qBounce:

Ramsey gravity resonance spectroscopy explained

FOR SOCIETY

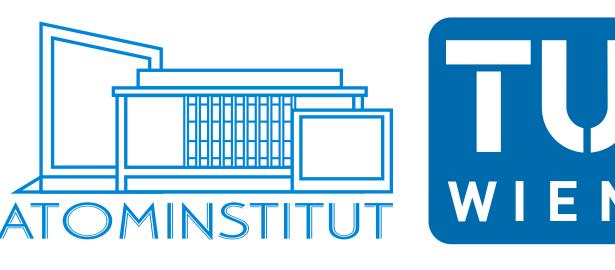




Jakob Micko^{1,2}, J.Bosina^{1,2}, C. Killian^{1,2}, T. Jenke¹, M. Pitschmann², A.N. Ivanov², P. Geltenbort¹, R.I.P. Sedmik², H. Abele²



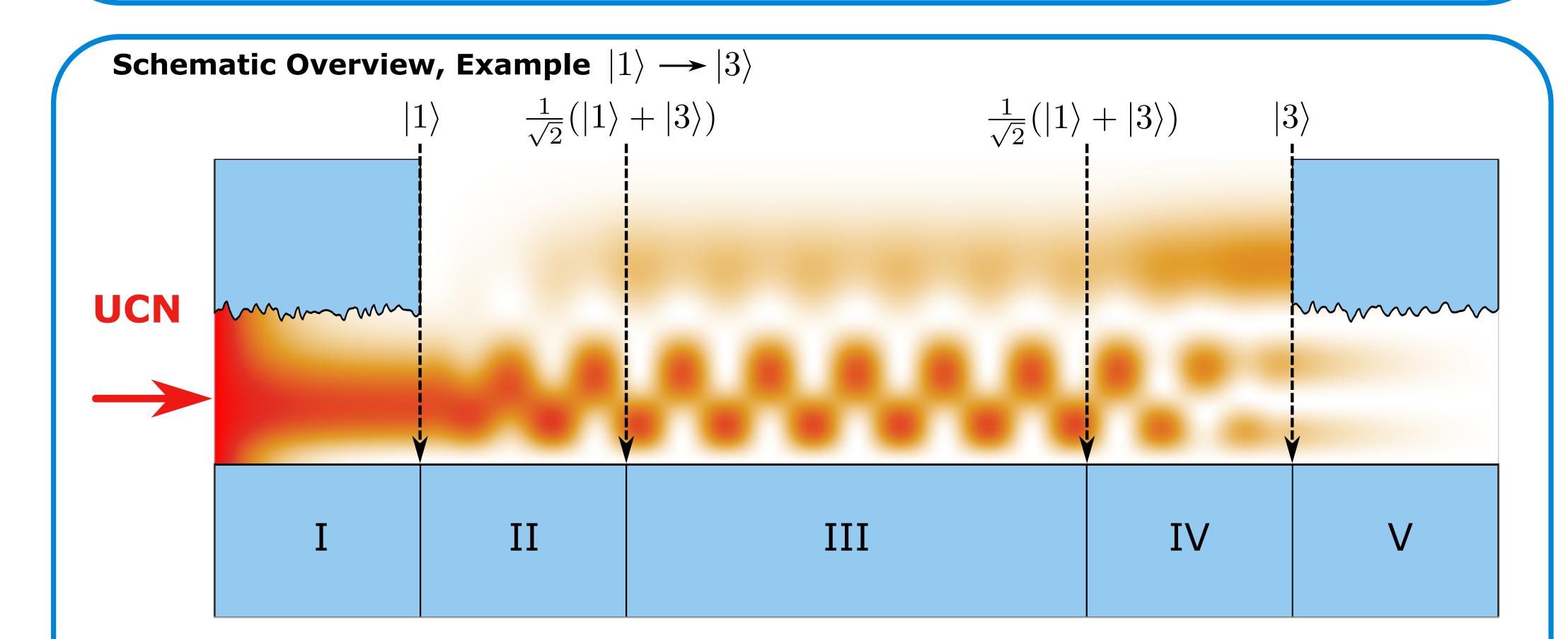
² Atominstitut, Technische Universität Wien, Stadionallee 2, 1020 Vienna, Austria



Abstract

The **qBounce** experiment investigates gravity at small distances. This is done using high precision frequency based spectroscopic methods. Ultracold neutrons (UCNs) form macroscopic bound states above a flat surface in the gravity potential of the Earth, connecting the quantum mechanical neutron wavefunction and gravity. Using this system we developed techniques for Gravity Resonance Spectroscopy (GRS) [1,2]. We realized a proof of principle with Ramsey's method of separated oscillating fields with gravitationally bound UCNs [3]. This method can be used to probe any interaction that couples to the neutron and shifts the eigenenergies of the neutron. Previous iterations of the qBounce experiment implemented GRS in a Rabi configuration and set limits on chameleon dark energy, axion-like dark matter and symmetron dark energy scenarios [4].

The qBounce experiment is located at the UCN source PF2 at the Institut Laue-Langevin (ILL) in Grenoble. After achieving the proof of principle for Ramsey-GRS in 2018 and optimising experimental parameters during the first half of 2019, we are prepared to take data at a projected sensitivity of $5x10^{-16} eV/day$.



Section I/V, Preparation/Selection: UCNs enter the setup from the left and only the first states in the gravity potential are transmitted. Section I is used to prepare the system in the ground state and section V to scatter all excited neutrons out of the system.

Section II/III, Excitations: Mechanical oscillation of the neutron wavefunction mirrors perturb the neutrons. In-phase excitation in sections II and IV leads to a transition from the initial to the target state.

Section III, Propagation: The superposition of states from section II propagates **freely** and the relative phase accumulates according to the energy difference between the states. This phase is sensitive to all pertubations of the gravitational eigenstates and allows an anlysis of the perturbing potential.

Experimental key facts

- high vacuum $p = 3x10^{-5}$ mbar
- µ-Metal magnetic shielding with attenuation factor 100
- experiment leveled relative to local gravitational accelleration ≤ 100 nrad
- relative height alignment of mirrors ≤ 500 nm
- detector (neutron counter tube): background (0.63 \pm 0.03) $\times 10^{-3}$ s-1
- sensitivity:

reached: 2x10⁻¹⁵ eV (Rabi) short term: 7x10⁻¹⁷ eV long term: 5×10⁻²¹ eV (ideal)

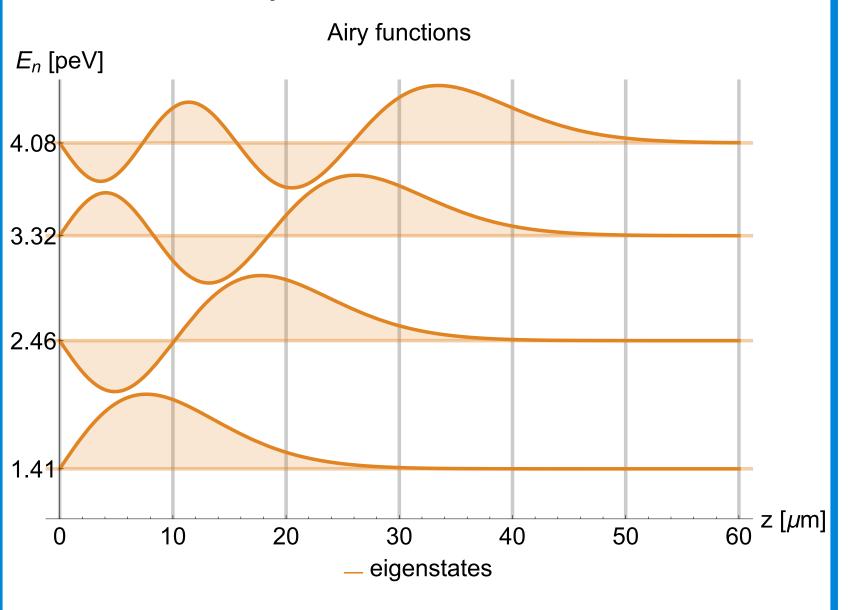
Theory

The neutron in the gravity potential is described by the Schrödinger equation with a linear potential

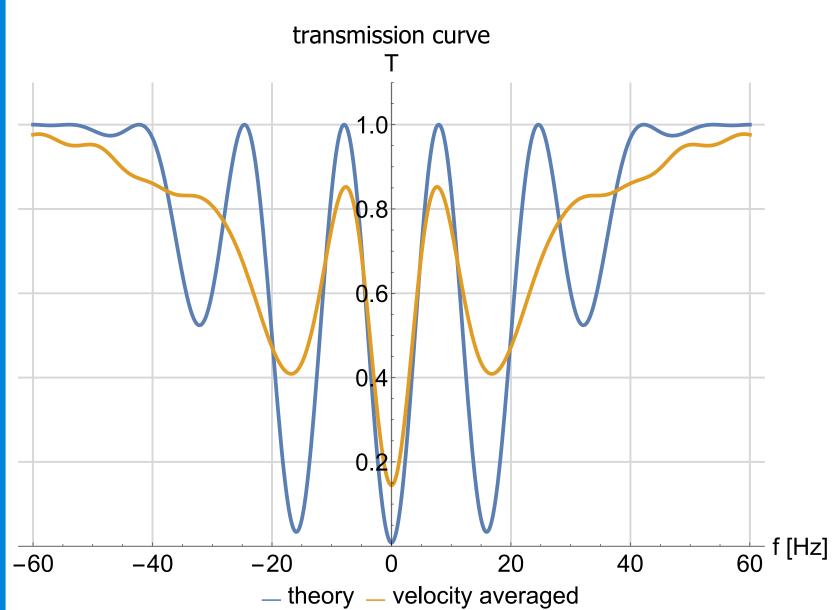
$$E_n \psi_n = \left(-\frac{\hbar^2}{2m_i} \frac{\partial^2}{\partial z^2} + m_g gz \right) \psi_n$$

with eigenenergies E_n .

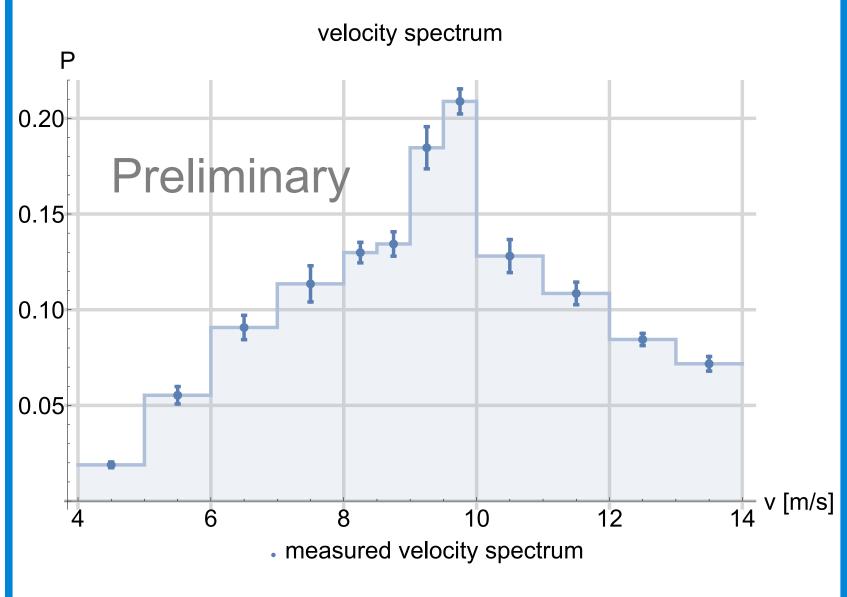
The solutions to this equation are the so called Airy functions.

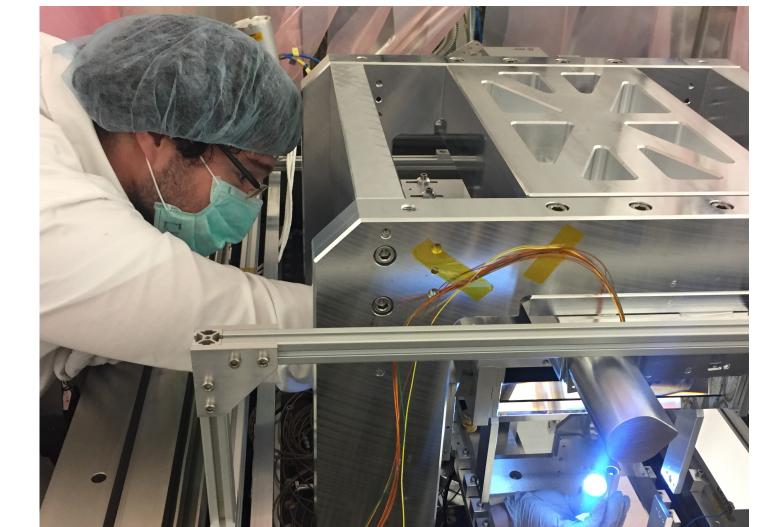


Driving transitions leads to a resonance in the transmission curve.



where the velocity spectrum used for the average was measured in 2019.





- [1] Jenke et al., Nature Physics 7, 468-472 (2011)
- [2] Jenke et al., Phys. Rev. Lett. 112, 151105 (2014)
- [3] Sedmik et al., arXiv:1908.09723[phys.ins-det]
- [4] Cronenberg et al., Nature Physics 14, 1022–1026 (2018)





