

# Searches for Exotic Interactions with Slow Neutrons



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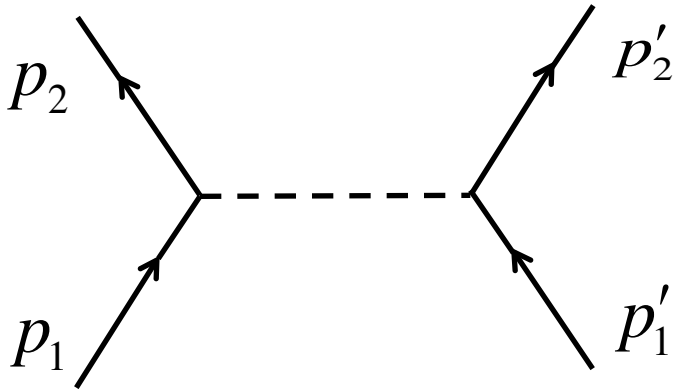
0. (General) Motivations to search for weakly-coupled, long range interactions
1. Searches for new Yukawa interactions at the nanometer scale
2. Searches for dark energy using neutron interferometry
3. Searches for exotic spin dependent interactions of the neutron

Thanks for slides to: H. Kamiya H. Abele, H. Shimizu, G. Pignol,...

Related talks: Wursten, Jaeckel, Abele, Kamiya

Related Posters: Ayres/Rawlik, Jenke, Rechberger, Thalhammer

# Searches for light, weakly interacting particles: complementary to LHC



$$V(\vec{r}) = g^2 \frac{1}{r} e^{-\frac{r}{\lambda}}$$

(Most) high energy physics explores:  $g \sim 1$ ,  $\lambda$  as small as possible

*This work emphasizes a different regime:*

*$g$  small,  $\lambda$  “large” (millimeters-microns) but not infinite*

# New interactions with ranges from millimeters to microns... “Who ordered that?”

1. *Weakly-coupled, long-range interactions are a generic consequence of spontaneously broken continuous symmetries (Goldstone theorem)*
2. *Specific theoretical ideas (axions, extra dimensions for gravity) imply new interactions at  $\sim\text{mm}-\mu\text{m}$  scales*
3. *Dimensional analysis: dark energy- $>100$  microns*

*Experiments should look!*

*Antionadis et al, Comptes Rendus Physique 12, 755-778 (2011)*

*J. Jaeckel and A. Ringwald, [Ann. Rev. Nucl. Part. Sci. 60, 405 \(2010\).](#)*

# Why use slow neutrons to search?

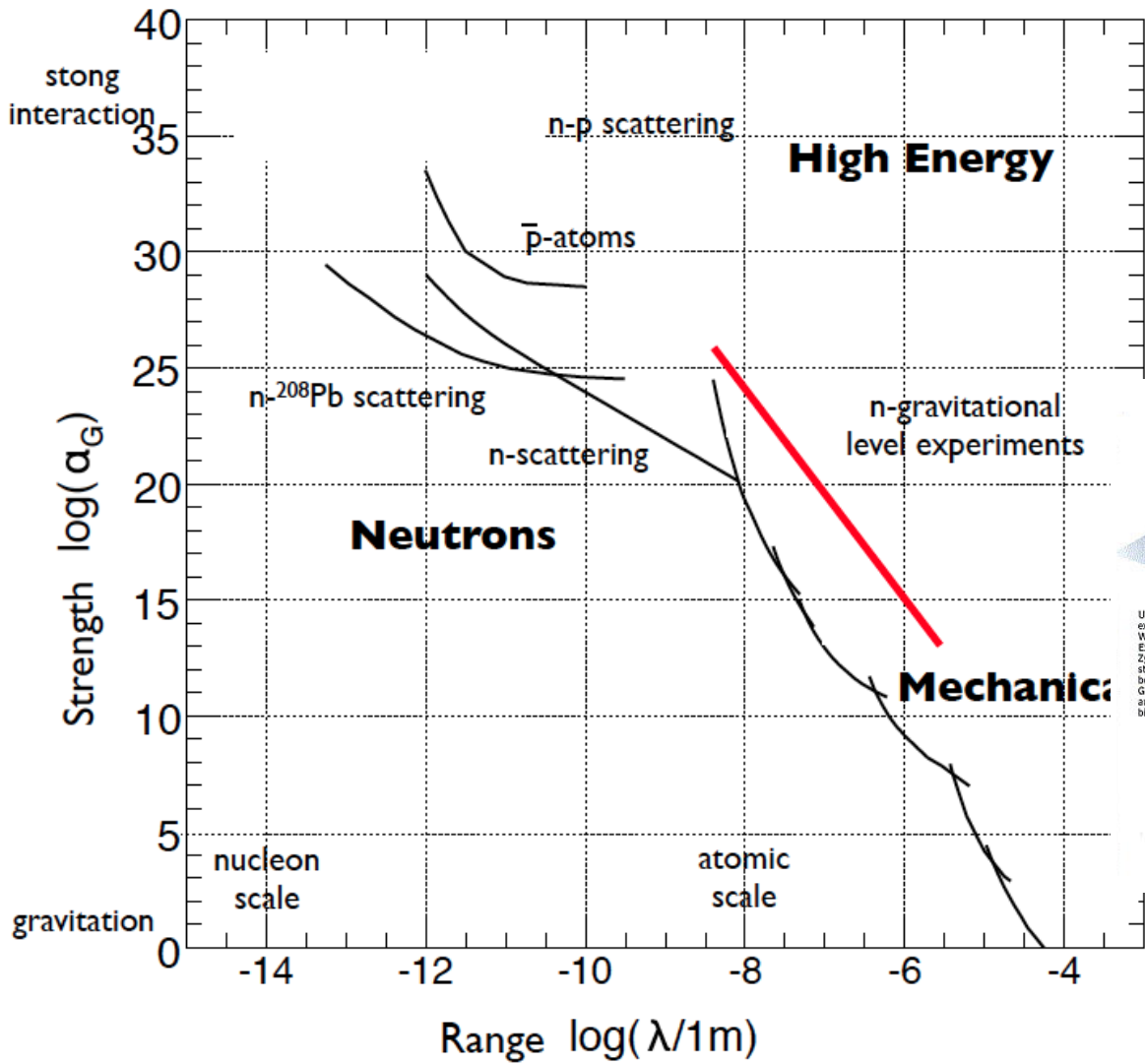
1. *Zero electric charge, small magnetic moment, very small electric polarizability->low "background" from Standard Model interactions*
2. *Deep penetration distance into macroscopic amounts of matter*
3. *Coherent interactions with matter->phase sensitive measurements possible*
4. *High neutron polarization (>~99%) routine for slow neutrons  
->important in searching for spin-dependent interactions*
4. *A broad set of facilities for experimental work is available*

*J. Nico and W. M. Snow, Annual Reviews of Nuclear and Particle Science **55**, 27-69 (2005).*

*H. Abele, Progress in Particle and Nuclear Physics **60**, 1-81 (2008).*

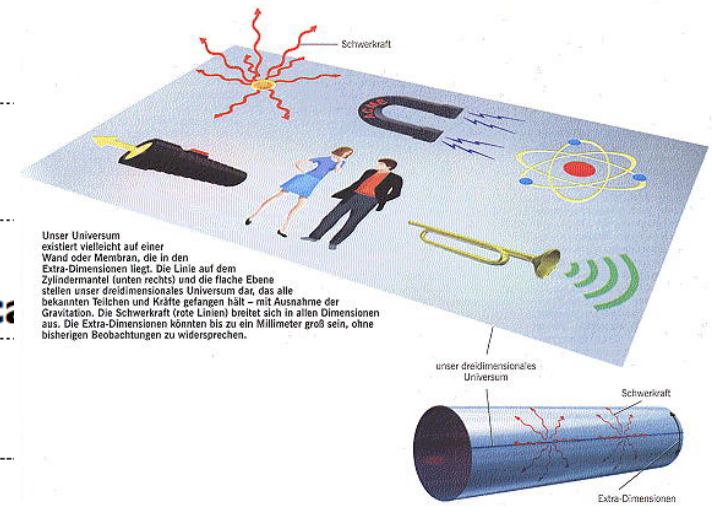
*D. Dubbers and M. Schmidt, Reviews of Modern Physics (2011).*

# Searches for new Yukawa interactions



Neutron measurements are the most sensitive from atomic to subnuclear scales

Continued work at atomic scales important to probe the idea of extra spacetime dimensions (Murata/Tanaka arXiv:1408.3588)



Recent theory to explain muon g-2 and p radius puzzle gives new Yukawa interaction with a boson mass as low as 100 keV (Liu/Miller PRL)

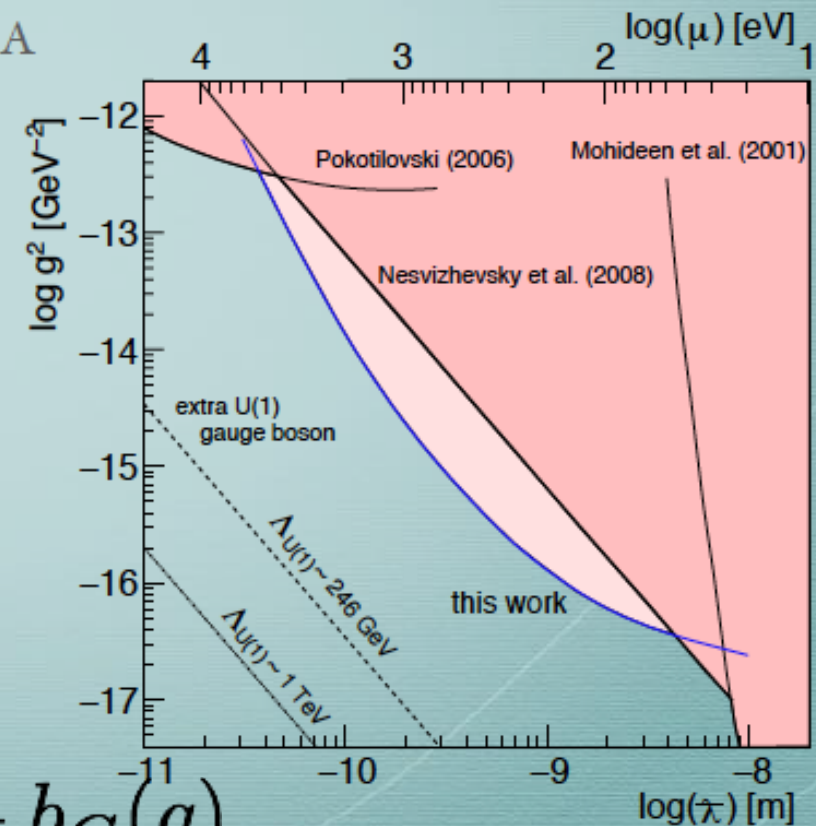
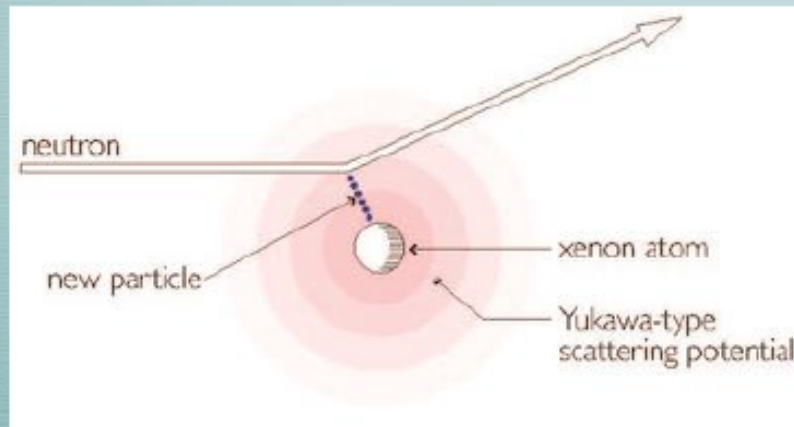
$$V(r) = G \frac{m_1 \cdot m_2}{r} (1 + \alpha \cdot e^{-r/\lambda})$$

# Search for new gravity-like interactions and test of the equivalence principle using slow neutrons

Yoshio Kamiya, Koji Yamada, Kenta Uchida, Yoshihiro Sasayama, Keita Itagaki, Misato Tani, Sachio Komamiya, and Guinyun Kim

*The Univ. of Tokyo / Kyngpook Nat. Univ.*

We report on a new constraint on gravity-like interactions obtained by measuring the angular distribution of 5 Å neutrons scattering off atomic Xe gas.



$$b_c(q) = (b_c - b_e Z(1 - f(q))) + b_G(q)$$

*Neutron-Xenon Gas Scattering Search for Yukawa Interaction at J-PARC Spallation Neutron Source* H. M. Shimizu, K. Hirota, M. Kitaguchi, C. Haddock, W. M. Snow, K. Mishima, T. Yoshioka, T. Ino, S. Matsumoto, T. Shima

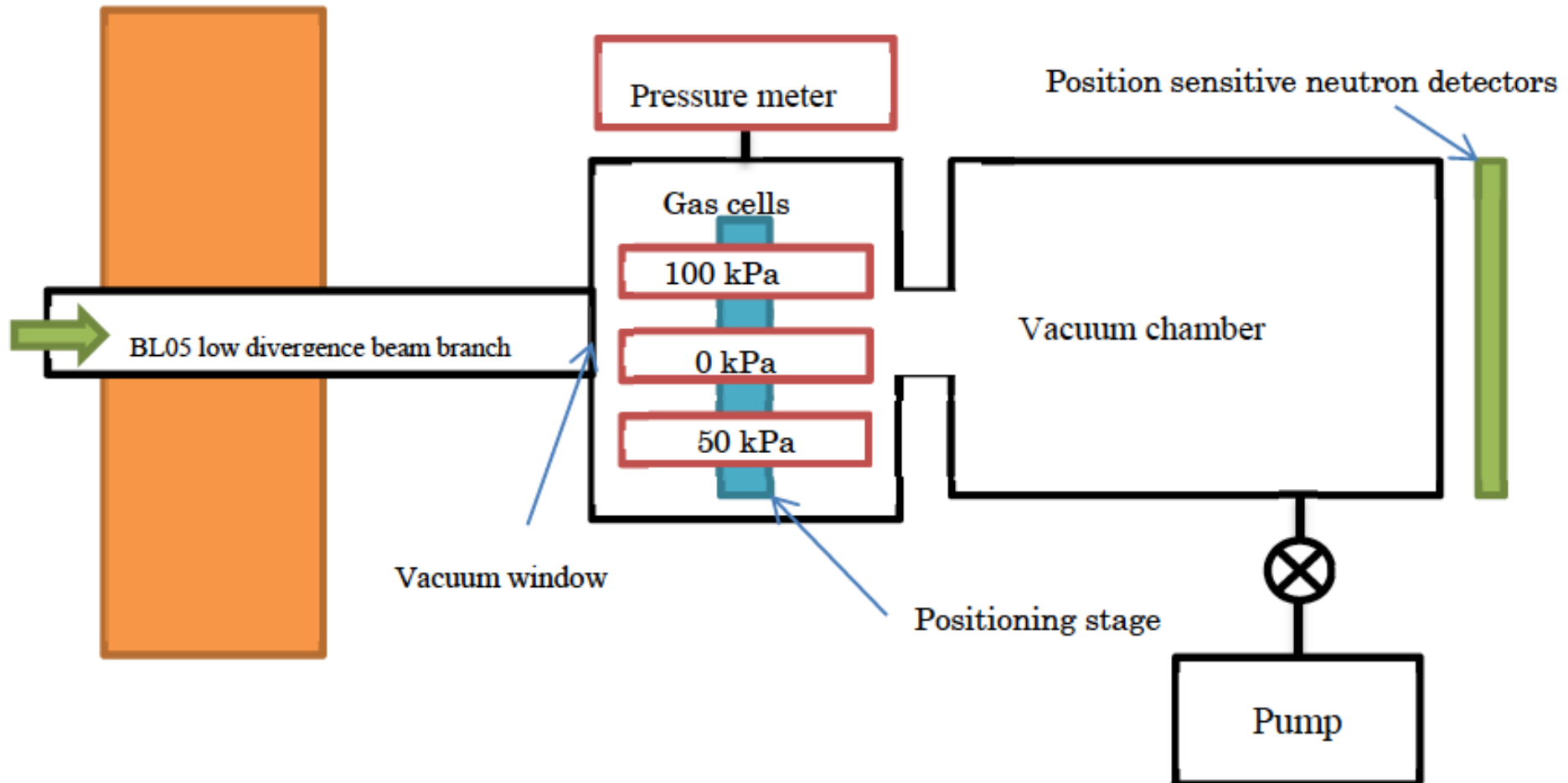


Figure.4. Experimental apparatus.

*Use angular distribution on n-Xe scattering to search for Yukawa Interaction at very short ranges Experiment in progress at JPARC*

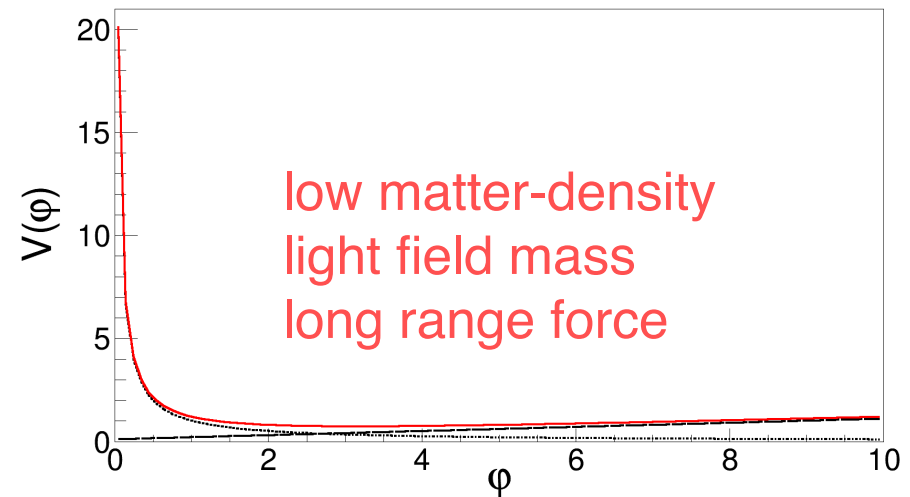
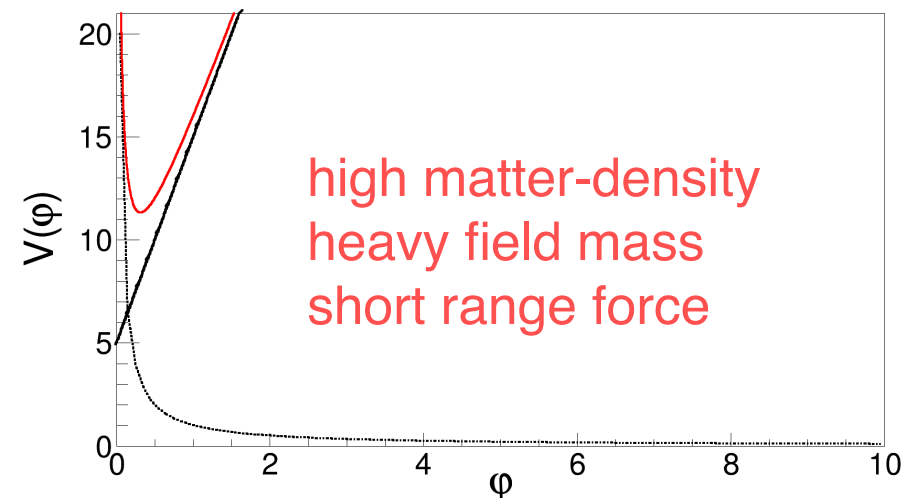
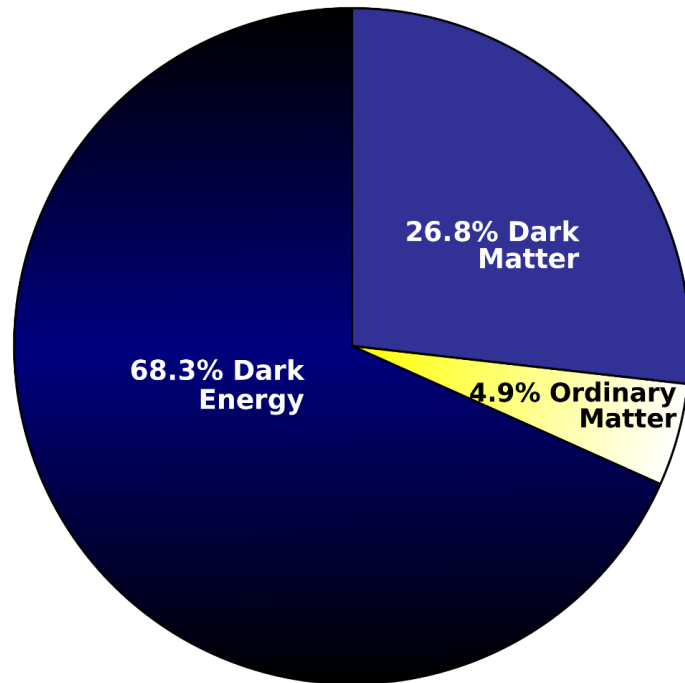


# Chameleon scalar field

J. Khoury and A. Weltman,  
PRL **93**, 044026 (2004)

One proposed form of Dark Energy: a new scalar field (quintessence)

- Chameleon field is engineered to be:
- Massless on cosmological scales
  - Massive field in laboratory





# Chameleon field

Effective potential:

$$V(\varphi)_{eff} = V(\varphi) + \rho A(\varphi)$$

Runaway potential

Coupling to matter

*Ratio of pressure to energy density evolves to  $w=-1$  dynamically*

$$w_q = \frac{p_q}{\rho_q} = \frac{\frac{1}{2}\dot{Q}^2 - V(Q)}{\frac{1}{2}\dot{Q}^2 + V(Q)}$$

$$A(\varphi) = \sum e^{\beta_i \varphi / M_{PI}} \approx \beta \varphi / M_{PI}$$

*Parameters:*

$$M_{PI} = 2.44 \times 10^{18} \text{ GeV}$$

$$\Lambda = 2.4 \text{ meV}$$

*Implemented by a Ratra/Peebles potential*

$$V(\varphi) = \Lambda^4 + \frac{\Lambda^{n+4}}{\varphi^n}$$

# Chameleon field on neutron interferometer

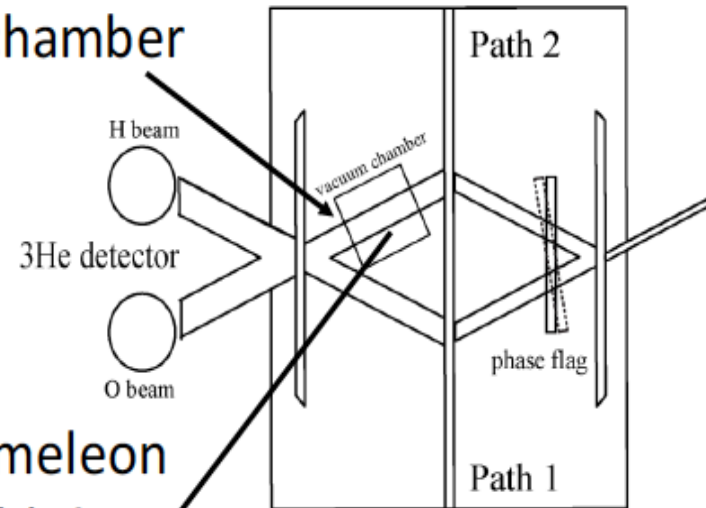
The phase shift will be

$$\Delta\Phi = \int dx \frac{m^2 \beta \varphi(x)}{M_{Pl} k}$$

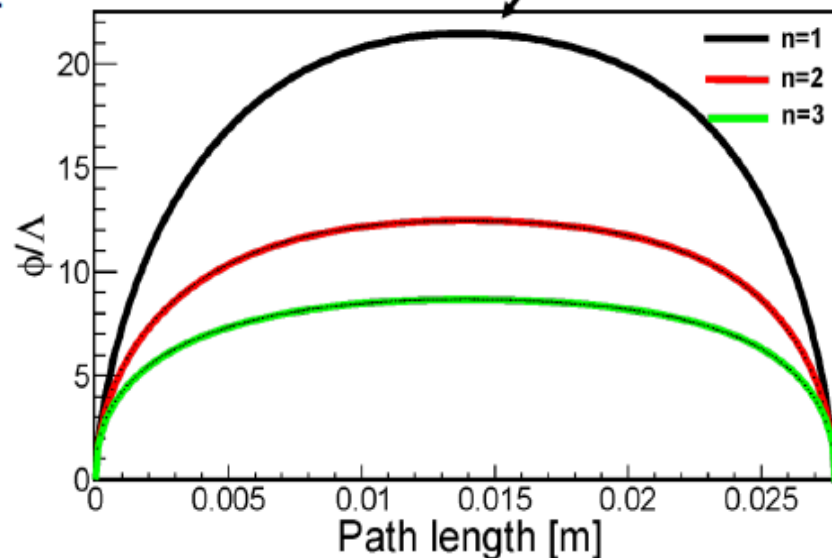
$m$  is the mass of neutron

$k$  is the magnitude of neutron wave-vector

vacuum chamber



Chameleon bubble!



*Vary gas pressure to make chameleon field appear and disappear. A few mbar of helium is enough to kill the field*

*Look for phase shift on the neutron.  $\beta$  is neutron-matter coupling,  $n$  is Ratra-Peebles index*

# INDEX Collaboration

*K. Li, W. M. Snow, CEEM, Indiana University*

*M. Arif, M. G. Huber, NIST Center for Neutron Research*

*D. G. Cory, D. Pushin, Institute for Quantum Computing,  
University of Waterloo*

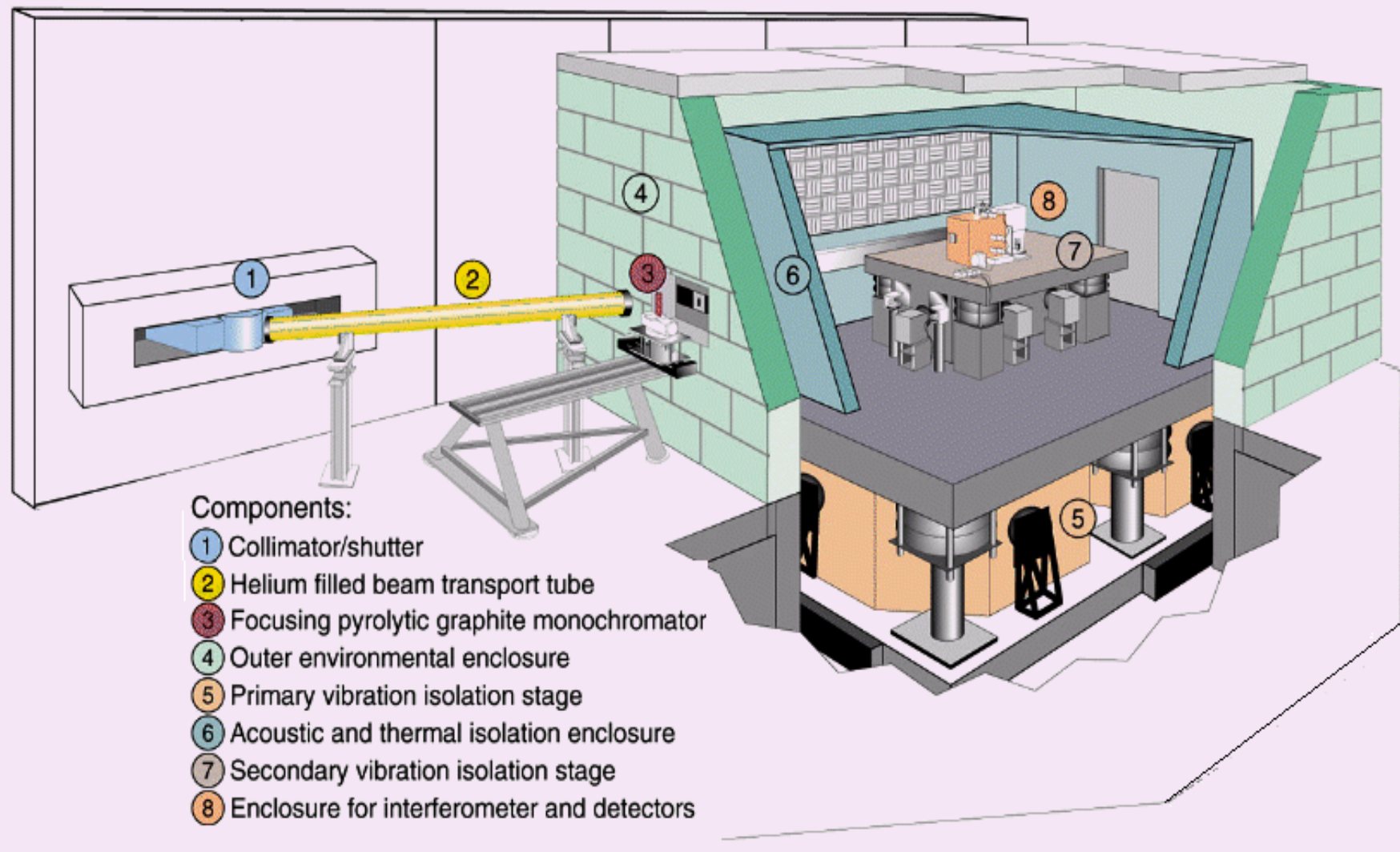
*R. Haun, C. B. Shahi, Tulane University*

*B. Heacock, V. Skavysh, A. R. Young, North Carolina State  
University*

*J. Nsofini, P. Saggu, D. Sarenac, University of Waterloo*

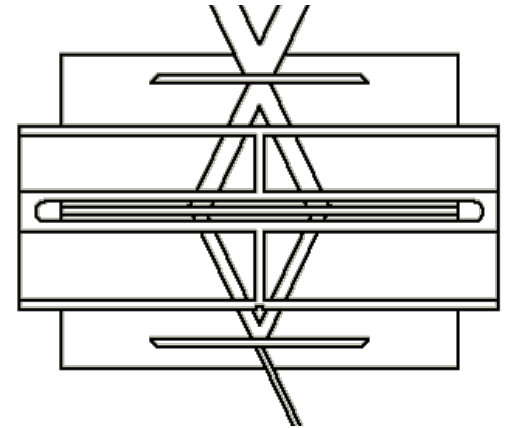
# NIST

## Neutron Interferometer and Optics Facility

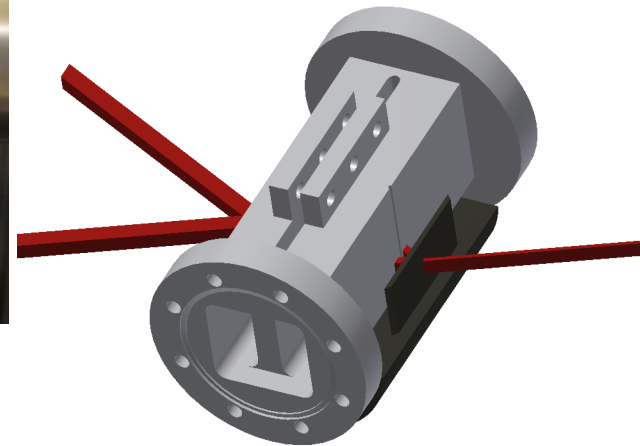




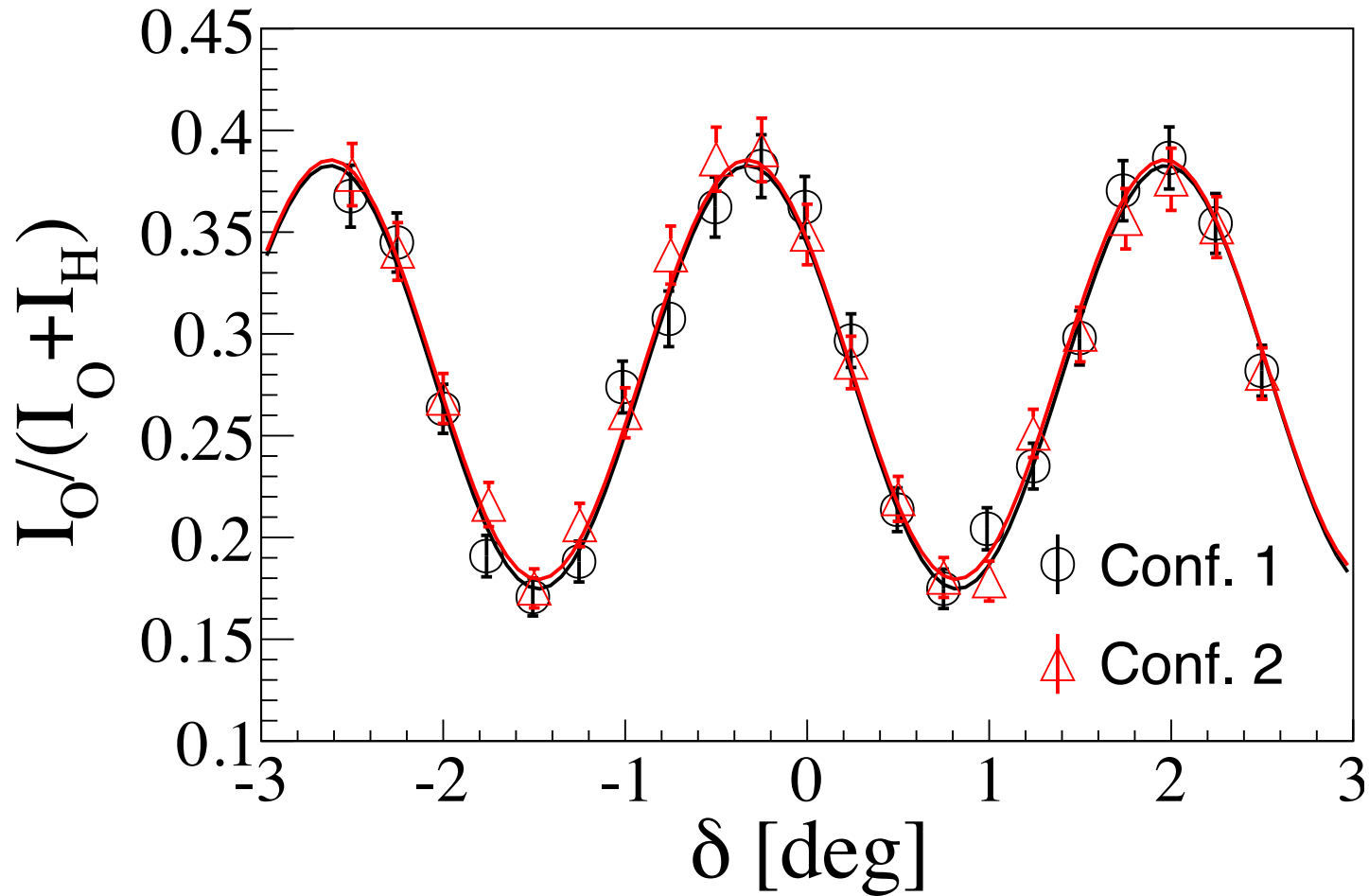
# Vacuum cell for Chameleon Experiment



*Allows addition of gas at different pressures on each subbeam*

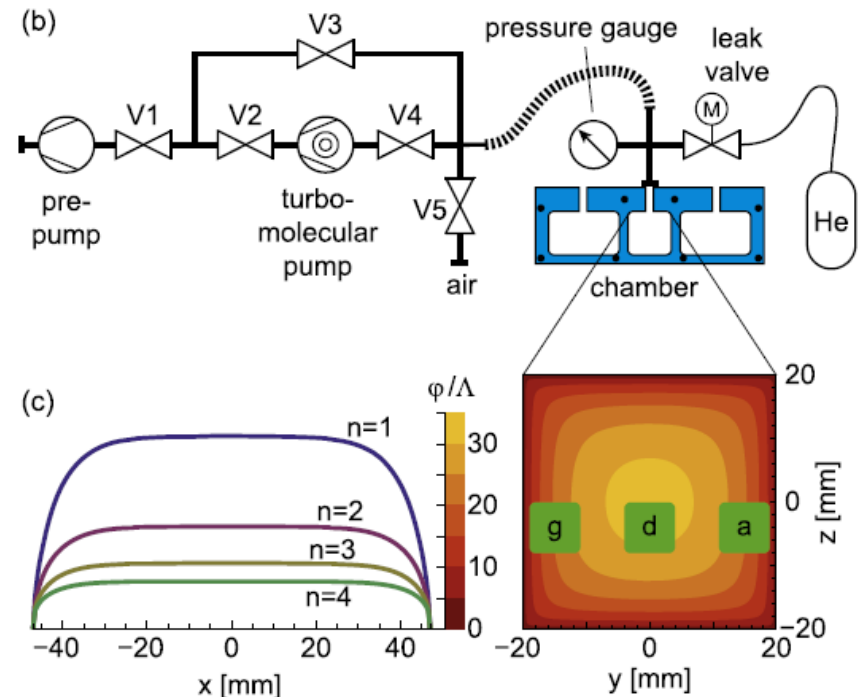
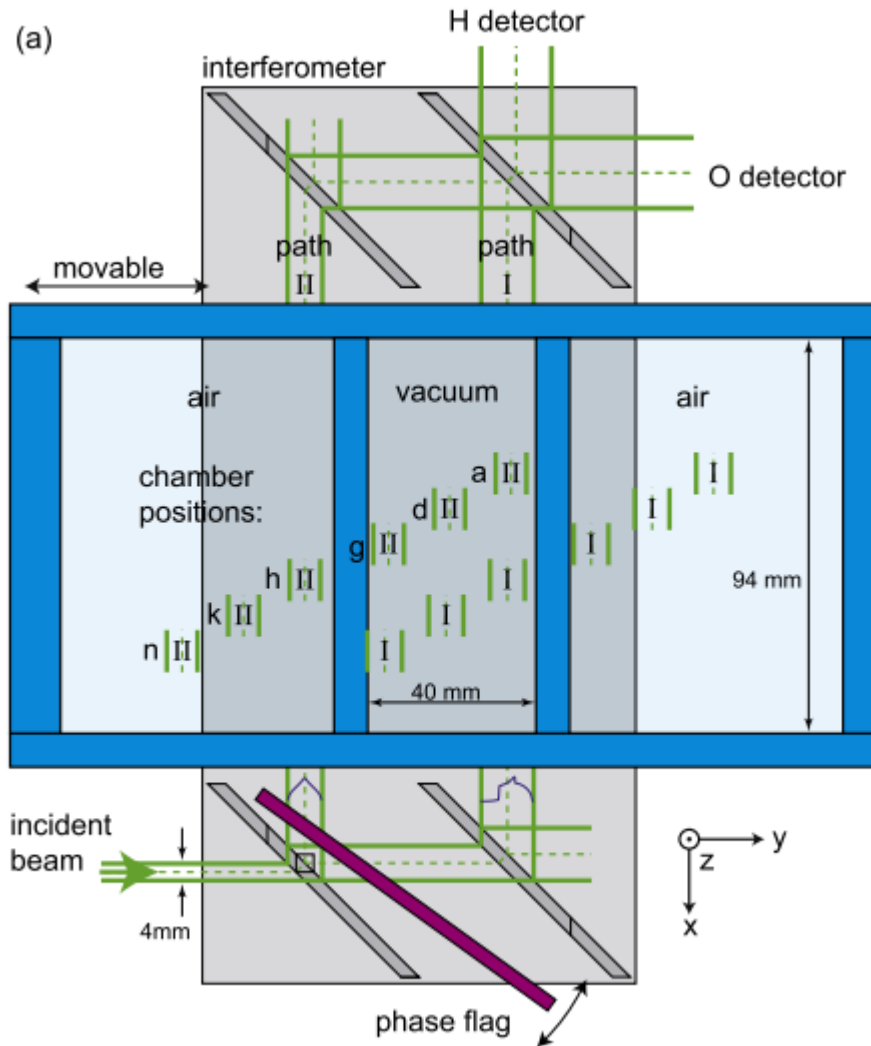


# Some Experimental Data from NIST



*No evidence for chameleon dark energy*

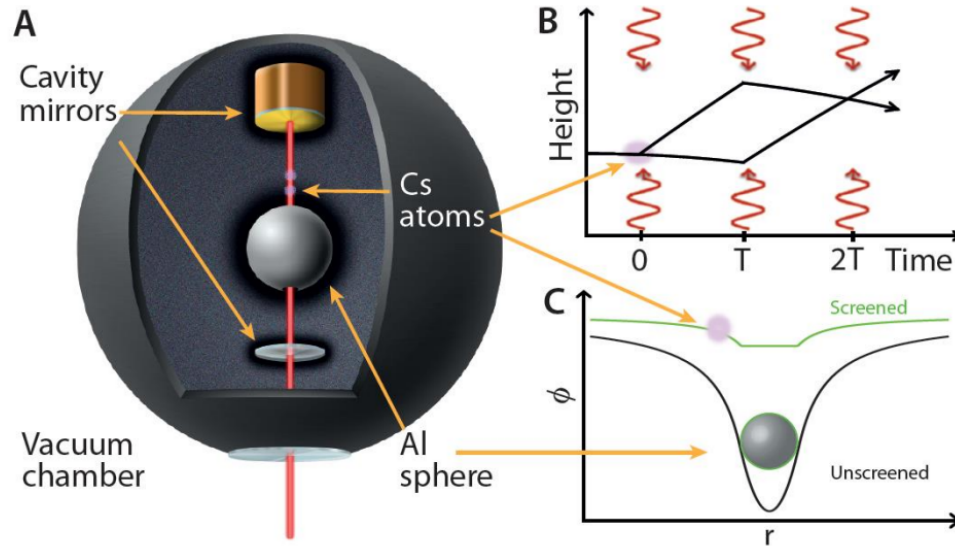
# Neutron interferometer experiment at ILL



*Physics Letters B 743 (2015)*  
*See poster by T. Jenke*



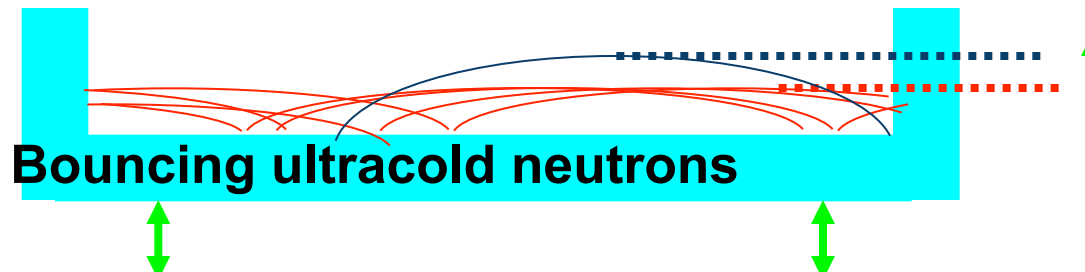
# Other recent chameleon experiments:



**Fig. 1. Screened fields in our experiment** (A) The vacuum chamber (radius 5 cm, pressure  $\sim 6 \times 10^{-10}$  Torr, mostly hydrogen) holds a pair of mirrors forming a Fabry-Perot cavity and the aluminum (Al) source sphere. Laser beams pass a 1.5-mm radius hole in the  $r_s = 9.5$ -mm radius sphere. A Mach-Zehnder interferometer is formed using cold cesium atoms at an effective distance of 8.8 mm from the sphere surface from a magneto-optical trap (not shown). (B): Photons in three flashes of laser radiation resonant in the cavity impart momentum to the atoms, directing each atomic matter wave on two paths. (C) Potential generated by a macroscopic sphere as function of distance from the center.

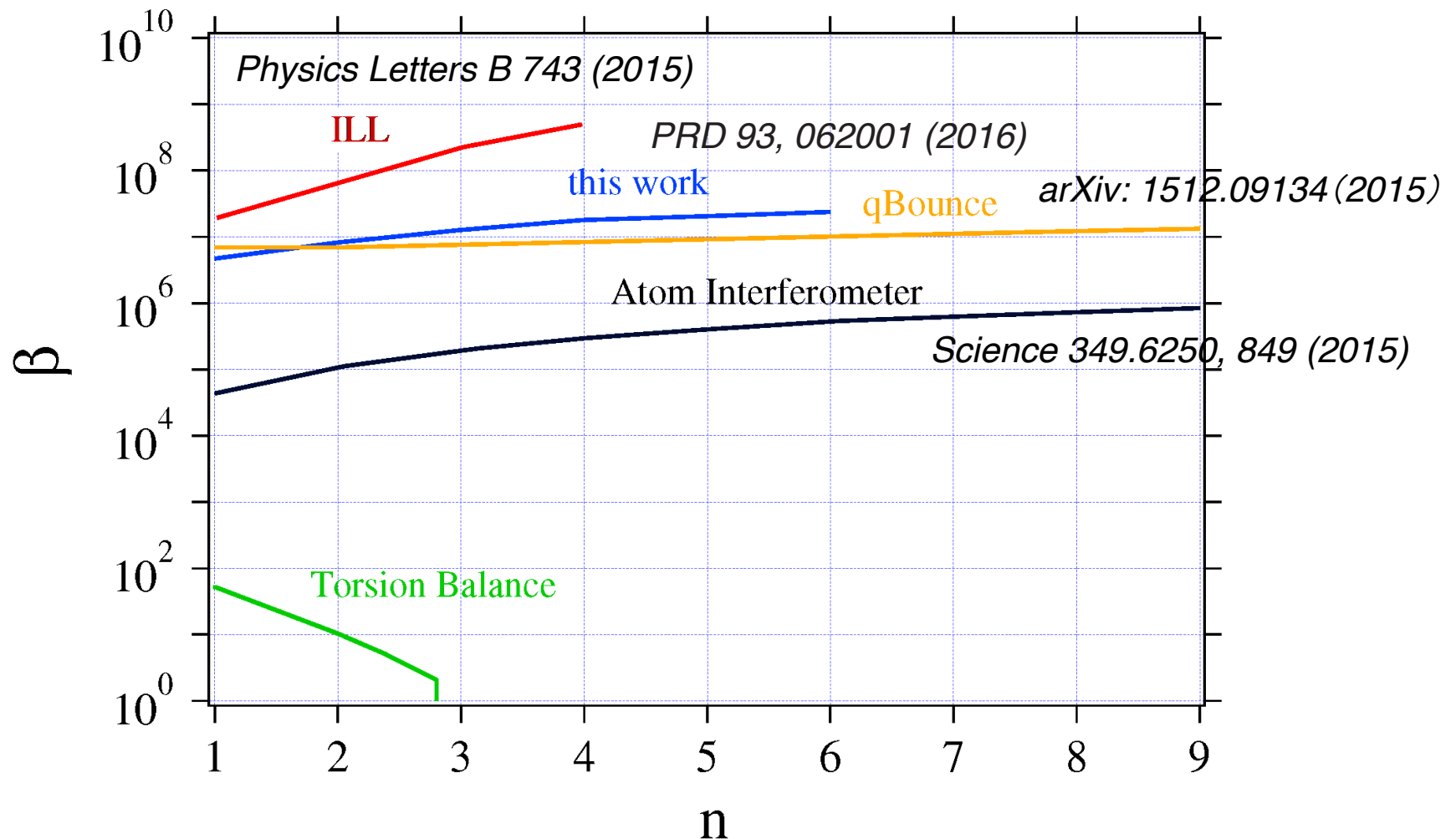
Atom interferometer  
(see H. Mueller talk)

Gravitational spectroscopy  
with ultracold neutrons  
(see H. Abele talk,  
posters by T. Rechberger,  
M. Thalhammer )



**Bouncing ultracold neutrons**

# Constraints in Chameleon-Matter coupling/ Ratra-Peebles index space



# Spin-dependent macroscopic interactions mediated by light bosons: general classification

$$\mathcal{O}_1 = 1 ,$$

$$\mathcal{O}_2 = \vec{\sigma} \cdot \vec{\sigma}' ,$$

$$\mathcal{O}_3 = \frac{1}{m^2} (\vec{\sigma} \cdot \vec{q}) (\vec{\sigma}' \cdot \vec{q}) ,$$

$$\mathcal{O}_{4,5} = \frac{i}{2m^2} (\vec{\sigma} \pm \vec{\sigma}') \cdot (\vec{P} \times \vec{q}) ,$$

$$\mathcal{O}_{6,7} = \frac{i}{2m^2} \left[ (\vec{\sigma} \cdot \vec{P}) (\vec{\sigma}' \cdot \vec{q}) \pm (\vec{\sigma} \cdot \vec{q}) (\vec{\sigma}' \cdot \vec{P}) \right] ,$$

$$\mathcal{O}_8 = \frac{1}{m^2} (\vec{\sigma} \cdot \vec{P}) (\vec{\sigma}' \cdot \vec{P}) .$$

$$\mathcal{O}_{9,10} = \frac{i}{2m} (\vec{\sigma} \pm \vec{\sigma}') \cdot \vec{q} ,$$

$$\mathcal{O}_{11} = \frac{i}{m} (\vec{\sigma} \times \vec{\sigma}') \cdot \vec{q} ,$$

$$\mathcal{O}_{12,13} = \frac{1}{2m} (\vec{\sigma} \pm \vec{\sigma}') \cdot \vec{P} ,$$

$$\mathcal{O}_{14} = \frac{1}{m} (\vec{\sigma} \times \vec{\sigma}') \cdot \vec{P} ,$$

$$\mathcal{O}_{15} = \frac{1}{2m^3} \left\{ [\vec{\sigma} \cdot (\vec{P} \times \vec{q})] (\vec{\sigma}' \cdot \vec{q}) + (\vec{\sigma} \cdot \vec{q}) [\vec{\sigma}' \cdot (\vec{P} \times \vec{q})] \right\}$$

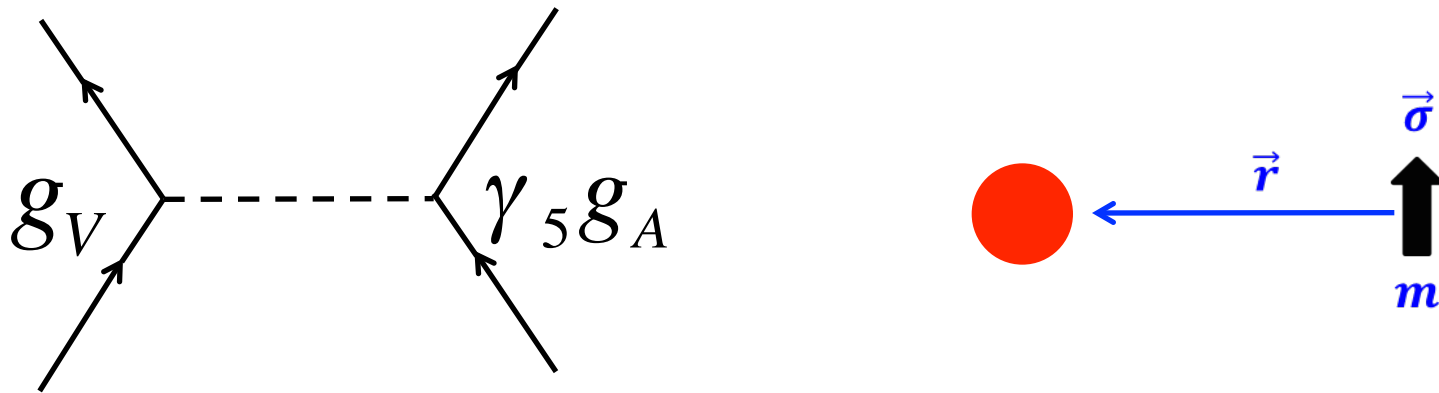
$$\mathcal{O}_{16} = \frac{i}{2m^3} \left\{ [\vec{\sigma} \cdot (\vec{P} \times \vec{q})] (\vec{\sigma}' \cdot \vec{P}) + (\vec{\sigma} \cdot \vec{P}) [\vec{\sigma}' \cdot (\vec{P} \times \vec{q})] \right\} .$$

- 16 independent scalars can be formed: 8 P-even, 8 P-odd
- 15/16 depend on spin
- Traditional “fifth force” searches constrain  $\mathcal{O}_1$

*B. Dobrescu and I. Mocioiu, J. High Energy Phys. 11,005 (2006)*

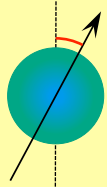
# *Example of a nonstandard P-odd interaction from spin 1 boson exchange:*

*[Dobrescu/Mocioiu 06, general construction of interaction between nonrelativistic fermions]*

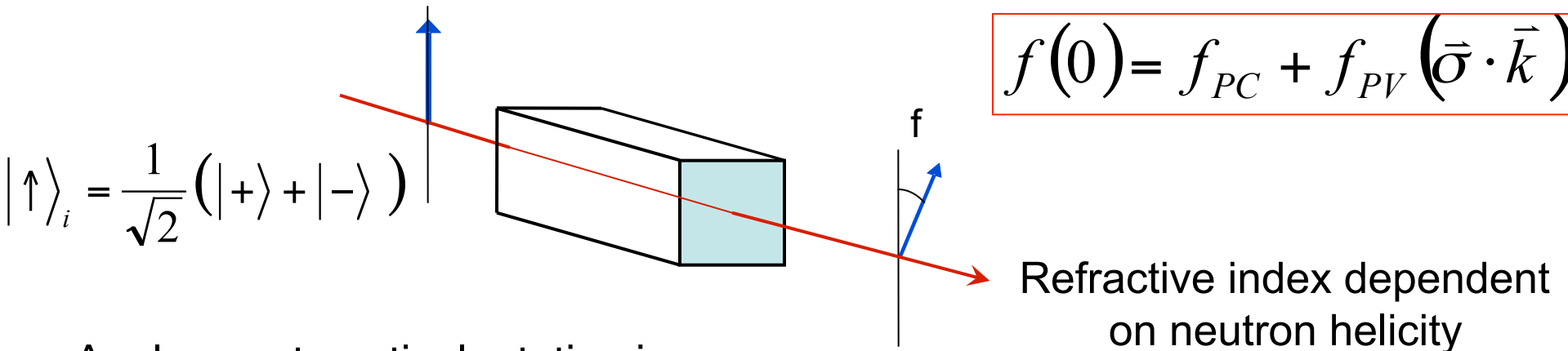


$$V(\vec{\sigma}, \vec{r}, \vec{v}) = \frac{\hbar}{8\pi m c^2} g_A g_V \vec{\sigma} \cdot \vec{v} \frac{1}{r} e^{-\frac{r}{\lambda}}$$

- Induces an interaction between polarized and unpolarized matter
- Violates  $P$  symmetry
- Not very well constrained over “mesoscopic” ranges (millimeters to microns)
- Best investigated using a beam of polarized particles



# Parity-odd Neutron Spin Rotation



◆ Analogous to optical rotation in an “handed” medium.

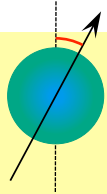
◆ Transversely-polarized neutrons corkscrew from any parity-odd interaction

◆ **PV Spin Angle** is independent of incident neutron energy in cold neutron regime,

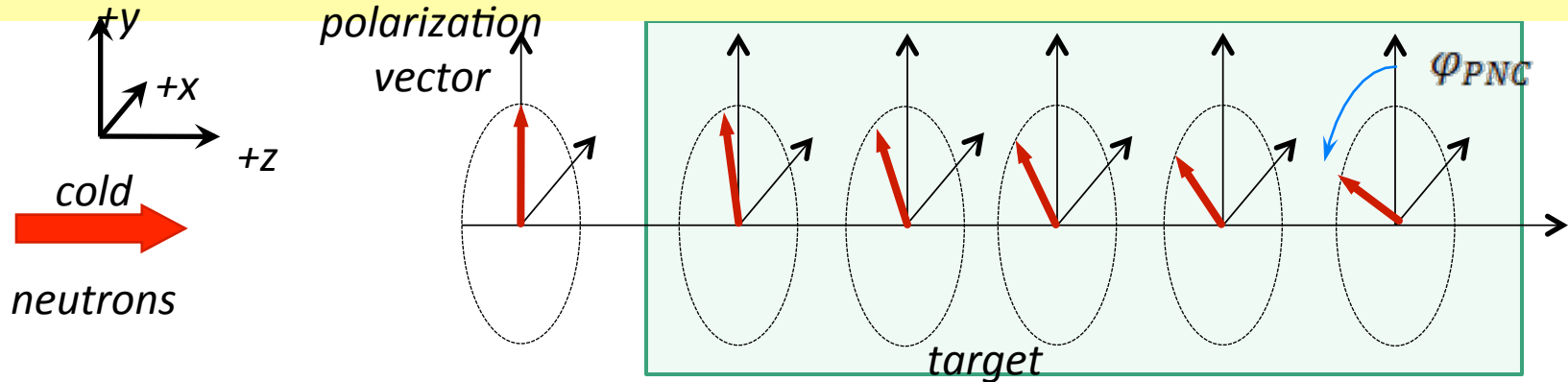
◆  $d\phi_{PV}/dx \sim 10^{-6}$  rad/m sensitivity achieved so far

$$\frac{1}{\sqrt{2}} \left( e^{-i(\phi_{PC} + \phi_{PV})} |z\rangle + e^{-i(\phi_{PC} - \phi_{PV})} |-z\rangle \right)$$

$$\varphi_{PV} = \phi_+ - \phi_- = 2\varphi_{PV} = 4\pi l \rho f_{PV}$$



*Parity-odd interaction of neutron with matter will produce neutron spin rotation:*



$$f(0) = f_{strong} + f_{P-odd} (\vec{\sigma} \cdot \vec{p})$$

*Forward scattering amplitude of neutron in matter sensitive to all neutron-matter interactions*

$$f_{P-odd} = g_A g_V \lambda^2$$

$$\phi_{\pm} = \phi_{strong} \pm \phi_{P-odd}$$

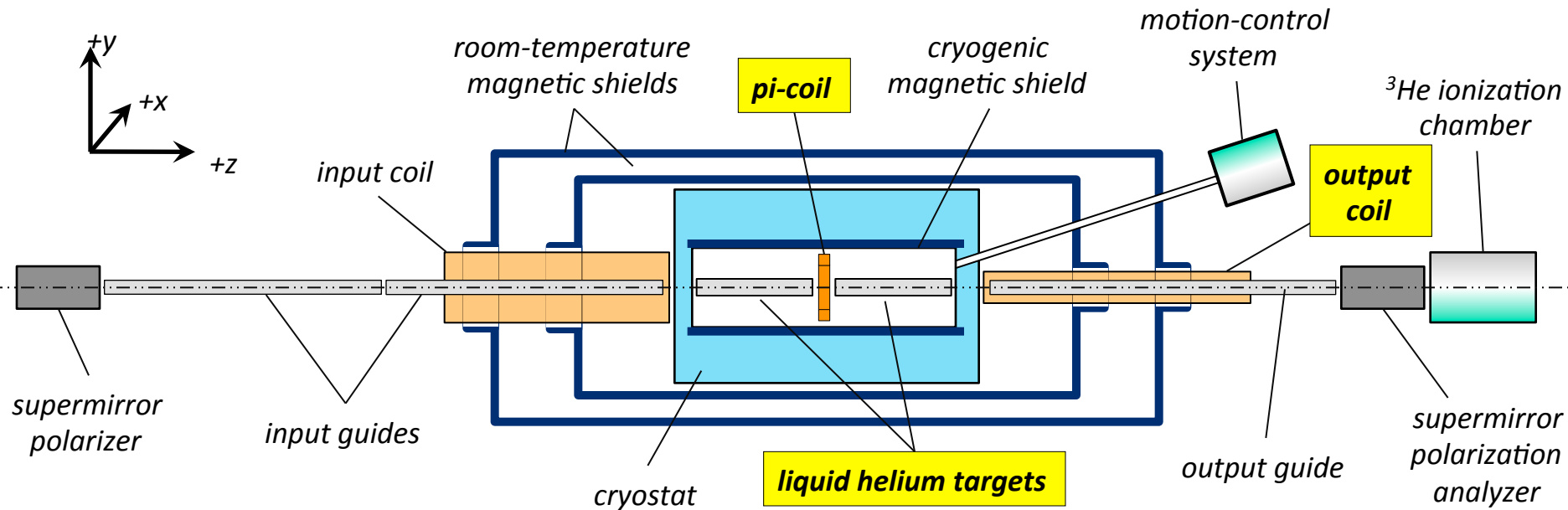
Parity-odd interaction gives helicity-dependent phase shift and therefore rotation of plane of polarization vector

$$\frac{d\phi_{P-odd}}{dL} = 4g_A g_V \rho \lambda^2$$

*An upper bound on  $f_{P-odd}$  places a constraint on possible new P-odd interactions between neutrons and matter over a broad set of distance scales*

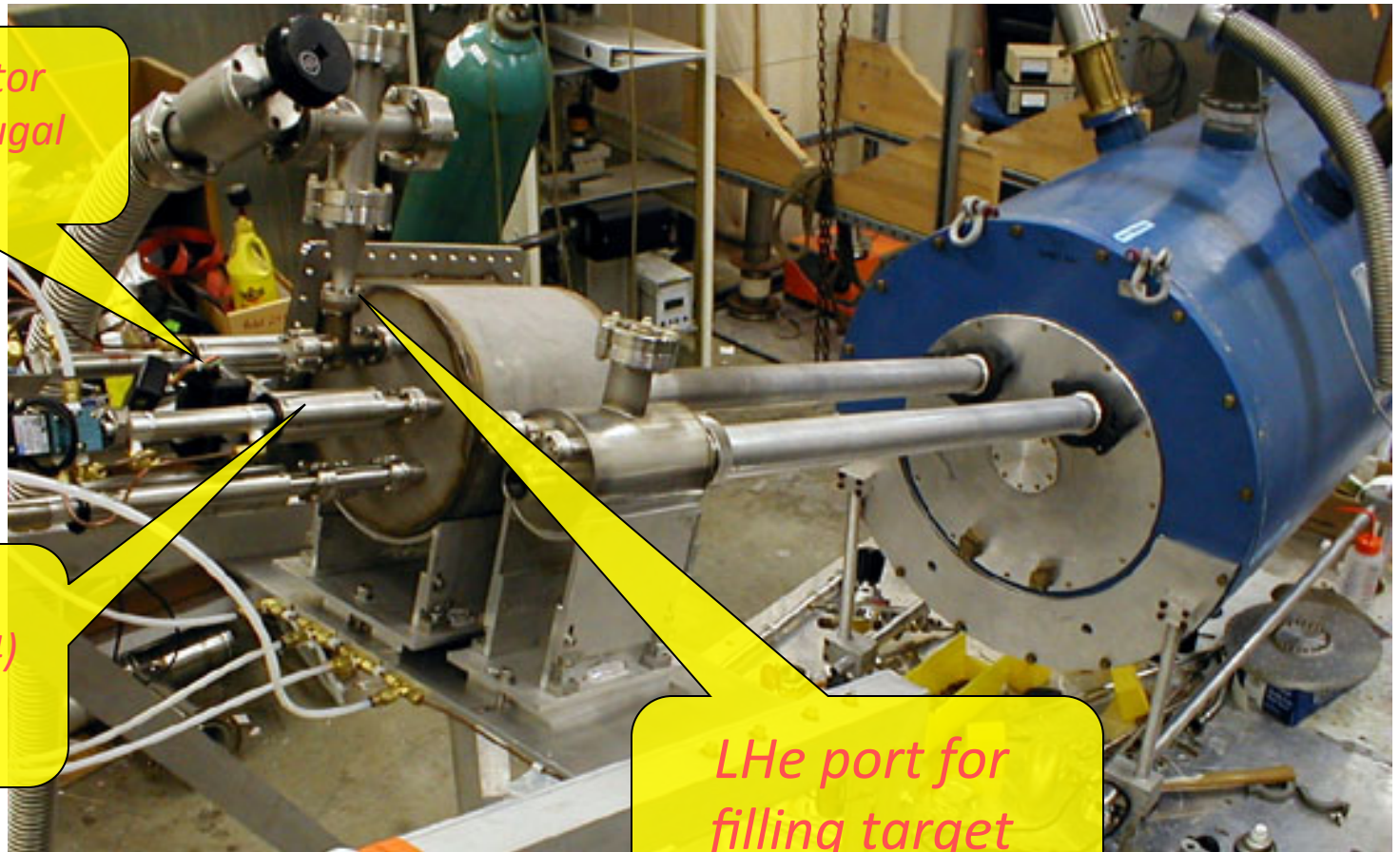
# Neutron Spin Rotation in Liquid Helium

*Apparatus measures the horizontal component of neutron spin generated in the liquid target starting from a vertically-polarized beam*





# Liquid Helium Cryostat and Motion Control



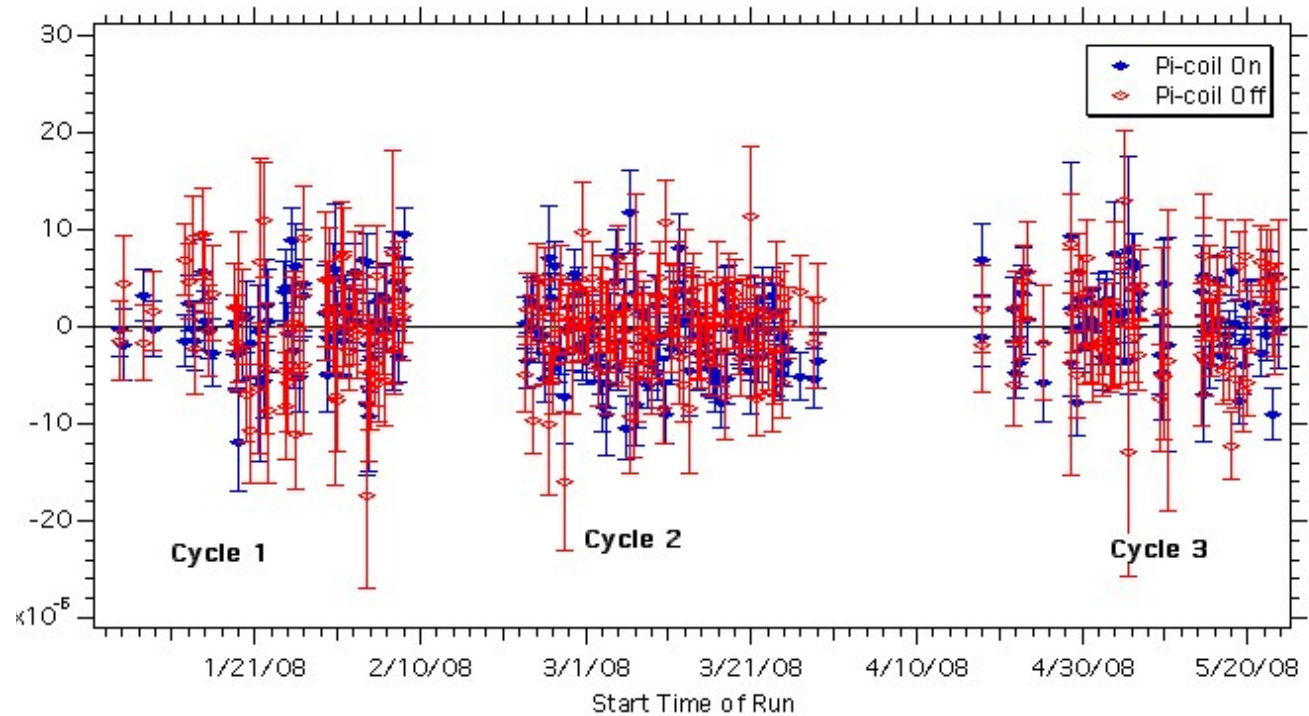
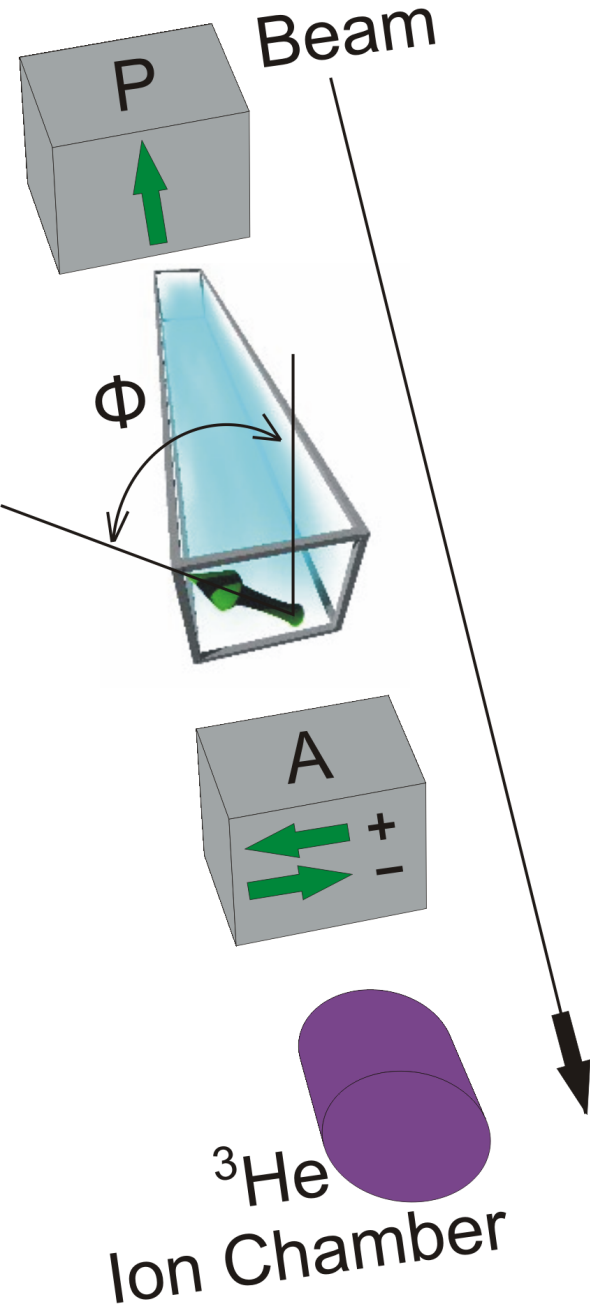
*Stepper motor  
turns centrifugal  
pump*

*Pneumatic  
actuators (x4)  
raise/lower  
drainpipes*

*LHe port for  
filling target  
helium bath*

- Nonmagnetic movement of liquid helium.
- Cryogenic target of 4K helium, volume~10 liters

# Neutron Spin Rotation in $n+4\text{He}$

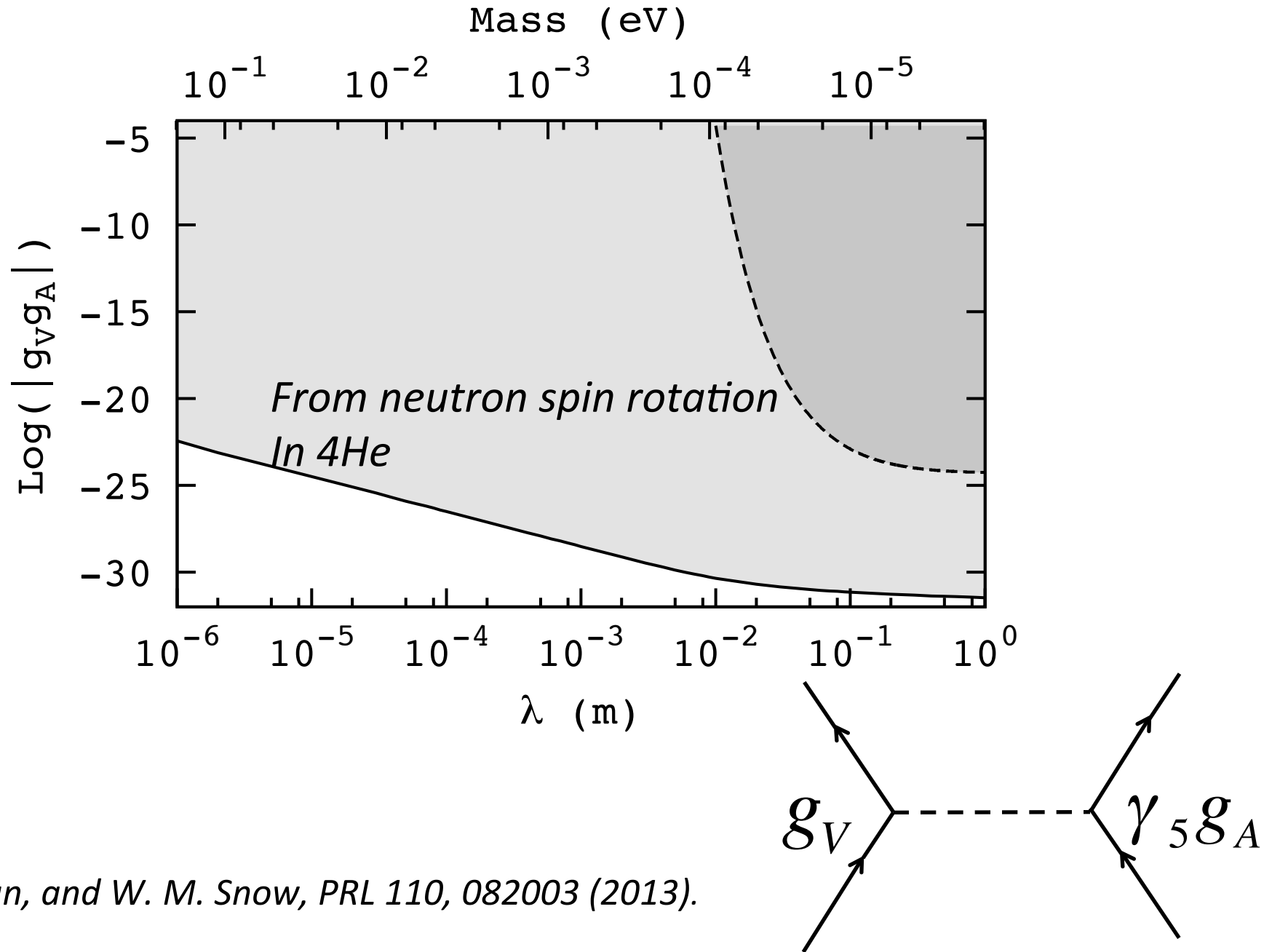


$$\phi_{\text{PNC}} = [+1.7 \pm 9.1 (\text{stat}) \pm 1.4 (\text{sys})] \times 10^{-7} \text{ rad/m}$$

W. M. Snow et al., *Phys. Rev. C* **83**, 022501(R) (2011).

Result analyzed to constrain short-range gravitational torsion: R. Lehnert, H. Yan, W. M. Snow, *Phys. Lett* **B730**, 353 (2014), **B744**, 415 (2015), [arXiv:1311.0467](https://arxiv.org/abs/1311.0467)

# Constraints on exotic V-A interactions



*H. Yan, and W. M. Snow, PRL 110, 082003 (2013).*

# More Constraints on exotic V-A interactions

Searching for New Spin-Velocity Dependent Interactions by Spin Relaxation of Polarized  $^3\text{He}$  Gas

Y.Zhang,<sup>1,2</sup> G.A.Sun,<sup>1</sup> S.M.Peng,<sup>3</sup> C.Fu,<sup>4</sup> Hao Guo,<sup>5</sup> B.Q.Liu,<sup>1</sup> and H.Y.Yan<sup>1,\*</sup>

<sup>1</sup>Key Laboratory of Neutron Physics, Institute of Nuclear Physics and Chemistry, CAEP, Mianyang, Sichuan, 621900, China

<sup>2</sup>School of Nuclear Science and Technology, University of Science and Technology of China, Hefei, 230026, China

<sup>3</sup>Institute of Nuclear Physics and Chemistry, CAEP, Mianyang, Sichuan, 621900, China

<sup>4</sup>Department of Physics, Shanghai Jiaotong University, Shanghai, 200240, China

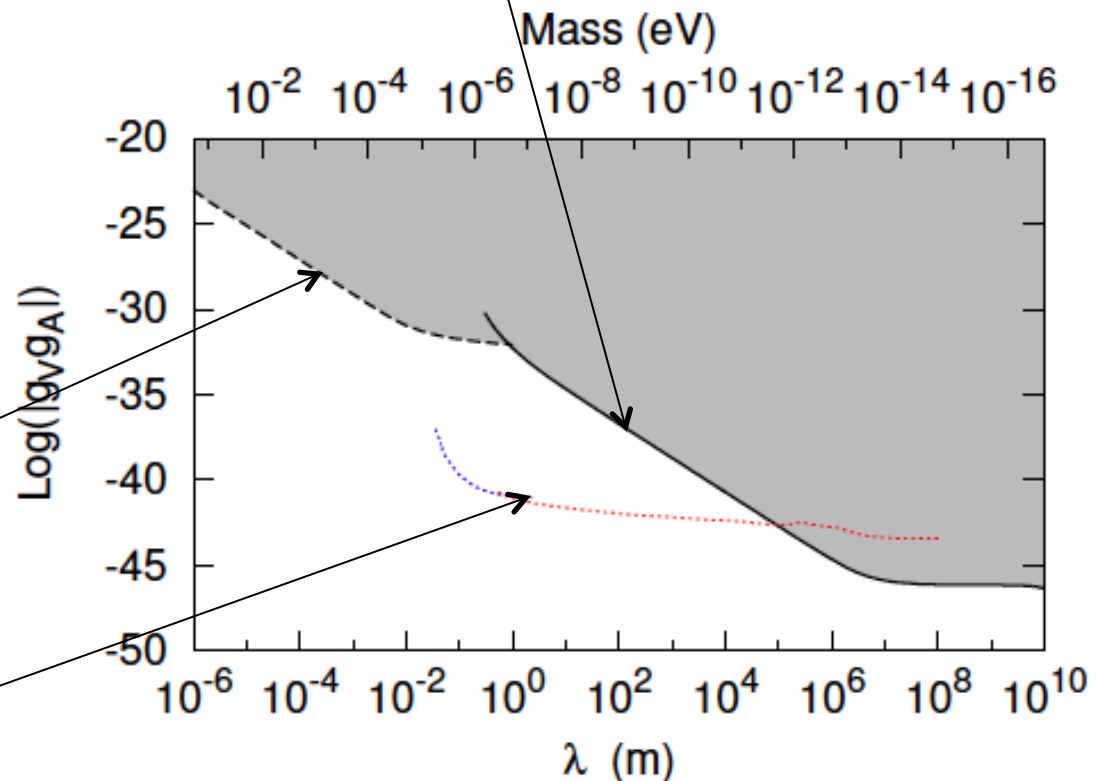
<sup>5</sup>Department of Physics, Southeast University, Nanjing, 211189, China

(Dated: August 12, 2015)

*This led to more work to constrain parity-odd interactions of the neutron*

*H. Yan and W. M. Snow, PRL 110, 082003 (2013)*

*E. G. Adelberger and T. A. Wagner, PRD 88, 031101 (2013)*



# A Spin-1 Boson Axial Coupling Search

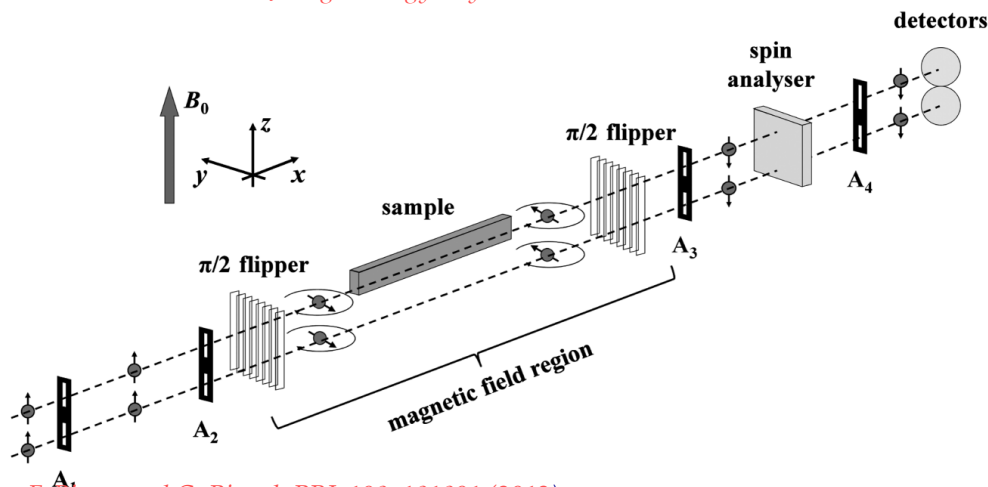
*F. Piegsa and G. Pignol placed a first upper bound on the axial coupling constant for a beyond-the-Standard-Model light spin-1 boson in the millimeter range by passing polarized neutrons near one side of a non-magnetic mass and looking for an induced rotation of the polarization direction.*

*F. Piegsa and G. Pignol, PRL 108, 181801 (2012)*

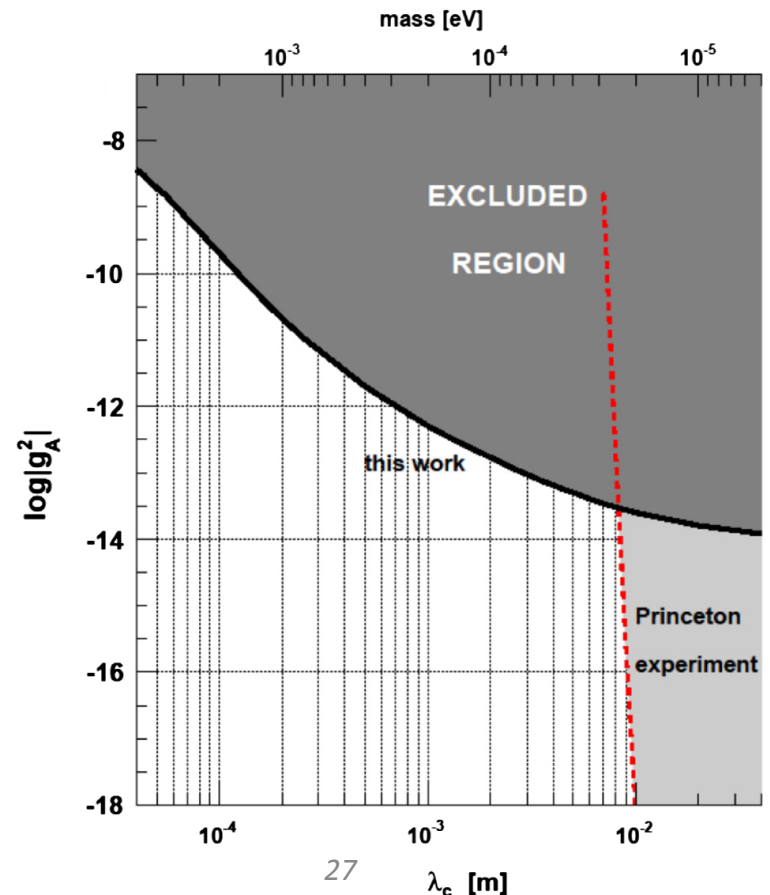
$$\mathcal{L} = \bar{\psi} (g_V \gamma^\mu + g_A \gamma^\mu \gamma^5) \psi X_\mu$$

$$V_{AA} \propto g_A^2 \vec{\sigma} \cdot (\vec{v} \times \hat{r}) \left( \frac{1}{\lambda} + \frac{1}{r} \right) \frac{e^{-r/\lambda}}{r}$$

*B. Dobrescu and I. Mocioiu, J. High Energy Physics. 11, 005 (2006)*



*F. Piegsa and G. Pignol, PRL 108, 181801 (2012)*



*Measurement performed at PSI*



# Neutron Spin Rotation (NSR) Collaboration

W.M. Snow<sup>1</sup>, E. Anderson<sup>1</sup>, L. Barron-Palos<sup>2</sup>, C.D. Bass<sup>3</sup>, B.E. Crawford<sup>4</sup>, C. Crawford<sup>5</sup>, W. Fox<sup>1</sup>, J. Fry<sup>1</sup>, C. Haddock<sup>1</sup>, B.R. Heckel<sup>6</sup>, M. Maldonado-Velazquez<sup>2</sup>, H.P. Mumm<sup>7</sup>, J.S. Nico<sup>7</sup>, C. Paudel<sup>8</sup>, S. Penn<sup>9</sup>, M.G. Sarsour<sup>8</sup>, S. van Sciver<sup>10</sup>, H.E. Swanson<sup>6</sup>, J. Vanderwerp<sup>1</sup>

Indiana University / CEEM<sup>1</sup>

Universidad Nacional Autonoma de Mexico<sup>2</sup>

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Gettysburg College<sup>4</sup>

University of Kentucky<sup>5</sup>

University of Washington<sup>6</sup>

National Institute of Standards and Technology<sup>7</sup>

Georgia State University<sup>8</sup>

Hobart and William Smith College<sup>9</sup>

Florida State University<sup>10</sup>

Support: NSF, NIST, DOE, CONACYT, LANL/LANSCE

# A search for very light exotic vector bosons using slow neutrons

$$\mathcal{L} = \bar{\psi} (g_V \gamma^\mu + g_A \gamma^\mu \gamma^5) \psi X_\mu$$

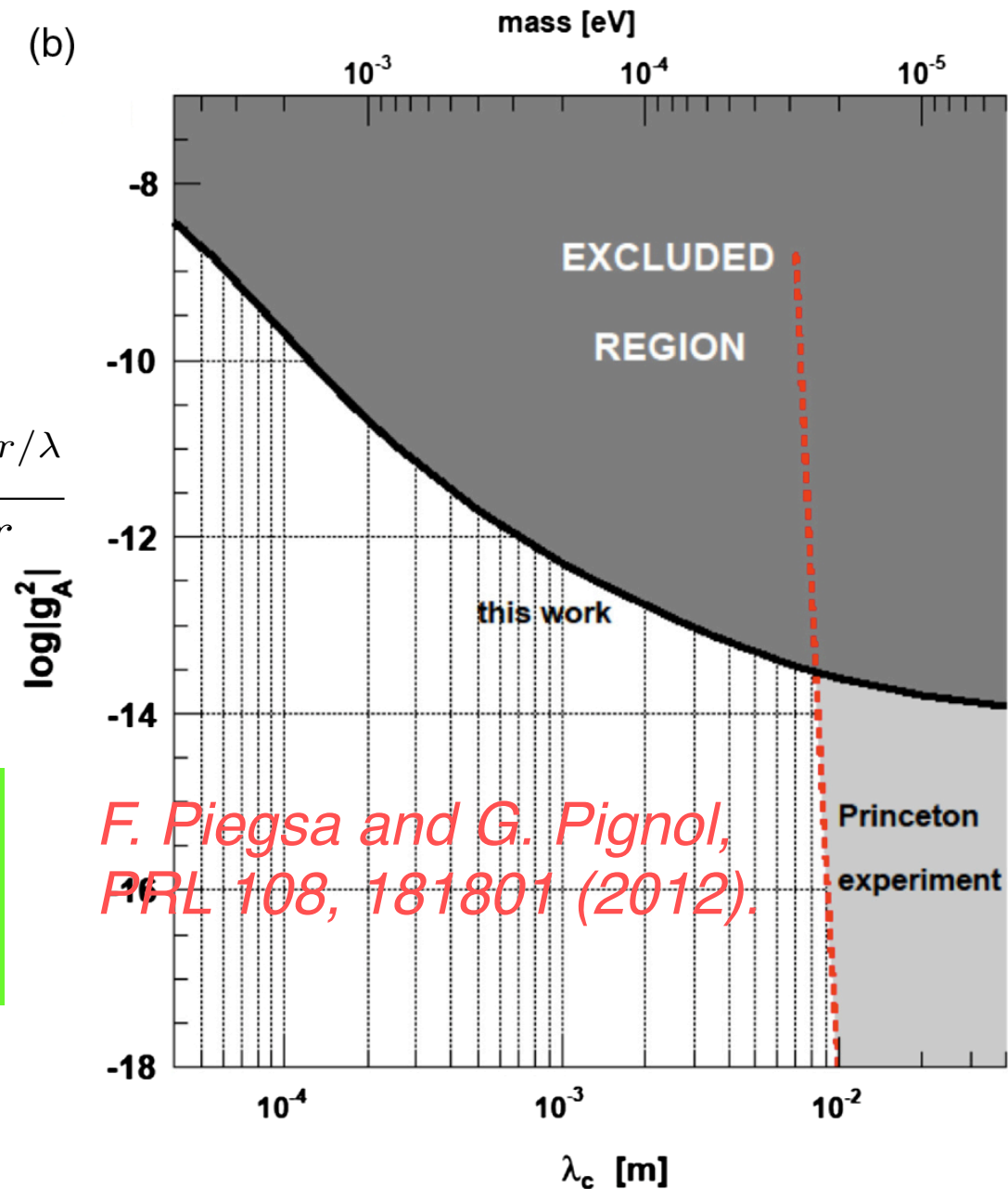
$$V_{AA} \propto g_A^2 \vec{\sigma} \cdot (\vec{v} \times \hat{r}) \left( \frac{1}{\lambda} + \frac{1}{r} \right) \frac{e^{-r/\lambda}}{r}$$

*Rotates neutron spin forward by an angle*

$$\phi = l \frac{g_A^2}{4} N \frac{\hbar}{mc} \lambda_c e^{-\frac{\Delta y}{\lambda_c}}$$

*Goal for fall 2016: complete experiment, improve previous search by ~3 orders of magnitude*

(b)



*F. Piegsa and G. Pignol, PRL 108, 181801 (2012).*

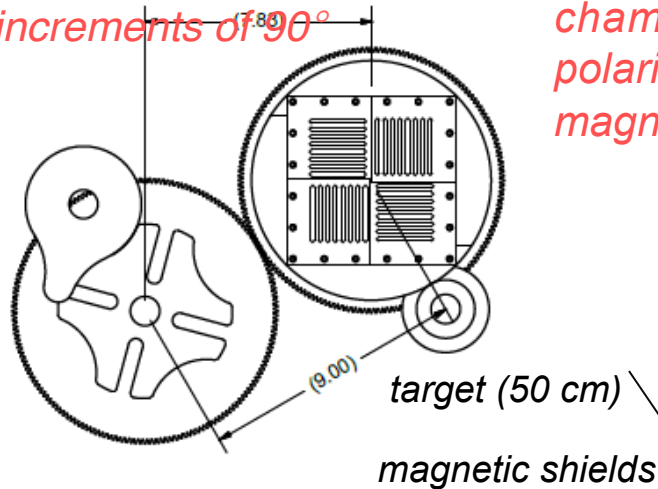


# Neutron Polarimetry Apparatus on FP12 at LANSCE

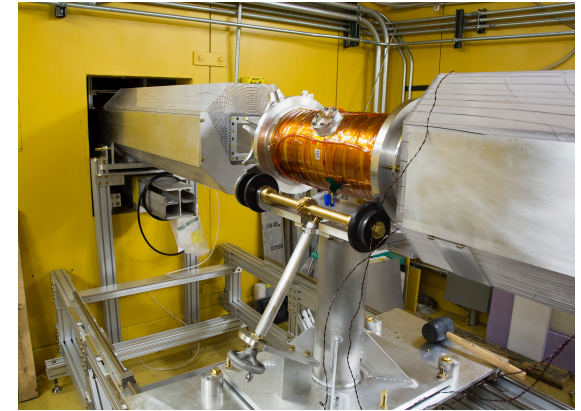
Geneva

Mechanism:

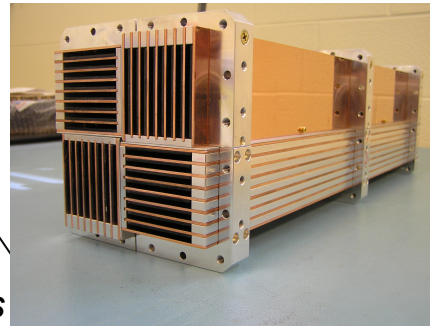
Rotate the target by increments of  $90^\circ$



View inside FP12 cave showing input/output supermirror guides and coils and target vacuum chamber. Neutron supermirror polarizer/analyzer, ion chamber, magnetic shielding not shown.



Plates of different nucleon density  $N$  are assembled so that the polarized neutrons traveling between the gaps will always see a density gradient.



AFP spin flipper

supermirror polarizer (45 cm)

non-magnetic SM ( $m=2$ ) input guide (125 cm)

input coil

$\pi/2$  turner

output coil

non-magnetic SM ( $m=2$ ) output guide (200 cm)

Segmented  $^3\text{He}$  ionization chamber

supermirror polarization analyzer (45 cm)

# Conclusions

*Experimental searches for weakly-coupled interactions with ranges from the millimeter to the atomic scale are actively pursued experimentally and appear in various theoretical scenarios*

*The properties of slow neutrons are well-suited to search for new interactions in this regime*

*Rapid experimental progress has occurred over the last few years, with the first measurements for certain spin-dependent interactions over sub-millimeter ever conducted.*

*New results and ongoing experiments have appeared at several neutron facilities throughout the world*