



WIR SCHAFFEN WISSEN – HEUTE FÜR MORGEN

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# Applied Particle Physics - Irradiation Facilities for Components Testing at PSI

LTP Monday Seminar 25.06.2018

# Outline

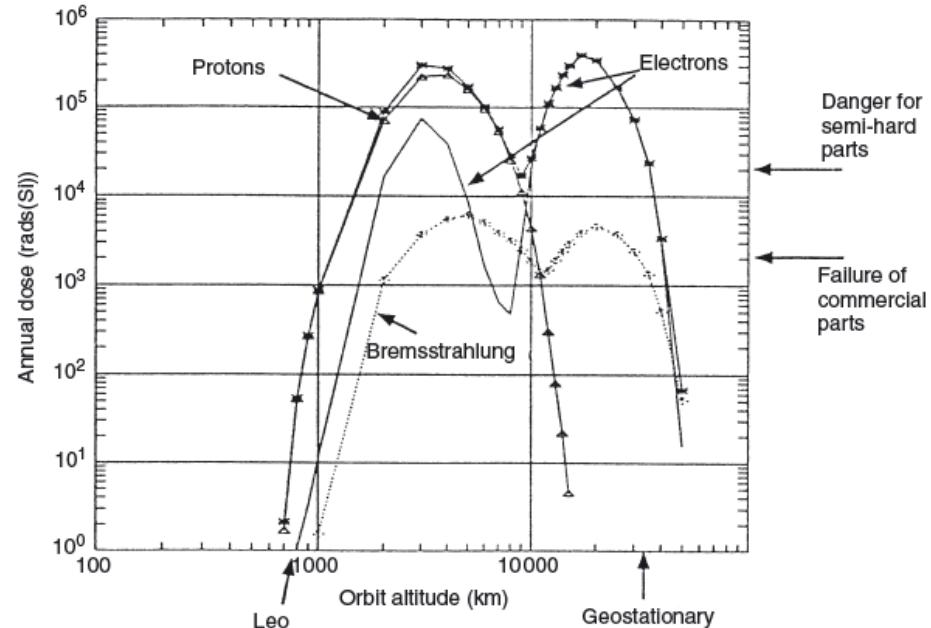
1. Spacecraft environments and radiation in space
  - a. European Components Irradiation Facilities
  - b. Why PSI and why protons? 1<sup>st</sup> latch-up in space
2. Radiation effects ...general
3. Radiation hardness test definition
4. ESA test methods / radiation hardness characterization
5. Effects examples from two missions: Galileo and POLAR
6. Space radiation environment: solar, cosmic, trapped, Jovian
7. Models vs. reality / example from RHESSI
8. PSI exposure facilities for radiation effects
  - a. Accelerators
  - b. Protons PIF
  - c. Electrons EMON
  - d. Electrons piM1
9. Detection system and dosimetry
  - a. Standard detector technology
  - b. New developments / Propix II (m)
  - c. PSI detectors for space RADEM (m)
10. Operation and users

# Spacecraft environment

## Spacecraft environment

- Pre-operational and Operational
- Launch phase
  - Vibration
  - Acoustics
  - Shock and acceleration
  - Thermal
  - Pressure
- Operational
  - Solar radiation
  - Ionizing radiation
  - Charging
  - Meteorites
  - Debris
  - Thermal
  - Earth orbit environment

Annual doses (SI) in **circular equatorial orbits**  
computed with SHIELDOSE and AEBMAX, APBMAX models  
4 mm spherical aluminium shielding.



# European Components Irradiation Facilities

## ECIF for studies of radiation effects in lab

- HIF / Belgium
- RADEF / Finland
- PIF / Switzerland
- CASE / The Netherlands

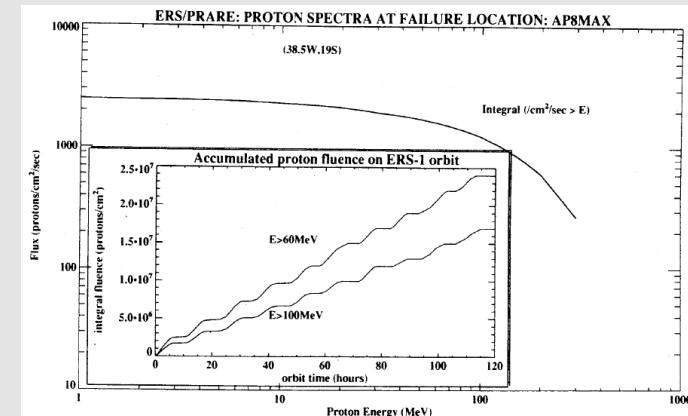
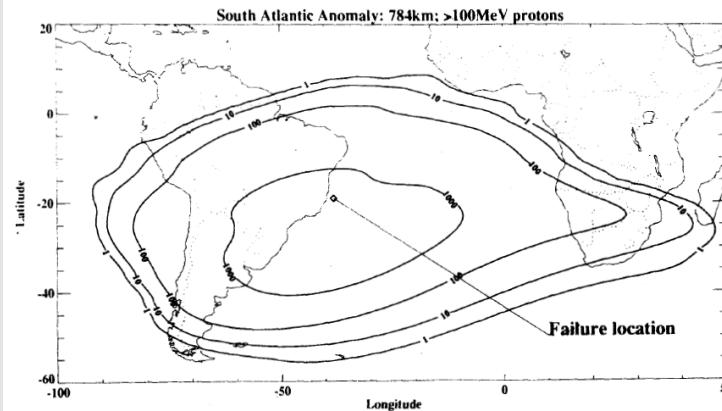
### *Other European Sites:*

- *TSL / Uppsala / Sweden*
- *GANIL / France*
- *GSI, COSY / Germany*
- *AGOR / The Netherlands*
- *Catania*
- ...



# Why PSI, why protons? 1<sup>st</sup> Latch-up in space

- Earth Resource Satellite launched in July 1991 into 784 km orbit
- The Precision Range and Range Rate Equipment shuts down 5 days later
- Reason - 9W overcurrent lasting 32s; the instrument did not restart anymore
- Switch-off occurred at the South Atlantic Anomaly
- ESA, PSI (R. Henneck) and University of Stuttgart exposed PRARRE to protons at OPTIS
- System behavior confirmed and latch-up found in CMOS RAM ( at  $F \sim 3 \cdot 10^7$  p/cm<sup>2</sup>! )

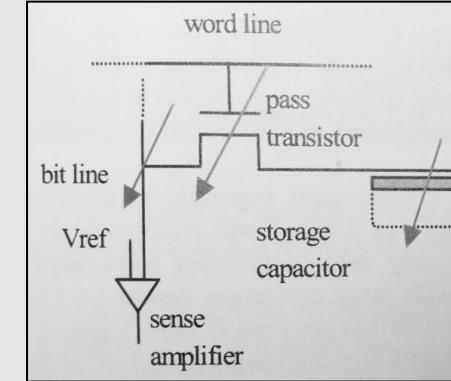


# Radiation effects ... broad view

## Relevant for tests at PSI

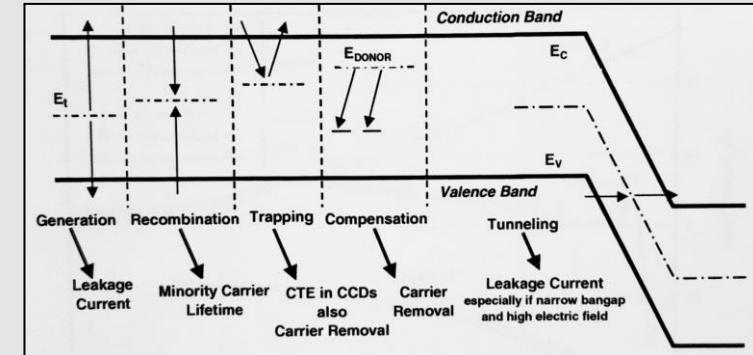
### Single Event Effects SEE

- Mostly in nuclear processes
- Critical charge
- Sensitive volume



### Displacement Damage DD

- Lifetime damage
- Carrier removal
- Mobility



# Radiation hardness test definition

Radiation effect	Parameter	Test means
Electronic component degradation	Total ionising dose	Radioactive sources (e.g. $^{60}\text{Co}$ ), particle beams: electron, <b>proton</b>
Material degradation	Total ionising dose	Radioactive sources (e.g. $^{60}\text{Co}$ ), particle beams: electron, <b>proton</b>
Material degradation (bulk damage)	Non-ionising dose (NIEL)	<b>Proton</b> beams
CCD and sensor degradation	Non-ionising dose (NIEL)	<b>Proton</b> beams
Solar cell degradation	Non-ionising dose (NIEL) & equivalent fluence.	<b>Proton</b> beams (~ low energy)
Single Event Effects SEE SEU, MEU, SEL, SEFI, SET, SEB, SEGR, etc. .	LET spectra (ions), proton energy spectra, explicit SEU/L rate.	Heavy ion particle beams <b>Proton</b> particle beams
Sensor interference - background signals	Flux above energy threshold, explicit background rate.	Radioactive sources, particle ( <b>e.g. proton</b> ) beams
Internal Electrostatic Charging	Electron flux, fluence dielectric E-field.	Electron beams Discharge characterisation

# ESA test methods

**ESTEC - EUROPEAN SPACE RESEARCH AND TECHNOLOGY CENTRE**

**ECSS - EUROPEAN COOPERATION FOR SPACE STANDARDIZATION**

- Standard documents

Radiation test specifications and procedures

**ECSS-Q-ST-70-06C Annex B**

Radiation test report

**ECSS-Q-ST-70-06C Annex C**

Request for radiation test

**ECSS-Q-ST-70-06C Annex A**

- Applicable documents

– Single Event Effects Test Method and Guidelines

– Total Dose Steady-state Irradiation Test Method

– Radiation Design Handbook

# Example-1 of effects in space – Galileo

Galileo - NASA mission to Jupiter  
 Launch 1989, operation 1995 – 2003  
 Most instruments affected by radiation

Displacement Damage:

- *Dark current*
- *Response uniformity*
- *Electronic chain degradation*
- *Sensor current consumption etc.*

Single Event Effects:

- *Upsets*
- *Transients*
- *Latch-ups*
- *Stacked bits etc.*

In the meantime things just ... evaluated



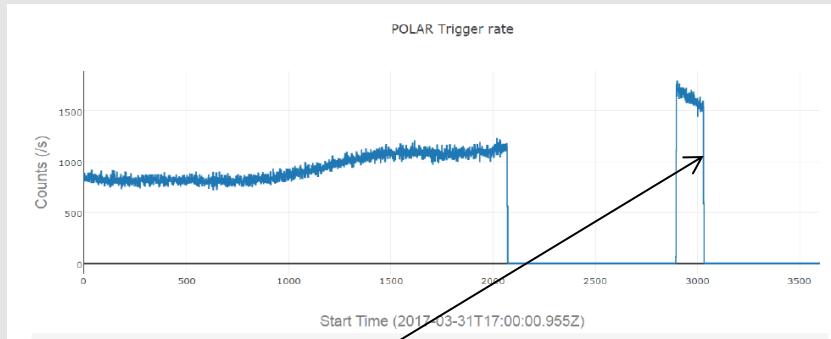
SYMPTOM	CAUSE	FIX
Spurious signals at slip rings.	++,1A	Reprogram software to ignore signals
Camera returns white images.	++,1A	Drop signal input to sensitive FET.
Infrared spectrometer (NIMS) memory resets.	++,1B	Scheduled software reloads in radiation.
Instrument (EPD) memory resets.	++,1C, 4C	Scheduled software reloads in radiation.
Quartz oscillator frequency changes.	++,2A,3A	Receivers widen bandwidths.
Spin detector signal noise increase.	++,2A	Reprogrammed to output a constant spin rate determined by other means.
Gyro electronics suffer signal bias.	++,2B	Frequent characterization tests. Less use of gyros.
Star Scanner sees false stars, blinded.	++,3A	Use bright stars.
Visible camera (SSI) image noise.	++,3A	Adjacent pixel averaging.
Polarimeter (PPR) signal noise.	++,3A,1C	Strip out "impossible" values from data set.
Infrared spectrometer (NIMS) signal noise.	++,3A,1C	Hand removal of noise from data set.
Dust detector (DDS) signal noise.	++,3A,1B	Instrument design allows noise/data discrimination.
Voltage controlled oscillator frequency jump.	++,1C, 2C	Pulse current to neutralize ion drift in electronic device.
Particle detector (EPD) sensitivity loss.	++,2B	Park detector behind nearby mass to provide shielding. Loss of channel in one case.
Spectrometer (UVS) grating failure.	++,2B	None - loss of instrument.
Photomultiplier tube (Star Scanner) gain loss.	++,2B	Use bright stars, adjust predicted intensities.
Camera (SSI) image compression failure.	++,2B	None - some forms of on-CCD compression lost.
S-band fr degradation in Io torus.	++,3B	De-weight data for navigation
Magnetometer processor lock-up.	+,1C, 4C	Scheduled power-cycles & memory reloads.
Voltage controlled oscillator frequency drift.	+,,2C	Input frequency adjusted to VCO's new base frequency.
Dust detector sensitivity decrease.	+,,2C	None
Analog to digital converter shift.	+,,2C	None
CMOS Memory cell failures.	+,,4C	Reprogram around failed cells.

# Example-2 of effects in space – POLAR

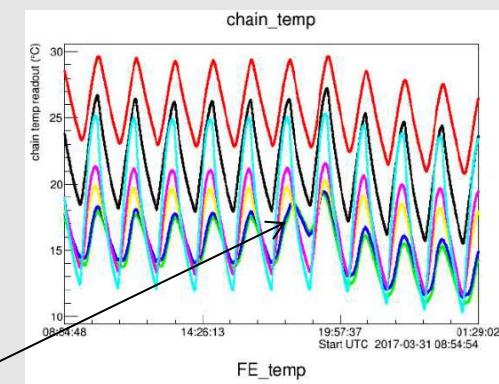
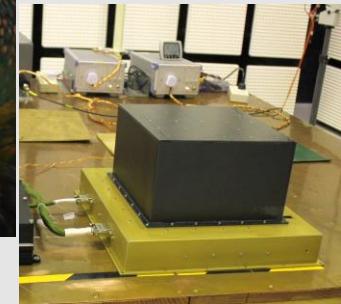
## POLAR – novel hard X-ray polarimeter (CH-CN)

- Launched in Sep 2016 on Chinese TG2
- Goal - polarimetry of Gamma Ray Bursts (linked to GW)
- PSI electronics onboard: Frontends and Central Unit
- High Voltage unit lost on 31.03.2017 – catastrophic failure
- Location at SAA boarder
- Reasons unclear: charging, protons ...

(see [polar.psi.ch](http://polar.psi.ch) , POLAR PSI Data Center by Hualin Xiao)

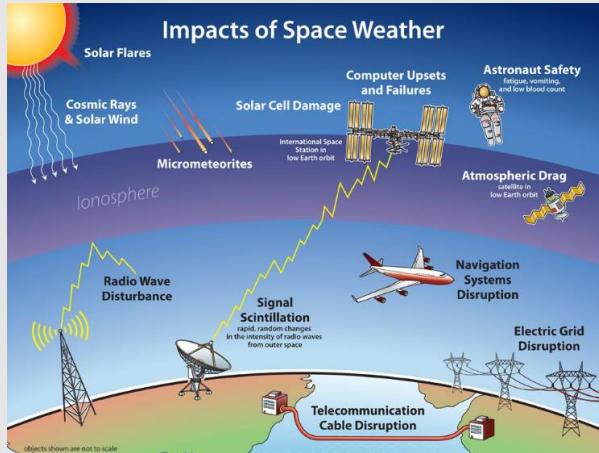


*Sudden rate break with HV loss and temperature rise*



# Space radiation environment

Real, dynamic, wide-ranging and costly



## Solar Events

Solar energetic particles  
Geo-storms; particle injections

## Trapped Radiation

Proton belt  
Electron belts  
South Atlantic Anomaly  
Human made

## Cosmic Rays

Galactic  
Anomalous

## Jupiter Radiation Environment

## Hot Plasma

*Testing on ground is indispensable*

# Solar events

## Solar flares

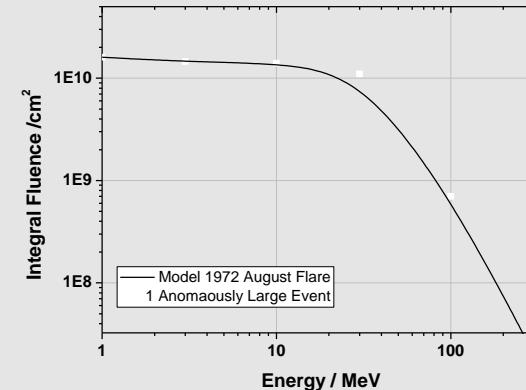
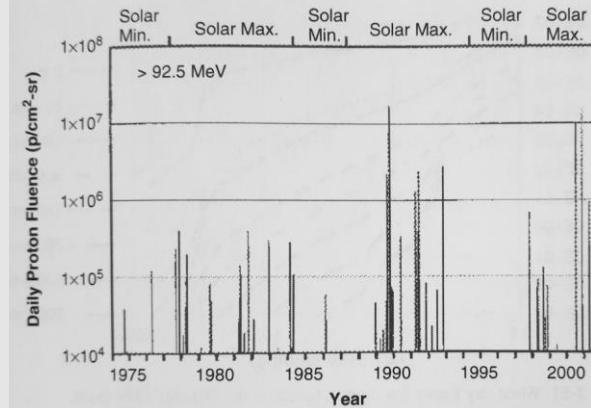
Electron rich  
Lasting hours

## Coronal mass ejections

Proton rich  
Lasting days

*Solar energetic particles can cause doses of even ~10 krad*

See e.g. [srem.web.psi.ch](http://srem.web.psi.ch) for observations of PSI designed ESA Standard Radiation Environment Monitor SREM



# Cosmic rays

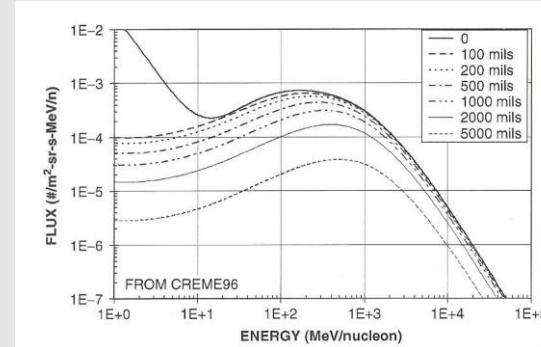
## Galactic Cosmic Rays GCR

- Low intensities
- High penetration
- High energies > GeV/n
- High LET >  $10^4$  MeV/g/cm<sup>2</sup>

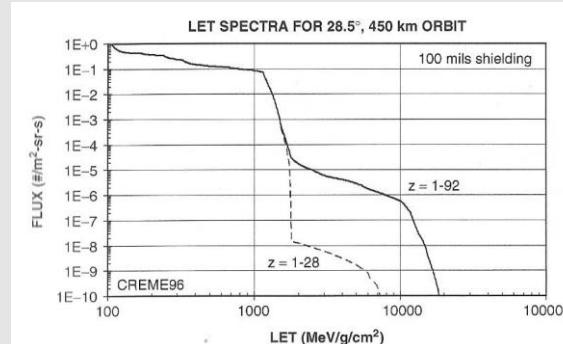
## Anomalous Cosmic Rays ACR

- Reaching tens MeV/n
- Weakly ionized

- GCR mainly responsible for Single Events Effects
- In longer missions also for Displacement Damage and Total Ionizing Dose effects
- ACR effects often negligible



*GCR iron spectra behind Al-shielding*

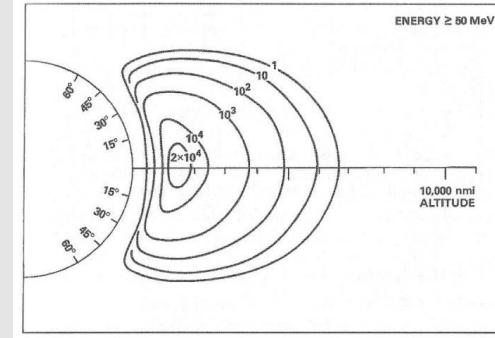


*GCR LET spectra with two different ranges of elements*

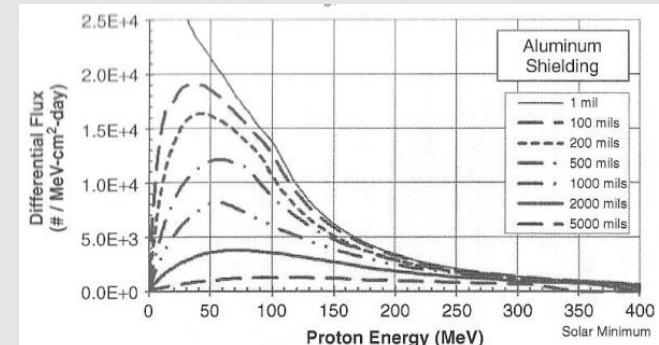
# Trapped protons I.

## Proton Radiation Belt and South Atlantic Anomaly SAA

- Partially responsible for Single Event Effects
- Major agent for Displacement Damage
- Energies above hundred MeV
- Shielding partly effective
- Typical doses  $\sim 10$  rad/day
- Peak energies after shielding  $\sim 50\text{-}60$  MeV
- Proton belt is relatively stable



*AP8MIN omnidirectional proton flux ( $p/\text{cm}^2/\text{s}$ )*

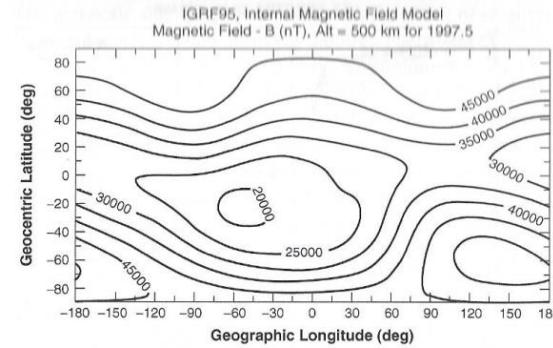
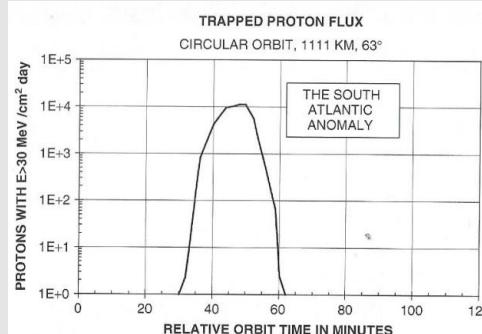


*Proton spectra behind Al-shielding (97° inc., 500 km)*

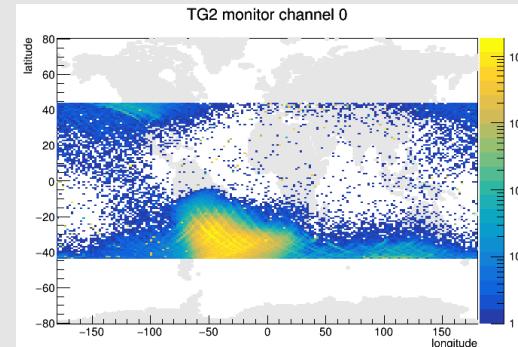
# Trapped protons II.

## Proton Radiation Belt and South Atlantic Anomaly SAA

- SAA dominates at low altitudes and inclinations
- Caused by offset and tilt between geo-magnetic axis and Earth rotation axis
- The belt is intruding to lower altitudes
- Both protons and electrons are present
- *ISS also crosses SAA (fast passage ~20%)*



Magnetic field contours for 500 km altitude

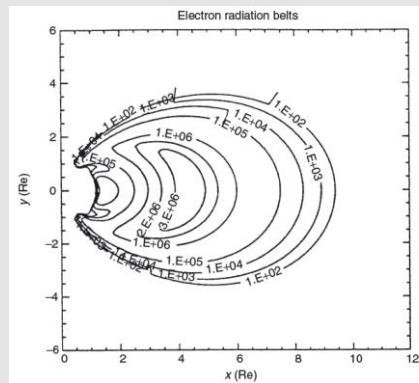


TG2 Radiation monitor rate map with SAA (6.3.2017)

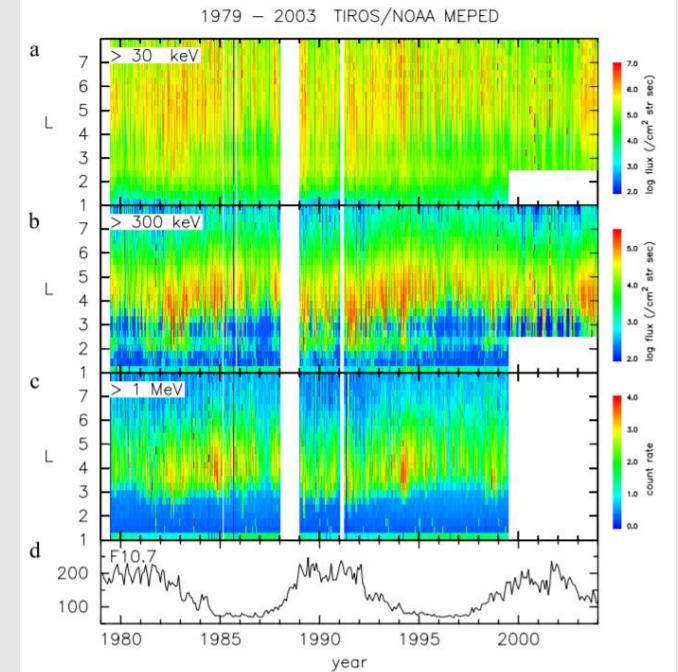
# Trapped electrons

## Radiation Belts: inner, outer and SAA

- Major agent for total ionizing dose
- Energies up to few MeV
- Highly dynamic, sensitive to storms
- Extends behind GEO orbit
- Shielding effective (local, thin plates)
- Typical doses  $\sim 10$  rad/day
- Causes bremsstrahlung



*AE8 model omnidirectional electron flux ( $\text{e}/\text{cm}^2/\text{s}$ )*

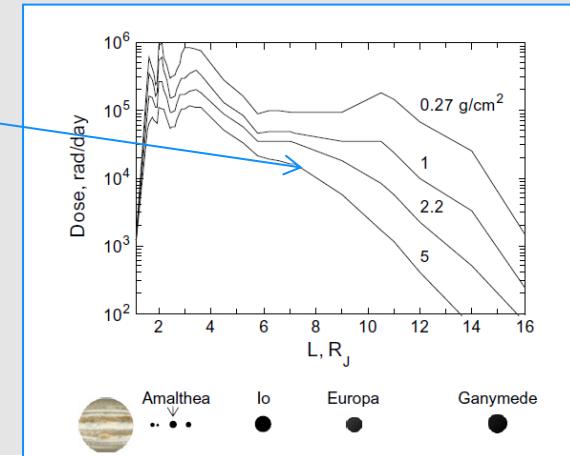
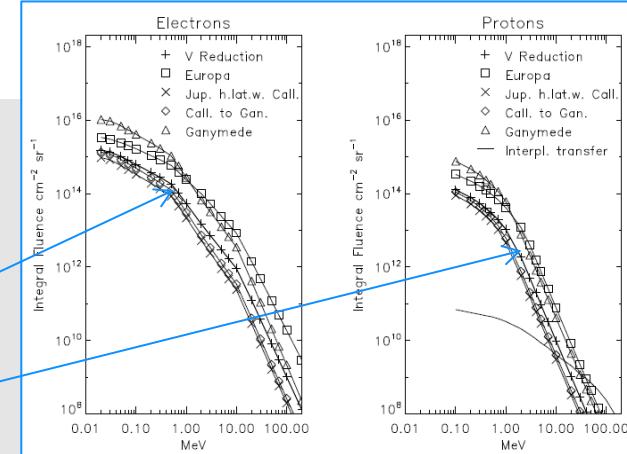


*Electrons flux data from NOAA satellites ( $E > 1\text{MeV}$ )*

# Jupiter radiation environment

## Challenge for satellites and subsystems

- Intense radiation belts
- Highly variable environment
- Very high fluxes / doses
- Examples for ESA JUICE mission (GIRE model)
  - $F_e \sim 10^{14}-10^{15}/\text{cm}^2/\text{sr}$  ( $E > 300$  keV)
  - $F_p \sim 10^{11}-10^{12}/\text{cm}^2/\text{sr}$  ( $E > 5$  MeV)
  - Dose rate  $\sim 10^4 - 10^5$  rad/day
- High energy tails ( $> 100$  MeV)
- Presence of heavier elements
- Strong magnetic fields
- *Not well studied yet*
- *PSI participates in JUICE radiation monitor*



# Models vs. reality / example from RHESSI

## ECSS Standard models

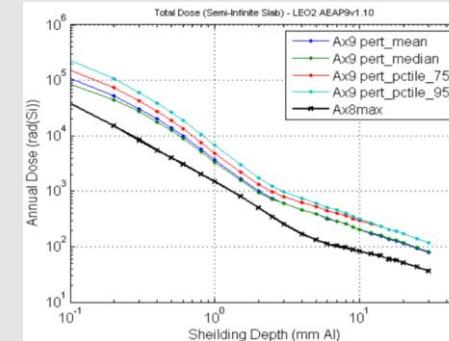
- AE8 trapped electrons
- AP8 trapped proton
- *Static, isotropic models based on limited dataset for solar minimum and maximum*
- *Not suited for low altitude and inclination*

ESA estimated accuracy at factor of 2-3

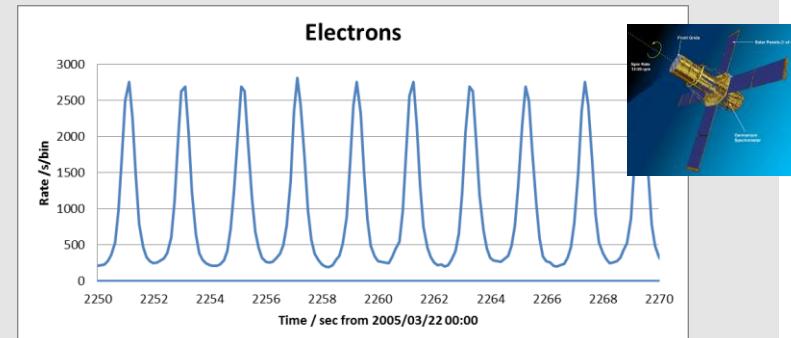
## New development – AE9 and AP9

- with uncertainties and errors in data
- based on probabilities with Monte Carlo

Tend to predict higher fluxes than AP8/AE8



Comparison of AX8 and AX9 models for LEO orbit

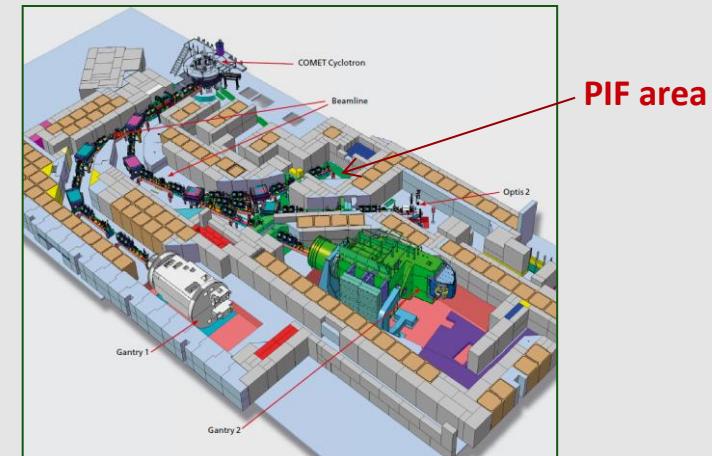


Electron rate in RHESSI monitor vs. rotation angle

# PSI exposure facilities for radiation effect studies

- Proton Irradiation Facility operates continuously since 1992
- Connected to COMET cyclotron of the Proton Therapy Center
- Priorities given by patient exposure plan
- Other exposure sites and particles are also utilized:
  - piM1 secondary beam area with electrons, pions and protons
  - Electron mono-chromator with mono-energetic electrons from beta sources

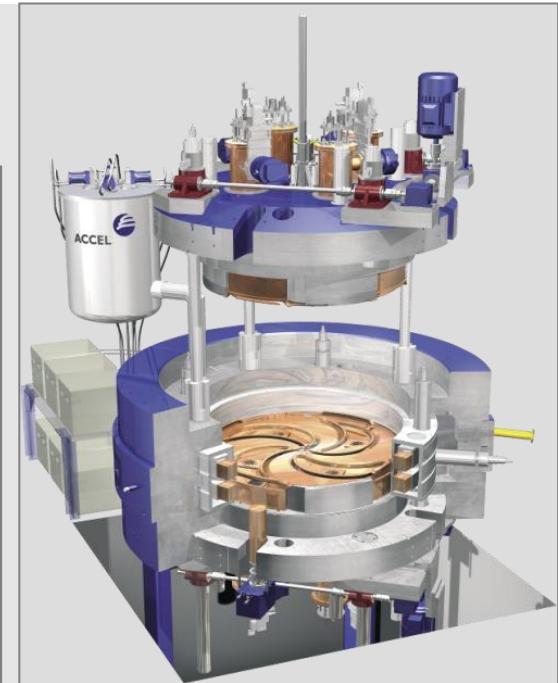
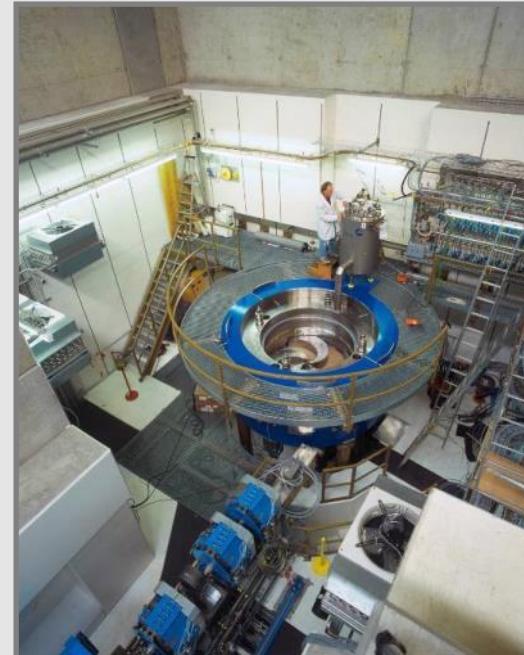
- Main functions:
  - User-lab for radiation effects studies in electronics
  - Realistic simulator of space radiation environment
  - Source of mono-energetic particles for rad-tests
  - Calibration station for monitors and detectors
  - Radiation qualification for space technologies



NA area with COMET cyclotron and exposure sites

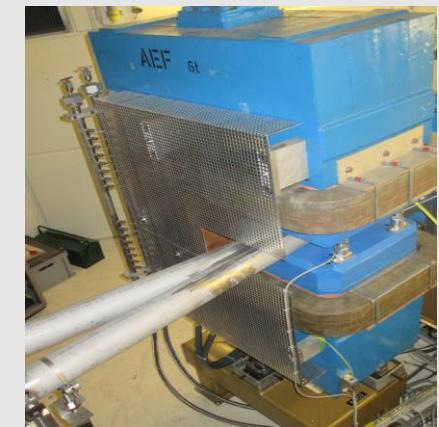
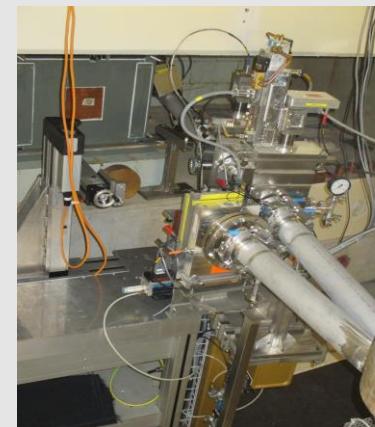
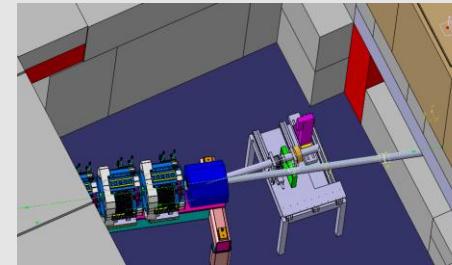
# COMET Compact Medical Therapy Cyclotron

- Facility of the PSI Proton Therapy Center
- Designer - H. Blosser, MSU/USA
- Delivered by VARIAN
- 250 MeV fixed energy
- Mass 90 tons
- Intensity range 0-1000 nA
- In operation since 2007



# Proton Irradiation Facility

- Initial energies: 230, 200, 150, 100, 74 MeV
- Energies after degrader: 230 MeV to 6 MeV
- Max intensity: 2 nA ( $E>200$ ) – 10 nA ( $E<100$ )
- Flux range  $10^2$  –  $2 \cdot 10^8$  p/sec/cm $^2$
- Profiles Gaussian-like: FWHM 9 cm
- Max beam diameter of 90 mm
- Options: focused (6 mm  $\phi$ ) or flat beam
- User adapted dosimetry and test flow
- Standard calibrations runs and checks / fluxes, profiles, scaling

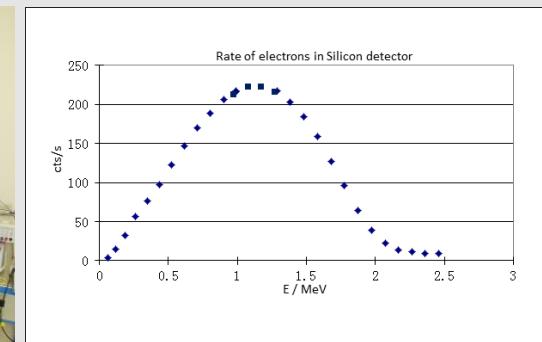


# Electron monochromator

- Simple bending magnet and strong electron source in large vacuum chamber
- Flux control system constructed; Si-detector with dedicated DAQ
- XY-table with remote sample control; support with TV cameras and illumination
- Two units:
  - PSI
  - ESTEC



*Si-sensor and DAQ system (left)  
Monochromator chamber (right)*



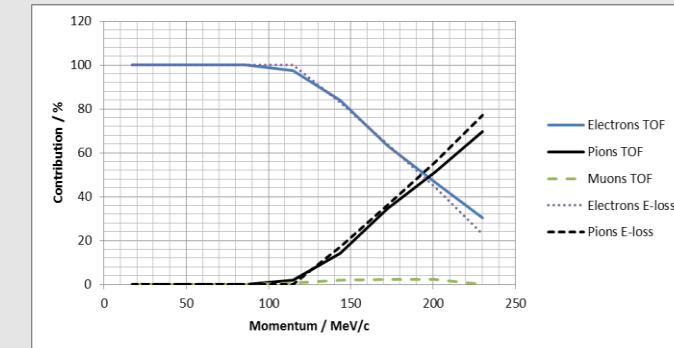
*Intensity curve for <sup>90</sup>Sr source; Si-detector 2 cm from beam exit*

# High energy electrons in piM1 I.

- Adapting secondary beam area piM1 of PSI large cyclotron
- Positive and negative particles possible
- Clean electrons beams from about 10 MeV up to 100 MeV
- Protons available up to 70 MeV
- Pions and muons from 100 MeV/c to 350 MeV/c



*piM1 Test area: beam exit and PIP-JUICE setup*



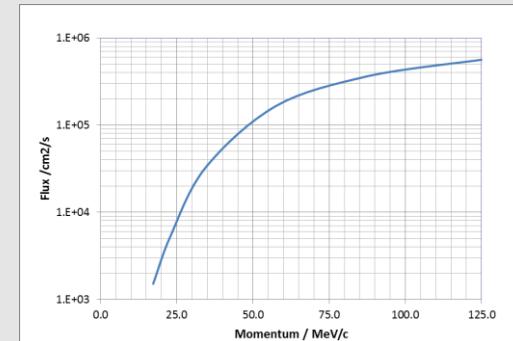
*Beam contamination level as function of momentum*

# High energy electrons in piM1 II.

- Typical intensities:  $2 \cdot 10^5 - 1 \cdot 10^7$  /s and fluxes:  $2 \cdot 10^3 - 5 \cdot 10^5$  /cm<sup>2</sup>/s
- FWHM between 4 cm and 10 cm
- Well suited for studies of instrument shielding and calibration
- Too low fluxes for TID tests; other test areas studied

*Electron beam parameters in piM1 area.*

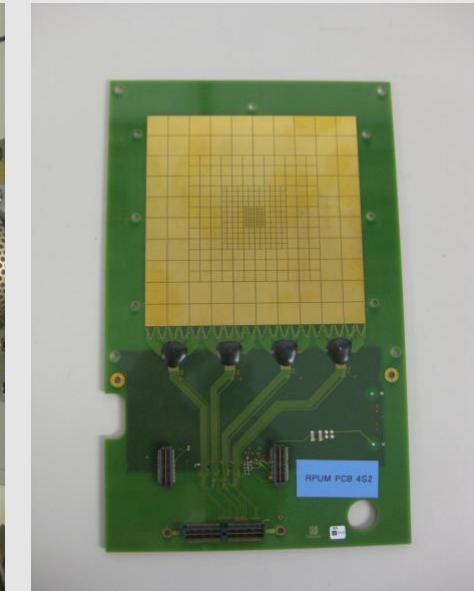
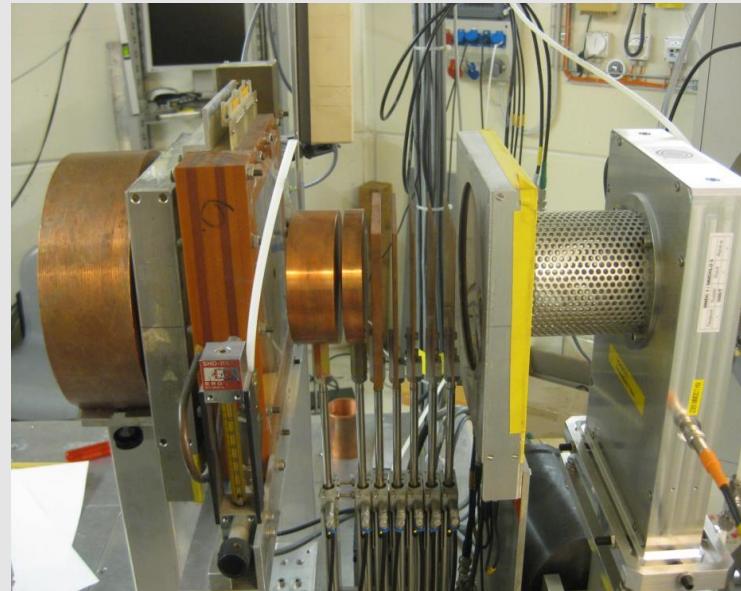
Momentum MeV/c	Intensity s/mA	Flux cm <sup>2</sup> /s/mA	FWHMx cm	FWHMy cm
17.3	1.16E+05	7.21E+02	10.4	13.2
23.0	3.28E+05	2.57E+03	9.0	12.9
34.5	1.16E+06	1.56E+04	6.6	9.6
57.5	3.08E+06	7.88E+04	5.2	6.6
86.3	5.13E+06	1.69E+05	4.2	5.1
115.0	5.18E+06	2.42E+05	4.4	4.3



*Electron flux vs. momentum*

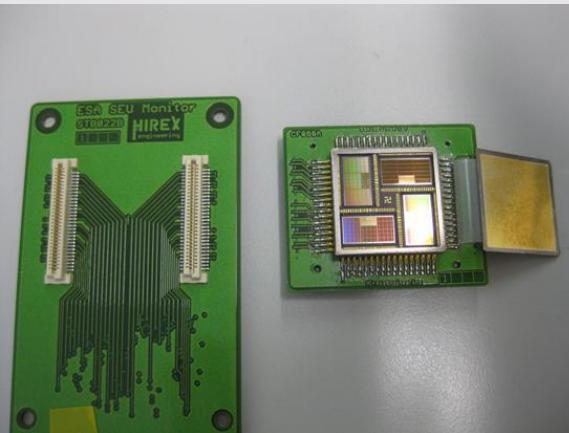
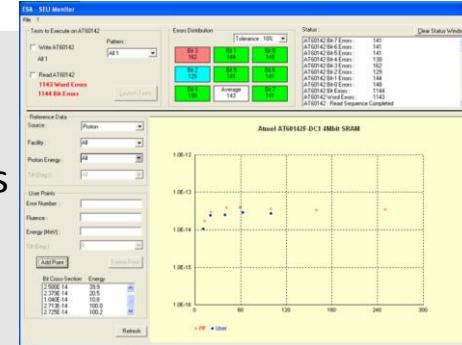
# Detector technology and dosimetry

- Plastic or Silicon detector to calibrate the beam
- Real time dosimetry uses two ionization chambers IC
- Profiles are measured with pixelated IC, plastic scans or luminescence foils



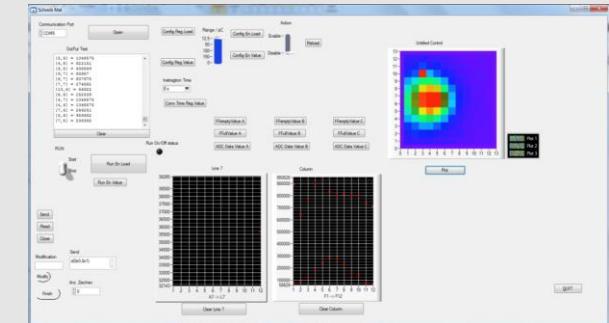
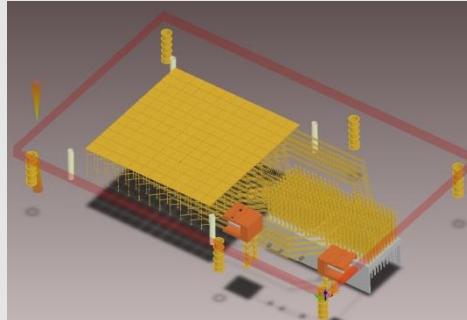
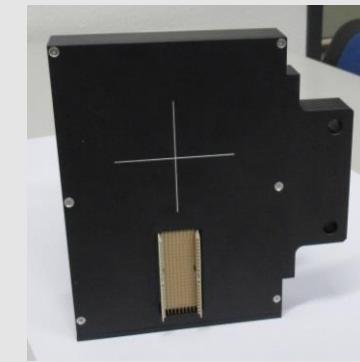
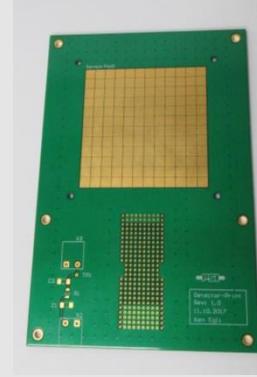
# Dosimetry I. ESA SEU Monitor

- Developed in 2005 by R.H. Sorensen (ESA-ESTEC) and HIREX
  - Based on ATMEL AT60166F 16Mbit SRAM, version 2009
  - Carefully calibrated at ESA facilities; tens of units at different test sites
  - Easy comparison of measured and expected flux values



# Dosimetry II. new PIF pixel ionization chamber

- Standard ionization chamber components
- Sensitive, improved readout electronics
- Easy data acquisition system
- Easy setup and operation
- Fast beam profiling
- Very wide dynamic range

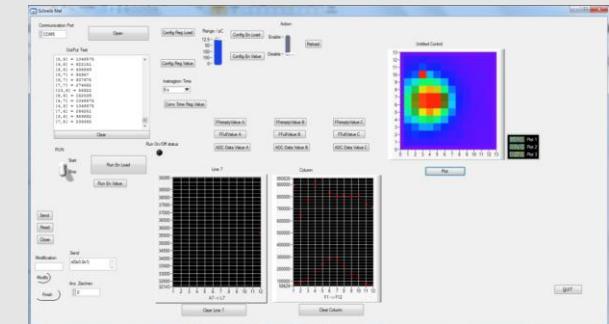
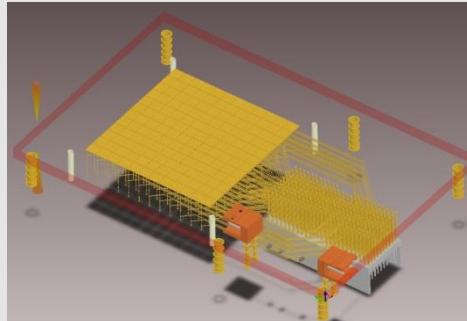
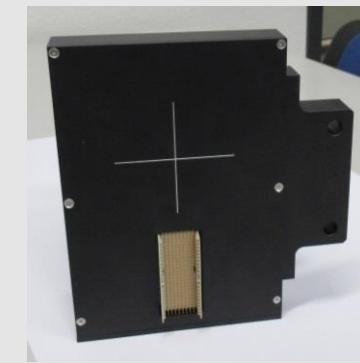
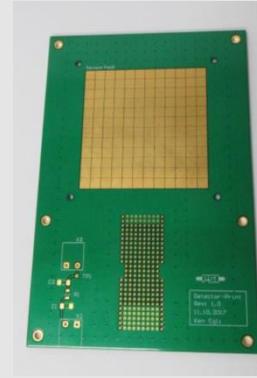


# Dosimetry III. new PIF pixel ionization chamber



# Dosimetry II. new PIF pixel ionization chamber

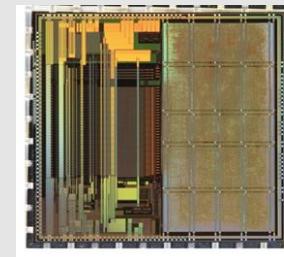
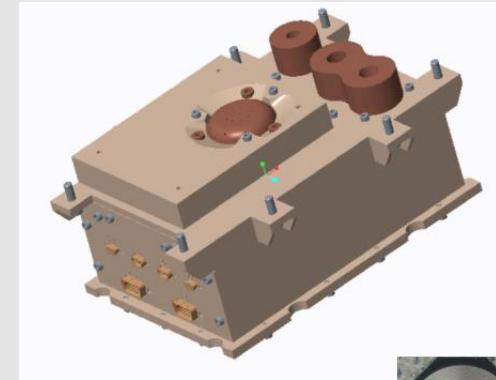
- Standard ionization chamber components
- Sensitive, improved readout electronics
- Easy data acquisition system
- Easy setup and operation
- Fast beam profiling
- Very wide dynamic range



# Detector technology I. RADEM for ESA JUICE

## Radiation Hard Electron Monitor RADEM

Electron energy range	0.3 - 40 MeV
Proton energy range	5 - 250 MeV
Energy resolution	8 log bins for e and p
Peak flux	$10^9 /cm^2 s^1$
Ion sensitivity	LET 0.1-10 MeV/(mg/cm <sup>2</sup> )
Directionality	31 directions; $\Delta\Theta = \pm 75^\circ$
Mass; Volume	$\sim 4$ kg; $1000$ cm <sup>3</sup>
Power; Temperature	$\sim 4$ W; $-30/+50$ °C
Lifetime	11 years
Dedicated readout ASIC	VATA466



*RADEM EM currently at PSI for particle tests  
Final delivery – December 2019*

# Detector technology II. RADEM for ESA JUICE



# Detector technology II. RADEM for ESA JUICE



- Structure integration
- Propulsion integration
- Harness integration
- MAG off-line pre-assembly
- Optical bench off-line pre-assembly
- Hardware units mechanical and electrical integration
- Walls mounting
- External units mounting**
- Optical bench coupling
- MAG boom coupling
- HGA integration
- MLI installation
- Solar generator integration



RIME antenna  
PEP/NU  
PEP/JENI  
RADEM  
RIME/Matching network  
SAS +Z

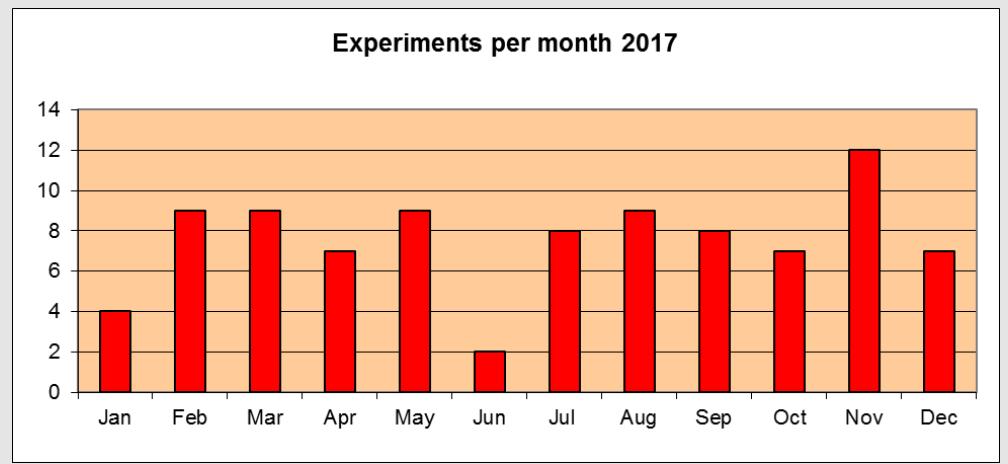
# Operation and users I.

- Operation: weekends and nights; usually 2-3 weekends/month
- Flexible, user-specific test arrangement
- Fast, uncomplicated set-up and operation
- Automated irradiations
- Standard sample frame  
(as at HIF, RADEF and Brookhaven)
- Irradiation usually in air
- Typical laboratory apparatus available  
(also vacuum chambers)

## Statistics for 2017

Shifts	> 200
Number of tests	- 91
Institutions	- 29
Users	- 160
ESA related shifts	- 142
ESA pool shift	- 30

# Operation and users II.



Nr	Institution	Project	Days	Start	Stop	Shifts	ESA	Project	ESA Pool
1	PSI	Maintenance services of the proton and electron facility	3	18.1	20.1	3	Y	N	
2	PSI	Test of beam optics and detector winter shutdown	2	22.1	23.1	1.5	Y	N	
3	PSI	Maintenance services of the electron facility	1	23.1	22.1	1	Y	N	
4	E2V	DD and TID tests of new CCDs	2	29.1	30.1	1	Y	N	
5	SpaceTech	Tests of components for MERLIN mission / TID and DD	2	3.2	4.2	2	Y	N	
6	PSI	Test of optical sensors and transistors	2	1.2	2.2	1	Y	N	
7	ESA	Test of OpAmps and transistors	2	4.2	5.2	3	Y	Y	
8	E2V	Various optical sensors exposures	2	10.2	11.2	1.5	Y	N	
9	PSI-EFAC-E2V-LIP	Measurements active area and the RADEM Directionality Sensor	1	12.2	12.2	2	Y	N	
10	Open University	Exposure series of beam optics from E2V	2	12.2	12.2	2	Y	N	
11	CERN	Beamline dosimetry electronics	3	16.2	18.2	5	N	N	
12	E2V	Irradiation of CCDs	2	24.2	25.2	2	Y	N	
13	Heise Scale Arc	Irradiation of CCDs	2	25.2	26.2	2	Y	N	
14	ESA, TRAD	Tests of various SEE components	2	3.3	4.3	3	Y	Y	
15	Thales-Alenia	Test and calibration of NGRM EM / noise thresholds, responses	3	1.3	4.3	3	Y	N	
16	Thales-Alenia	Test and calibration of NGRM EM / noise thresholds, areas and deadtime	2	4.3	5.3	2	Y	N	
17	ETEC	Testing of NEO-MIST GNSS receiver	1	15.3	15.3	2	Y	N	
18	E2V	Exposure and tests of optical sensors from E2V	1	18.3	18.3	2	Y	N	
19	CERN	Beamline dosimetry electronics	2	25.3	26.3	4	N	N	
20	EFACCO-PSI-LIP	Characterization of RADEM Detectors / area, sensitivity	2	1.2	2.2	2	Y	N	
21	PSI-EFAC-E2V-LIP	Test and calibration of NGRM EM with electrons	1	22.3	22.3	2	Y	N	
22	Thales-Alenia	Tests of active area of DD sensor (signal line contribution) / RADEM	1	23.3	23.3	2	Y	N	
23	SAT	Test of various optical components and detector sensitivity	2	1.3	2.3	2	Y	N	
24	Tsinghua University, PSI	TID / DD characterisation of CASCA ASIC	2	7.4	7.4	1.5	N	N	
25	Thales-Alenia	Calibration of NGRM EM accurate deadtime tests	2	7.4	8.4	2	Y	N	
26	CERN	Beamline dosimetry electronics	2	22.4	23.4	4	N	N	
27	ESR, Uni Monash, ESCC, JPL	Test of various optical components	1	23.4	24.4	3	Y	N	
28	E2V	Optical devices for Project	1	30.4	30.4	1	Y	N	
29	PSI, ESCC and Un Montpellier	Characterization of Beam / noise of beam settings	1	30.4	30.4	1	Y	N	
30	PSI	Test of proton beam at 375 MeV/c at PM1	1	16.3	16.3	5	Y	Y	
31	Tsinghua University, PSI	Beam optimization for CASCA ASIC SEE test at pM1	2	21.5	21.5	1	N	N	
32	E2V	Test of optical sensors at pM1	2	22.5	23.5	3	Y	N	
33	Tsinghua University, PSI	SEE tests of the CASCA ASIC at pM1	2	23.5	24.5	3	Y	N	
34	PSI	Test of electron beam at 375 MeV/c at pM1	1	24.6	25.6	2	Y	Y	
35	PSI-EFAC-E2V-LIP	Proton tests of RADEM BB with ASIC VATA466 at pM1	2	25.5	26.5	2	Y	N	
36	PSI-EFAC-E2V-LIP	Electron test of RADEM BB with ASIC VATA466 at pM1	1	27.5	27.5	1.5	Y	N	
37	PSI-EFAC-E2V-LIP	RADEM BB with ASIC VATA466 at pM1	2	28.5	29.5	3	Y	N	
38	Thales-Alenia	NGRM EG, test of electron discrimination in EDSS and SDSS	2	29.5	30.5	3	Y	N	
39	PSI-EFAC-E2V-LIP	RADEM BB with ASIC and detector test with electrons from EMON	2	1.6	2.6	2	Y	N	
40	PSI	Beam tests of various types of systems after upgrade	2	28.6	29.6	2	Y	N	
41	Tsinghua University / PSI	Year of Si-OfS / SiS	1	1.2	2.7	2	N	N	
42	ESA ESTEC	Tests of optical tracelens	3	7.7	9.7	4.5	Y	Y	
43	CERN	Various dosimetry electronics tests	3	14.7	16.7	5	N	N	
44	ESTEC	Test of various ICs	2	21.7	22.7	2	Y	N	
45	PSI PIF	Software and hardware optimization for dosimetry	1	22.7	22.7	1	Y	N	
46	PSI / LIP / EFAC	Test of various IC	1	23.7	23.7	2	N	N	
47	PSI / LIP / EFAC	Optimization of beam optics and sensors installation	3	10.7	12.7	3	Y	N	
48	PSI	RADEN EM tests of Si-diode energy resolution	2	25.8	26.8	3	N	N	
49	PSI	Test of millions of electrons	3	26.8	28.8	5	N	N	
50	Xinjiang Institute of CAS	Further tests of beam optics and sensors installation	3	26.8	28.8	1.5	N	N	
51	Xinjiang Institute of CAS	Xinjiang Institute of CAS / continuation	1	28.8	28.8	1	N	N	
52	PSI / LIP / EFAC	RADEN FM diodes electronic resolution tests	1	2.8	2.8	1.5	Y	N	
53	PSI / LIP / EFAC	RADEN FM diodes electronic resolution tests	1	4.8	4.8	1.5	N	N	
54	PSI / LIP / EFAC	RADEN FM diodes electronic resolution tests	2	17.8	18.8	2.5	Y	N	
55	PSI / LIP / EFAC	RADEN FM diodes electronic resolution tests	1	28.8	28.8	1.5	N	N	
56	PSI	Replacement of IC channel and new Plastic detector tests	1	23.8	23.8	1	Y	N	
57	INRA	Test of various beam optics	1	1.2	2.2	2	Y	N	
58	PST	Electron beam tests at pM1	2	1.9	5.9	2	N	N	
59	PSI	Test of various optical sensors	4	13.9	16.9	4	N	N	
60	SiLab	Test of PsiASIC / FGAs	1	16.9	17.9	2	N	N	
61	AIRBUS	Test of various events in beam memories	30.9	30.9	2	N	N		
62	PSI / LIP / EFAC	RADEN FM diodes electronic resolution re-tests	2	6.9	8.9	2	Y	N	
63	PSI / LIP / EFAC	RADEN FM diodes electronic resolution re-tests (3 mm diam)	1	11.9	11.9	1.5	N	N	
64	PSI / LIP / EFAC	RADEN FM Si-diodes resolution re-test of electronic resolution	2	20.9	20.9	2	Y	N	
65	AIRBUS	Test characterization of flash memories	2	1.0	2.0	2	Y	N	
66	EASII	Test of various components	3	6.1	8.1	4	Y	N	
67	E2V-E2V-Tel	DO and DZ tests of new CCDs	2	1.2	2.2	1	Y	N	
68	TRAD	Radiation tolerance of fast memories	3	14.0	16.0	6	Y	N	
69	PSI	Test of Si-PMTs	1	28.0	28.1	1	N	N	
70	PSI	Beam optimization and calibration of RADEM / JUICE	1	29.10	29.10	1	Y	N	
71	PSI	First commissioning of RADEM EM	2	29.11	29.11	3	Y	N	
72	EFACCO-PSI-LIP	TPC and absorbers tests	2	3.11	4.11	2.5	Y	Y	
73	Selsendorf Lab	Test of various components for ESA projects	3	4.11	6.11	4.5	Y	Y	
74	HREX	Test of various components for ESA projects	2	2.11	3.11	2	Y	N	
75	E2V-E2V-Tel	Test of various beam optics	2	10.11	11.11	4	N	N	
76	CERN	Test of various parts for beam line dosimetry	2	17.11	18.11	2	Y	N	
77	PSI	Test of new SiPM and Plastic detector automated beam scanner	1	18.11	18.11	0.5	N	N	
78	PSI	Beam pre-selection with EASII	1	19.11	19.11	0.5	N	N	
79	ESA / NANOEPIX	Test of focused beams for accelerated DO tests	1	20.11	20.11	1	Y	N	
80	PSI / LIP / EFAC	Test of various components for LHC accelerator	3	1.11	2.11	4	N	N	
81	EASII / ST	Test of various detector types and test structures	2	22.11	23.11	3	Y	N	
82	ESA / ISEL	OBOT / PTO / PTO-2	2	25.11	26.11	4.5	Y	Y	
83	E2V-E2V-Tel	Test of optical sensors for PLATO mission	2	17.11	27.11	1	Y	N	
84	Surrey Satellite Technology	Test of various integrated circuits	1	30.11	30.11	2	Y	N	
85	PSI	Test of various components for LHC accelerator	3	1.11	2.11	4	N	N	
86	Surrey Satellite Technologies	Test of various components for LHC accelerator and SEL effects	1	1.11	3.12	2	N	N	
87	E2V-Tel	Test of optical parts for Eddington	2	3.12	4.12	2	Y	N	
88	PSI / CERN / E2V-Tel	Pre-commissioning of beam test detectors	1	14.12	14.12	1	N	0	
89	PSI	Test of various beam optics	3	1.11	2.11	4	D	D	
90	E2V-Tel	Test of optical sensors for MTG and Flex	2	17.12	18.12	1	Y	N	
91	PSI	test of new pixelated ionization chamber (PROPIX upgrade)	1	19.1	19.1	1	Y	Y	

# Wir schaffen Wissen – heute für morgen

**My thanks go to you**

