

# New Results from NOvA

Giulia Brunetti on behalf of the NOvA Collaboration





#### Neutrinos

Neutrino mix: flavors eigenstates are linear combinations of mass eigenstate

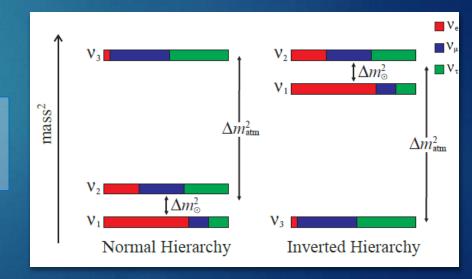
$$|v_{\alpha}\rangle = \sum_{k=1}^{n} U_{\alpha k} |v_{k}\rangle \quad (\alpha = e, \mu, \tau)$$

- Non-zero probability of detecting a different neutrino flavor than that produced at the source
  - depends on: squared mass difference, mixing angles, CP-violating phase, hierarchy....
- Mixing matrix for the three-flavor case:

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{-i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$\theta_{23} \sim 45^{\circ} \qquad \theta_{13} = 8.5^{\circ} \qquad \theta_{12} = 33.5^{\circ}$$

$$\Delta m_{23}^{2} \sim \pm 2.5 \times 10^{-3} eV^{2} \qquad \delta_{CP}? \qquad \Delta m_{21}^{2} = +7.5 \times 10^{-5} eV^{2}$$



#### Neutrinos

- Open questions:
  - Maximal mixing in the atmospheric sector?  $(\theta_{23})$
  - ightharpoonup CP-violation? ( $\delta_{CP}$ , P( $v_{\mu}$ ) vs P( $\overline{v}_{\mu}$ ) , matter/antimatter asymmetry in the universe)
  - ► Hierarchy?  $(sign(\Delta m_{23}^2))$ , matter effects)
  - Majorana or Dirac? (IH & no 0νββ decays)
  - Absolute masses?

NOvA (NuMI Off-Axis v<sub>e</sub> Appearance) Experiment

200+ collaborators41 institutions7 countries

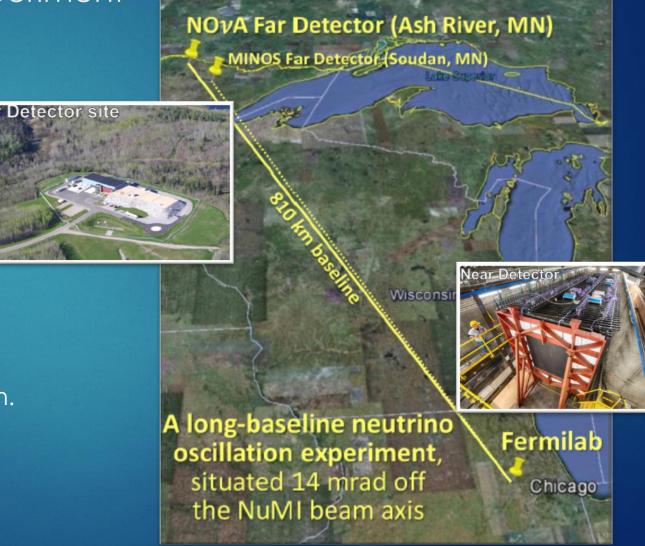
b designed to answer the next generation of  $\nu$  questions: tuned for  $v_e$  appearance in an almost pure  $v_u$  beam

#### NOVA

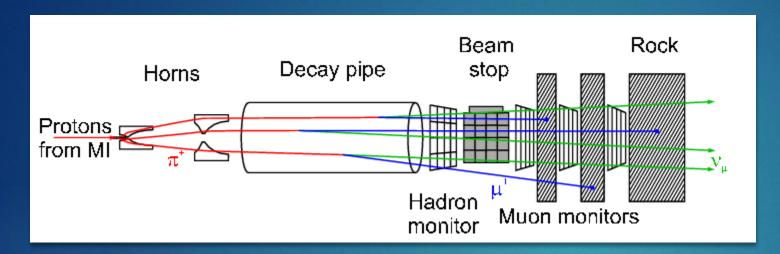
NuMI Off-Axis  $v_e$  Appearance Experiment

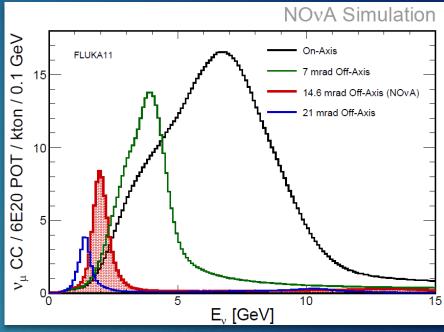
- NOvA is a long baseline (810 km), off-axis (14.6 mrad) neutrino oscillation experiment
- NuMl beam at Fermilab
- Energy peak @ 2 GeV
- 2 functionally identical detectors:
  - ▶ ND underground at Fermilab. 290-ton.
  - Used to predict event rate at the FD
  - ▶ FD on surface in Ash River, MN. 14-kton.

To look for oscillations



#### The NuMI beam

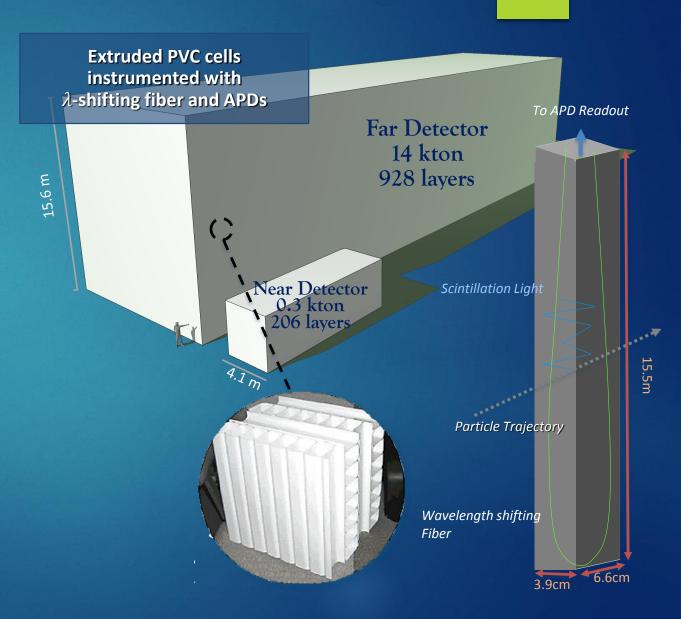




- 120 GeV protons onto a graphite target
- Secondary mesons charge-selected and focused by two magnets
- Pions decay into neutrinos/antineutrinos
- ▶ 6.05 10<sup>20</sup> POT in 14 kton equivalent detector
- Currently running at 560 kW, achieved 700 kW design goal in tests on June 13

#### NOvA Detectors

- Functionally identical, PVC cells filled with 10.2M Liters liquid scintillator
- Low-Z, 65% active volume, DAQ runs without deadtime (beam trigger, cosmic calibration samples, SNEWS, exotics)
- Read-out using WLS to APDs
- Cells organized in horizontal and vertical planes
- ▶ FD is 14 kton, ND is 0.3 kton



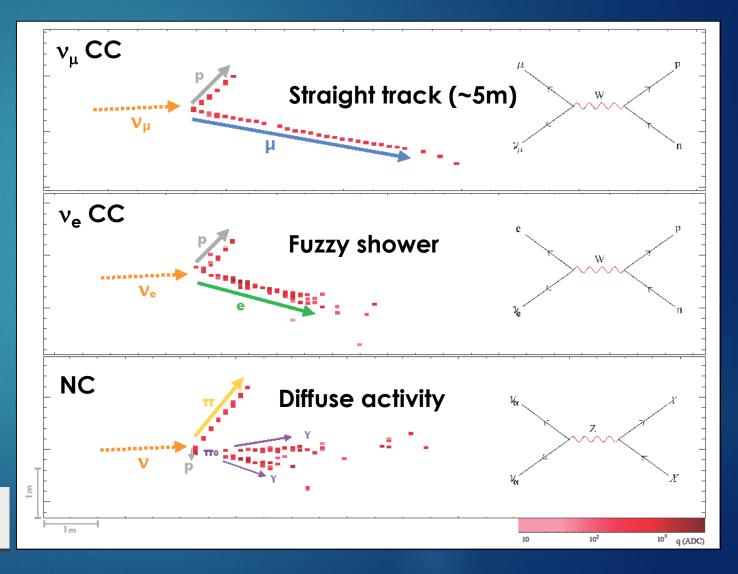
### NOVA Physics

#### ▶ 3-flavor oscillation analyses

- **▶** DISAPPEARANCE:  $\nu_{\mu}$  ( $\overline{\nu}_{\mu}$ )  $\rightarrow \nu_{\mu}$  ( $\overline{\nu}_{\mu}$ )
  - $ightharpoonup \Delta m_{23}^2$ ,  $\sin^2 2\theta_{23}$
- ► APPEARANCE:  $v_{\mu} (\overline{v}_{\mu}) \rightarrow v_{e} (\overline{v}_{e})$ 
  - $\triangleright$   $\theta_{13}$ ,  $\theta_{23}$ ,  $\delta_{CP}$ , mass hierarchy
  - ▶ Matter effects over 810 km  $\rightarrow \pm 30\%$

- Good granularity
- X0 =38cm (6 cells depths, 10 cells widths)

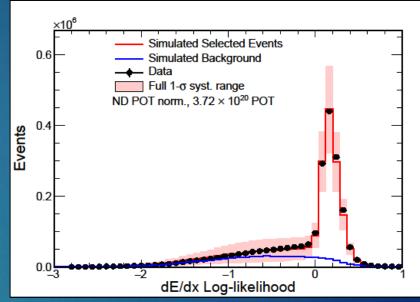


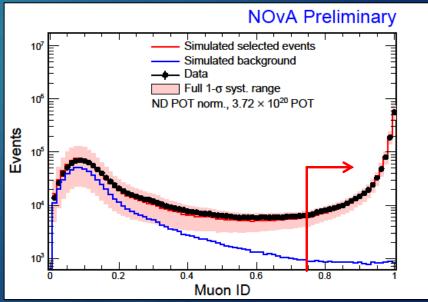


- The principle:
  - ▶ Select  $v_{\mu}$  CC sample: events with long tracks and distinctive dE/dx
- Extrapolation of the ND spectrum to the FD and measurement of the deficit
- ▶ 2-flavor oscillation approximation works well in this case:

$$P_{\mu\mu}\sim 1-sin^2 2\theta_{23}sin^2\left(\frac{\Delta m_{23}^2L}{4E}\right)$$
  
 $\theta_{23}\sim 45^\circ \rightarrow$  at the oscillation max almost all  $v_\mu$  disappear

- NC and cosmic background suppression, containment cuts to remove events with activity close to the detector walls
- $\nu_{\mu}$  ID: Multivariate kNN classifier using 4 variables:
  - ▶Track length
  - ▶dE/dx
  - Scattering along the track
  - ▶Track only fraction of planes
- ▶ 81% selection efficiency for signal with 95% purity

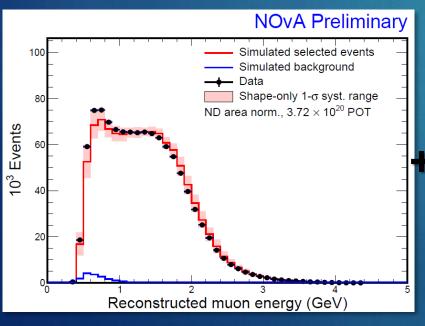


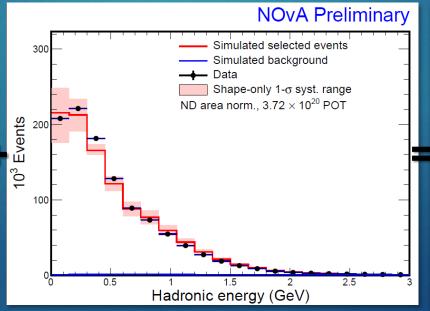


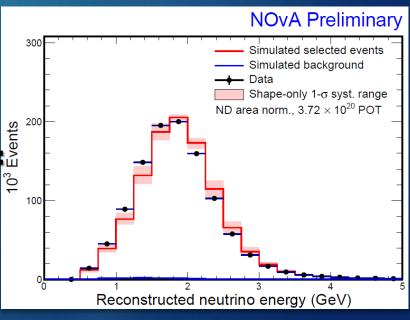
#### $\triangleright$ $\nu_{\mu}$ ND events

- Hadronic energy scale uncertainty from 14% to 5% with the addition of MEC events to the simulation (w.r.t. NOvA 2015 results)
- ▶ ND reconstructed energy spectrum unfolded and extrapolated to FD using Far/Near true ratio for prediction

 $Ev = E\mu$  (L track) +  $E_{had}$  (7% res)





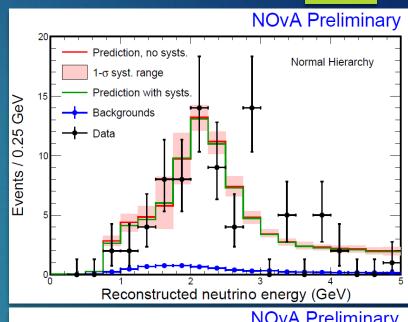


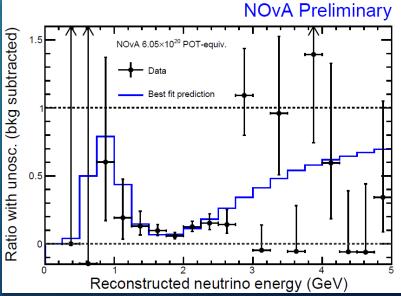
# ν<sub>μ</sub> Disappearance

- $\mathbf{v}_{\mu}$  FD events: **78** events observed
  - ▶ No oscillation prediction: 473±30
  - ▶ Best oscillation fit: 82 events
  - ▶ Beam BG: 3.7, Cosmics: 2.9

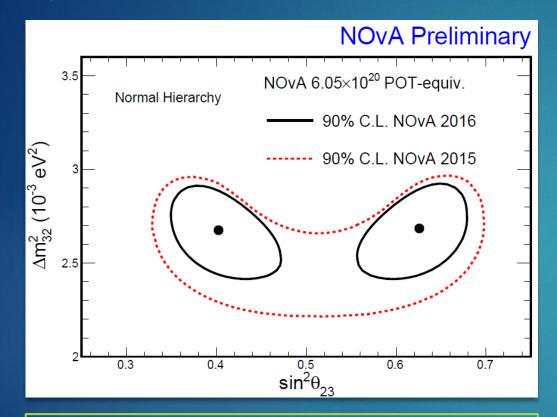
χ2/NDF=41.6/17

driven by fluctuations in the tail, no pull in oscillation fit





# ν<sub>μ</sub> Disappearance

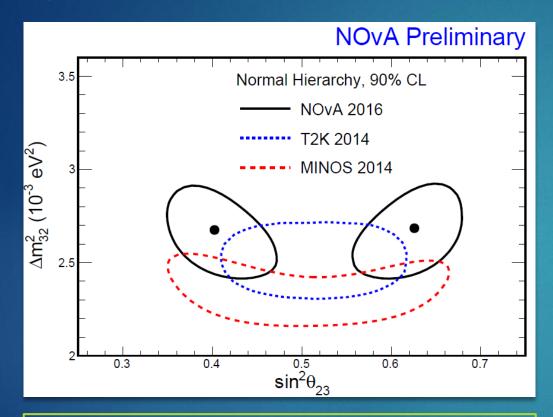


Our best fit (in NH):  $|\Delta m^2_{32}| = 2.67 \pm 0.12 \cdot 10^{-3} \text{ eV}^2$   $\sin^2 \theta_{23} = 0.40^{+0.03}_{-0.02} (0.63^{+0.02}_{-0.03})$ 

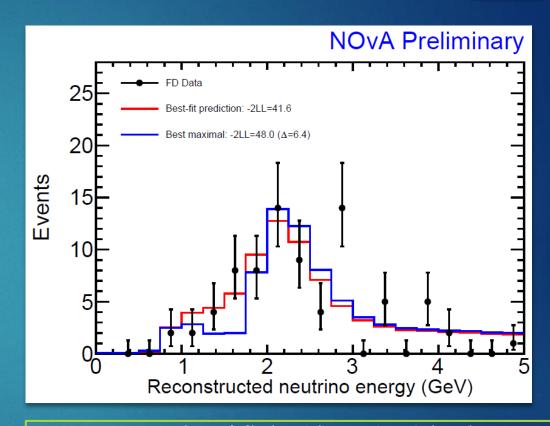
- Fit for  $\Delta m^2$  and  $\sin^2\theta_{23}$
- Dominant systematic effects included in fit:
  - Normalization
  - NC background
  - ► Flux
  - Muon and hadronic energy scale
  - Cross section
  - Detector response and noise

Maximal mixing ( $\theta_{23}$  =45°) excluded at 2.5 $\sigma$ 

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- Non-maximal fit is driven by bins in oscillation dip (1-2 GeV)
- Forcing maximal mixing gives:

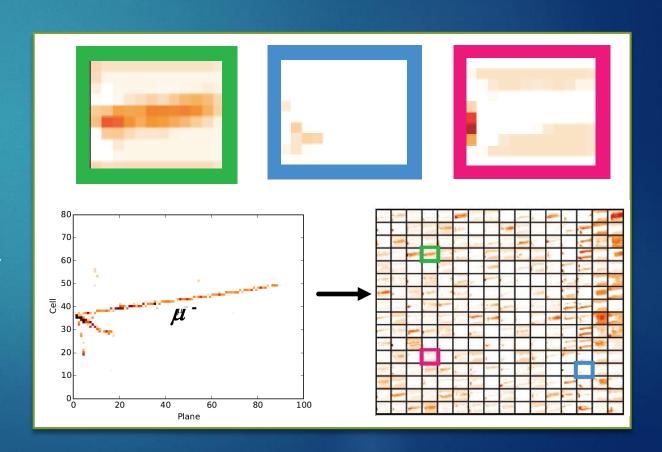
$$|\Delta m^2_{32}| = 2.46 \cdot 10^{-3} \, eV^2$$

### Improved event selection

CVN – Convolutional Visual Network: new event selection technique based on ideas from computer vision and deep learning

- Calibrated hit maps are inputs to the CVN
- Series of image processing transformations applied to extract abstract features
- Extracted features used as inputs to a conventional neural network to classify the event

Improved sensitivity equivalent to 30% more exposure

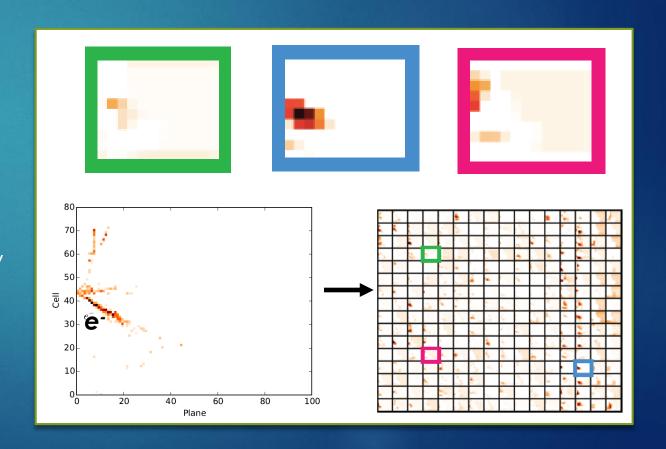


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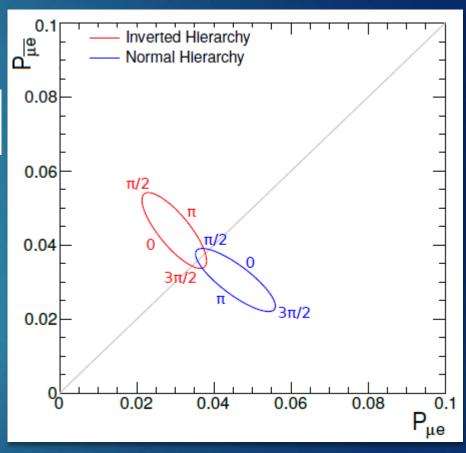
### v<sub>e</sub> Appearance

$$P(\nu_{\mu} \to \nu_{e}) \approx \sin^{2} 2\theta_{13} \sin^{2} \theta_{23} \frac{\sin^{2}(\Delta_{31} - aL)}{(\Delta_{31} - aL)^{2}} \Delta_{31}^{2}$$

$$\alpha \sin 2\theta_{13} \cos \delta \frac{\sin(aL)}{(aL)} \frac{\sin(\Delta_{31} - aL)}{(\Delta_{31} - aL)} \cos \Delta_{32} - \alpha \sin 2\theta_{13} \frac{\sin(\Delta_{31} - aL)}{(aL)} \sin(\Delta_{31} - aL)}{(\Delta_{31} - aL)} \sin \Delta_{32}$$

$$\Delta_{ij} \equiv \frac{1.27\Delta m_{ij}^2 [\text{eV}^2] L[\text{km}]}{E[\text{GeV}]}$$
$$a = G_F N_e \sqrt{2} \simeq (4000 \text{ km})^{-1}$$

- Depends simultaneously on θ<sub>13</sub>, θ<sub>23</sub>, δ<sub>CP</sub>, sign(Δm<sup>2</sup><sub>31</sub>)
- $\blacktriangleright$   $sin^2 2\theta_{13}$ =0.095  $\rightarrow$  most  $v_{\mu}$  go to  $v_{\tau}$
- Look for deviations due to hierarchy (matter effects) and CPviolation
- NOvA measures  $P(v_{\mu} \rightarrow v_{e})$  and  $P(\overline{v}_{\mu} \rightarrow \overline{v}_{e})$  at  $\overline{2}$ GeV, different dependence on  $sign(\Delta m_{32}^{2})$  and  $\delta_{CP}$



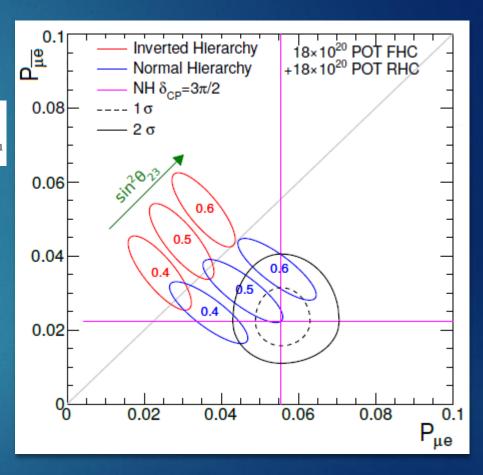
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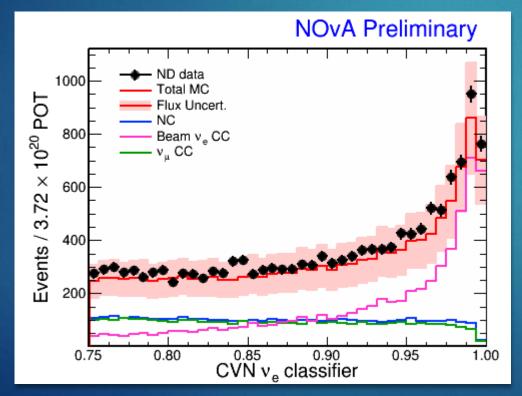
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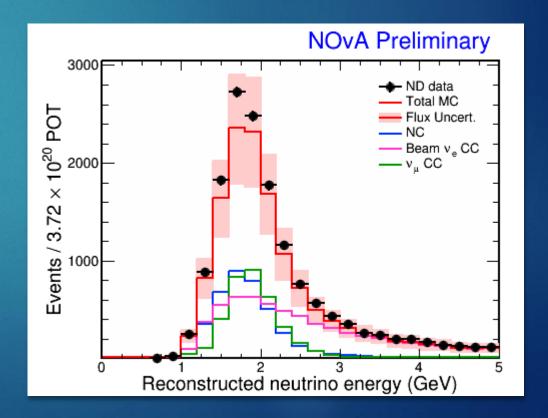
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- $P \propto sin^2 \theta_{23}$
- Constrain a space region



- CVN PID, loosen cut on Pid optimized to favor parameter measurement
- ▶ Separate  $v_e$  CC interactions from backgrounds, backgrounds evaluated in ND:
  - intrinsic beam  $v_{e}$ , Neutral Currents,  $v_{u}$  CC, each propagate differently
  - Use ND data to predict background in the FD
- Looking for an excess in the FD





Expected events depend on oscillation parameters:

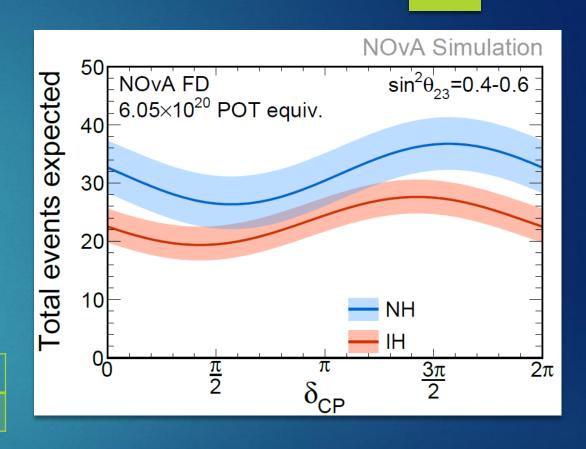
$$\sin^2\theta_{23} = 0.5, \pm 5\%$$
 syst.

**Total Prediction** (signal+background):

NH, 3π/2	IH, π/2
36.4	19.4

Background components (±10% syst):

Total BG	NC	Beam $v_{\rm e}$	$v_{\mu}CC$	$v_{\tau}CC$	Cosmics
8.2	3.7	3.1	0.7	0.1	0.5

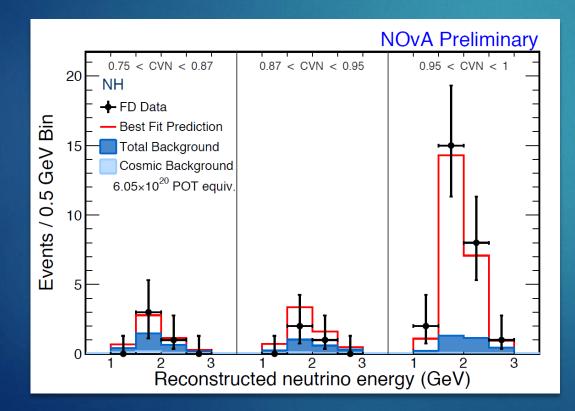


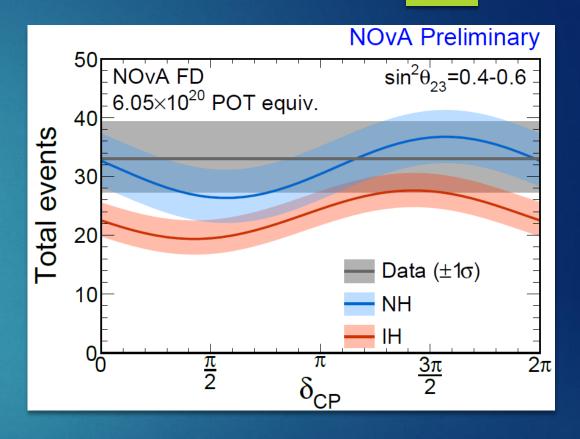
Each component extrapolated in bins of energy and CVN output

Total Prediction (signal+background):

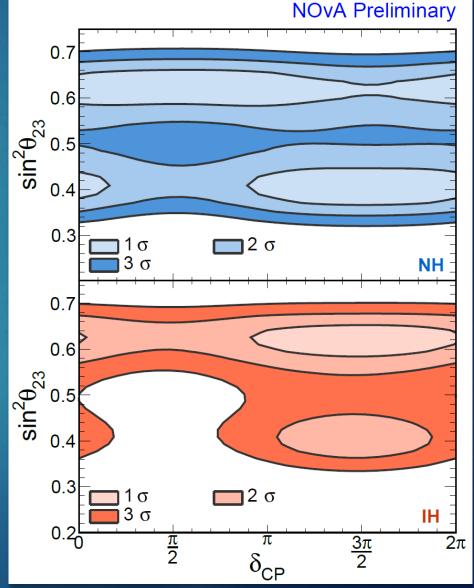
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Observed events in FD: 33

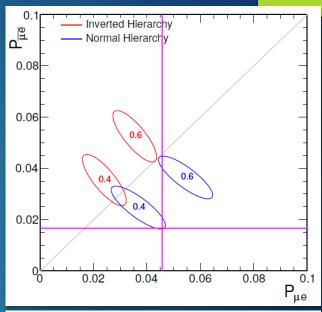


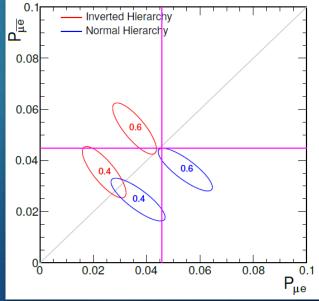


- Fit for hierarchy,  $\delta_{\rm CP}$ ,  $\sin^2\theta_{23}$ 
  - ► Constrain  $\sin^2(2\theta_{13})=0.085\pm0.05$  from reactor
  - ► Constrain  $\Delta$ m and  $\sin\theta_{23}$  with NOvA disappearance results
  - Not a full joint fit, syst and other oscillation parameters not correlated
- ► Global best fit, preference for NH,  $\Delta \chi^2 = 0.47$ 
  - $\delta_{CP} = 1.49\pi$ ,  $\sin^2(\theta_{23}) = 0.40$
  - ▶ Both octants and hierarchies allowed at 1σ
  - ► IH lower octant around  $\delta_{CP} = \pi/2$  excluded at  $3\sigma$



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  - $\blacktriangleright$  Both octants and hierarchies allowed at  $1\sigma$
  - ► IH lower octant around  $\delta_{CP} = \pi/2$  excluded at  $3\sigma$
- ► Antineutrino Run (planned for spring 2017) will help resolve degeneracies





### Summary

- ► Early days for NOvA, our baseline program is six times our current exposure NOvA collected 6.05·10<sup>20</sup> POT, oscillation results:
  - $\mathbf{v}_{u}$  disappear, maximal mixing is excluded at  $2.5\sigma$
  - $\triangleright$   $v_e$  appear:
    - slight preference for NH
    - ▶ IH lower octant around  $δ_{CP} = π/2$  is exlcuded (>3 σ)
- Antineutrino run in spring 2017

- Many other interesting NOvA analyses!
  - sterile neutrinos, cross section measurements, supernovæ...

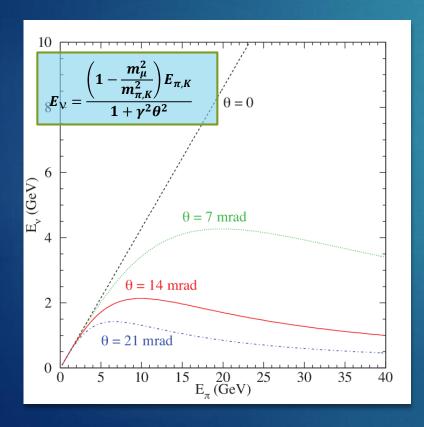
Argonne, Atlantico, Banaras Hindu, Caltech, CUSAT, Czech Academy of Sciences, Charles, Cincinnati, Colorado State, Czech Technical University, Delhi, Dubna, Fermilab, Goias, IIT-Guwahati, Harvard, IIT-Hyderabad, Hyderabad, Indiana, Iowa State, Jammu, Lebedev, Michigan State, Minnesota-Twin Cities, Minnesota-Duluth, INR Moscow, Panjab, SDMT, South Carolina, SMU, Stanford, Sussex, Tennessee, Texas-Austin, Tufts, UCL, Virginia, Wichita State, William and Mary, Winona State.

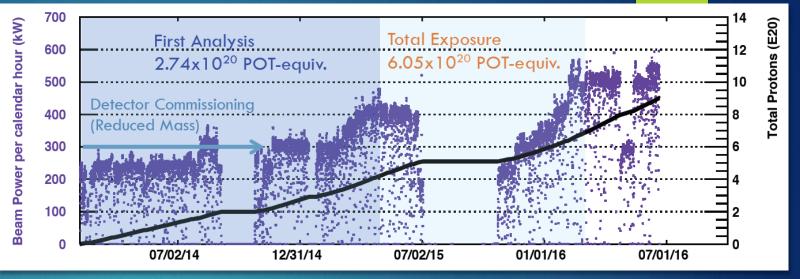


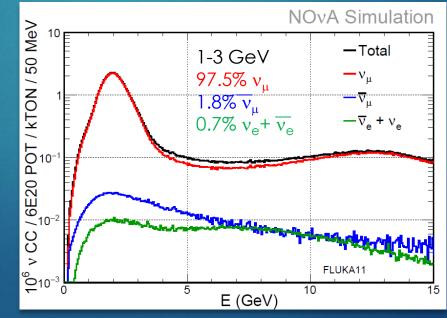
# Back up

#### NuMI beam

- Beam performance
- ▶ 14mrad Off-Axis:
  - ► Neutrino energy spectrum peaked at 2GeV, width~20%





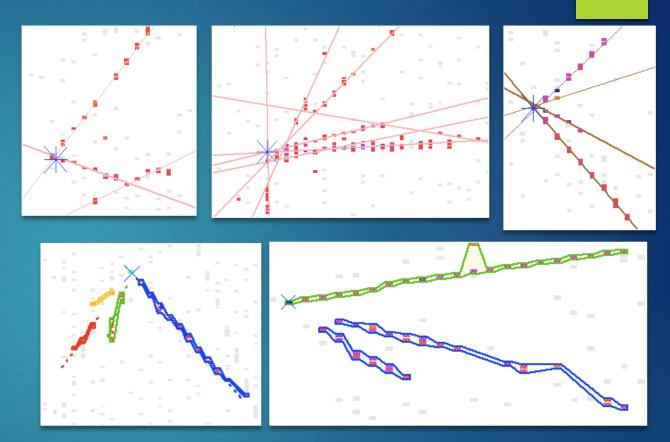


#### Reconstruction

Vertexing: Find lines of energy depositions w/ Hough transform CC events: 11 cm resolution

Clustering: Find clusters in angular space around vertex.

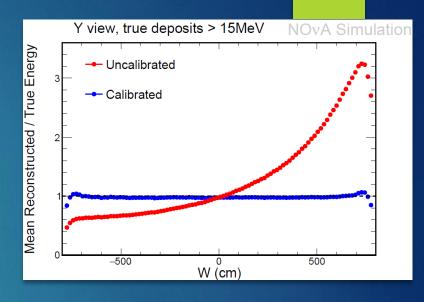
Merge views via topology and prong dE/dx

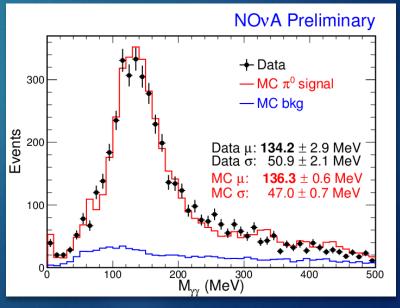


<u>Tracking:</u> Trace particle trajectories with **Kalman filter** tracker.

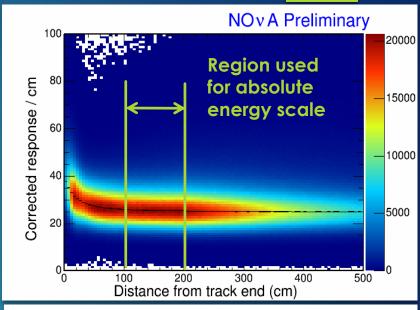
Also, **cosmic ray tracker**: lightweight, fast, and for large calibration samples, online monitoring.

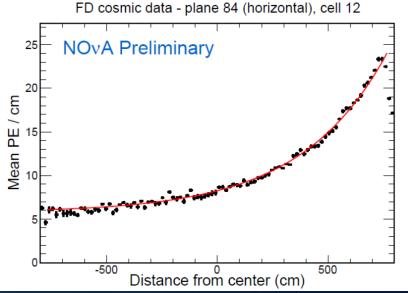
- Calibration and energy scale: Cosmic ray muons are the standard candle
- Cells individually corrected for
  - ► Llight attenuation along cell length
  - Shadowing due to detector bulk
  - Threshold effects far from readout
- Energy scale set by dE/dx near the end of stopping muons
  - Cross-check including π0 mass peak, Michel-e<sup>-</sup>, beam muon dE/dx
  - ▶ Take 5% absolute and relative errors





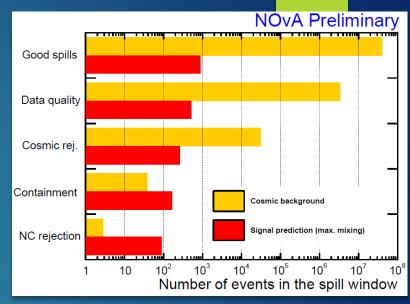
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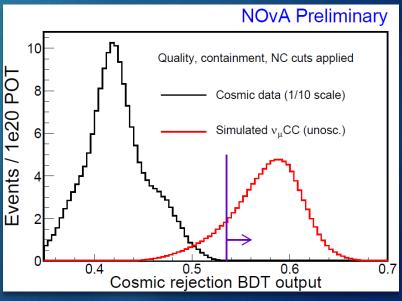




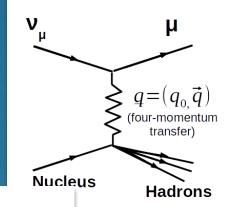
#### Cosmic rejection

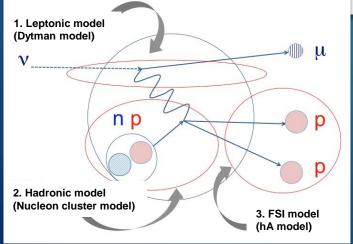
- 10µs spill window gives 10⁵ rejection
- Cosmic ray data in data are measured in time window adjacent to the spill
- Event topology+BDT provide additional O(10<sup>7</sup>) reduction
  - BDT inputs: track direction, track start and end point, track length, energy, number of hits

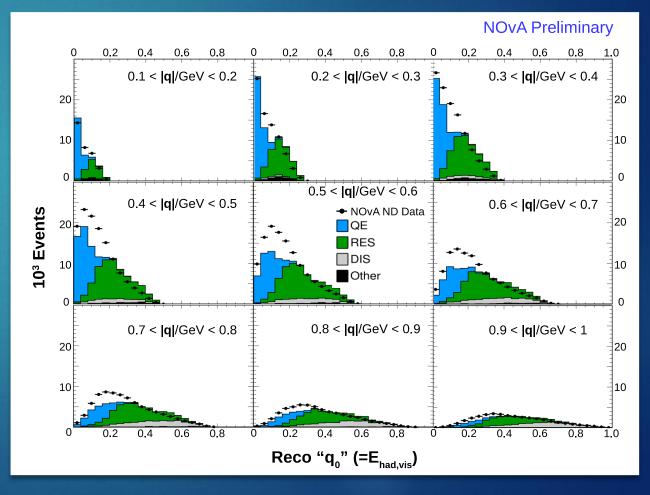




ND data suggest unsimulated process between QE and  $\Delta$  production (Minerva experiment reported similar excess)







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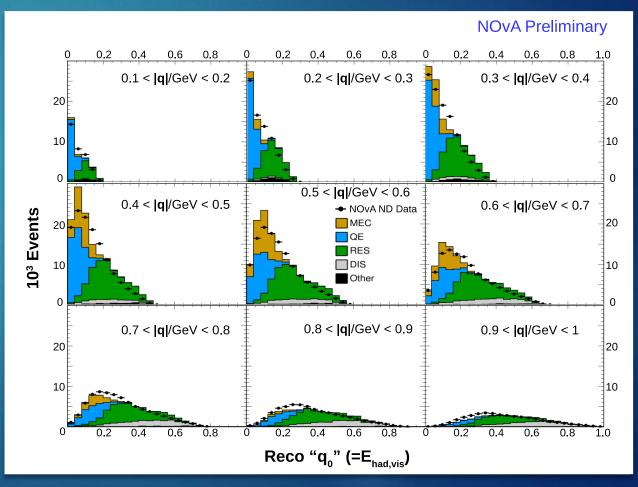
→enable GENIE empirical MEC (50% systematic on MEC component)

 $\rightarrow$ reweight the model to match observation as a function of  $\vec{p}$  transfer

Reduction of largest systematics

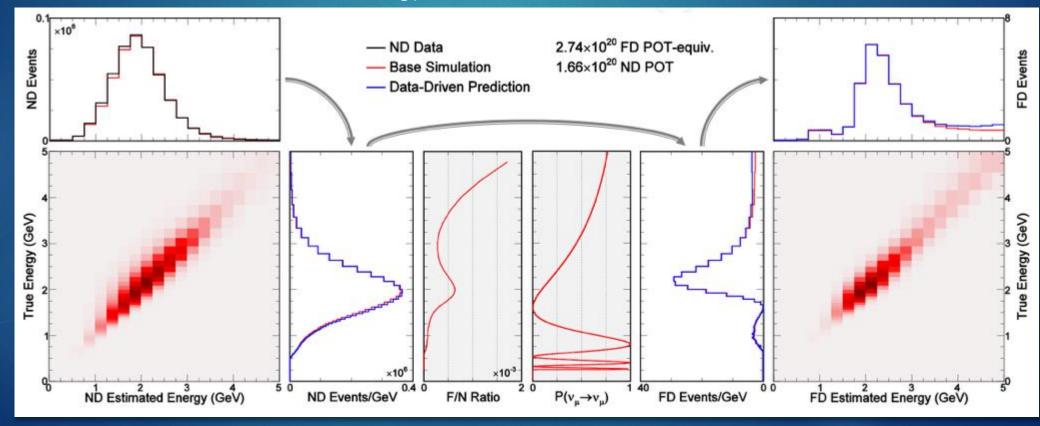
- -Hadronic energy scale
- -QE cross section modeling

Reduction of single non-RES pion production by 50%



Near-Far Extrapolation – 3 step process

- 1) Convert ND reconstructed energy to true energy
- 2) Use Near/Far ratio to convert to FD true energy spectrum
- 3) Translate back to reconstructed energy



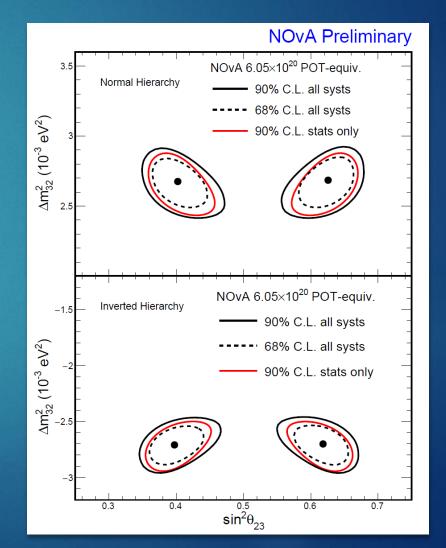
#### Systematic uncertainties

Systematic	Effect on sin²(θ <sub>23</sub> )	Effect on Δm <sup>2</sup> 32
Normalisation	± 1.0%	± 0.2 %
Muon E scale	± 2.2%	± 0.8 %
Calibration	± 2.0 %	± 0.2 %
Relative E scale	± 2.0 %	± 0.9 %
Cross sections + FSI	± 0.6 %	± 0.5 %
Osc. parameters	± 0.7 %	± 1.5 %
Beam backgrounds	± 0.9 %	± 0.5 %
Scintillation model	± 0.7 %	± 0.1 %
All systematics	± 3.4 %	± 2.4 %
Stat. Uncertainty	± 4.1 %	± 3.5 %

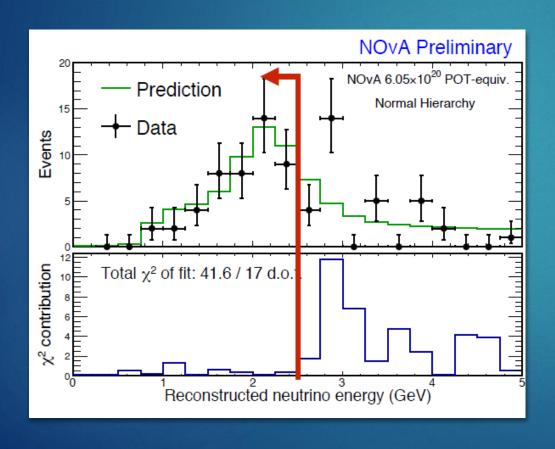
#### In each case:

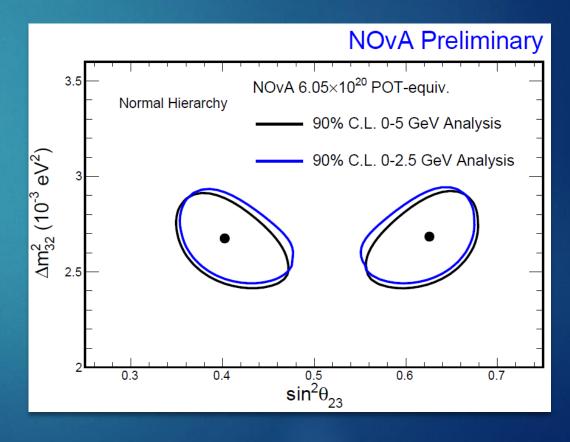
- The effect is propagated through the extrapolation
- We include those effects as pull terms in the fit
- The increase (in quadrature) of the parameter measurement error is recorded

#### Inverted hierarchy contours

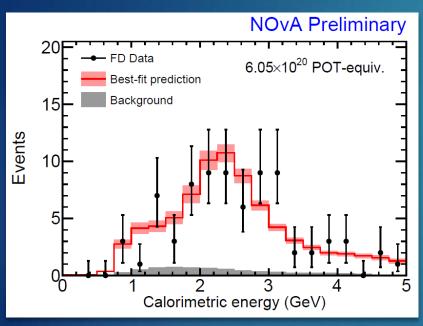


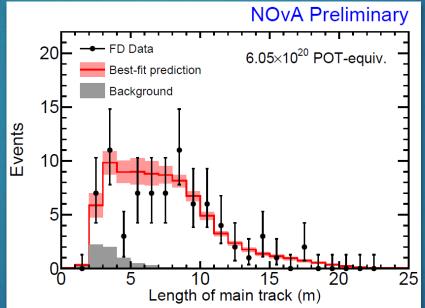
- Best Fit  $\chi^2/DOF = 41.5/17$  is driven by the tail
- ▶ There is no significant pull in the oscillation fit from bins in the tail

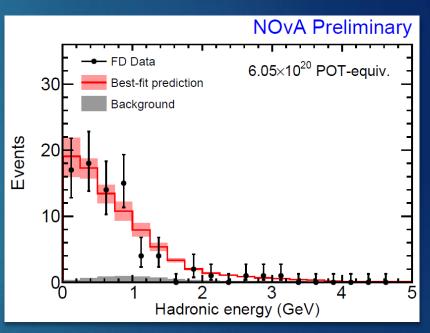




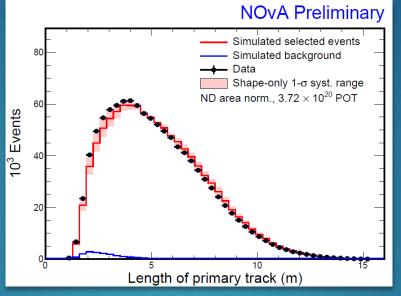
Fit-checks: best fit oscillation prediction matches other distributions well

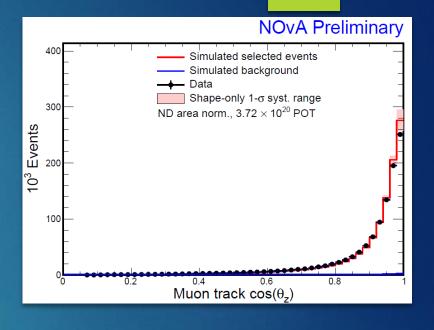




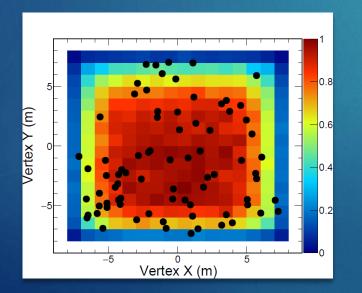


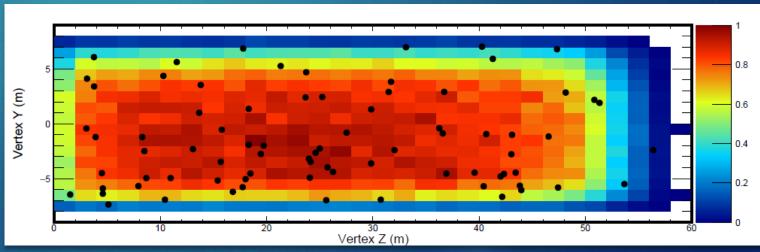
Muon Selection



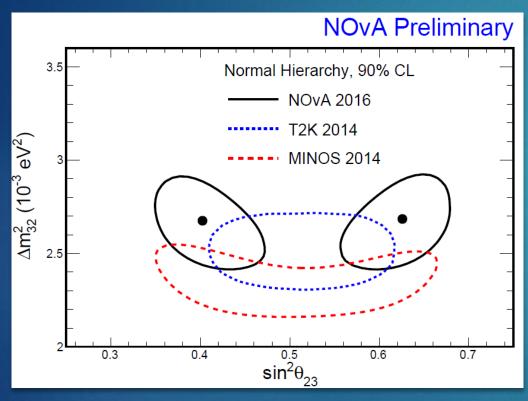


Muon Neutrino FD data

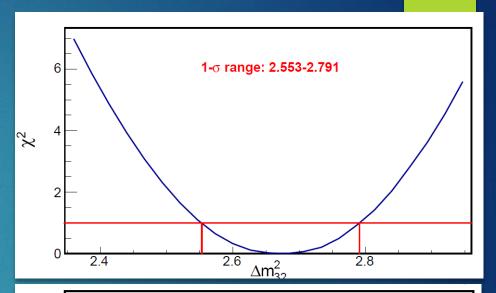


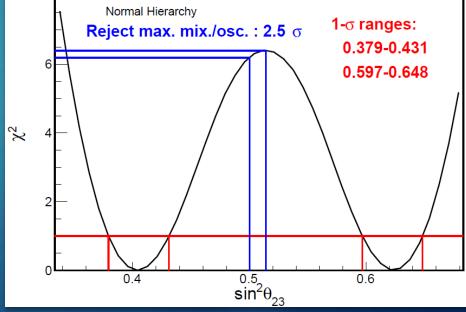


▶ 1-D profiles



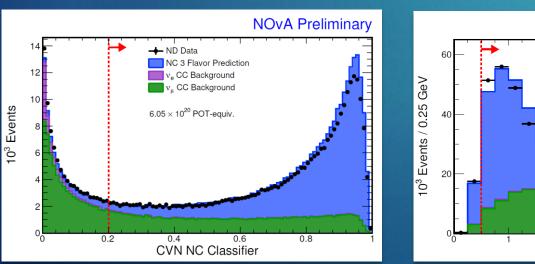
$$\left| \Delta m_{32}^2 \right| = 2.67 \pm 0.12 \times 10^{-3} \text{eV}^2$$
  
 $\sin^2 \theta_{23} = 0.40^{+0.03}_{-0.02} (0.63^{+0.02}_{-0.03})$ 

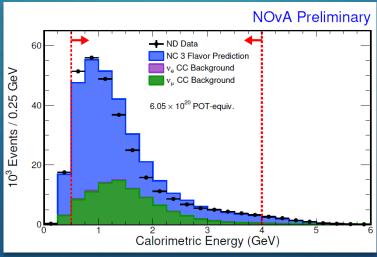


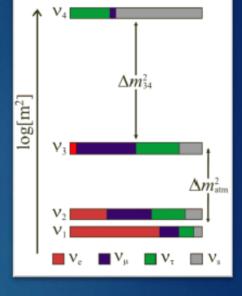


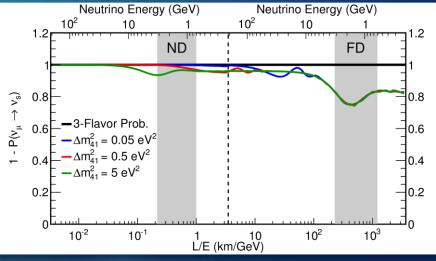
#### Neutral Current Results

- $\triangleright$  NC events in the ND with CVN classification, extrapolate to the FD  $\rightarrow$  prediction
- Count NC events in FD, compare to prediction
- For  $\Delta m_{41}^2 = 0.5 \text{ eV}^2$  rapid oscillations in FD, minimal in ND
- Normalization agrees well
- Data shifted to lower energy relative to MC
  - No MEC model for NC events
  - Large uncertainties on NC cross section









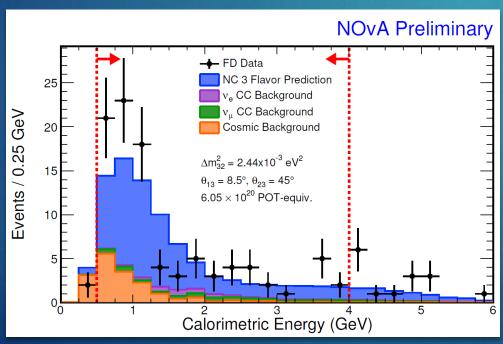
#### Neutral Current Results

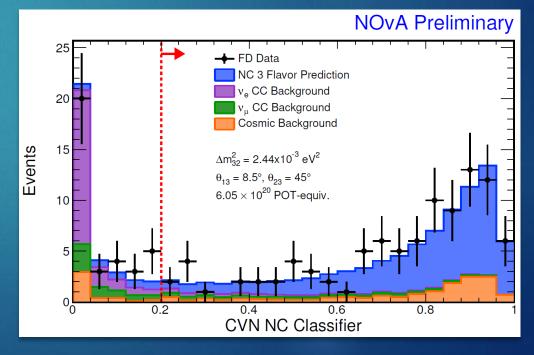
- Predicted events in the FD for 3-flavour mixing: 83.7 (60.6 NC, 4.8  $v_{\mu}$ CC, 3.6 beam  $v_{e}$ , 14.3 cosmics)
- Observed NC-like events in the FD: 95

No evidence of oscillations involving steriles, consistent within  $1\sigma$ 

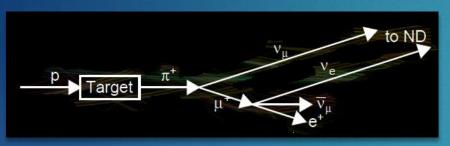
For  $0.05 \text{ eV}^2 < \Delta m_{41}^2 < 0.5 \text{ eV}^2$   $\theta_{34} < 35^\circ$ ,  $\theta_{24} < 21^\circ$  (90% CL)

Excellent NC efficiency (50%) and purity (72%) promise strong future limits on  $\theta_{34}$ 

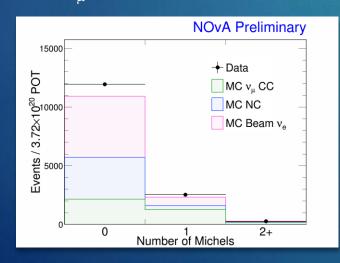


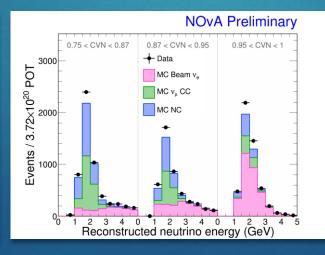


- $\triangleright$  CVN: 73%  $v_e$  signal efficiency, 76% purity
- Use ND data to predict FD background, every component propagate differently:
  - ▶ Beam v<sub>e</sub> CC



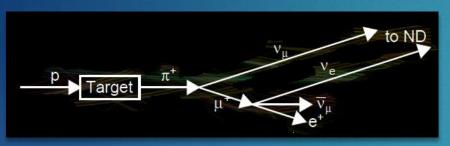
- low-E ν<sub>e</sub> and ν<sub>u</sub> trace back to the same π+ ancestor
- Use selected  $v_\mu$  CC events to constrain beam  $v_e$ : reweight Kaon and Pion component to match the  $v_\mu$  CC energy spectrum in the data
- Overall effect is a 4% increase  $\rightarrow$  Fix  $v_e$  CC to flux-reweighted in the ND
- $\mathbf{v}_{\mu}$  CC: use Michel-electron distribution to constrain



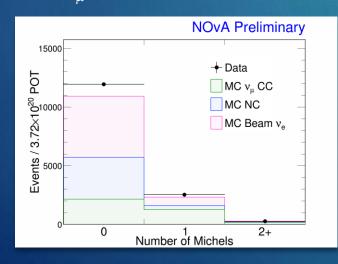


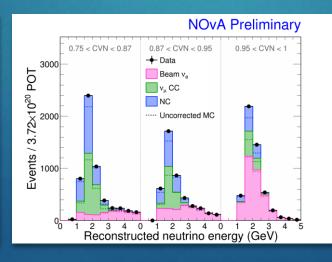
- Michel-e<sup>-</sup> are produced also in  $v_e$  CC and NC by pions but  $v_u$  has ~1more
- ▶ Fit observed N<sub>michel</sub> in each bin
- Data excess assigned between NC(+10%) and  $v_{\mu}$  CC (+10%)

- $\triangleright$  CVN: 73%  $v_e$  signal efficiency, 76% purity
- Use ND data to predict FD background, every component propagate differently:
  - ▶ Beam v<sub>e</sub> CC



- low-E ν<sub>e</sub> and ν<sub>μ</sub> trace back to the same π+ ancestor
- Use selected  $v_{\mu}$  CC events to constrain beam  $v_{e}$ : reweight Kaon and Pion component to match the  $v_{\mu}$  CC energy spectrum in the data
- Overall effect is a 4% increase  $\rightarrow$  Fix  $v_{\rm e}$  CC to flux-reweighted in the ND
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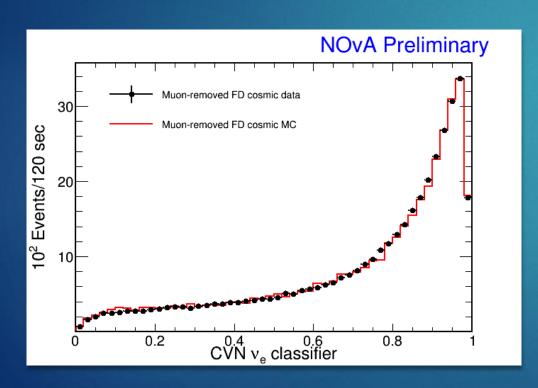




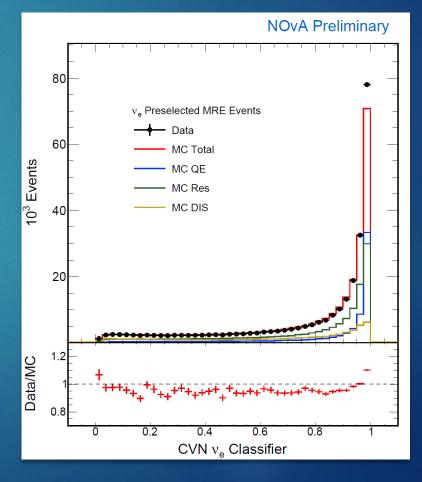
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#### Checking Signal Efficiency

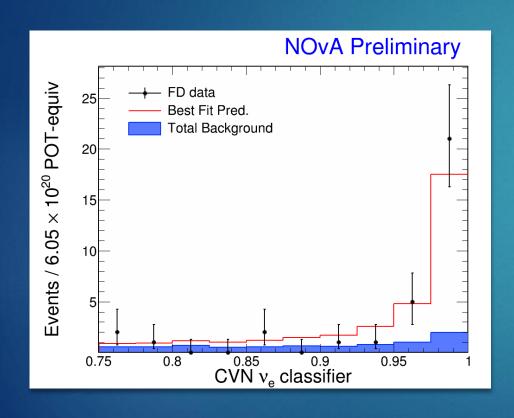
► Far detector: Remove muon track in cosmic rays to select Brem. Showers → simulation of EM showers matches well

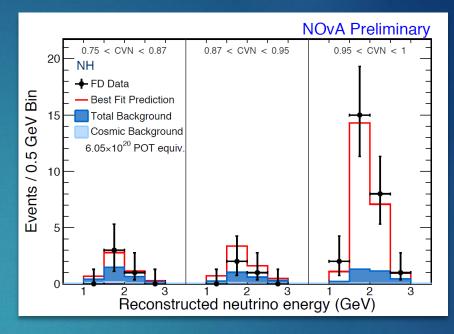


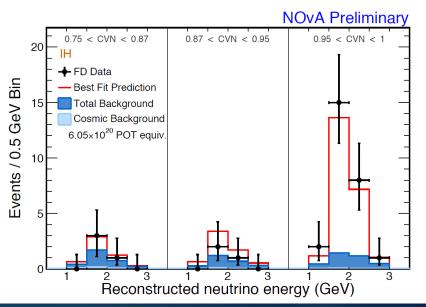
 Near Detector: replace muon tracks from ν<sub>μ</sub> CC data with simulated electron showers → data/MC difference < 1%</li>



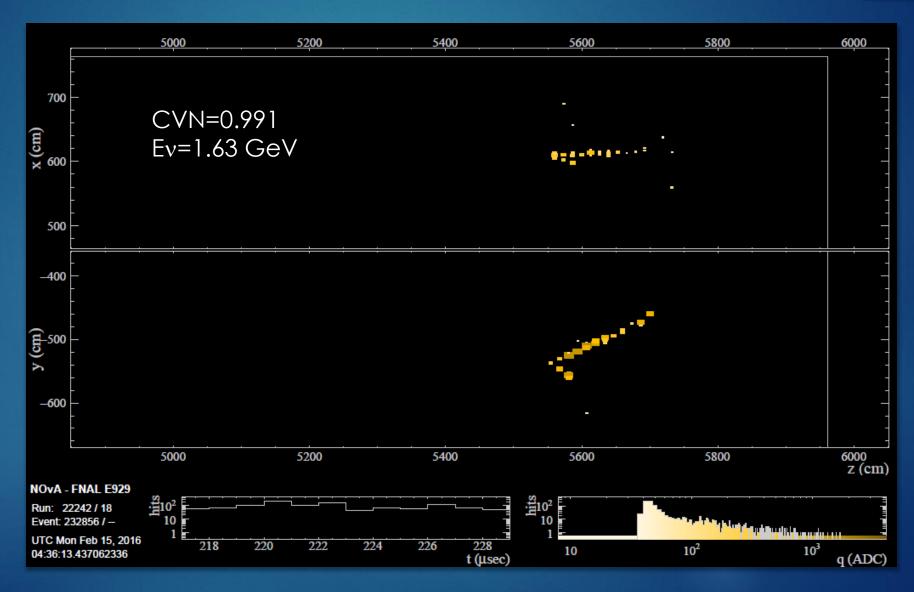
FD data



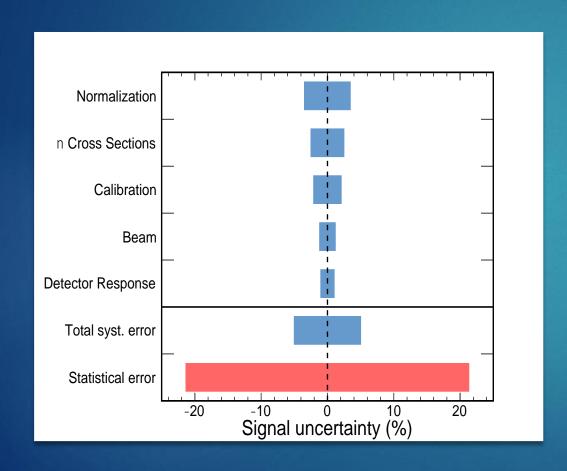


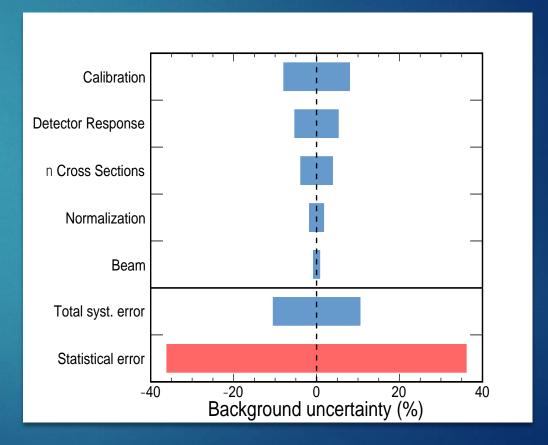


FD data



#### Systematics





Selection

