







### **Entering the Precision Era:** Antihydrogen Symmetry Tests with ALPHA

Physics of Fundamental Symmetries and Interaction PSI 2016

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Big questions

- MCF: arXiv:1309.7468
- Brief overview
- First precision results with anti-H
  - Implications of recent charge neutrality tests
- Status and Prospects



### "Big Questions"

## What is Particle Physics? (e.g. Grossman)



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### "Simple Answer"

### The Standard Model! is (technically) unnatural ...





Should be new physics at TeV scale, e.g SUSY

**Cosmological constant** 

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### Are we asking right question?

### "L=?" really right question to ask?

# Is Quantum Field Theory correct description of Nature?



### **Motivations: Symmetries**



- CPT: Fundamental property of QFT
  - Theorem: atomic spectra of H & anti-H identical



- Einstein's Equivalence Principle
  - Matter and Antimatter fall in same way

Any violation would force radical change in theory!



### **ALPHA Experiment**



### **ALPHA** brief history

- ATHENA: produced first cold Hbars (2002) (They were not trapped)
   Completed data taking in 2004
- Developed into new experiments (2005)
  - Trapping and Spectroscopy of Hbars





## **L**PHA

#### **Antihydrogen Laser Physics Apparatus**

Also Microwaves, Gravity, Charge...

### **ALPHA Collaboration**

#### University of Aarhus: G. Andersen, P.D. Bowe, J.S. Hangst

RIKEN: Y. Yamazaki
Federal University of Rio de Janeiro: C.L. Cesar, D. Miranda
<u>University of Tokyo</u> : R. Funakoshi, R.S. Hayano
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L. V. Jørgensen, M. Jenkins, D.P. van der Werf
<u>Auburn University:</u> F. Robicheaux
<u>University of California, Berkeley</u> : W. Bertsche, S. Chapman, J. Fajans, A. Povilus, J. Wurtele
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Université de Montréal: JP. Martin*
Simon Fraser University: M. Dehghani, M. Hayden* Canadian students
TRIUMF: P. Amaudruz*, M. Barnes, M.C. Fujiwara*, D.R. Gill*, & Postdocs
L. Kurchaninov*, K. Olchanski*, A. Olin*, S. Stracka J. Storey + Professional Support**
York University, S. Amole, A. Capra, H. Malik, S. Menary*
* Active faculty/staff in present phase
**P. Bennett, D. Bishop, R. Bula, S. Chan, B. Evans, D. Healey, T. Howland, H. Hui, K. Langton, C. Marshall, J. Nelson,
D. Rowbotham, P. Vincent, T. Meyer, M. Pavan ++ Undergrads: W. Lai, L. Wasilenko, C. Kolbeck, N. Evetts
Creat toom offart!

Great team effort!



### **Some ALPHA-1 Achievements**



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### A side project

#### Anti-H Charge Neutrality Tests (Nature Comm. 2014, Nature 2016)

### **Experimental Limits on** $|\delta q/q|$



JHF-Pbar Workshop, Feb 16 2002



### **2014: Anti-H Charge Neutrality Test**



**Biasing E field** 

Key: position sensitive detection



# Improved Limit on $|\delta Q/Q|$ (Nature 2016)



 $|q_{e^+} + q_{e^-}|/e$ 

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A test of CPT invariance. See also similar tests involving th

 $4 \times 10^{-8}$ 7 HUGHES92 RVUE• We do not use the following data for averages, fits, limits, etc. $2 \times 10^{-18}$ 8 SCHAEFER95 THEO Vito $2 \times 10^{-18}$ 9 MUELLER92 THEO Vito7 HUGHES92 uses recent measurements of Rydberg-energy and compared by the second second

<sup>8</sup>SCHAEFER 95 removes model dependency of MUELLER 92.

<sup>9</sup> MUELLER 92 argues that an inequality of the charge magnitudes order vacuum polarization, contribute to the net charge of atoms

Anti-H neutrality tests: 2014 (ALPHA-1): Q<~10<sup>-8</sup> 2016 (ALPHA-2): Q<0.7x10<sup>-9</sup> New e+ charge limit ~10<sup>-9</sup> (40 fold improv't over PDG)



### What about e+ mass?



#### I have issues with PDG and Fee, Chu et al.!

- 1. PDG "assumption that the Ps Rydberg is exactly half of the hydrogen one" does not make sense
- 2. It seems FEE93 assumed incorrect sensitivity between  $\Delta freq(1s-2s)$  and  $\Delta m_{e^+}/m_e$
- 3. e+ mass & charge should be treated independently (as for Pbars)
- 4. Not clear if the limit is 90% CL rather than  $1\sigma$



#### Positron charge & mass before ALPHA (MCF at LEAP 2016)



#### **Before ALPHA**

- $-\Delta m_{e^+}/m_{e^+} \sim 10^{-7}$
- $-\Delta Q_{e^+}/Q_{e^+} \sim 3x10^{-8}$ (Pbar mass, charge anomaly negligible)

Cf: PDG 2014 –  $\Delta m_{e^+}/m_{e^+}$ : 8 x10<sup>-9</sup> (x 10 overestimate of precision!)

 $-\Delta Q_{e^+}/Q_{e^+}$ : 4 x 10<sup>-8</sup>

### **Positron Charge & Mass after ALPHA-1**



- After ALPHA-1 [3]
  - Both  $\Delta m_{e^+}/m_{e^+}$  and  $\Delta Q_{e^+}/Q_{e^+}$  improved marginally ~ x2

#### **RIUMF**

### **Positron Charge & Mass after ALPHA-2**



- After ALPHA-2 [4]
  - Ignore pbar charge & mass anomaly (4x10<sup>-10</sup>)
  - $\Delta Q_{e^+}/Q_{e^+} \sim 7x10^{-10} (1\sigma)$ , 40-fold improvement over pre-ALPHA
  - $\Delta m_{e+}/m_{e+} \sim \pm 2 \times 10^{-8}$ , ~5 fold improvement
  - But central value shifted due to disagreement between theory and exp in Ps(1s-2s)



### **Antiproton Mass & Charge**

• Analysis so far assumed:

 $\delta m_{pbar}/m_{pbar}, \, \delta Q_{pbar}/Q_{pbar} << \delta m_{e^+}/m_{e^+}, \, \delta Q_{e^+}/Q_{e^+}$ 

- Next generation Anti-H exp'ts can no longer assume this.
- In general, need 4 independent measurements to determine m<sub>pbar</sub>, Q<sub>pbar</sub>, m<sub>e+</sub>, Q<sub>e+</sub>. Possibilities:

Measurement	Leading order dependence	Current precision (1σ)	Near future prospects
Pbar/p cyclotron	Q <sub>pbar</sub> / m <sub>pbar</sub>	7×10 <sup>-11</sup>	Base: 10 <sup>-11</sup> ?
Pbar He	m <sub>pbar</sub> Q <sub>pbar</sub> <sup>2</sup>	4×10 <sup>-10</sup>	ASACUSA: 10 <sup>-10</sup> ?
e+/e- cyclotron	Q <sub>e+</sub> /m <sub>e+</sub>	1.3×10 <sup>-7</sup>	Harvard ?
Ps(1s-2s)	(m <sub>e+</sub> /2) Q <sub>e+</sub> <sup>2</sup>	5×10 <sup>-9</sup>	ETH: 5×10 <sup>-10</sup> ?
Anti-H (charge)	Q <sub>pbar</sub> + Q <sub>e+</sub>	7×10 <sup>-10</sup>	ALPHA: 10 <sup>-12</sup> ?
Anti-H (1s-2s)	$m_{e^+} Q_{pbar}^2 Q_{e^+}^2$	-	ALPHA: 10 <sup>-11</sup> ?

#### **Anti-H studies entering precision era!**



### **Spectroscopy with ALPHA-2**

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### **ALPHA-2: Precision Spectroscopy Machine**



Laser access; Improved microwaves, 5 mirror coils Getting ready for first laser spectroscopy in 2016

# الله مركمة المحافظة Anti-H long term goal: Precision spectroscopy





### **Anti-H Laser Cooling**

#### Laser cooling

- Provides cold, dense, spatially confined sample
- Needed for high precision
   & gravity experiments
- 122 nm (Lyman-alpha) laser challenging!



Realistic proposal

[Donnan, MCF, Robicheaux, J. Phys. B. 46, 205302 (2013)]

- Pulsed laser cooling
- Cooling on 1 dimension
- Use coupling of deg. of freedom for 3-D cooling
- Cooling from ~500 mK to
  ~20 mK in few 100 s
- Laser built at UBC

First laser exp't attempts Fall 2016



### ALPHA-g



### **Antimatter Gravity Measurement**

- Gravity
  - Never measured with antimatter
- Very difficult experiment since gravity is so weak
- Now plausible due to long confinement time

nature physics PUBLISHED ONLINE: 5 JUNE 2011 | DOI: 10.1038/NPHYS2025

#### Confinement of antihydrogen for 1,000 seconds

The ALPHA Collaboration\*

Atoms made of a particle and an antiparticle are unstable, usually surviving less than a microsecond. Antihydrogen, made entirely of antiparticles, is believed to be stable, and it is this longevity that holds the promise of precision studies of matter-antimatter symmetry. We have recently demonstrated trapping of antihydrogen atoms by releasing them after a confinement time of 172 ms. A critical question for future studies is: how long can anti-atoms be trapped? Here, we report the observation of anti-atom confinement for 1,000 s, extending our earlier results by nearly four orders of magnitude. Our calculations indicate that most of the trapped anti-atoms reach the ground state. Further, we report the first measurement of the energy distribution of trapped antihydrogen, which, coupled with detailed comparisons with simulations, provides a key tool for the systematic investigation of trapping dynamics. These advances open up a range of experimental possibilities, including precision studies of charge-parity-time reversal symmetry and cooling to temperatures where gravitational effects could become apparent.

### **Antimatter Gravity Experiment**

### Does antimatter fall down?

- Many indirect constraints incl. EP tests
- Experimental question!
  - (e.g. Lykken et al, arXiv:0808.3929)
- Anti-H "gas" will sag due to gravity
- Need anti-H cooling to ~mK

1/2kT = mgh

Vertical trap:  $h \sim 1 m$ 

- Position sensitive detection via annihilations
- Laser cooling essential step: development at UBC
   NB: Cold atom tests of gravity: ~10<sup>-10</sup>





External Magnet

**Frigger/veto** 

TPC

e+

& cryostat

Meas.

Region

~1.5 m

Anti-H prod

pbar

& cooling

~0.5 m

External Magnet

MCP/

Probes RF

resonator

Η

Laser, uW

- A long (~ 2m) vertical trap
  - Anti-H production region Production, trapping, & cooling
  - Measurement region
    - Sagging of anti-H "gas"
    - Anti-atomic "fountain"
    - Anti-atomic interferometry
    - uW spectroscopy
- Major Canadian funding (thank you, referees!)

### **ALPHA-g in AD zone**



### Aiming for commissioning in 2017!



### **Radial TPC Construction at TRIUMF**

1/8 Prototype

#### **GEANT** simulation



2.3 m long, Radial thickness: 8 cm Unusual magnetic fields

→ Radial-drift TPC

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### Radial TPC Prototype, Cosmic Test



- Prototype: 1/8 length, full radial size
- Cosmic test: anode RO partially instrumente
- Cathode RO firmware under development



#### Reconstructed Cosmic Rays (no B field)



### **Future?**



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Hamilton et al, Phys. Rev. Lett. (2014)





Summary

- We hope to address some of most fundamental questions in physics with Anti-H
- After many years of efforts, anti-H studies finally entering precision era!
  - Improving the knowledge of positron charge and mass
  - First laser, and improved microwave spectroscopy
- Many exciting opportunities to come!
  - Anti-H gravity
  - Fountain, Interferometer, etc.



Canada's National Laboratory for Particle and Nuclear Physics Laboratoire national canadien pour la recherche en physique nucléaire et en physique des particules

# Thank you! Merci!

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