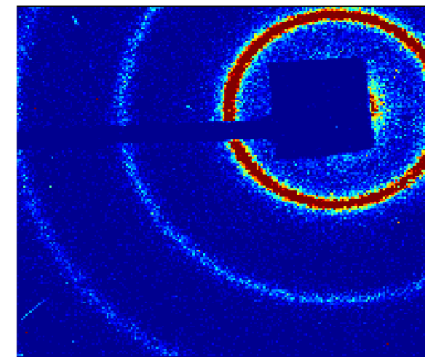
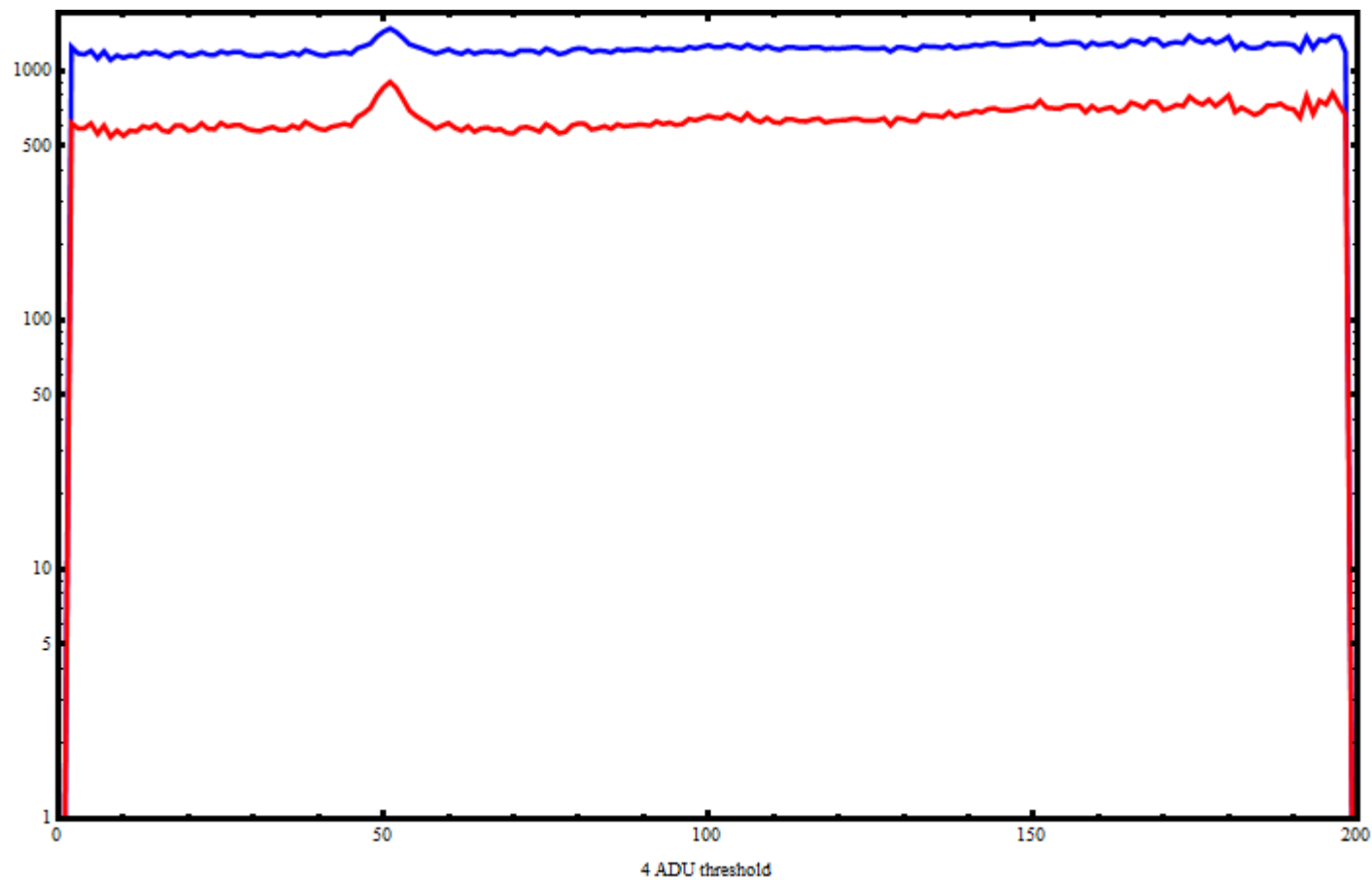
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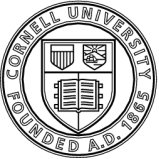
# *Varying Threshold*



(Azimuthal integration around detector)

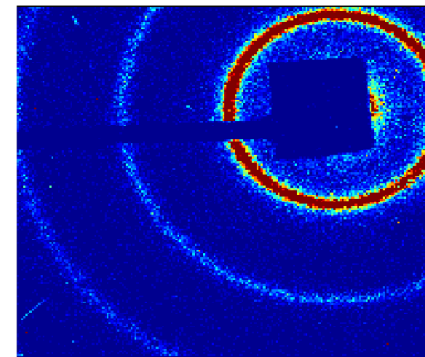
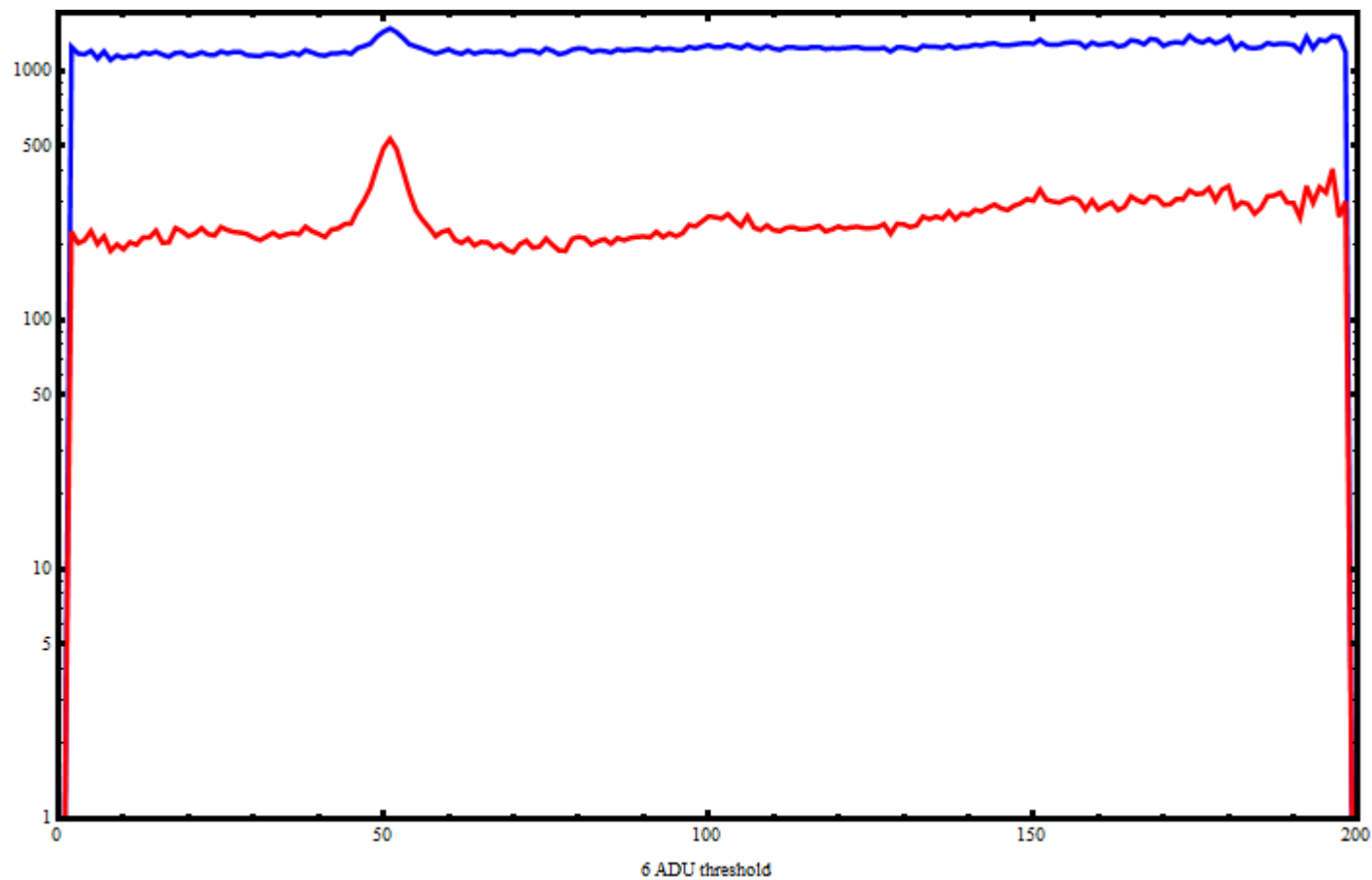
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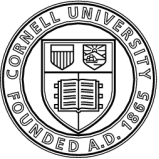
# *Varying Threshold*



(Azimuthal integration around detector)

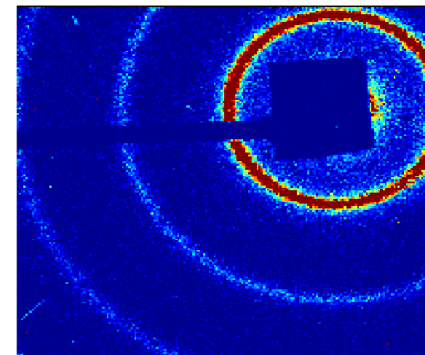
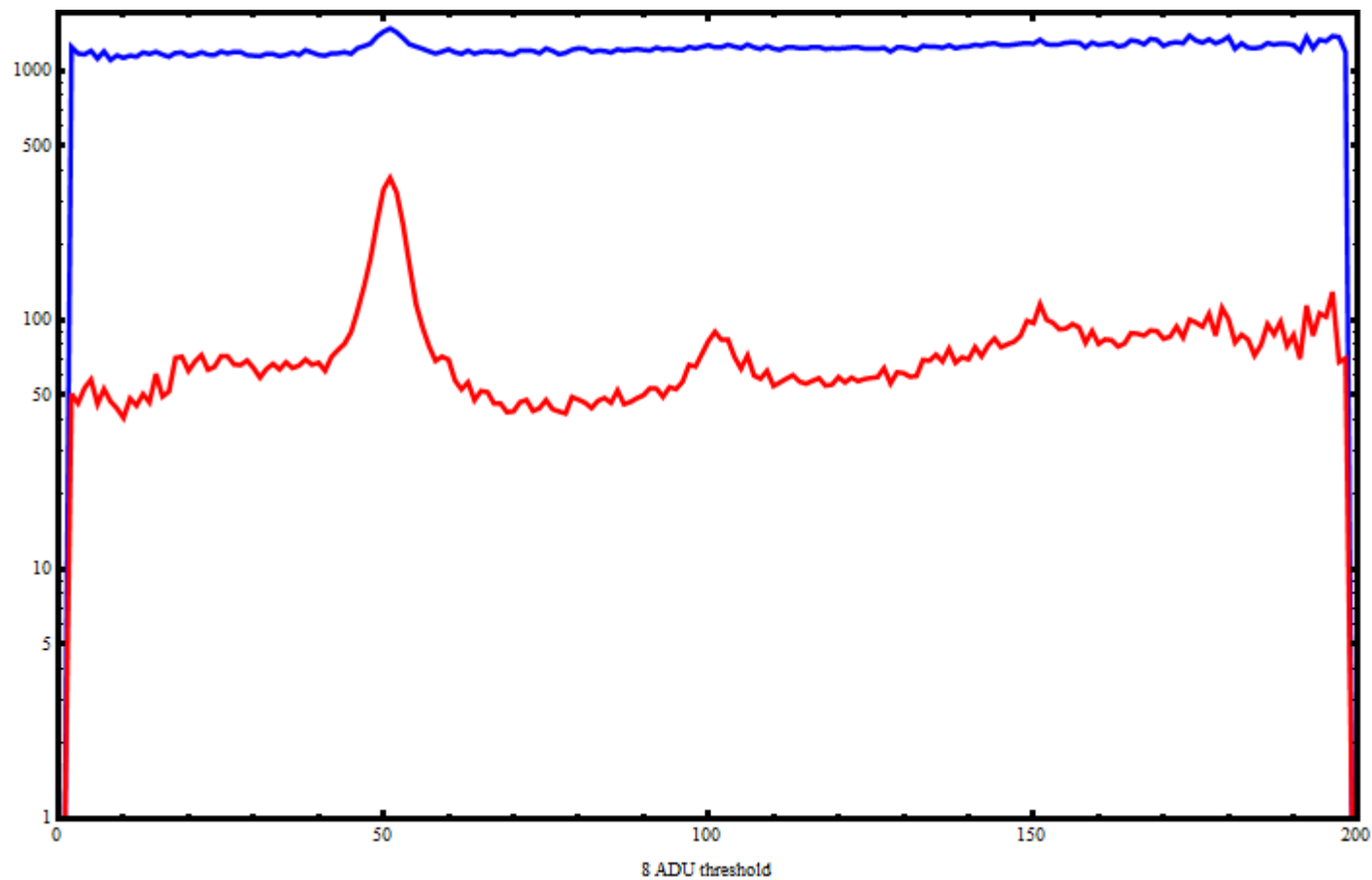
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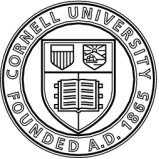
# *Varying Threshold*



(Azimuthal integration around detector)

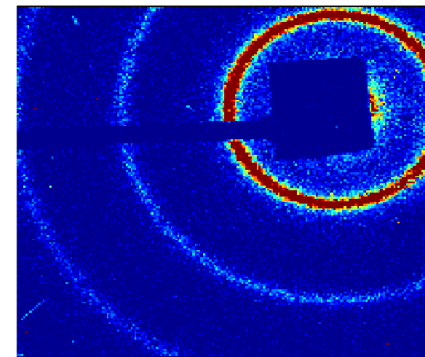
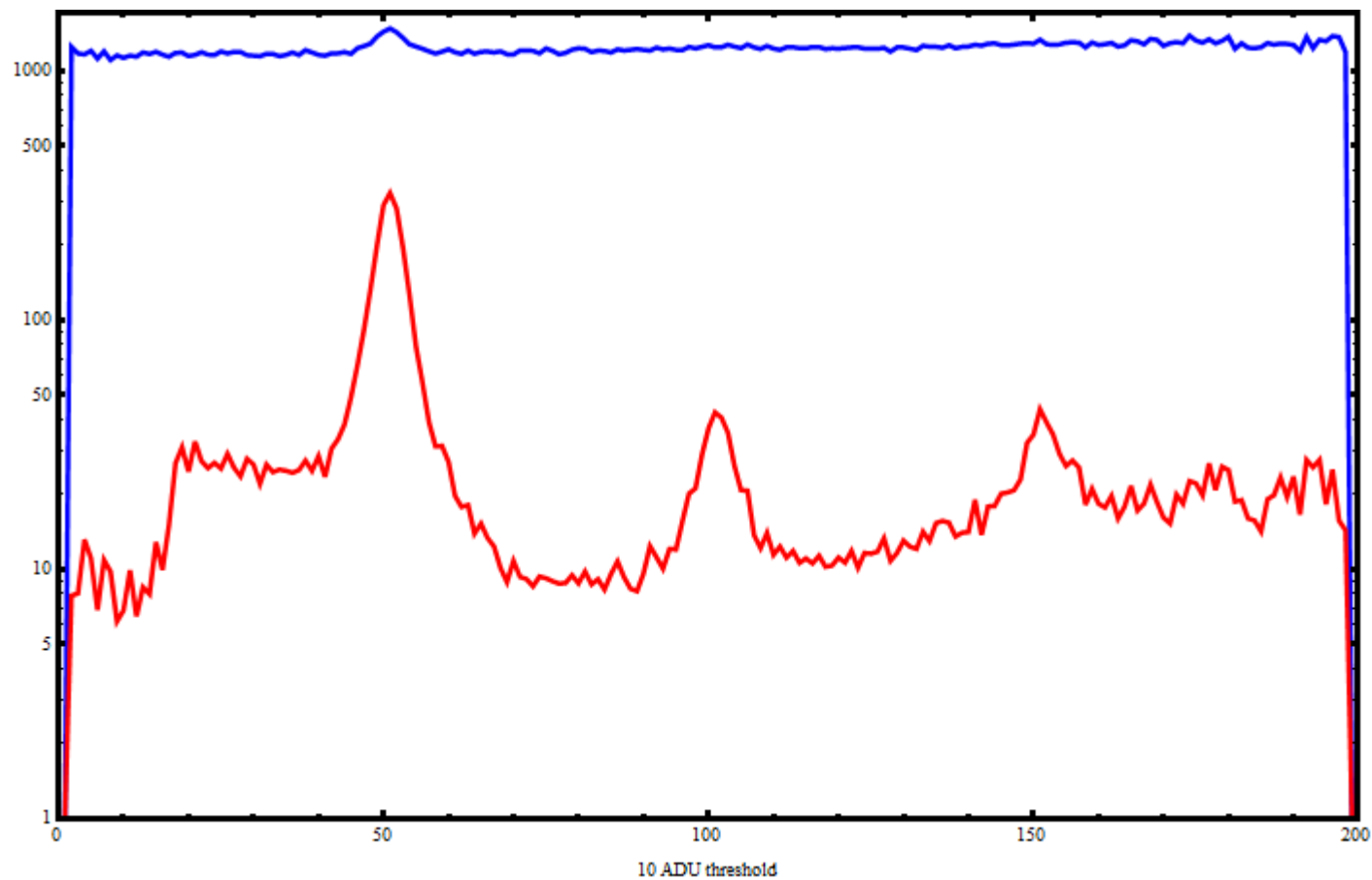
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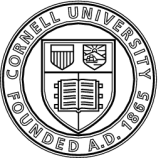
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(Azimuthal integration around detector)

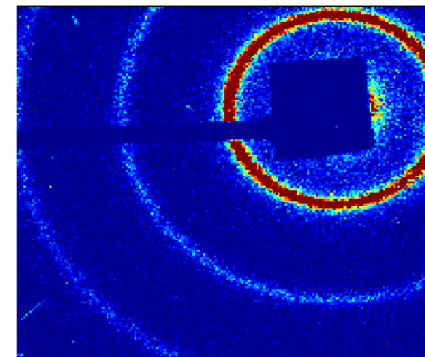
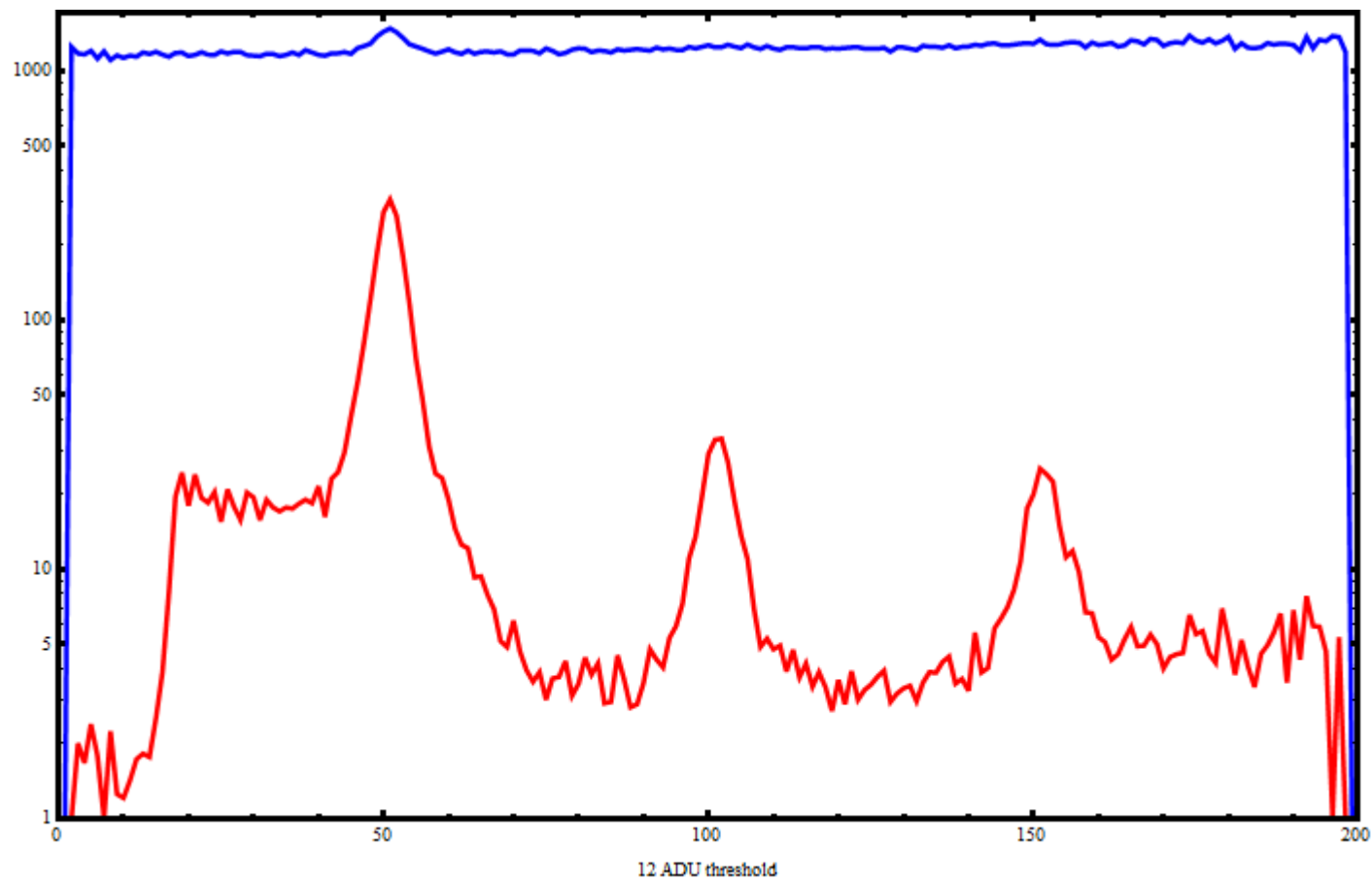
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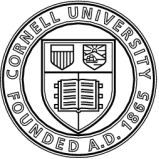
# *Varying Threshold*



(Azimuthal integration around detector)

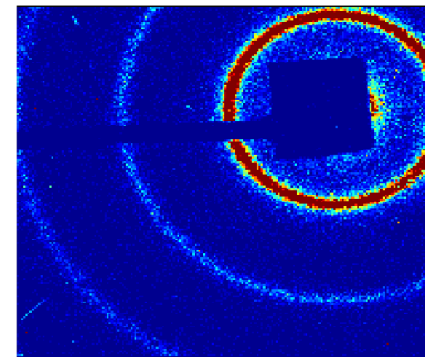
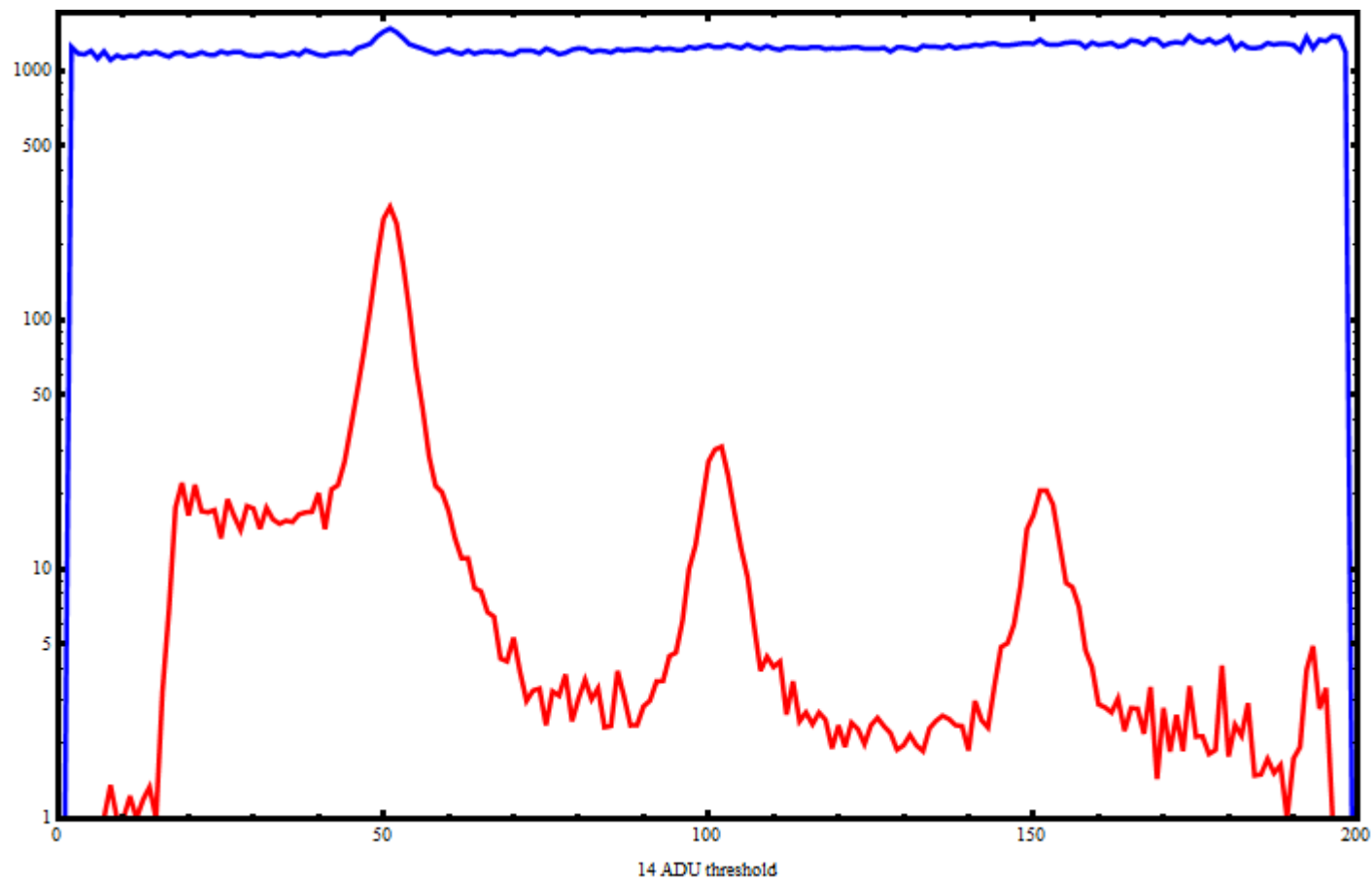
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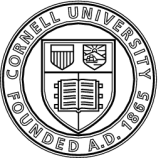


(Azimuthal integration around detector)

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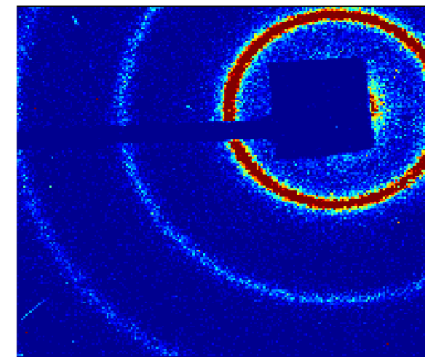
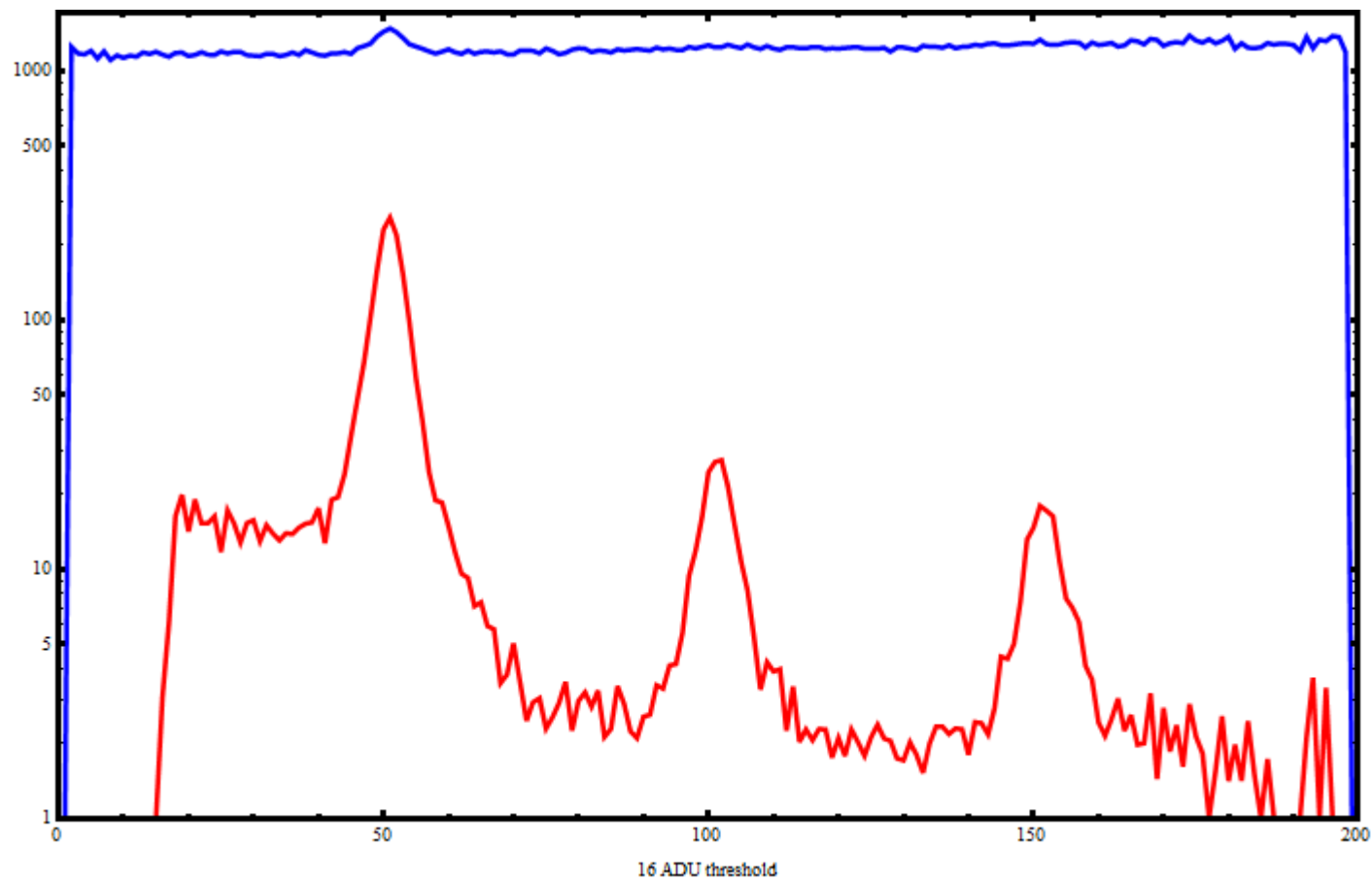
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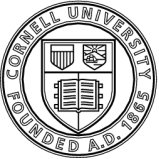
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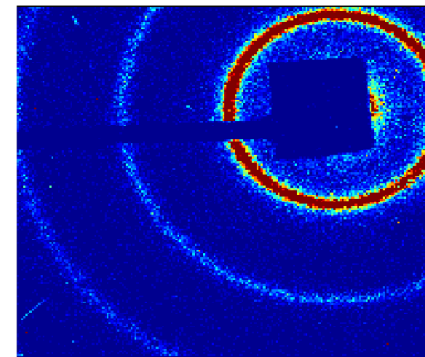
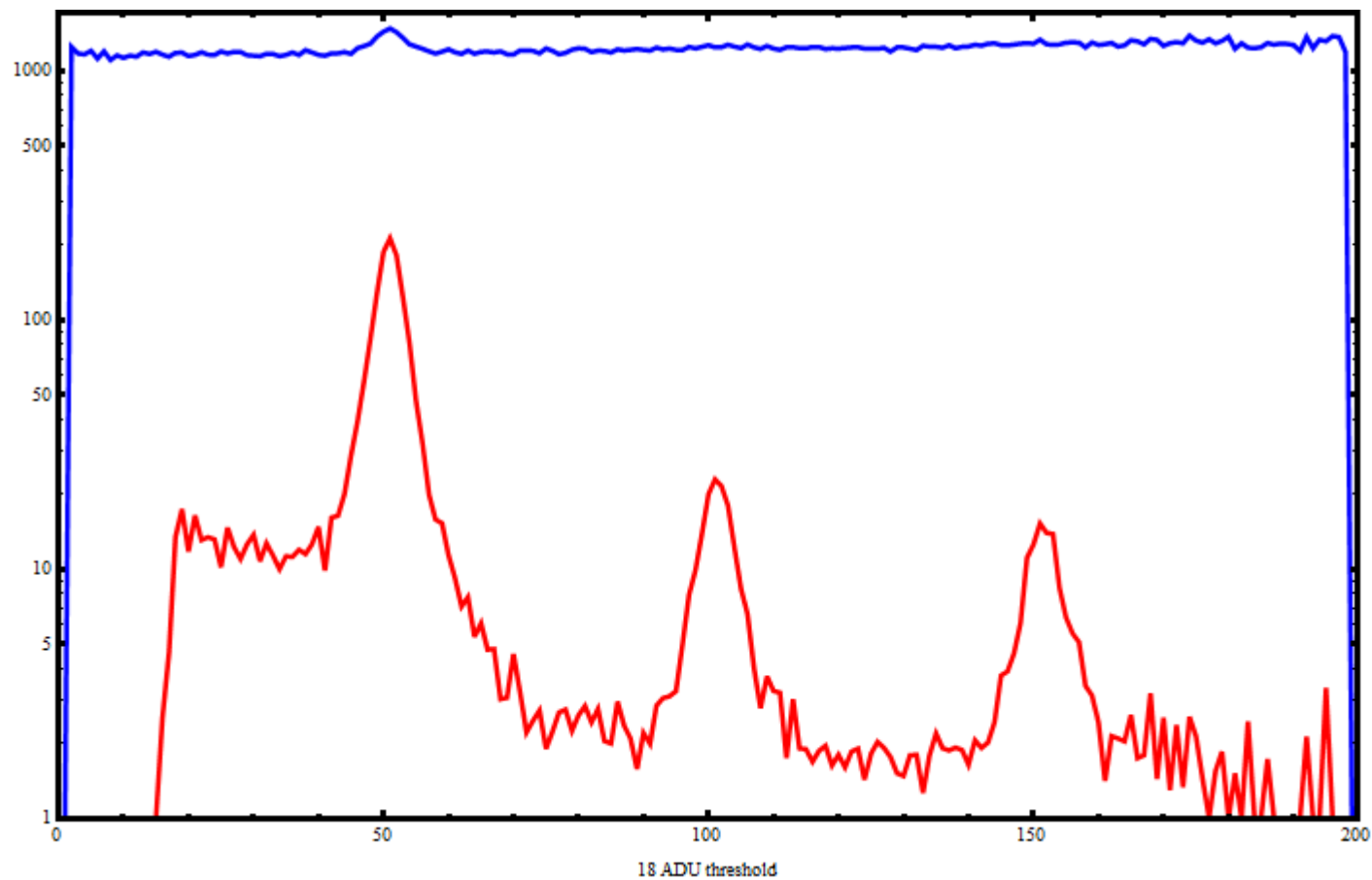
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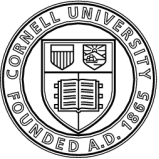
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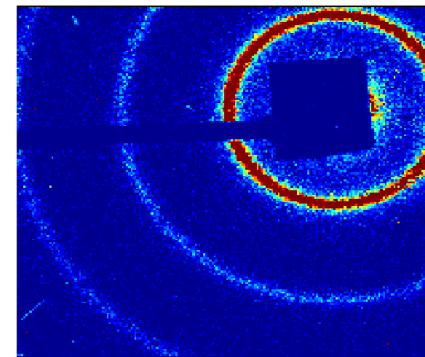
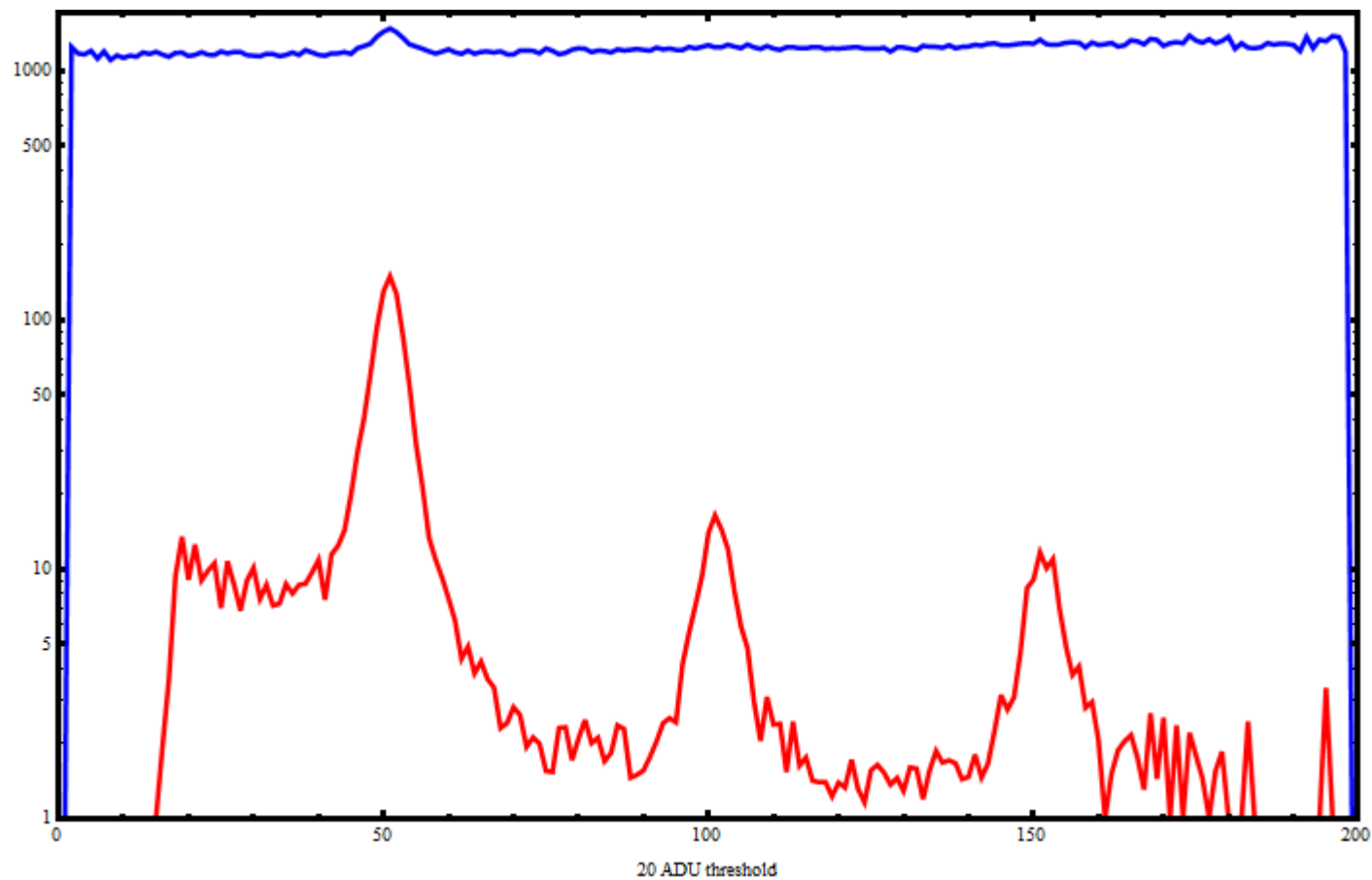
(Azimuthal integration around detector)

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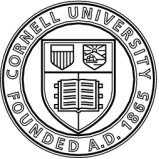
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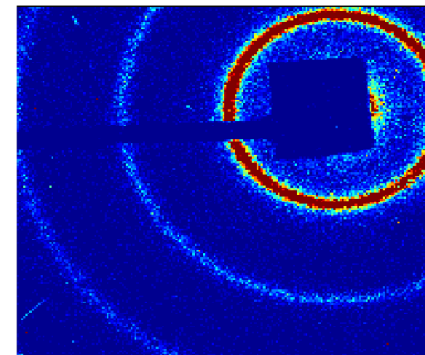
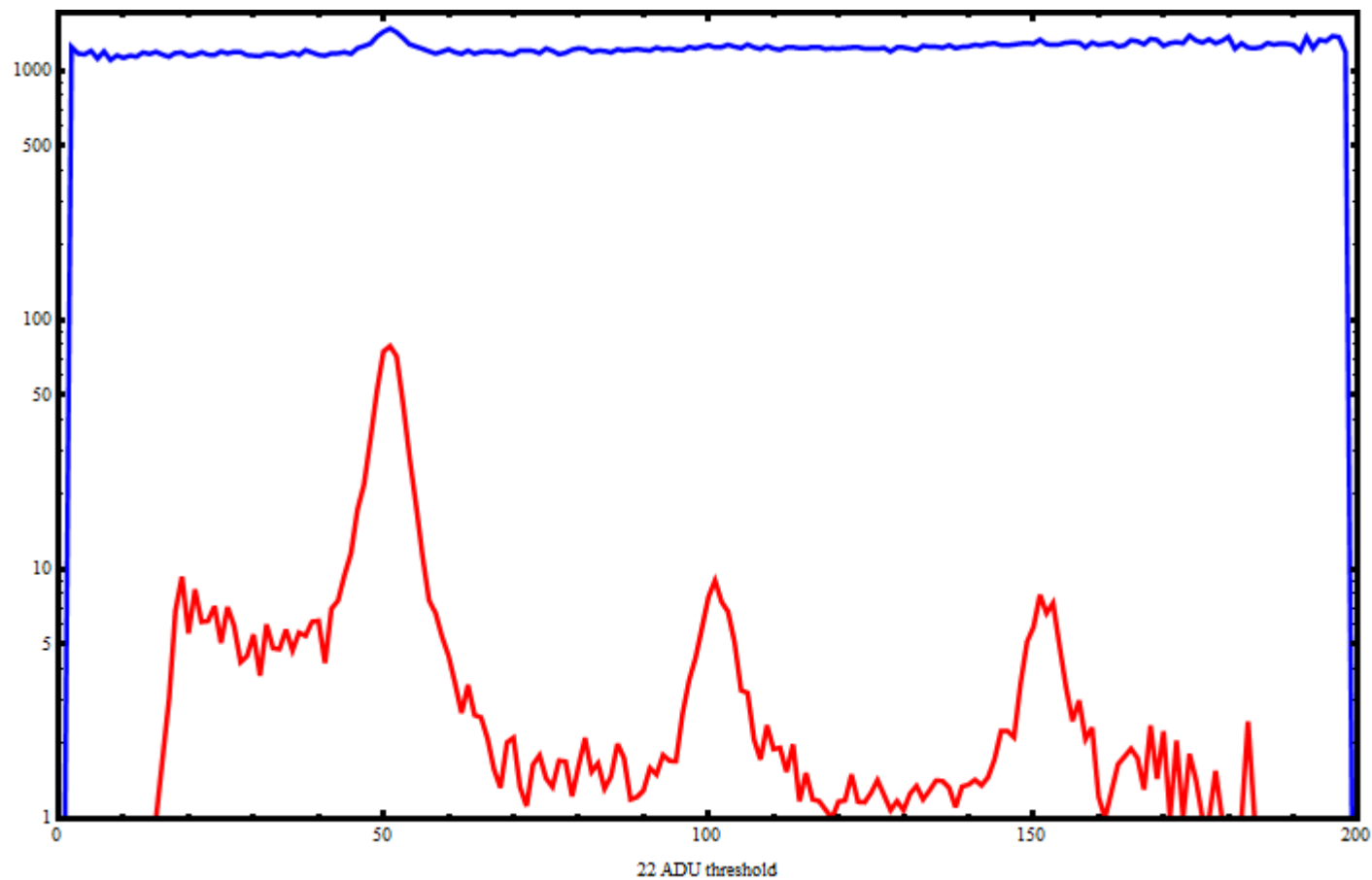


(Azimuthal integration around detector)



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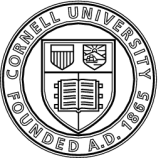
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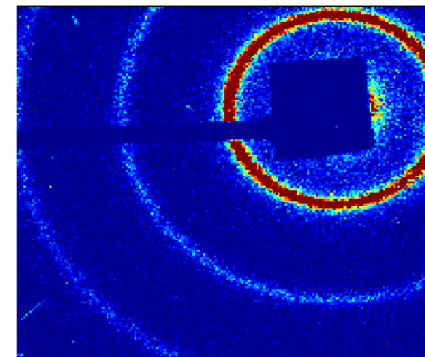
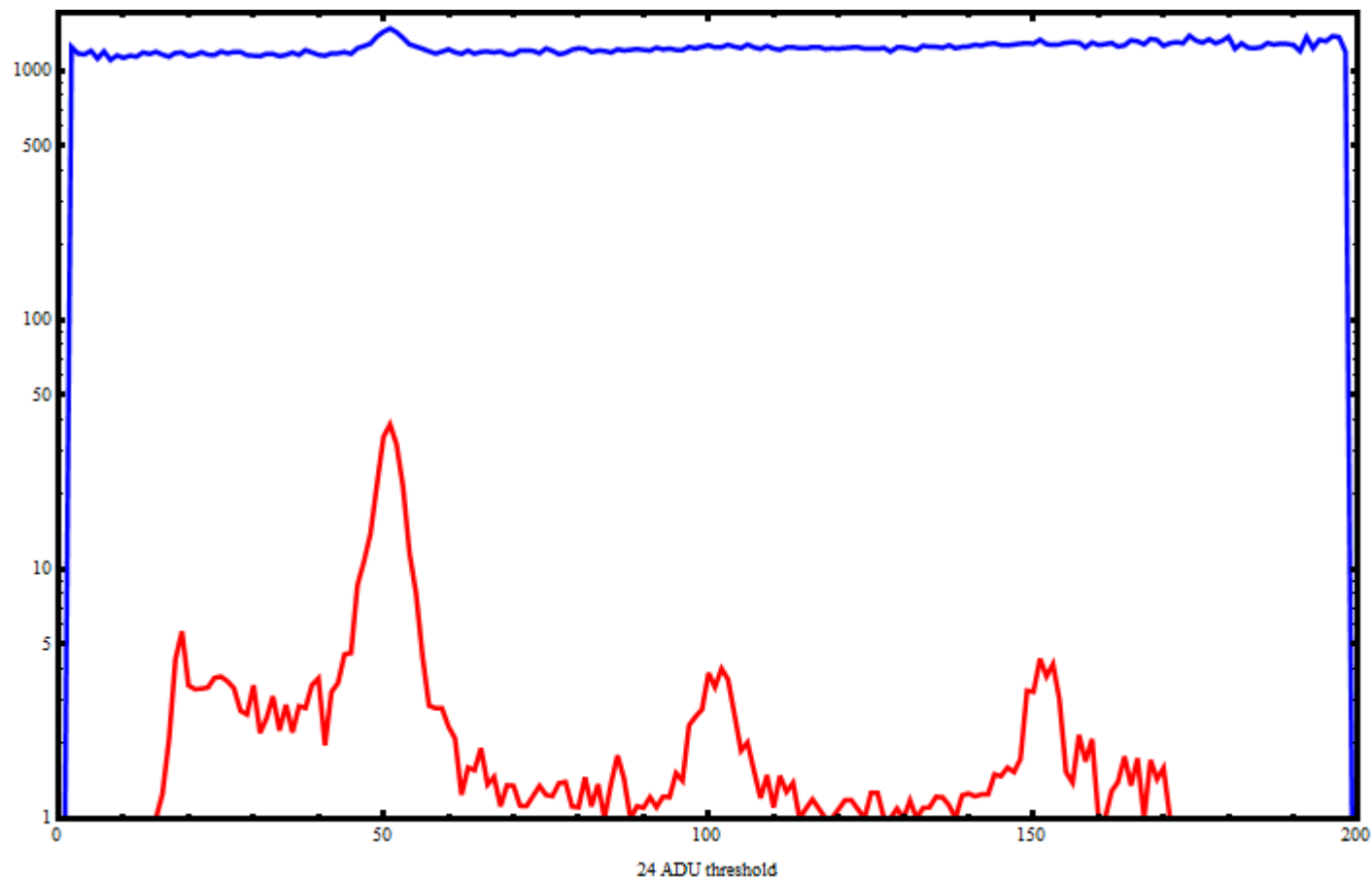
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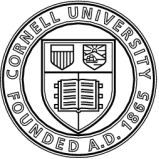
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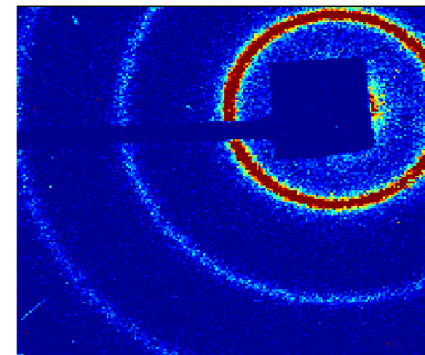
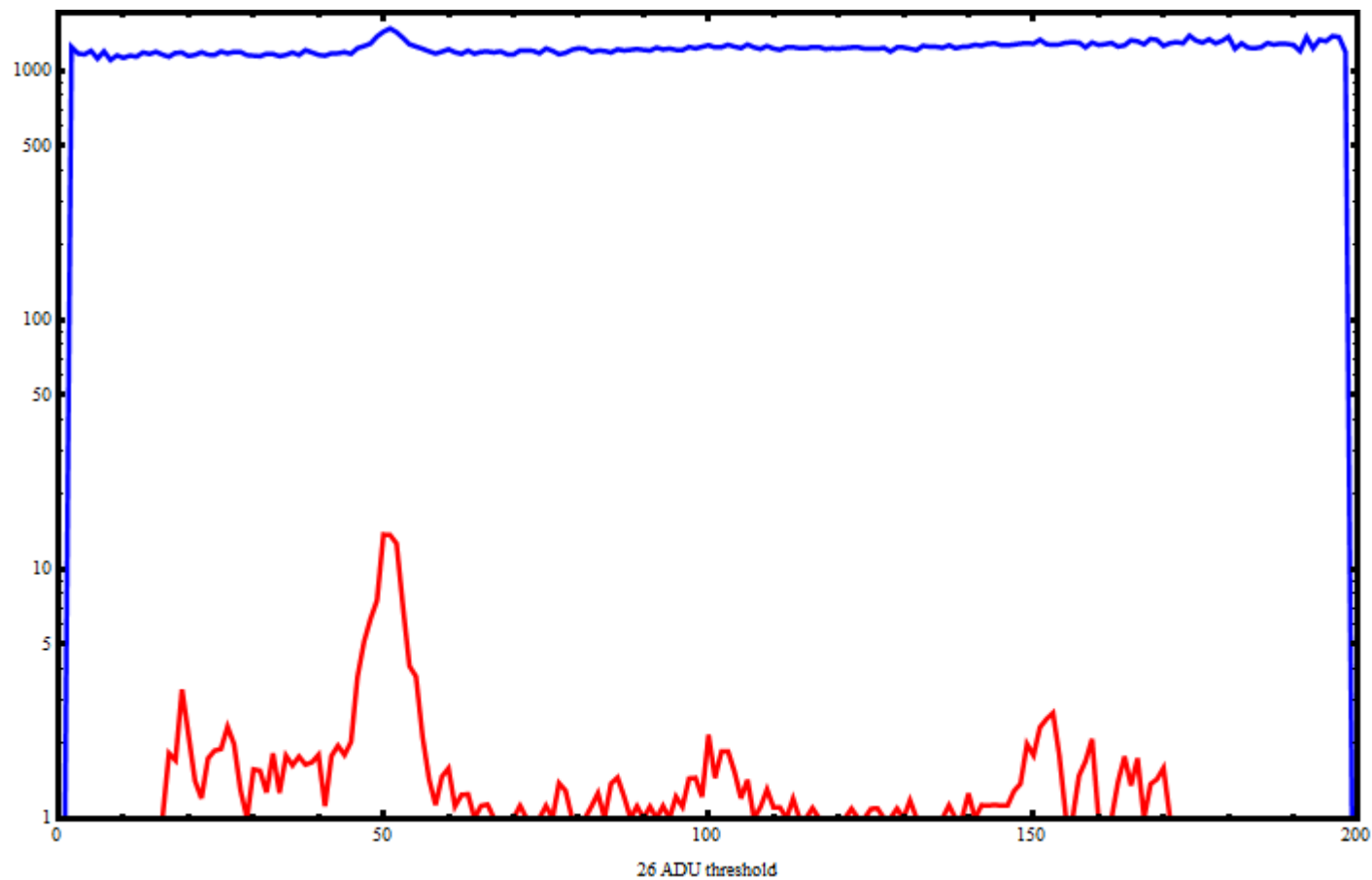
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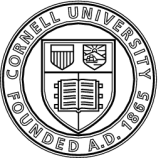


(Azimuthal integration around detector)

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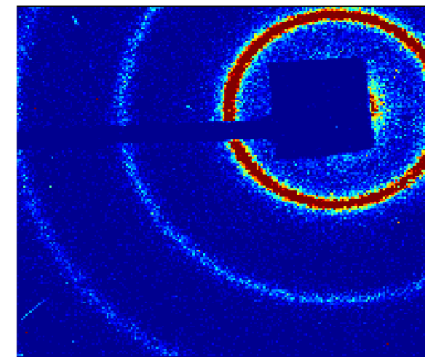
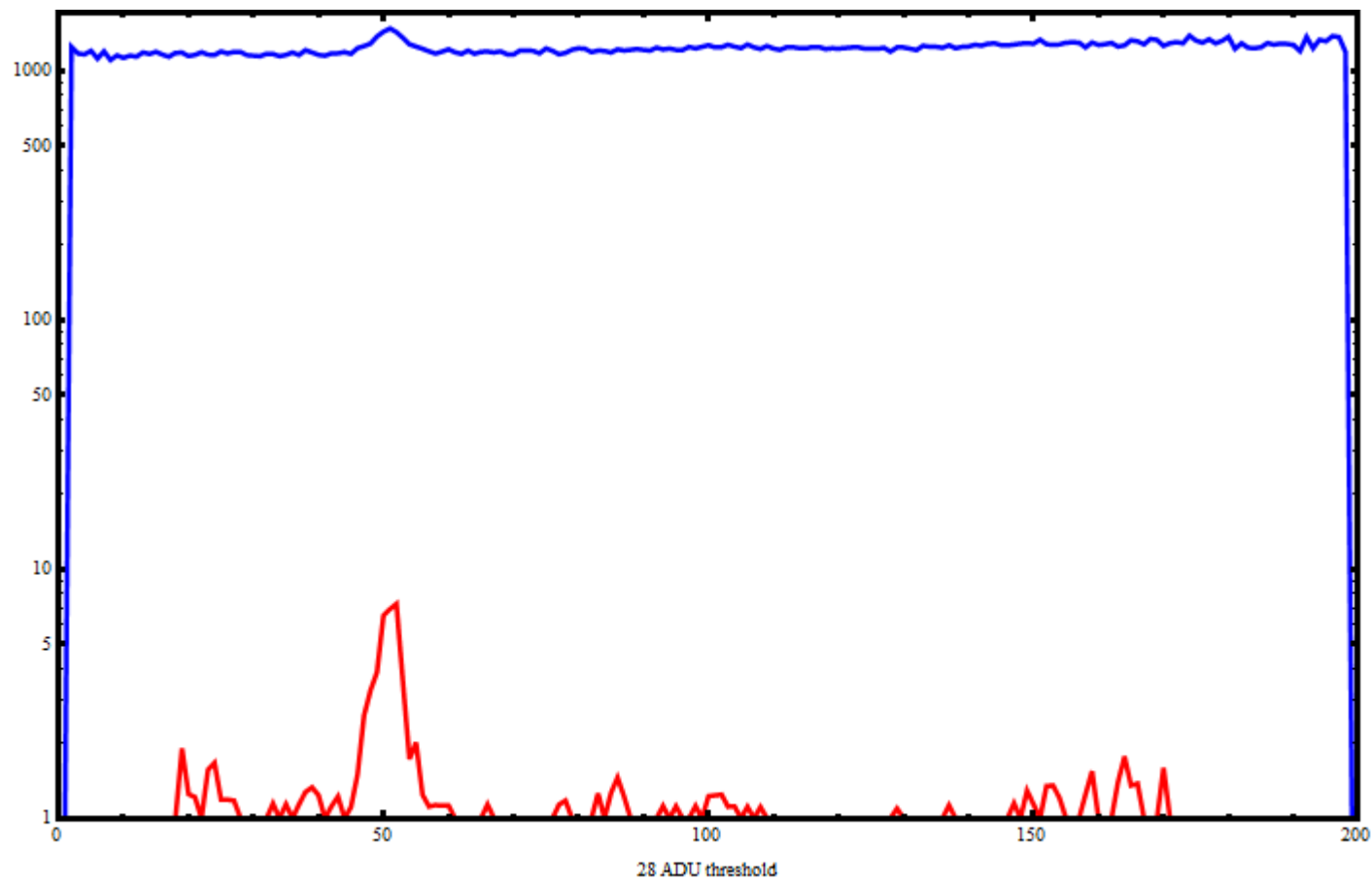
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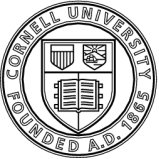
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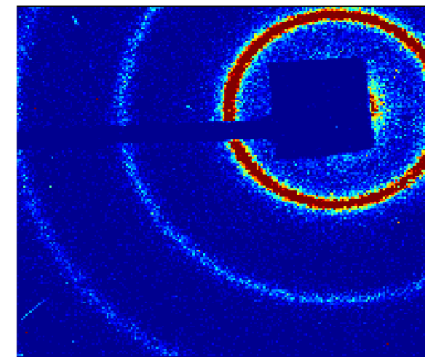
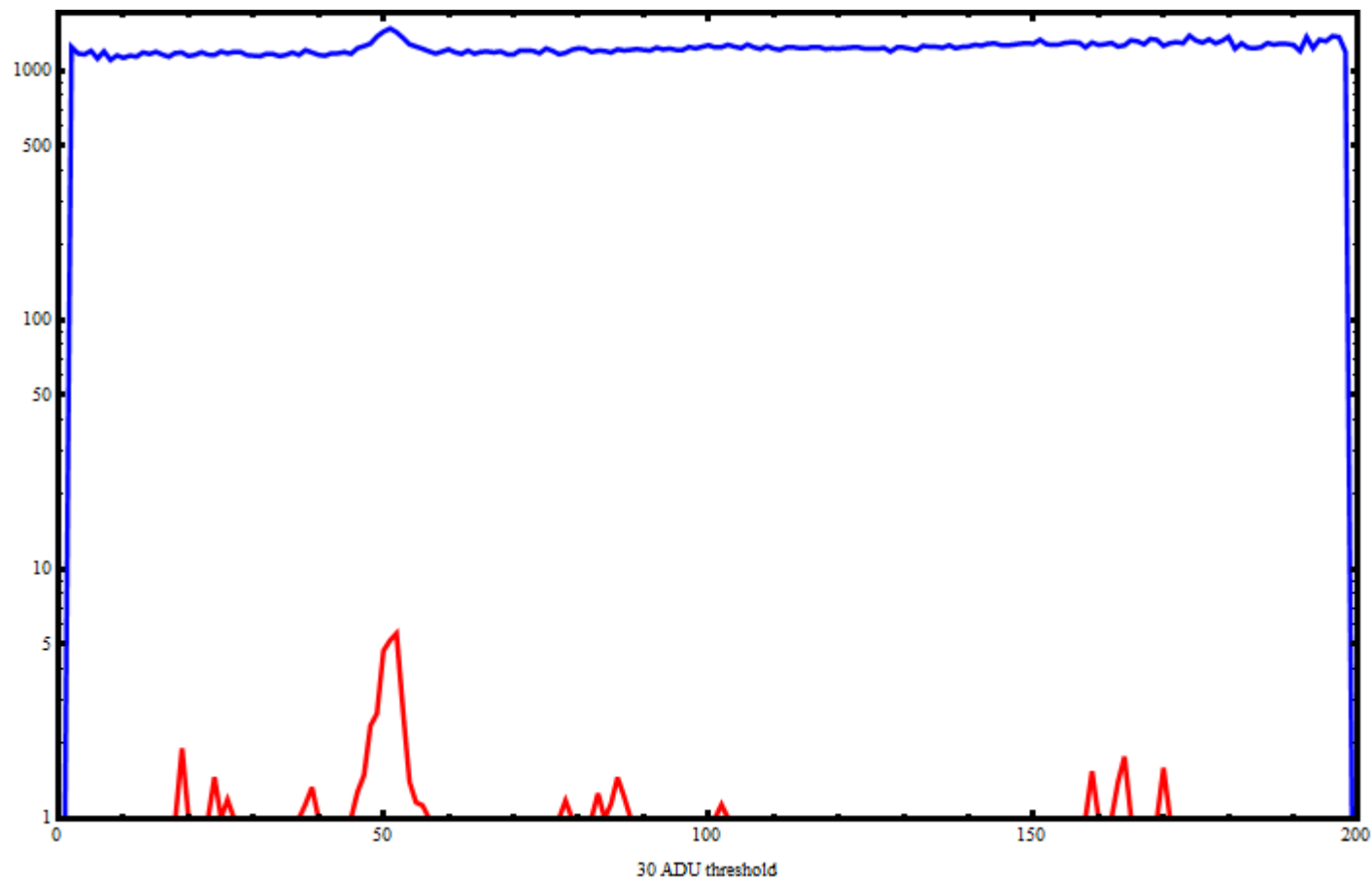
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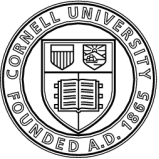
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(Azimuthal integration around detector)

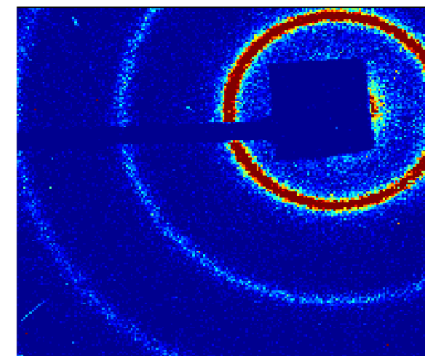
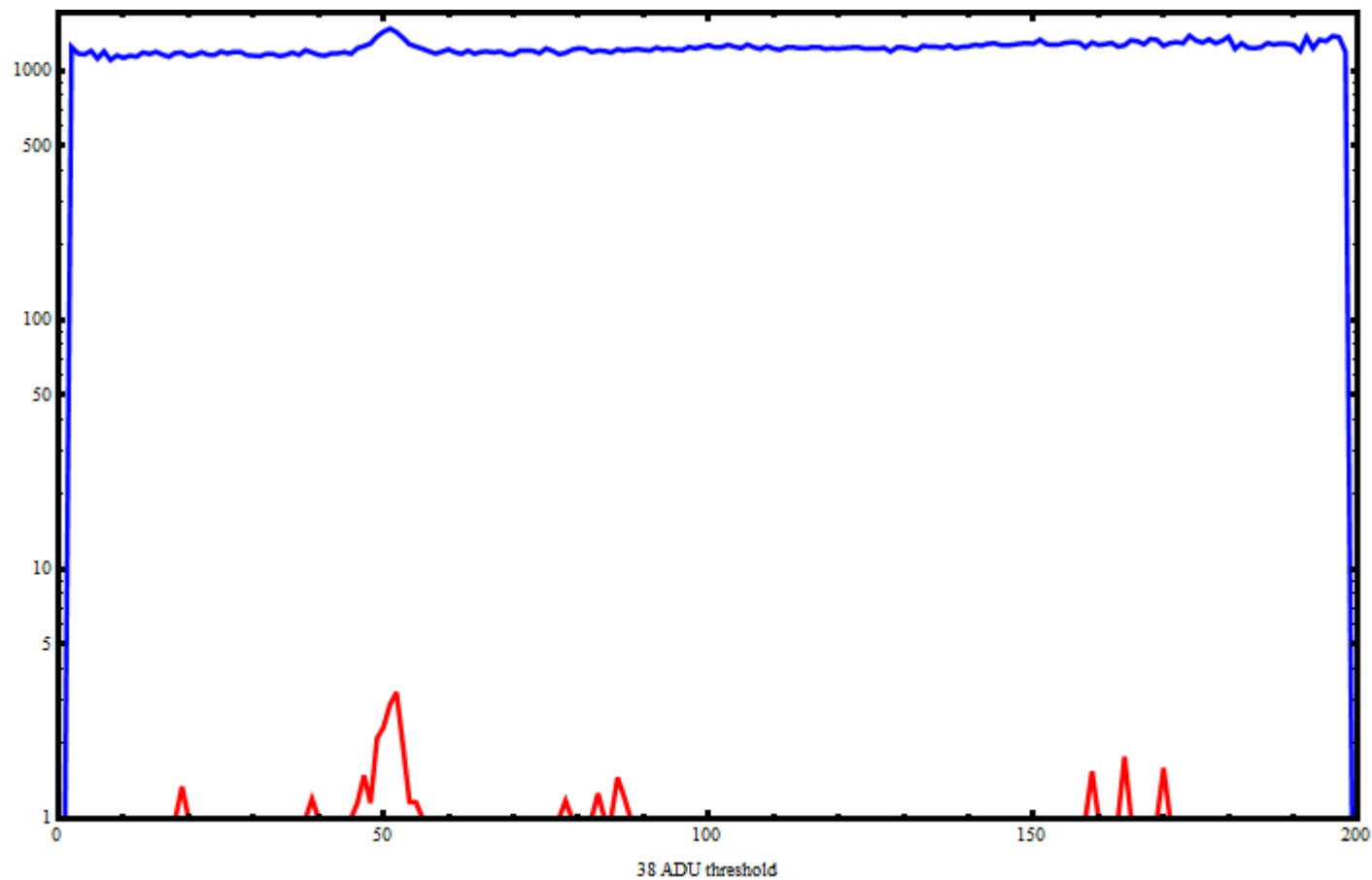
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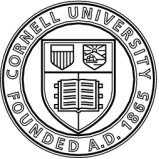
# *Varying Threshold*



(Azimuthal integration around detector)

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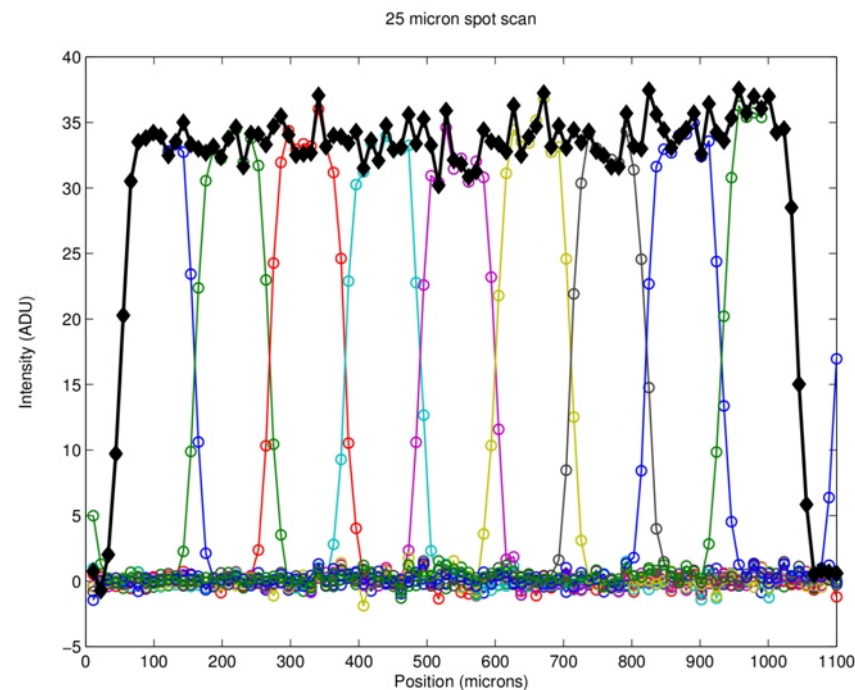
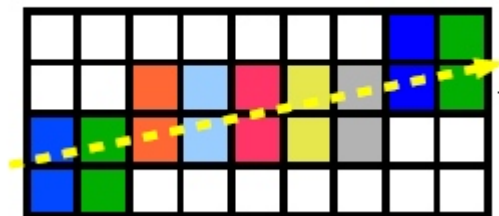
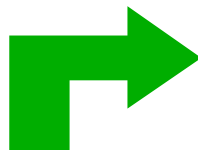
Hugh Philipp, iWoRID 13

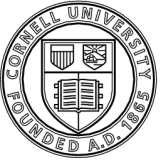


# Charge Sharing

One Advantage of integrating detectors:

- Summed charge proportional to absorbed energy. Don't worry about assigning photons to pixels.
- But thresholds cause problems.
- In fact, we see this.





# *Nearest Neighbors*

- ♦ Low flux data:

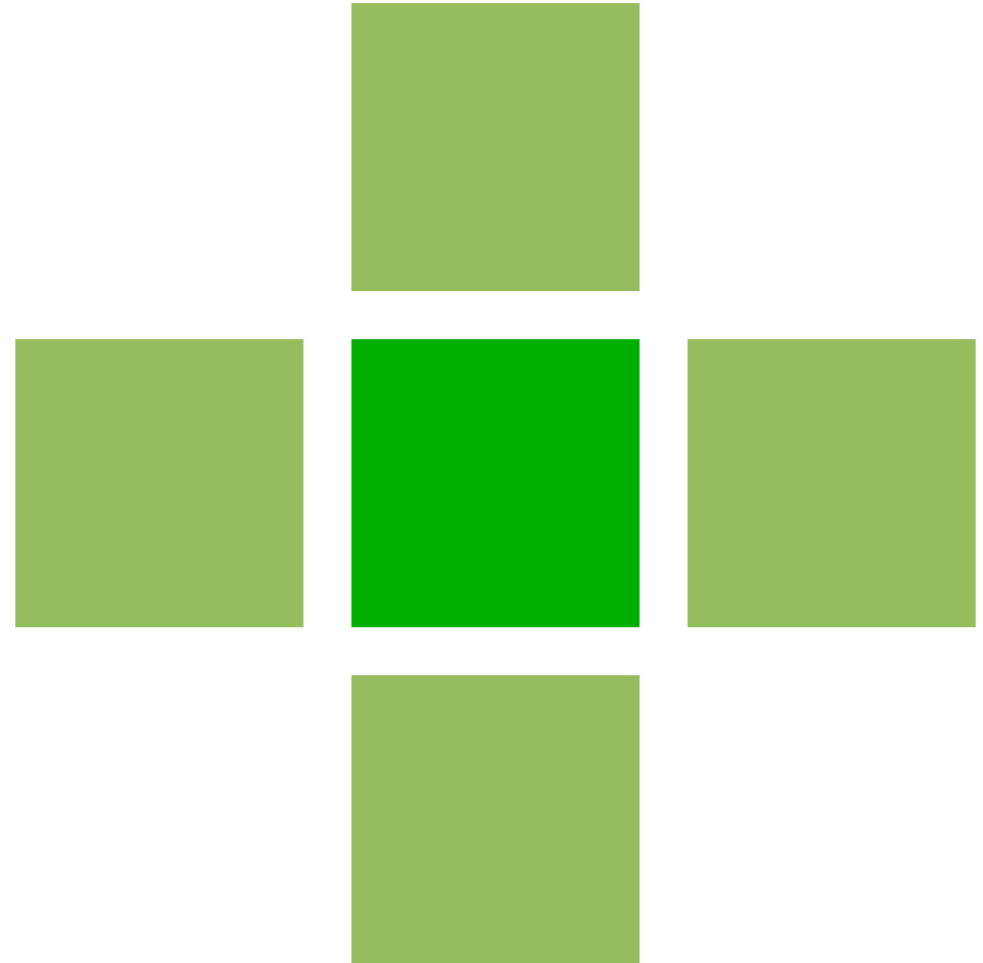
Threshold  $\sim 0.6$  photon

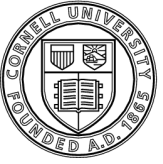


Signal reduction of 8% - just applied to central pixel.

But

Threshold allowing accumulation in nearest neighbors recovers 'most' Signal.





# *Nearest Neighbors*

- ♦ Low flux data:

Threshold  $\sim 0.6$  photon



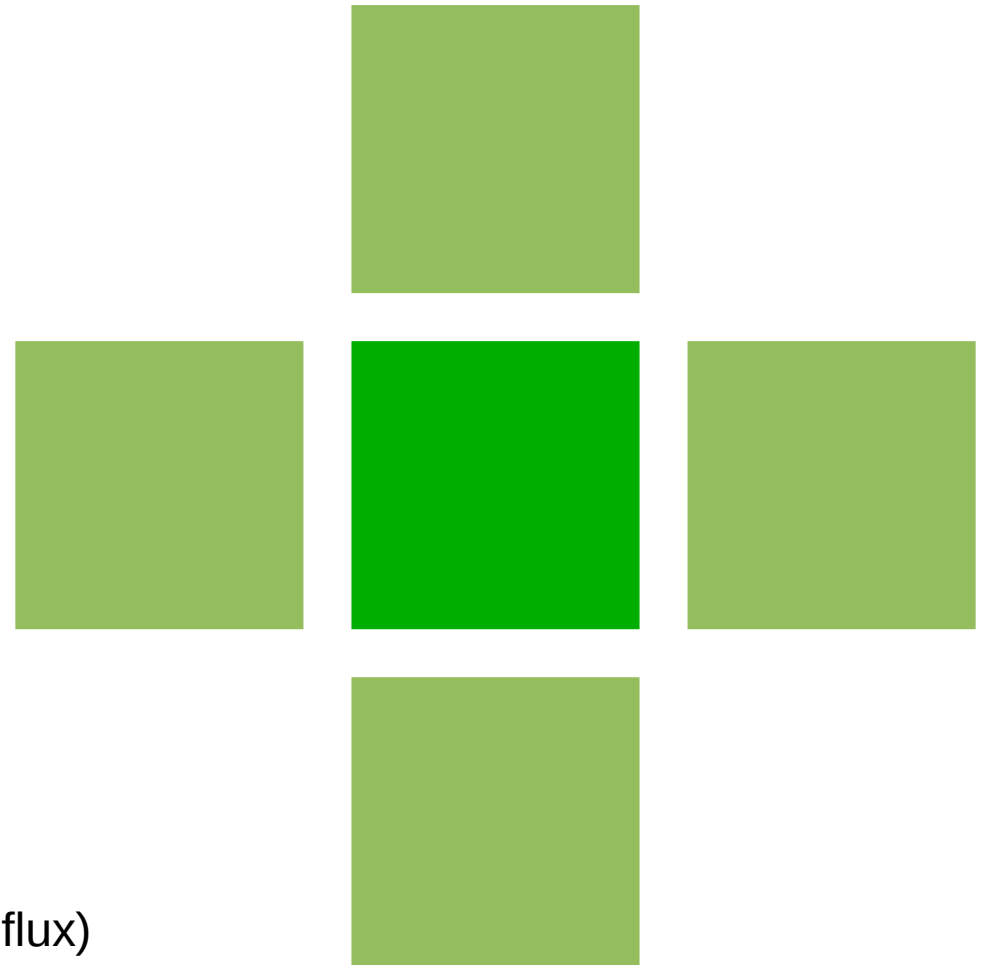
Signal reduction of 8% - just applied to central pixel.

But

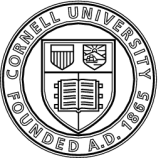
Threshold allowing accumulation in nearest neighbors recovers 'most' Signal.

Method has problems (e.g. High-flux next to low flux)

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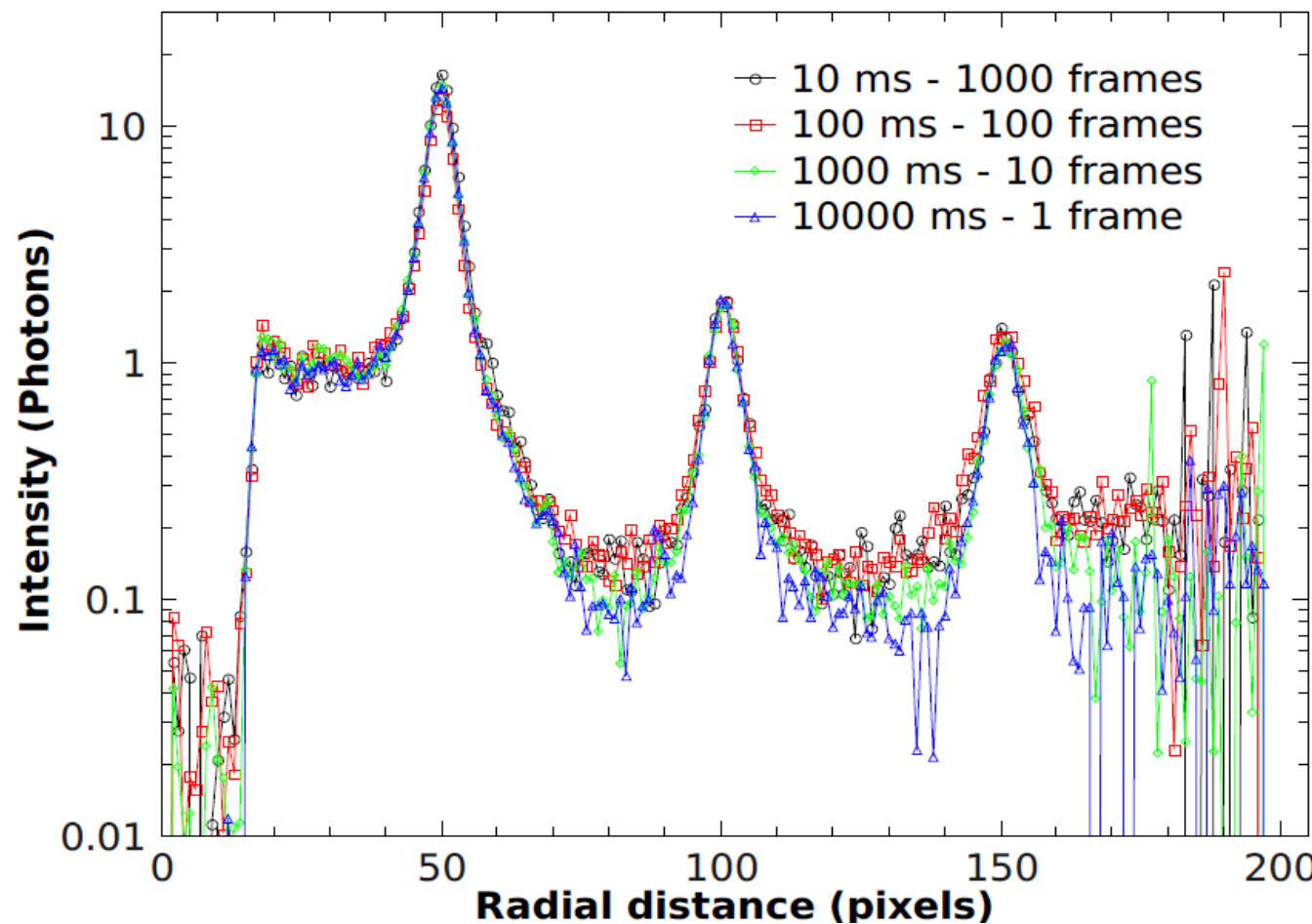






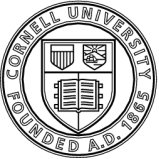
# *Is this the 'real' Diffraction Pattern?*

**10 s data acquisition  
azimuthal integration**



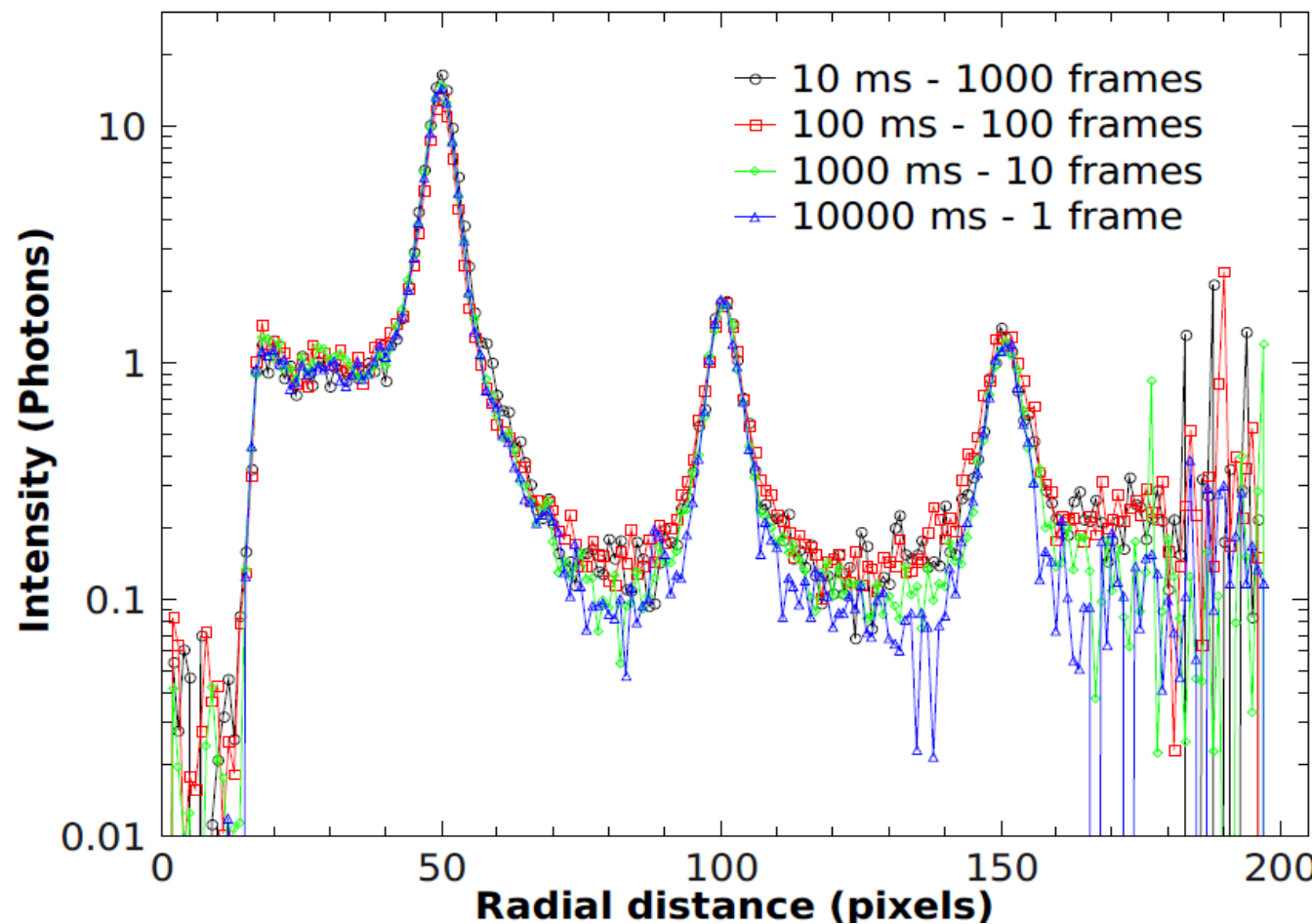
Threshold:  
12 ADU for lowest flux

Nearest Neighbors



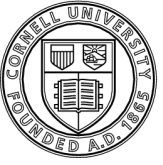
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**10 s data acquisition  
azimuthal integration**



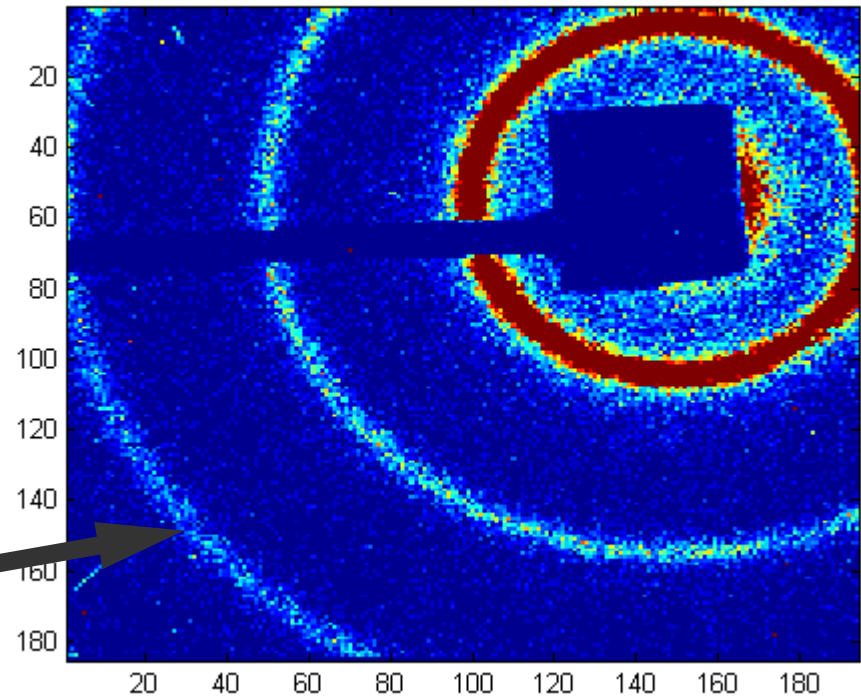
Threshold:  
12 ADU for lowest flux

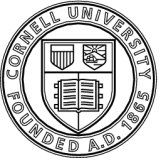
Nearest Neighbors



# *How Low?*

This ring less than  
 $10^{-3}$  photons/pixel/frame

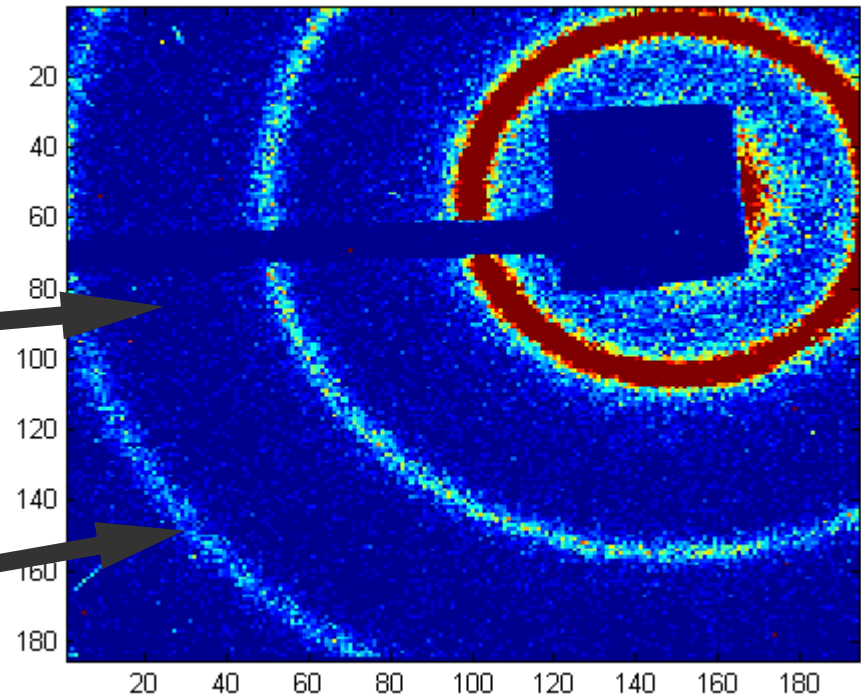


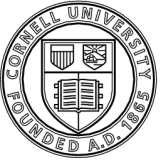


# *How Low?*

More than an order of magnitude less

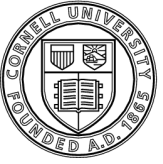
This ring less than  
 $10^{-3}$  photons/pixel/frame





## *Some Important Points*

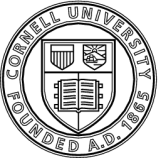
- ♦ How you reduce the detector data to recover meaningful data is not always straight forward.
- ♦ Giving users mountains (TBs) without direction about how to look at it = frustration, duplicated effort.
- ♦ Full and diligent calibration of detectors is needed before users spend time taking data.



# *Conclusions*

- ♦ For integrating detectors, thresholding (or something like it) is required for extremely low photon incidence.
- ♦ Cornell's pixel array detector (upon which the LCLS CXI detector is based) has demonstrated the capability of faithfully extracting diffractions patterns from extremely low flux (photons/pixel/frame).





### Launch of Fourth LCLS Instrument Reveals Crisp, Fine Molecular Detail

The first set of user experiments with the Linac Coherent Light Source's newest instrument is under way, and about 40 researchers are working very long hours this week to decipher the structures of proteins involved in photosynthesis, parasitic disease and other important life processes.

The results won't be known for months, after extensive analysis of the data. But near the end of the second 12-hour shift on Tuesday morning, scientists gathered in front of a bank of computer monitors in the CXI control room were beaming and pointing at the screens, to a chorus of "oohs" and "ahhs."

"You see this protein, all these rings here?" Petra Fromme of Arizona State University asked me, pointing to a printed image of what looked like tiny, bright stars arranged in circular patterns against a dark sky. "What's really nice is that each spot is so very fine. The diffraction pattern is so much cleaner" than the ones obtained with traditional structural analysis. What that implies, at a preliminary glance, is that the team may have captured the structure of a photosynthetic protein complex in very fine, crisp detail, approaching atomic resolution.

The LCLS is the world's most powerful X-ray free electron laser. Its strobe-light pulses—120 of them per second—could burn through steel,



The research team in the CXI control room.  
(Photo by Brad Plummer.)